



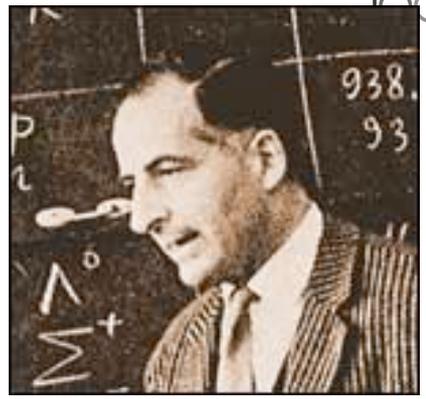
First $\nu_{\mu} \rightarrow \nu_e$ Oscillation Results from MiniBooNE

*Morgan Wascko
Imperial College London*

*Warwick Particle Physics Seminar
May 8, 2007*

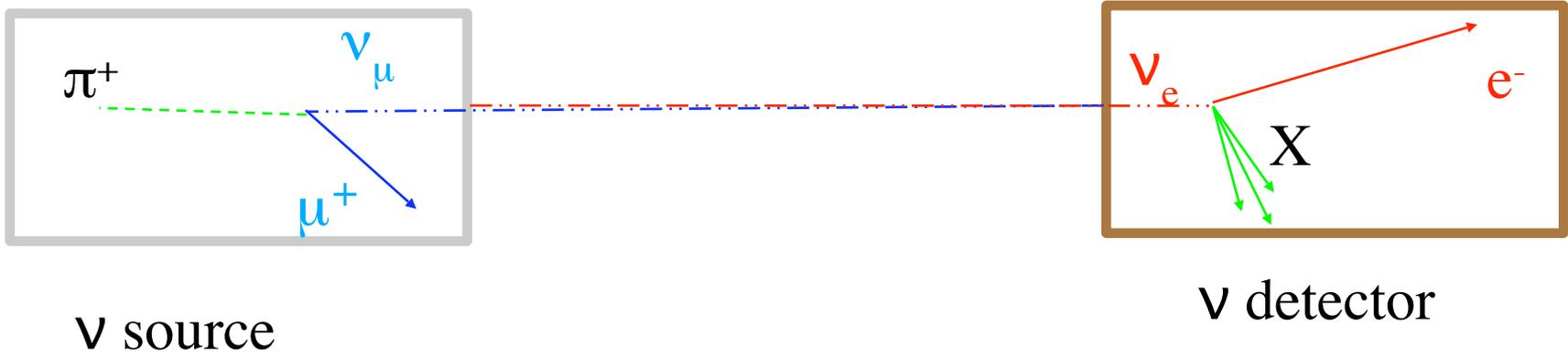
1. Motivation & Introduction
2. Description of the Experiment
3. Analysis Overview
4. Two Independent Oscillation Searches
5. First Results

Motivation: Neutrino Oscillations



Pontecorvo, 1957

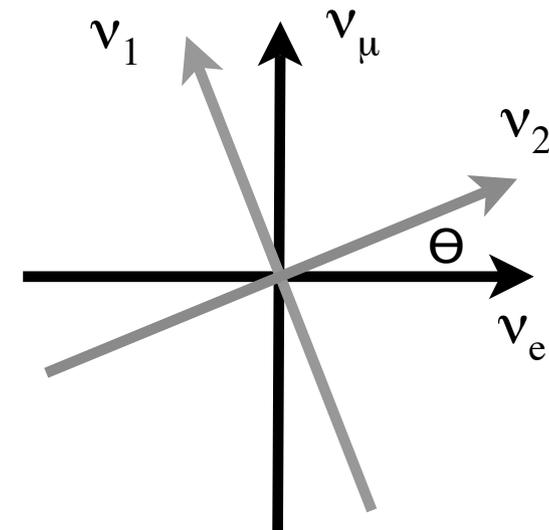
if neutrinos have mass, a neutrino that is produced as a ν_μ (e.g. $\pi^+ \rightarrow \mu^+ \nu_\mu$) has a non-zero probability to oscillate and some time later be detected as a ν_e (e.g. $\nu_e n \rightarrow e^- p$)



Motivation: Neutrino Oscillations

In a world with 2 neutrinos,
if the weak eigenstates (ν_e, ν_μ)
are different from the mass eigenstates (ν_1, ν_2):

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$



The weak states are mixtures of the mass states:

$$|\nu_\mu\rangle = -\sin\theta|\nu_1\rangle + \cos\theta|\nu_2\rangle$$

$$|\nu_\mu(t)\rangle = -\sin\theta (|\nu_1\rangle e^{-iE_1 t}) + \cos\theta (|\nu_2\rangle e^{-iE_2 t})$$

The probability to find a ν_e when you started with a ν_μ is:

$$P_{oscillation}(\nu_\mu \rightarrow \nu_e) = |\langle \nu_e | \nu_\mu(t) \rangle|^2$$

Motivation: Neutrino Oscillations

In units that experimentalists like:

$$P_{oscillation}(\nu_{\mu} \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 (eV^2) L (km)}{E_{\nu} (GeV)} \right)$$

Oscillation probability between 2 flavour states depends on:

1. fundamental parameters

$\Delta m^2 = m_1^2 - m_2^2$ = mass squared difference between states

$\sin^2 2\theta$ = mixing between ν flavours

2. experimental parameters

L = distance from ν source to detector

E = ν energy

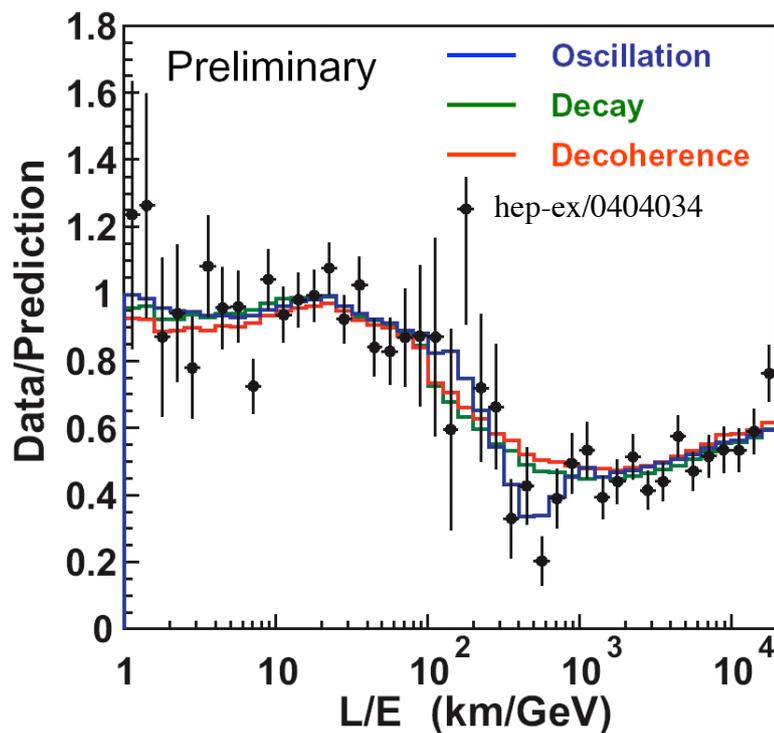
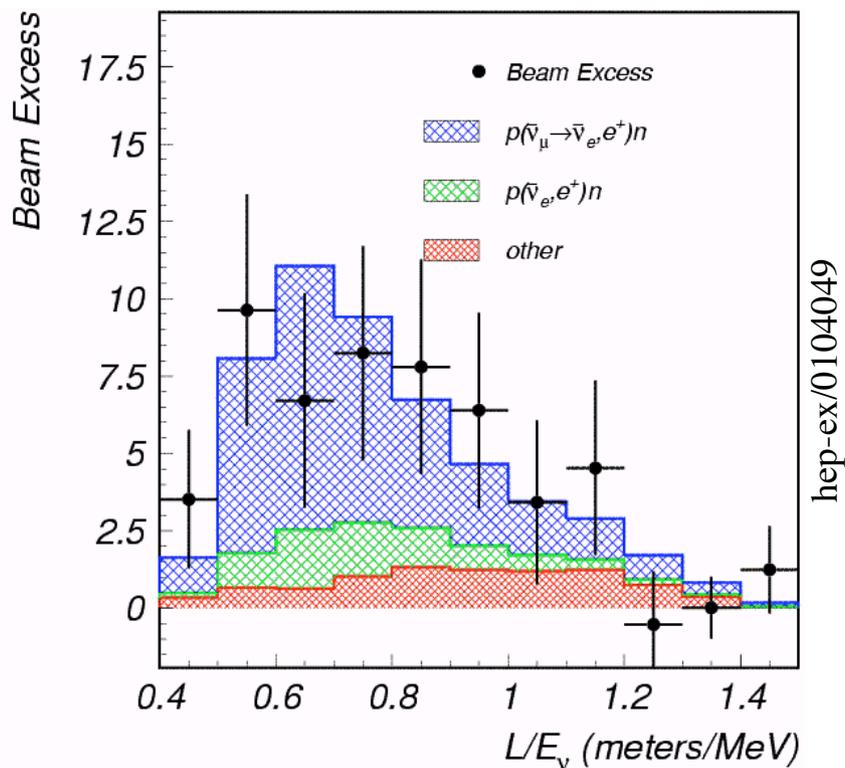
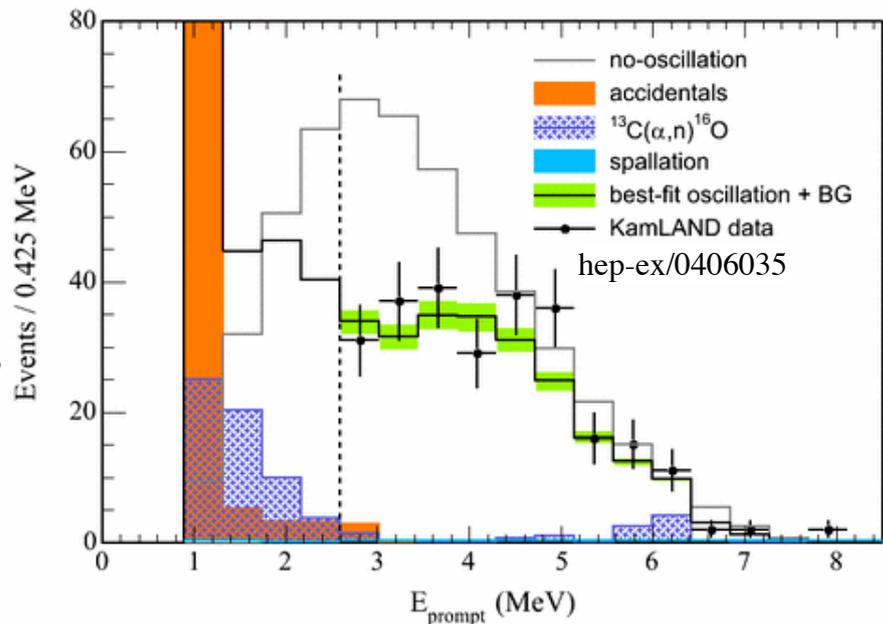


Motivation: Oscillation Signals

Solar ν : measured by Homestake, ..., SNO
confirmed by KamLAND

Atmospheric ν : measured by K-II, ..., Super-K
confirmed by Soudan2, MACRO, K2K, MINOS

