LEO OBJECT DETECTION USING BLIND STACKING

Ben Cooke

University of Warwick

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Motivation



- Not all of the objects are tracked, especially the smallest and faintest
- Optical observations can provide unique discovery and characterisation opportunities
- How do we detect LEO objects, without requiring prior information on their orbital properties?



Blind stacking – Introduction

- Method for detection of moving targets with unknown speed, direction and location
- Involves testing many potential paths on a pixel-by-pixel basis
- Procedure:
 - 1. Choose a path to test (i.e. rate and direction of motion between frames)
 - 2. Shift each image by the appropriate amount and combine the corresponding pixels (sum/mean/median)
 - 3. Compare stacked image with a master, replacing brighter pixels (and their meta data)
 - 4. Repeat steps 1-3 for each potential path



Individual frames containing a target moving (50,50) pixels per frame. Not all frames are shown.

To identify the target, multiple paths must be tested.



Tested path: 50 pixels South 50 pixels East (-50,50)

This is an incorrect path













Tested path: 50 pixels North 50 pixels East (50,50)

This is the correct path





- 8

- 6

- 4

- 2

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- -2

-4

- -6

Stack (3 frames)



This is the correct path











Blind stacking – Pros and Cons

• Pros

- Can test multiple orbits per dataset
- Can identify unknown objects or objects on unknown or imprecise orbits
- Can identify multiple targets per dataset, even if they have different motions/orbits
- Cons
 - Fast objects are seen as extended streaks compared to tracked observations
 - Therefore fainter
 - Computing time is significant

Optimisation avenues

• Exposure time

- Shorter streaks correspondingly fewer paths to test
- Lower background noise (to a point)
- Greater memory requirement
- Number of frames
 - More signal to stack
 - More memory required
- Binning
 - Increases SNR (to a point)
 - Fewer pixels
 - Less resolution
- Plenty of pipeline optimisations as well
 Path range, binning style, stacking method etc.

Simulation



- Targets faint enough to test on are hard to reliably schedule
- Need to simulate targets with random speed, direction, magnitude and location
- Based on real data from our telescope in La Palma. Simulated targets are then injected
- Can then run the blind stack code (for a range of optimisation parameters) and check for recoverability



Single frame limits

 Based on our observed data we can predict the limiting magnitude of a target in a single frame – function of exposure time and target velocity





Results – recovery

- Can look at the recoverability of simulated targets as a function of magnitude
- This has been run for multiple combinations of the three optimisation parameters
- In these results we'll look at each optimisation parameter separately



Exposure time



Number of frames



Binning





Results – runtime

- Runtime of the system is also very important
- Ideally want to be able to run in real time
- However, runtime significantly affected by optimisation parameters





Conclusions and next steps

- Blind stacking can recover LEO objects with little to no prior information
- Can detect fainter objects than are available in individual exposures
- Short exposures are better lower noise and reduced runtime
- ~5s of exposure total is best longer and more memory is required, shorter and can't stack as many frames to get the faint targets
- Binning=4 gives best recovery (but binning=8 gives shorter runtime)
- Further computational optimisation
- Continuing comparison with observed data

Thank you for listening