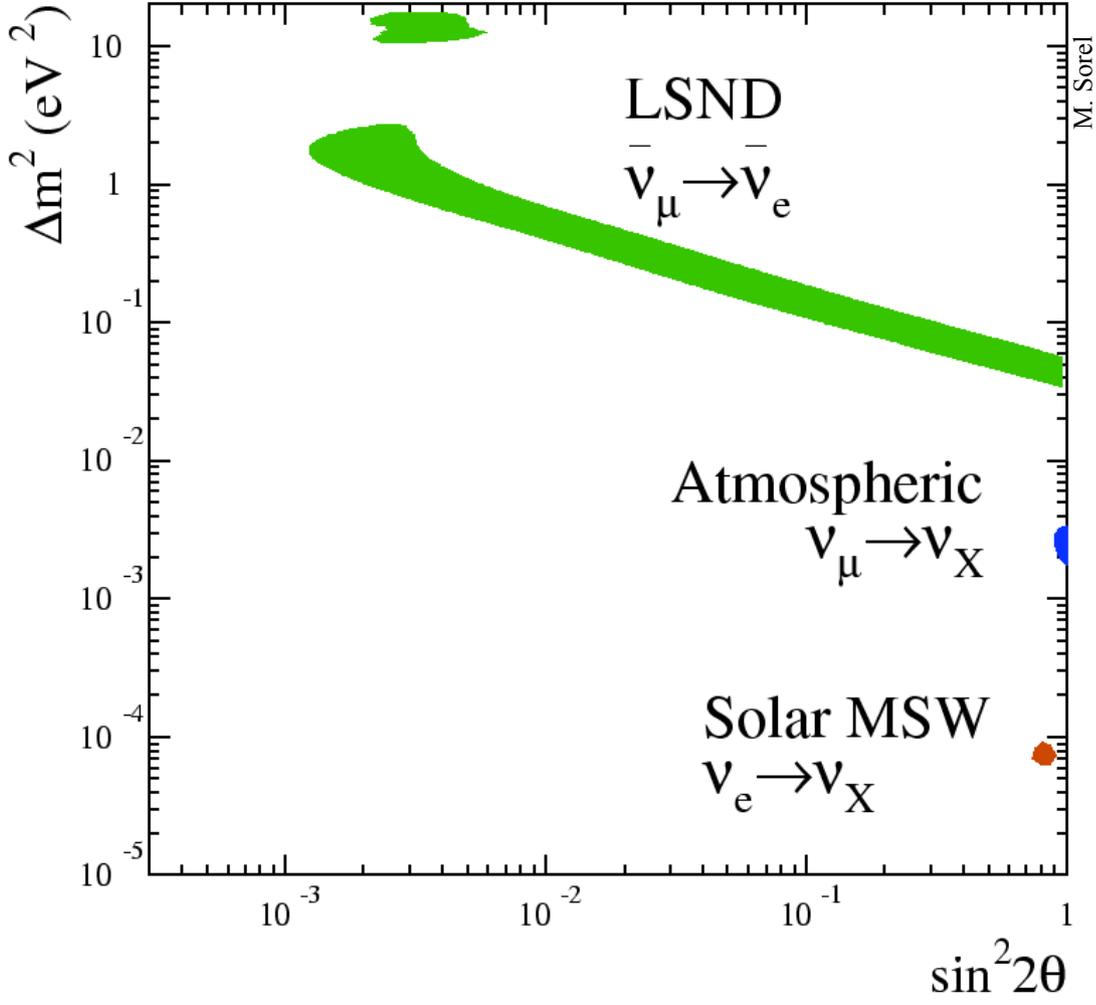
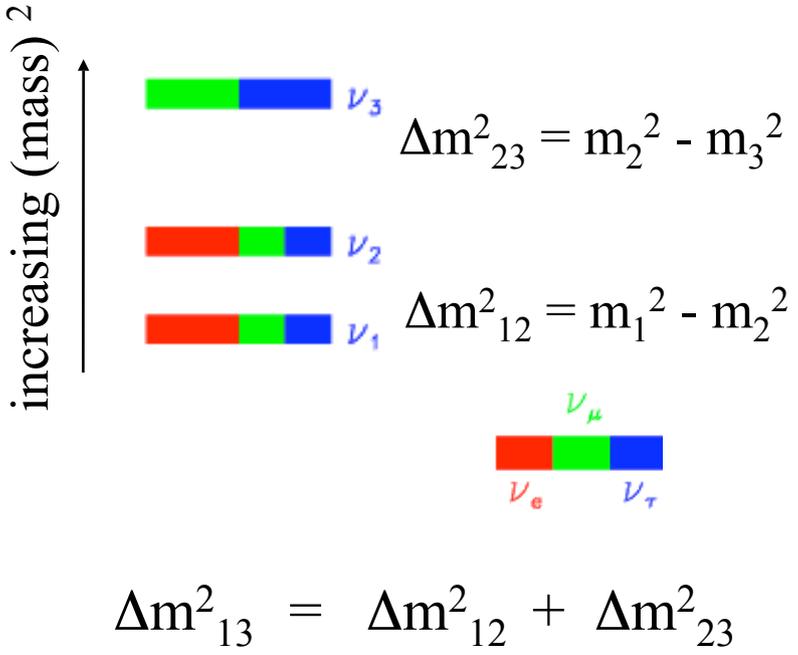
Accelerator ν : measured by LSND
unconfirmed



Motivation: The Problem

$$P_{oscillation}(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 (eV^2) L (km)}{E_\nu (GeV)} \right)$$

A standard 3 neutrino picture:

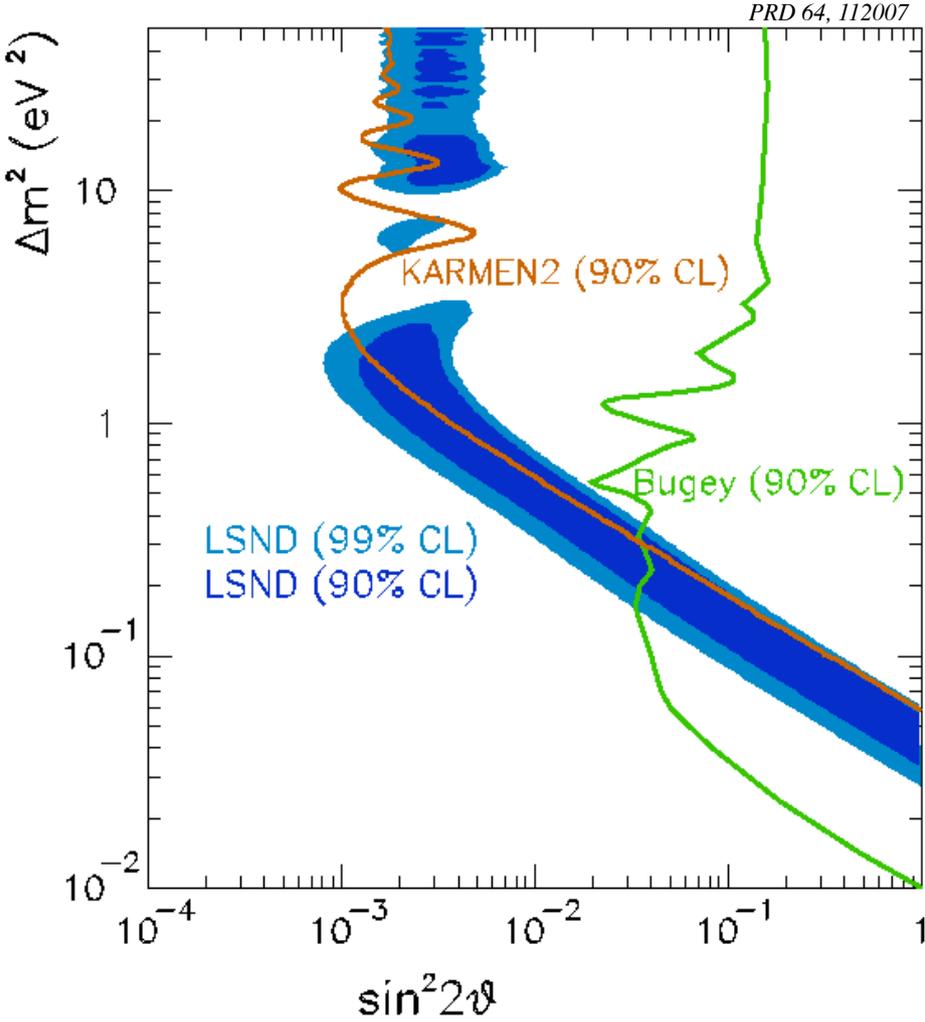


The oscillation signals cannot be reconciled without introducing physics (even farther) beyond the Standard Model.

Motivation: LSND

MiniBooNE was proposed in 1997 to address the LSND result.

LSND observed a 4σ excess of $\bar{\nu}_e$ events in a $\bar{\nu}_\mu$ beam: $87.9 \pm 22.4 \pm 6.0$
 interpreted as 2-neutrino oscillations, $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = 0.26\%$



$$P = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 (eV^2) L (km)}{E_\nu (GeV)} \right)$$

MiniBooNE strategy:

Keep (L/E_ν) same as LSND but change systematics, including event signature:

- Order of magnitude higher E_ν than LSND
- Order of magnitude longer baseline L than LSND
- Search for excess of ν_e events above background

Simple $\nu_\mu \rightarrow \nu_e$ oscillation

The MiniBooNE Collaboration

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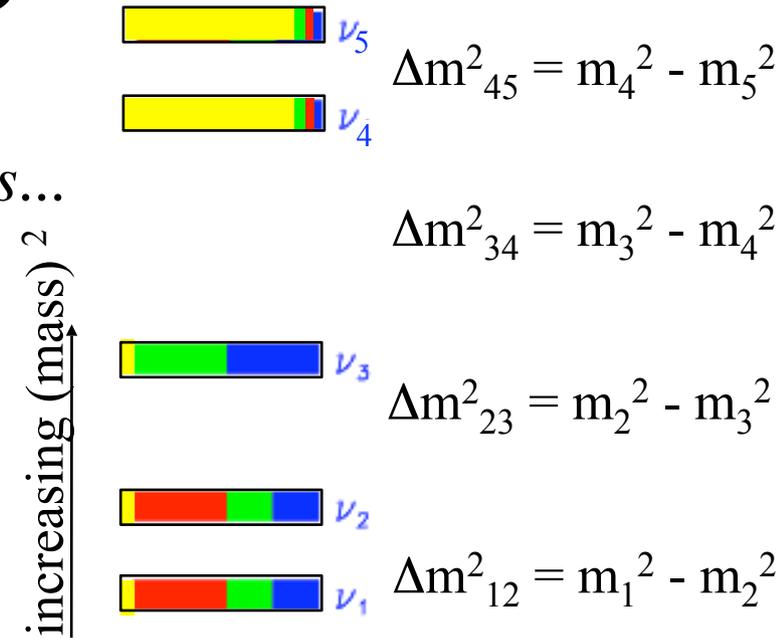
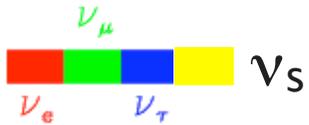
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Motivation: MiniBooNE and LSND

If MiniBooNE observes LSND-type ν oscillations...

The simplest explanation is to add more ν s,
to allow more independent Δm^2 values.

The new ν s would have to be **sterile**, otherwise
they would have been seen already.



If MiniBooNE does not observe LSND-type oscillations...

The Standard Model wins again!

Today: MiniBooNE's initial results on testing the LSND anomaly

- A generic search for a ν_e excess in our ν_μ beam,
- An analysis of the data within a $\nu_\mu \rightarrow \nu_e$ appearance-only context

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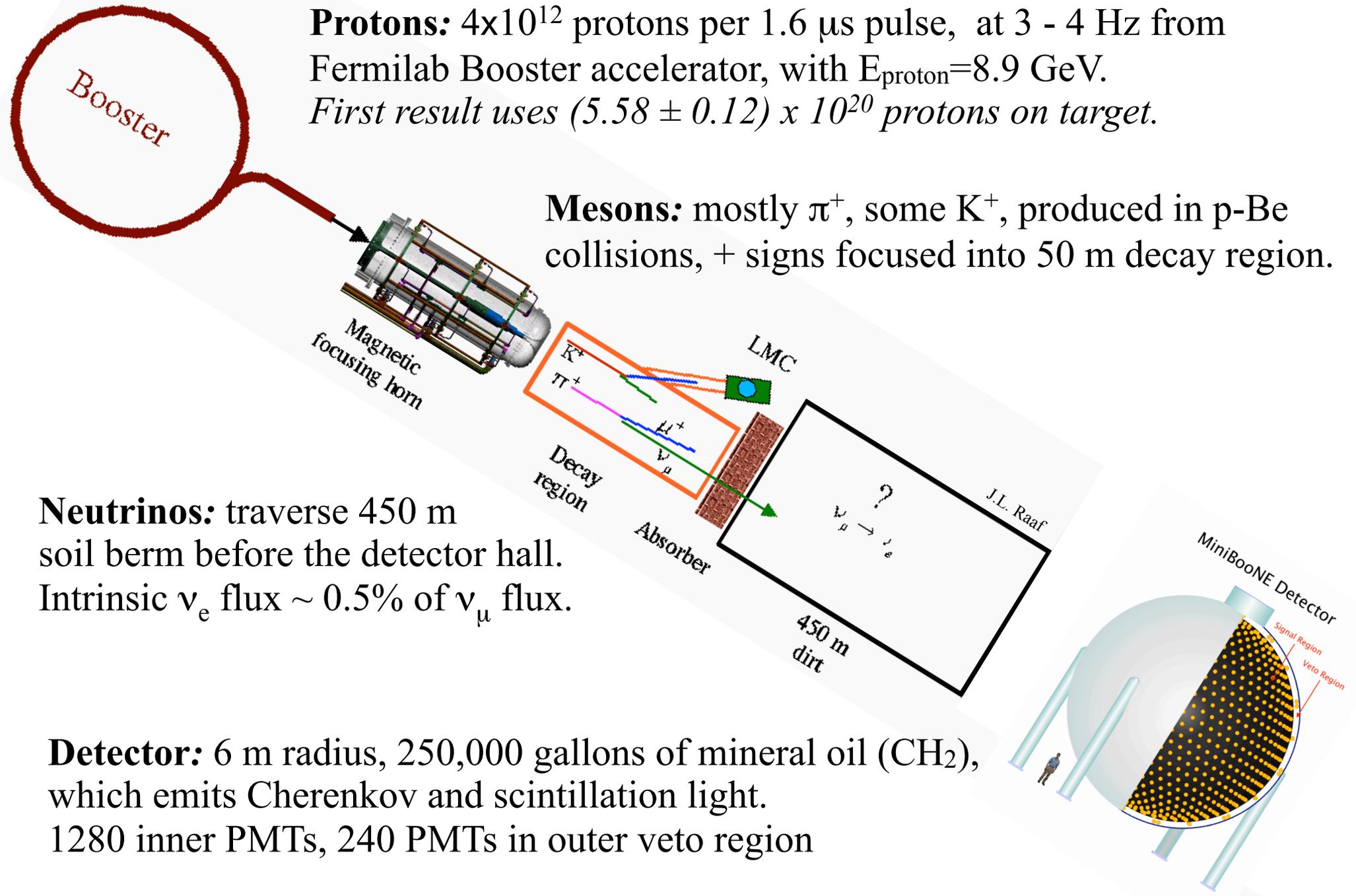
MiniBooNE Overview: Beam and Detector

Protons: 4×10^{12} protons per $1.6 \mu\text{s}$ pulse, at 3 - 4 Hz from Fermilab Booster accelerator, with $E_{\text{proton}} = 8.9 \text{ GeV}$.
First result uses $(5.58 \pm 0.12) \times 10^{20}$ protons on target.

Mesons: mostly π^+ , some K^+ , produced in p-Be collisions, + signs focused into 50 m decay region.

Neutrinos: traverse 450 m soil berm before the detector hall.
 Intrinsic ν_e flux $\sim 0.5\%$ of ν_μ flux.

Detector: 6 m radius, 250,000 gallons of mineral oil (CH_2), which emits Cherenkov and scintillation light.
 1280 inner PMTs, 240 PMTs in outer veto region

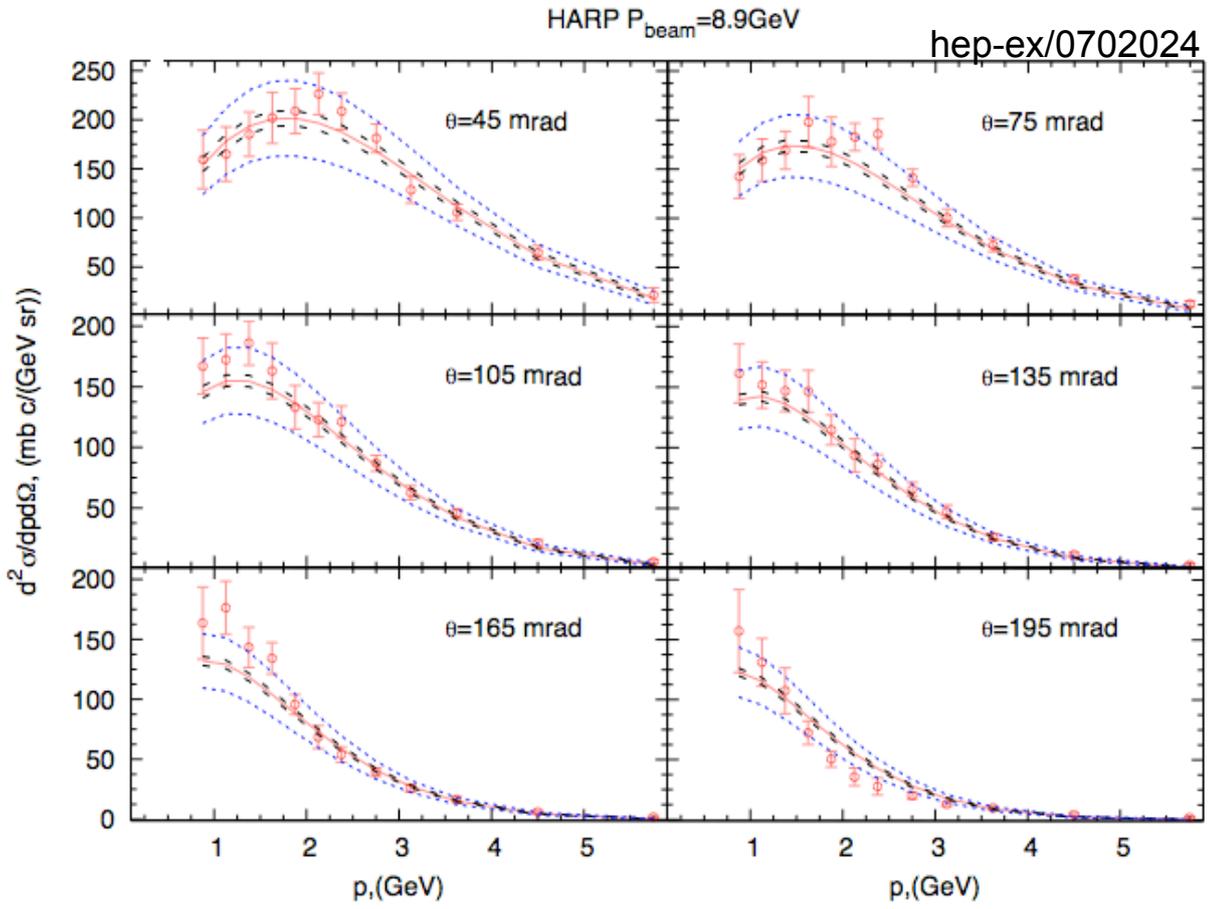
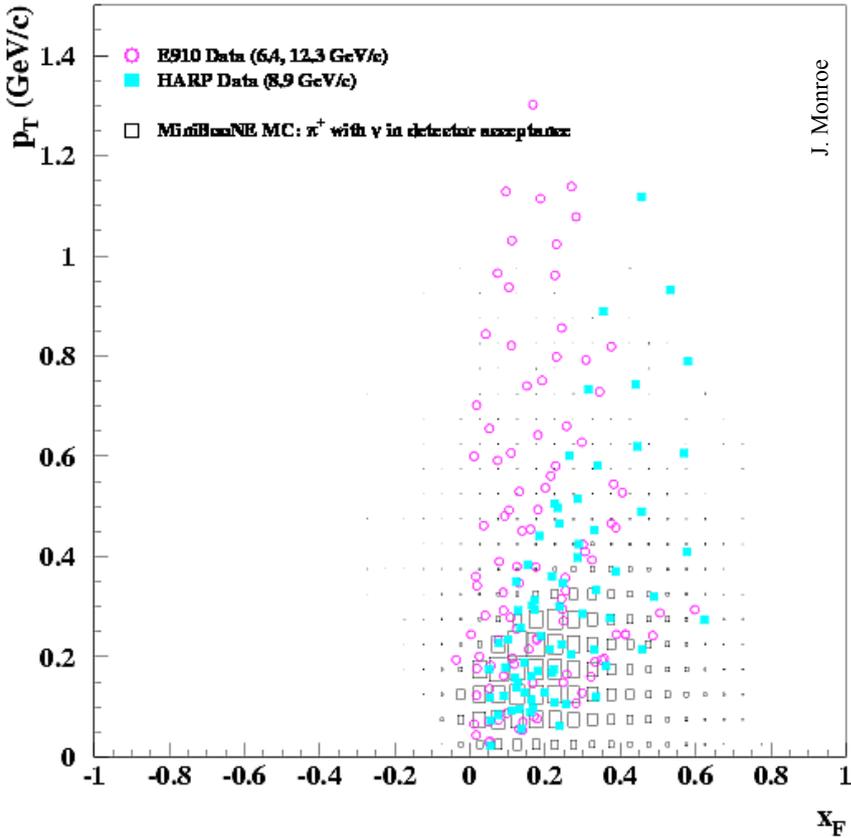


Booster Neutrino Beam: Modelling Meson Production

Prediction from a fit to $p Be \rightarrow \pi^+ X$ production data from E910 and HARP experiments ($p_p = 6-12 \text{ GeV}/c$, $\Theta_\pi = 0 - 330 \text{ mrad}$.)

Fit (shown at right) uses Sanford-Wang parametrisation

HARP has excellent phase space coverage for MiniBooNE



π^- similarly parametrised
Kaons flux predictions use a Feynman Scaling
parametrisation (no HARP data yet)

Booster Neutrino Beam: Neutrino Flux

MiniBooNE is searching for an excess of ν_e in a ν_μ beam

Modelled with a Geant4 Monte Carlo

“Intrinsic” $\nu_e + \bar{\nu}_e$ content: 0.5%

ν_e Sources:

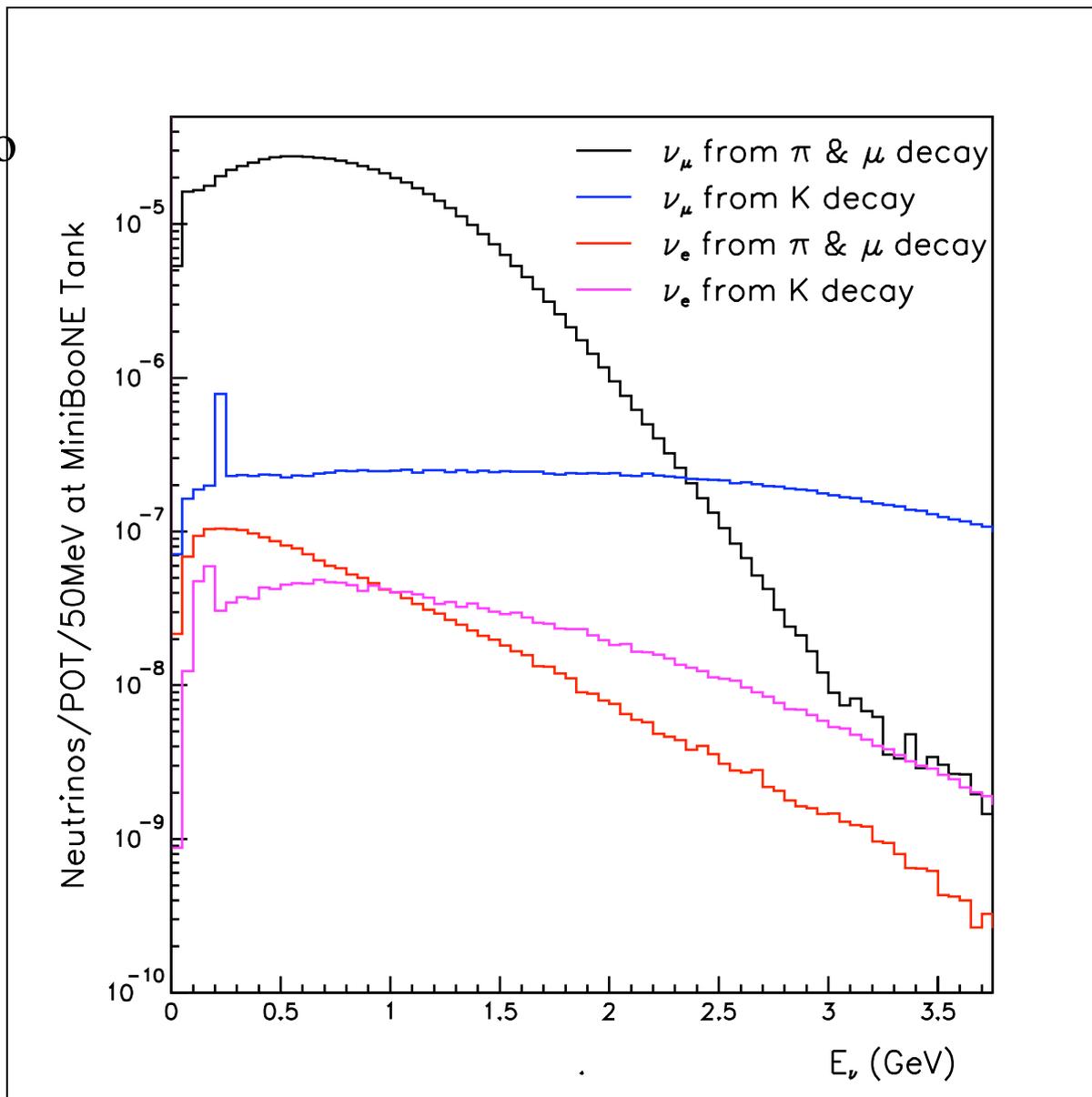
$$\mu^+ \rightarrow e^+ \bar{\nu}_\mu \nu_e \quad (42\%)$$

$$K^+ \rightarrow \pi^0 e^+ \nu_e \quad (28\%)$$

$$K^0 \rightarrow \pi^+ e^- \nu_e \quad (16\%)$$

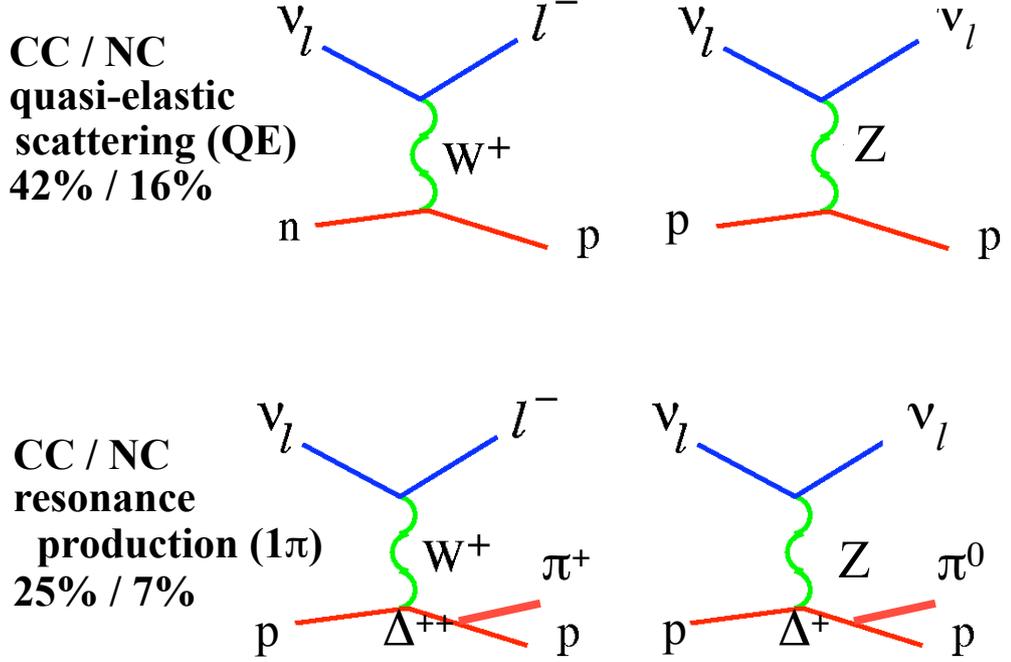
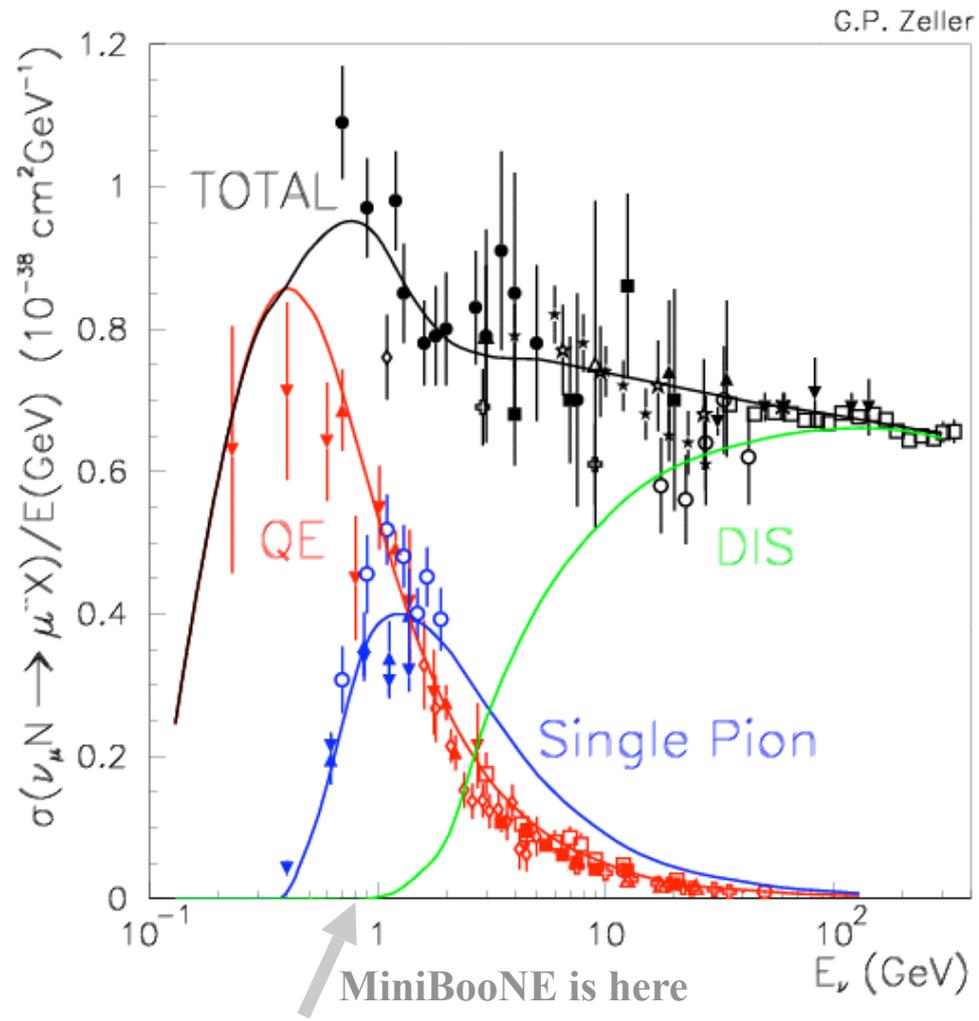
$$\pi^+ \rightarrow e^+ \nu_e \quad (4\%)$$

Antineutrino content: 6%



MiniBooNE Detector: Neutrino Cross Sections

Modelling what the neutrinos do in the detector



Cross section predictions from
NUANCE Monte Carlo

Use CCQE events for oscillation analysis signal channel:

$$E_{\nu}^{QE} = \frac{1}{2} \frac{2M_p E_{\ell} - m_{\ell}^2}{M_p - E_{\ell} + \sqrt{(E_{\ell}^2 - m_{\ell}^2) \cos^2 \theta_{\ell}}}$$

Only need lepton direction
and angle to find ν energy!

MiniBooNE Detector: Optics

charged final state particles produce Υ s

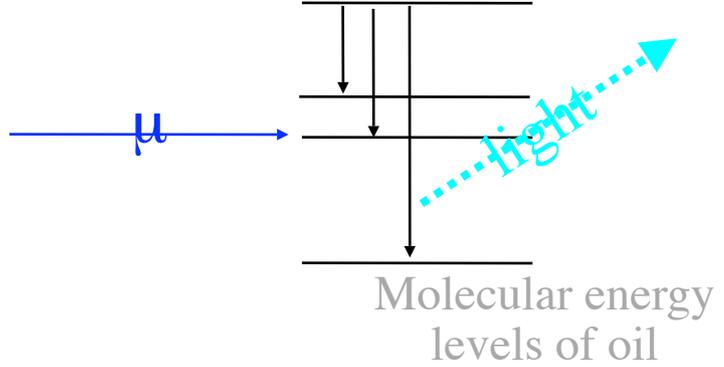
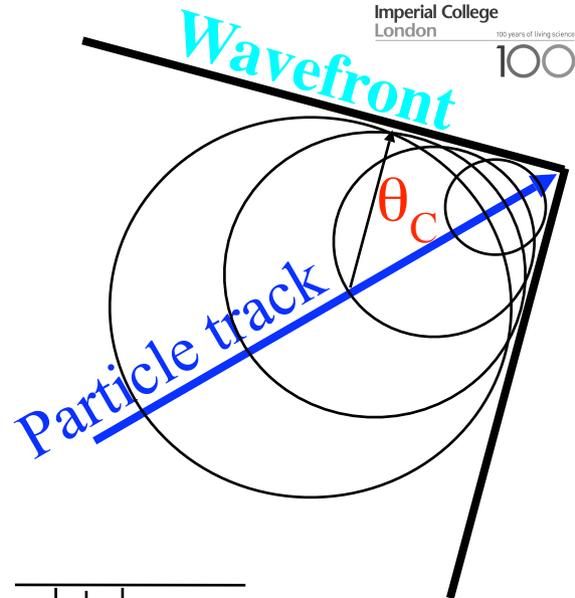
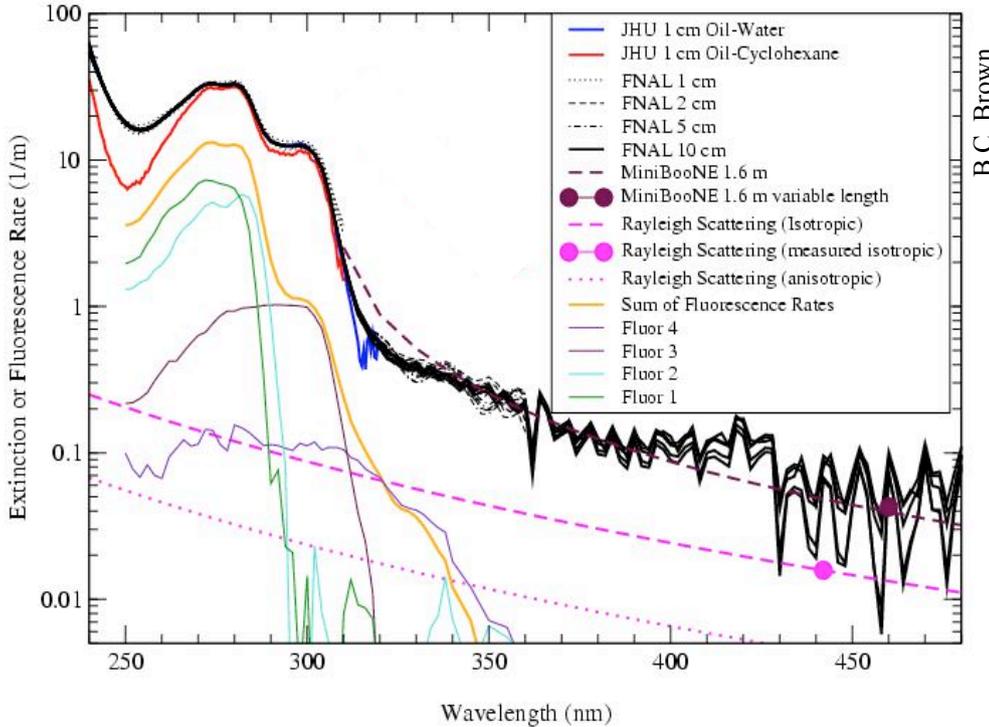
Cherenkov radiation

- Light emitted by oil if particle $v > c/n$
- forward and prompt in time

Scintillation

- Excited molecules emit de-excitation Υ s
- isotropic and late in time

Extinction Rate for MiniBooNE Marcol 7 Mineral Oil



Υ s are (possibly) detected by PMTs after undergoing absorption, reemission, scattering, fluorescence

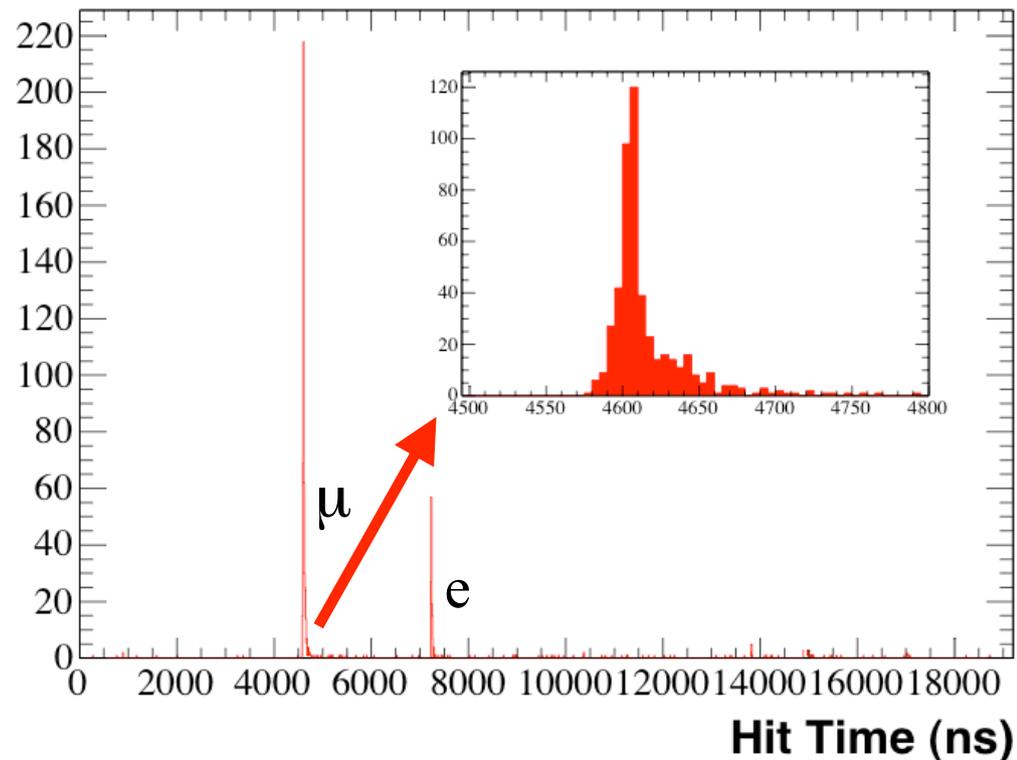
“the optical model”

MiniBooNE Detector: Hits

First set of cuts based on simple hit clusters in time: “sub-events.”

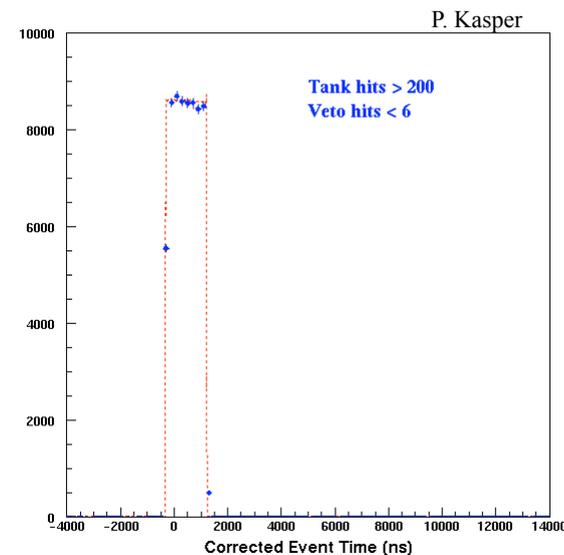
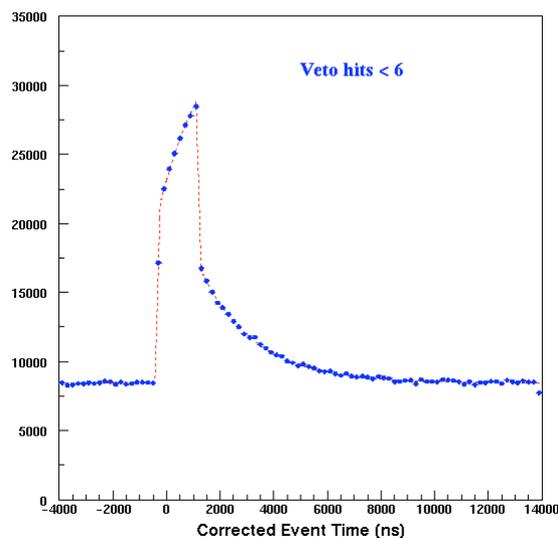
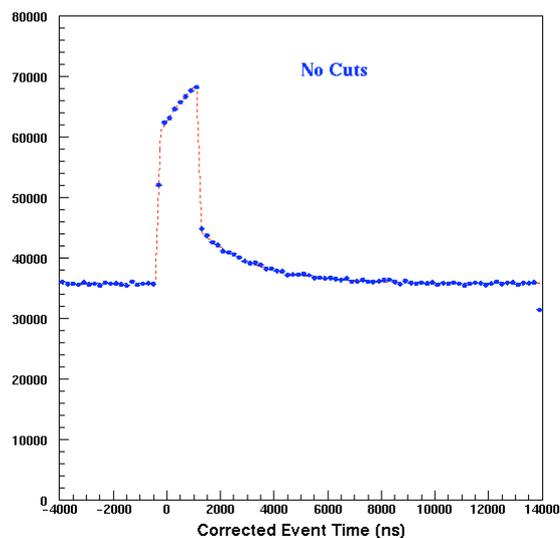
Most events are from ν_μ CC interactions, with characteristic two “sub-event” structure from stopped μ decay.

ν_e CC interactions have 1 “sub-event”.

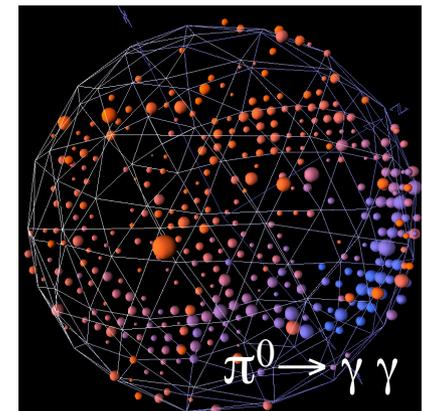
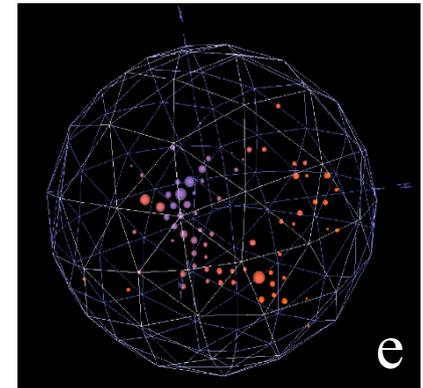
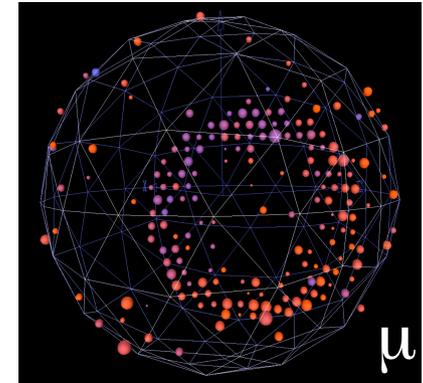
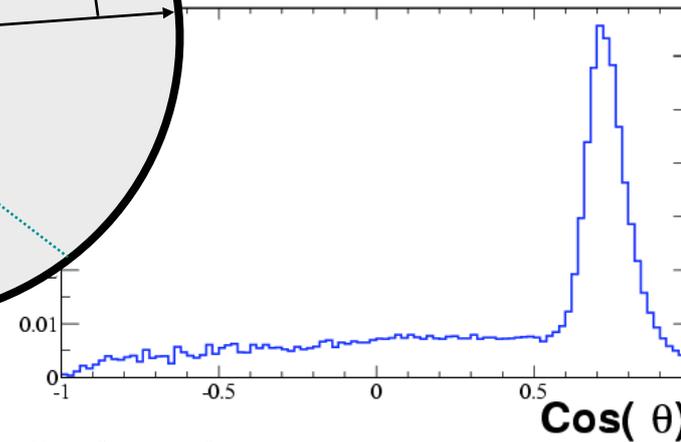
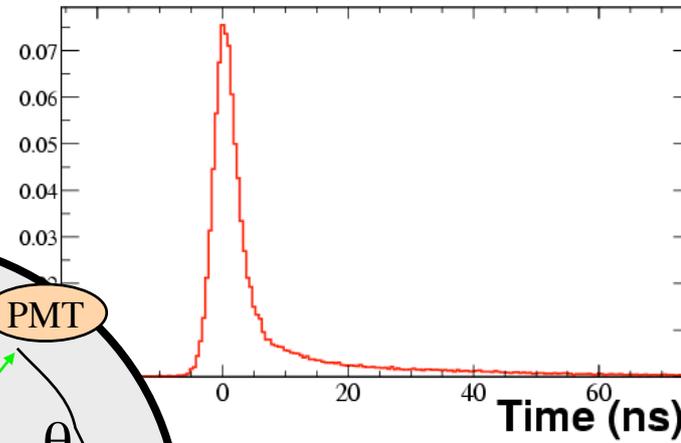
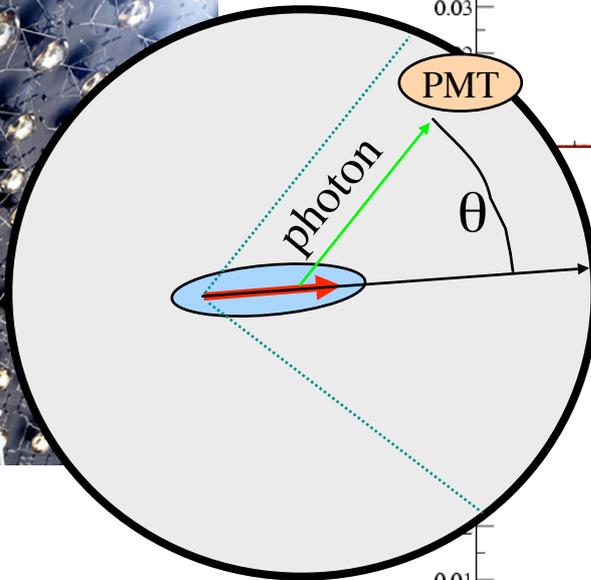
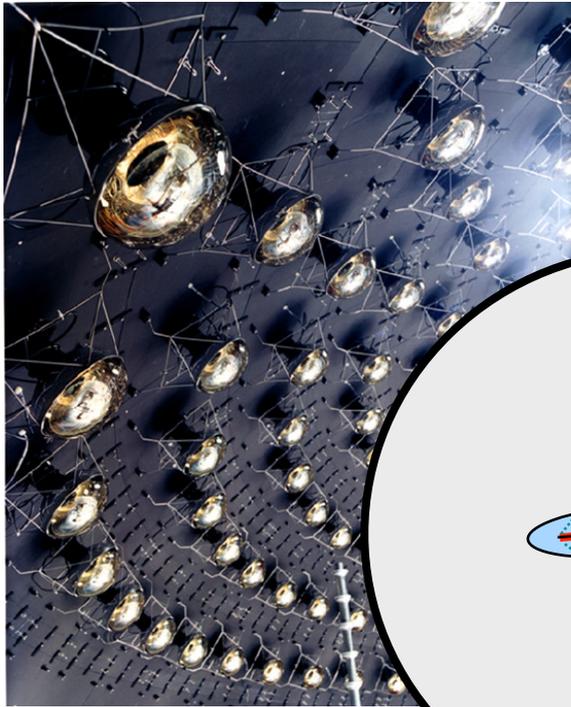


Simple cuts eliminate cosmic ray events:

1. Require < 6 veto PMT hits,
2. Require > 200 tank PMT hits.



MiniBooNE Detector: Reconstruction and Particle ID



Reconstruction:

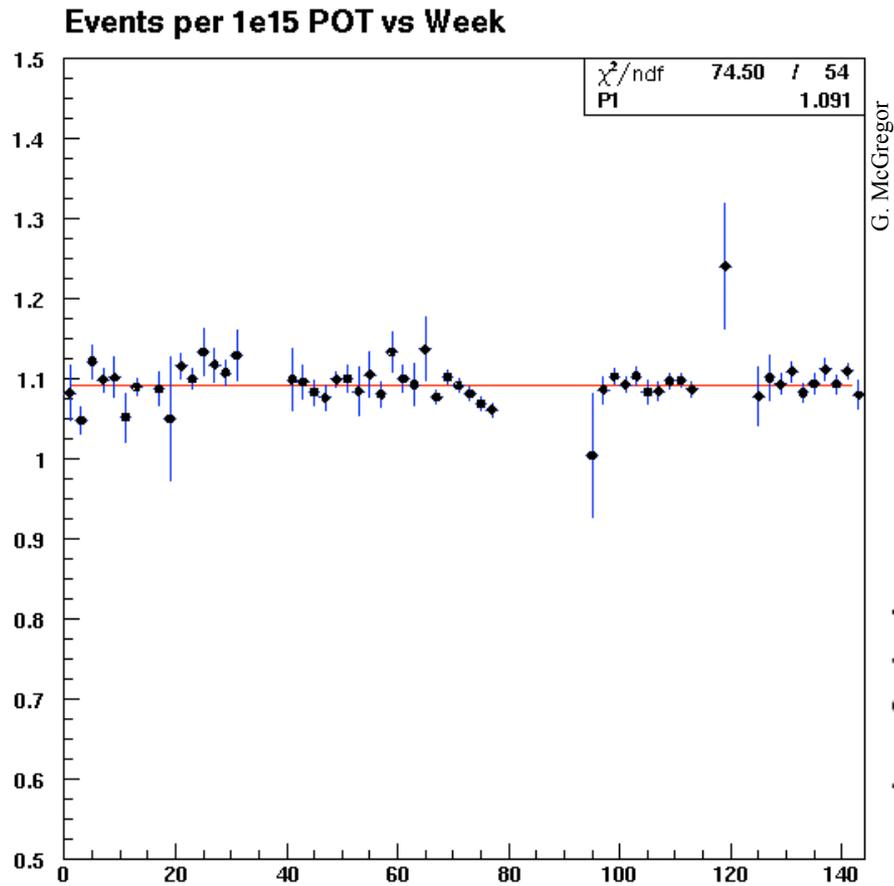
PMTs collect γ s, record t and q ,
fit time and angular distributions to find tracks

Final State Particle Identification:

muons have sharp Cherenkov rings and long tracks
electrons have fuzzy rings, from multiple scattering, and short tracks
neutral pions decay to 2 γ s, which convert and produce 2 fuzzy rings,
easily misidentified as electrons if one ring gets lost!

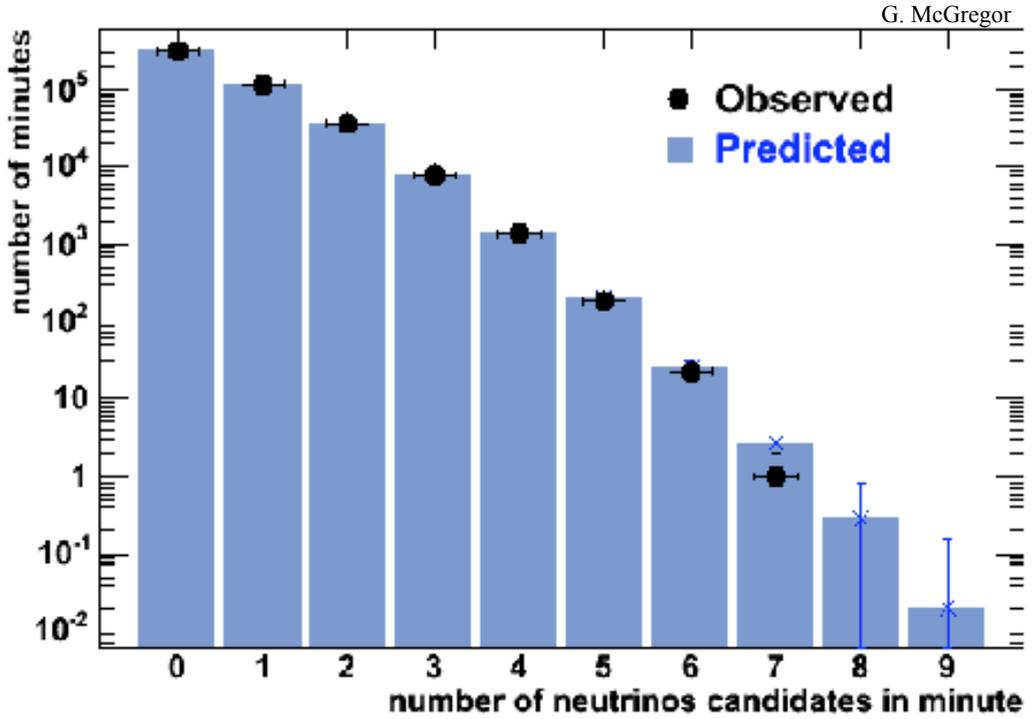
MiniBooNE Beam & Detector: Stability

Neutrinos per proton on target throughout the neutrino run:



MiniBooNE observes
~1 neutrino interaction
per 1E15 protons.

*Observed and expected
events per minute*



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Analysis Overview: Blind Analysis

To avoid bias, MiniBooNE has done a blind analysis.

“Closed Box” Analysis

To study the data, we defined specific event sets with $< 1\sigma \nu_e$ signal for analysis.

Initial Open Boxes

all non-beam-trigger data

0.25% random sample

ν_μ CCQE

ν_μ NC1 π^0

“dirt”

all events with $E_\nu > 1.4$ GeV

ν_μ CC1 π^+

ν_μ -e elastic

Use

calibration and MC tuning

an unbiased data set

measure flux, E_ν^{QE} , oscillation fit

measure rate for MC

measure rate for MC

check MC rate

check MC rate

check MC rate

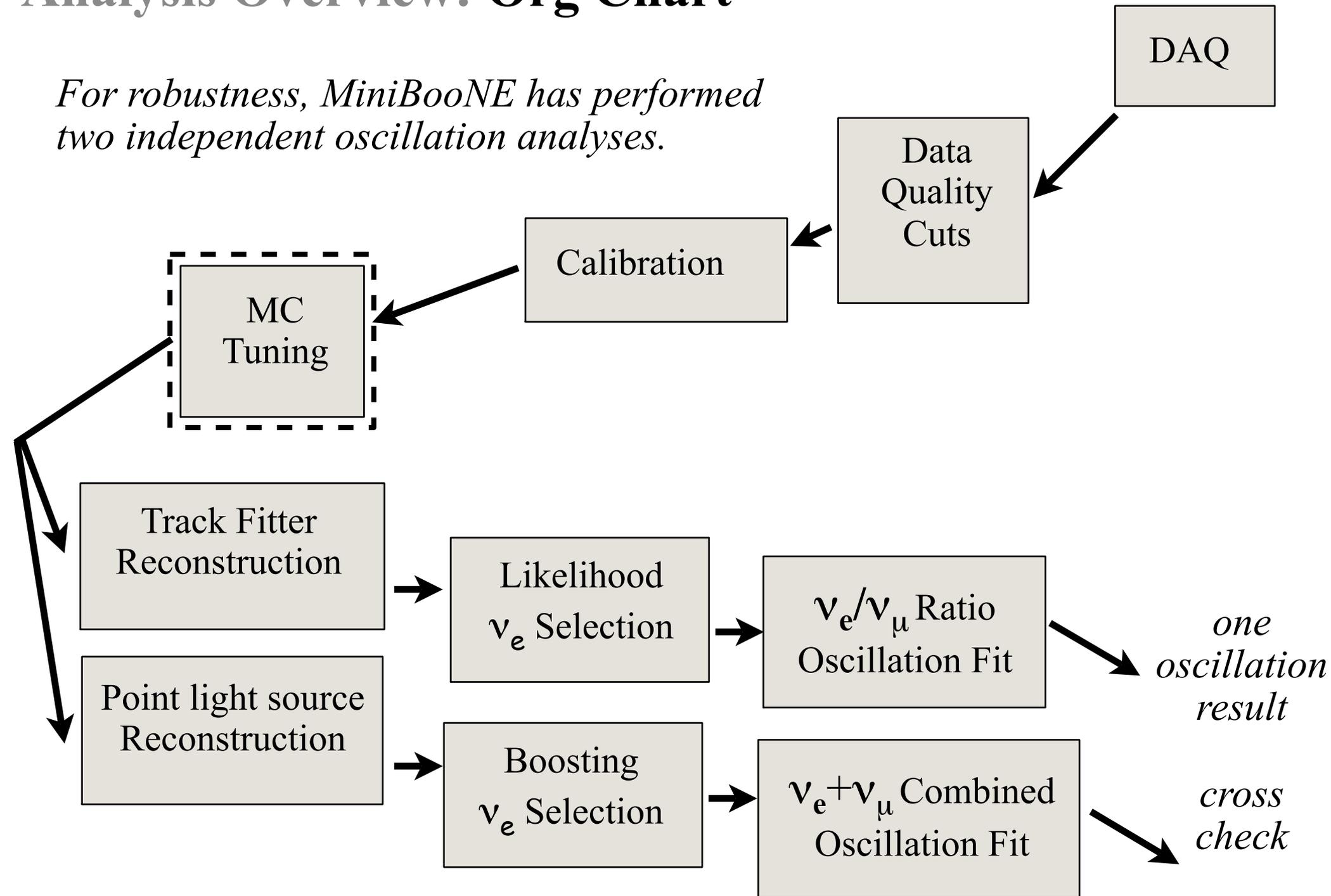
Second Step:

One closed signal box

explicitly sequester the signal,
99% of data open

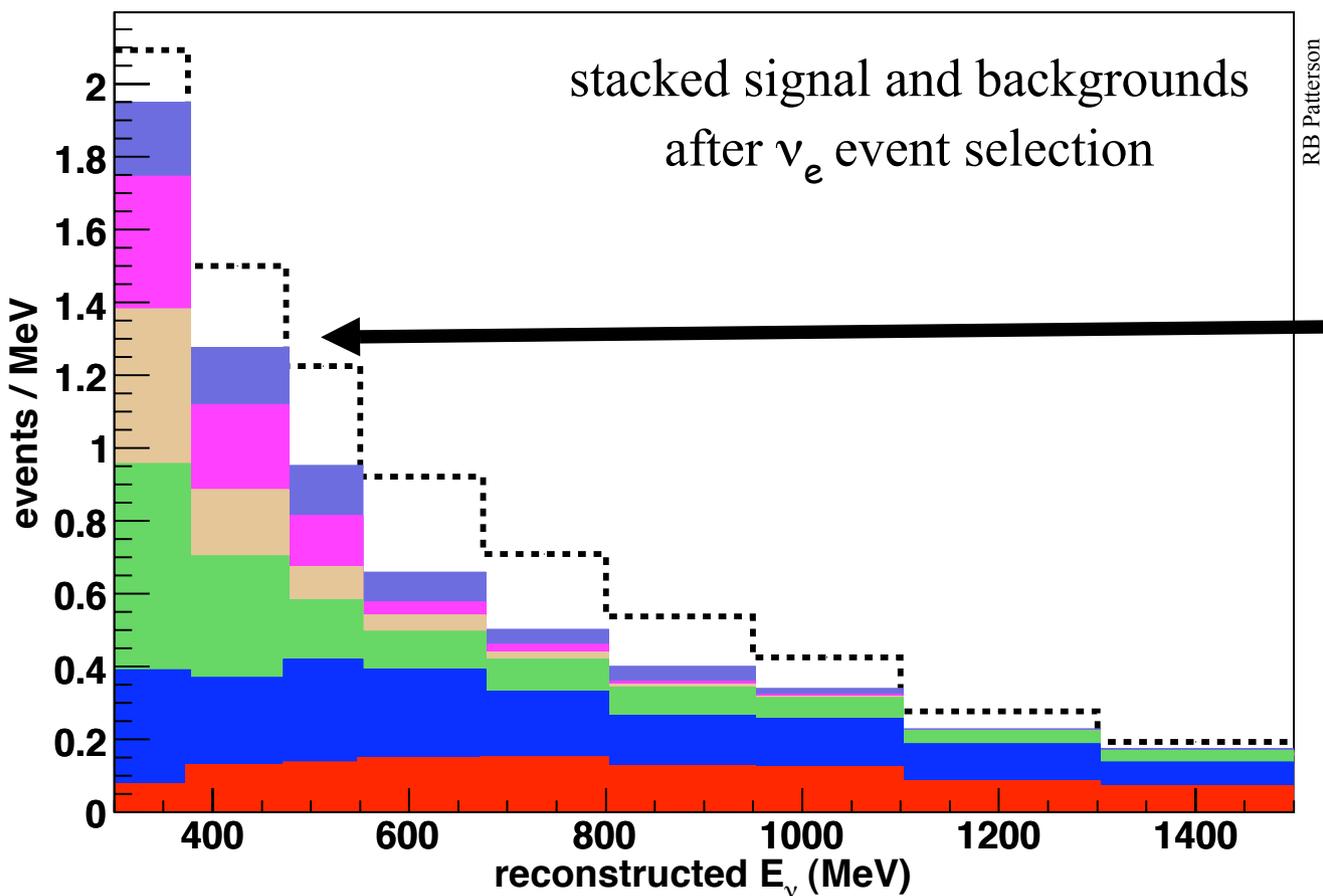
Analysis Overview: Org Chart

For robustness, MiniBooNE has performed two independent oscillation analyses.



Analysis Overview: Signal and Backgrounds

what we predict for the full ν data set ($5.6E20$ protons on target):



Oscillation ν_e

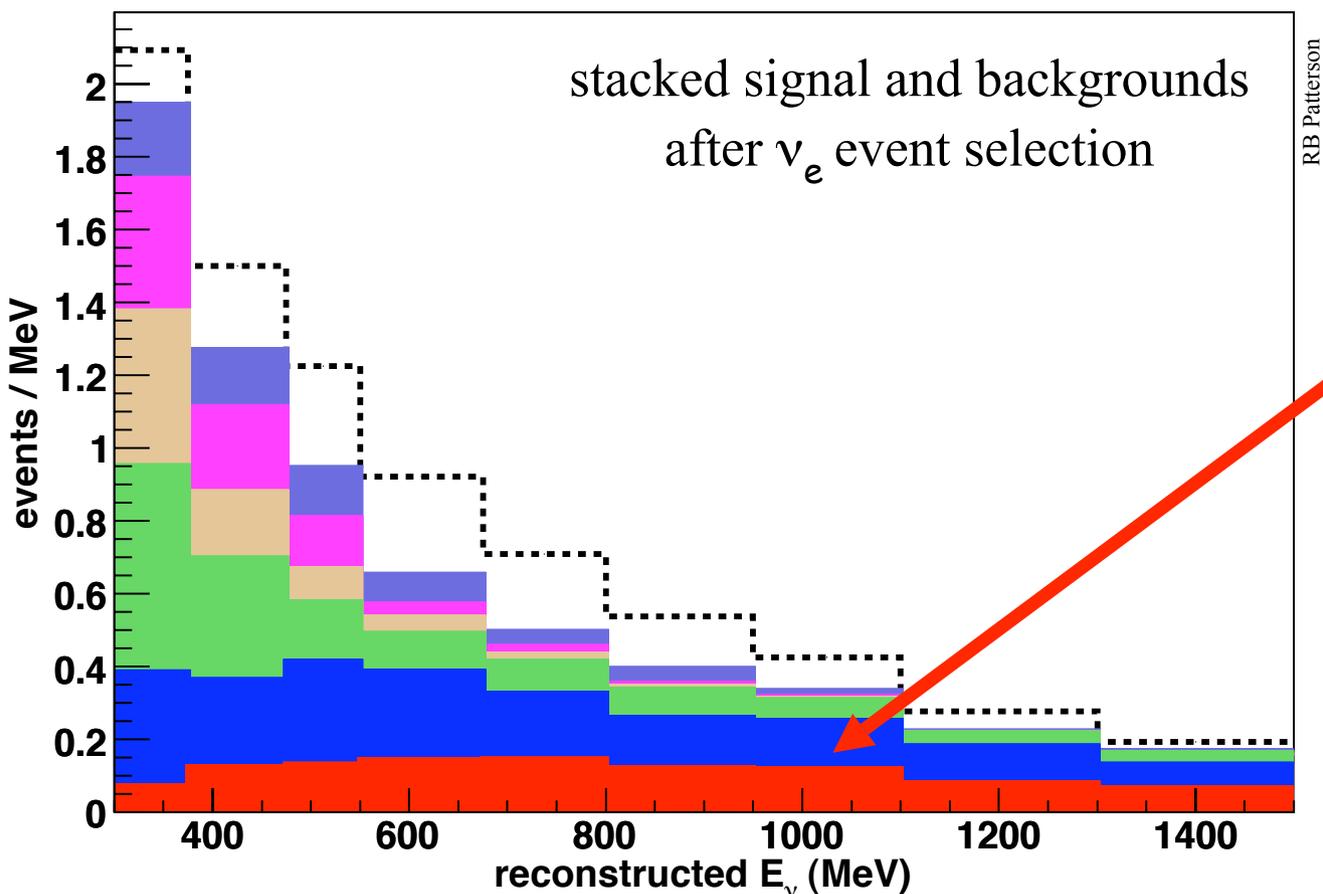
Example oscillation signal

- $\Delta m^2 = 1.2 \text{ eV}^2$
- $\text{SIN}^2 2\theta = 0.003$

Fit for excess as a function of
reconstructed ν_e energy

Analysis Overview: Signal and Backgrounds

what we predict for the full ν data set ($5.6E20$ protons on target):



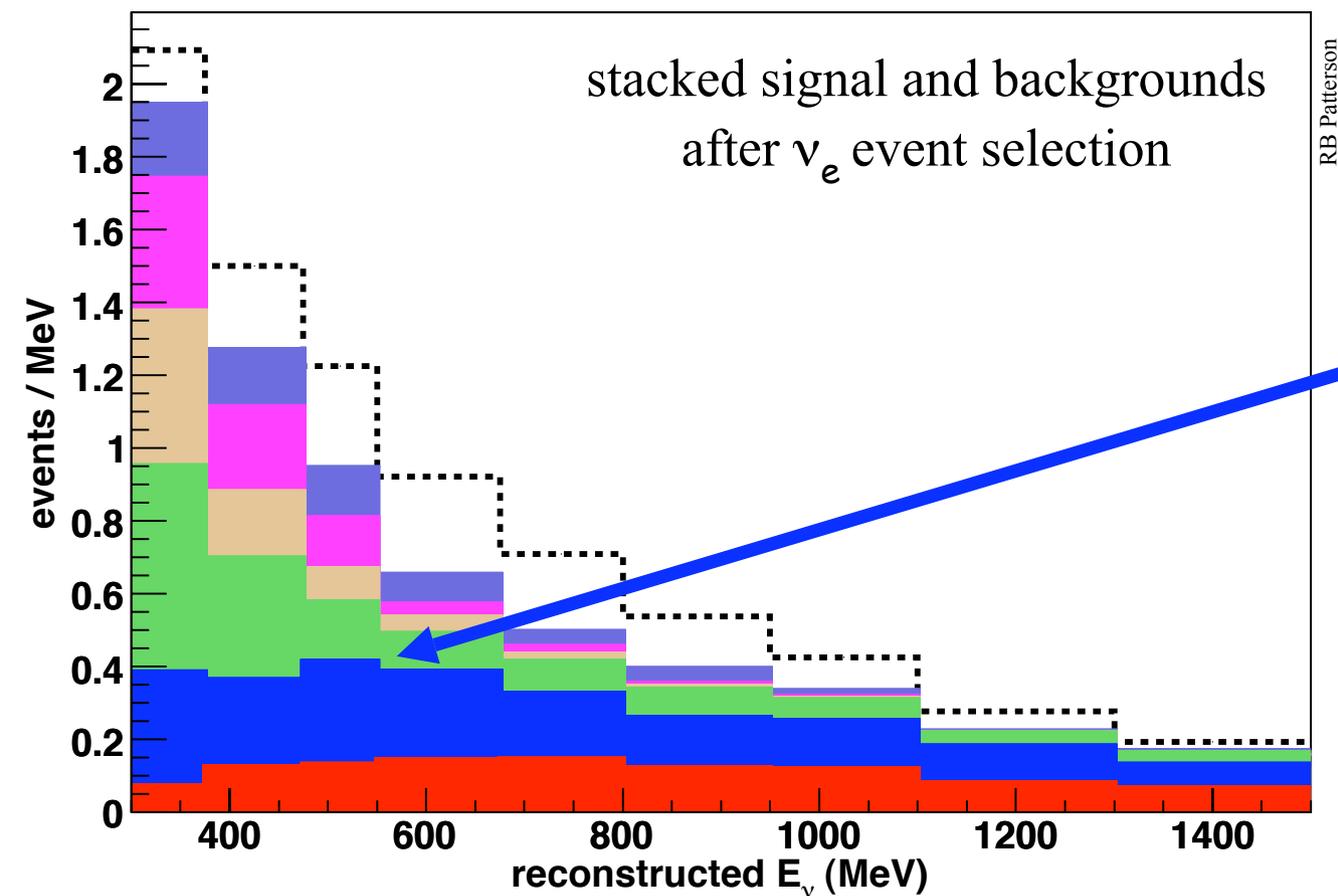
ν_e from K^+ and K^0

Use high energy ν_e and ν_μ
in-situ data for normalisation
cross-check

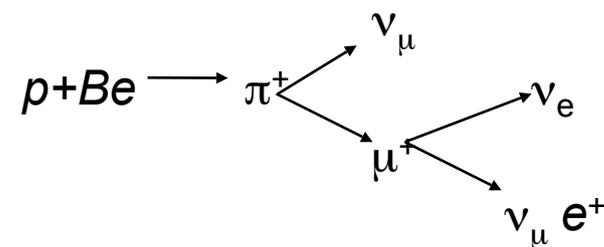
Use fit to kaon production
data for shape

Analysis Overview: Signal and Backgrounds

what we predict for the full ν data set ($5.6E20$ protons on target):



ν_e from μ^+



Measured with in-situ ν_μ
CCQE sample

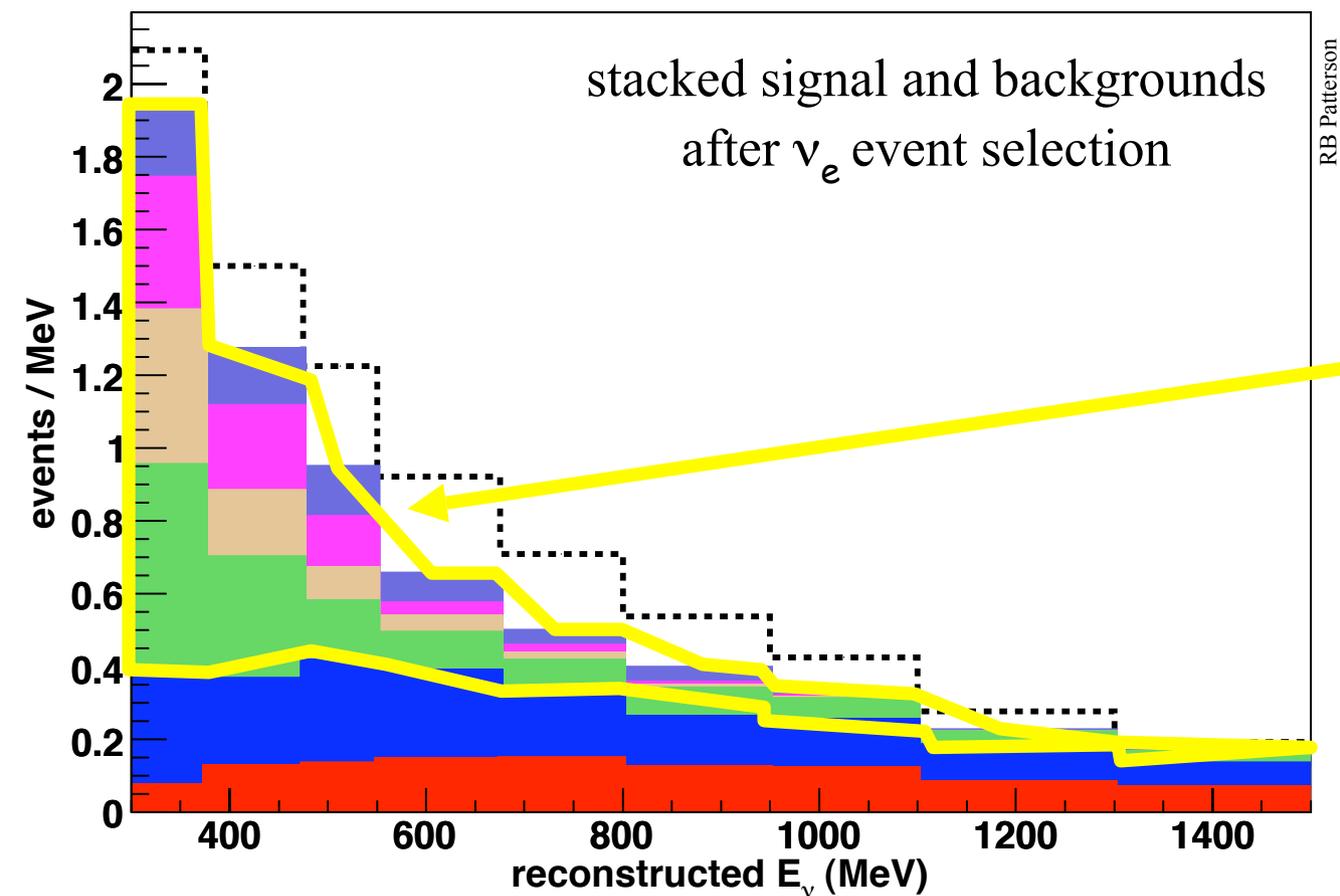
- Same ancestor π^+ kinematics

Most important background

- Constrained to a few %

Analysis Overview: Signal and Backgrounds

what we predict for the full ν data set ($5.6E20$ protons on target):



MisID ν_μ

~46% π^0

- Determined by clean π^0 measurement

~14% “dirt”

- Measure rate to normalise and use MC for shape

~16% $\Delta \gamma$ decay

- π^0 measurement constrains

~24% other

- Use ν_μ CCQE rate to normalise and MC for shape

Analysis Overview: Strategy

recurring theme: good data/MC agreement

in-situ data are incorporated wherever possible...

(i) MC tuning with calibration data

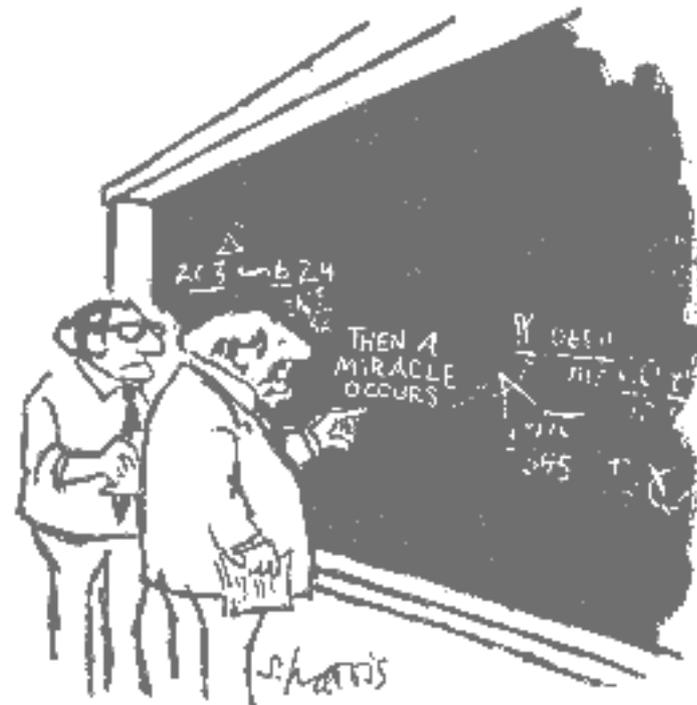
- energy scale
- PMT response
- optical model of light in the detector

(ii) MC fine-tuning with neutrino data

- cross section nuclear model parameters
- π^0 rate constraint

(iii) constraining systematic errors with neutrino data

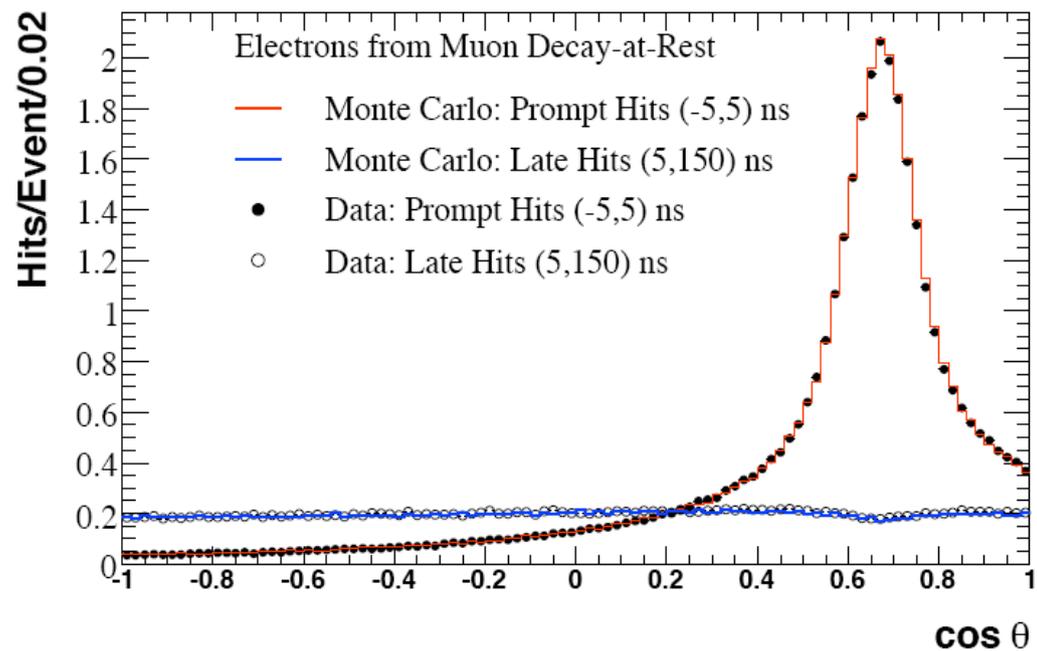
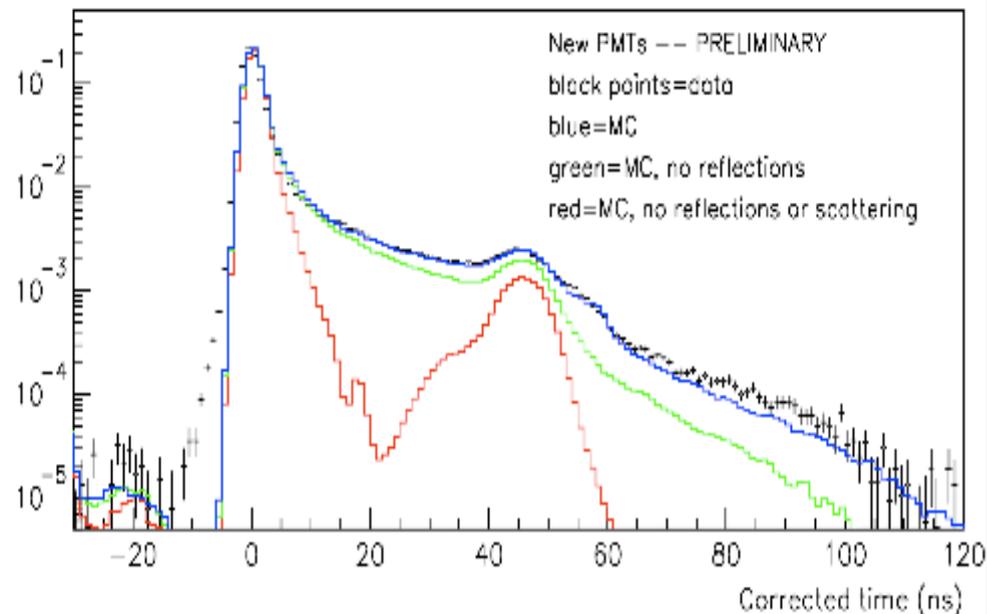
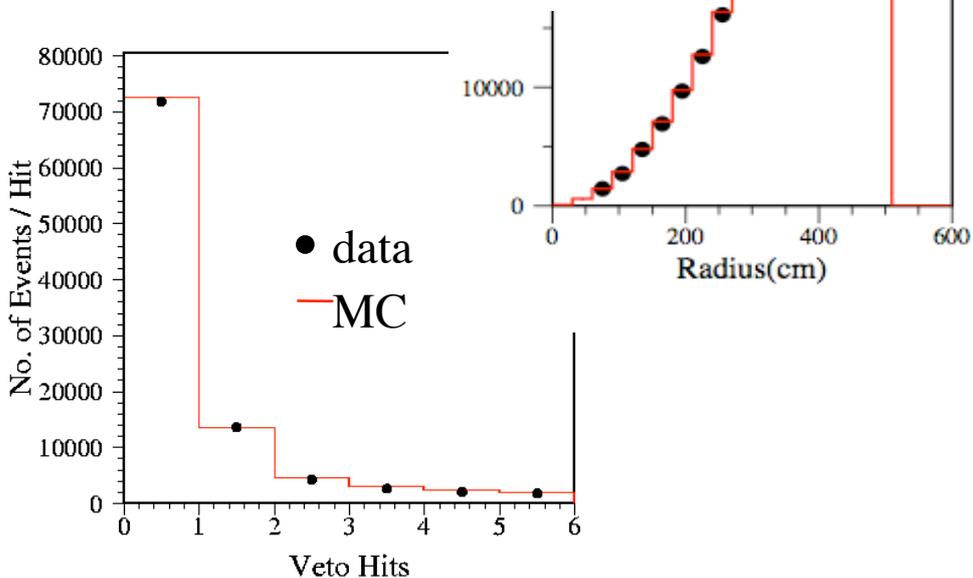
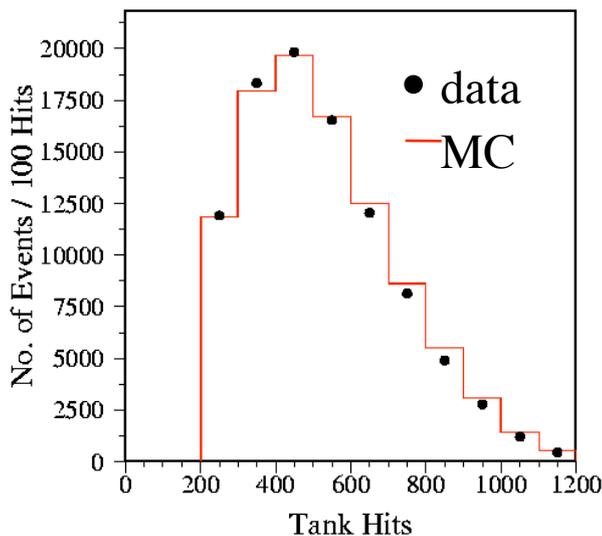
- ratio method example: ν_e from μ decay background
- combined oscillation fit to ν_μ and ν_e data



"I think you should be more explicit here in step two."

Analysis Overview: MC Tuning

MC tuning with calibration data



Analysis Overview: Strategy

in-situ data are incorporated wherever possible...

(i) MC tuning with calibration data

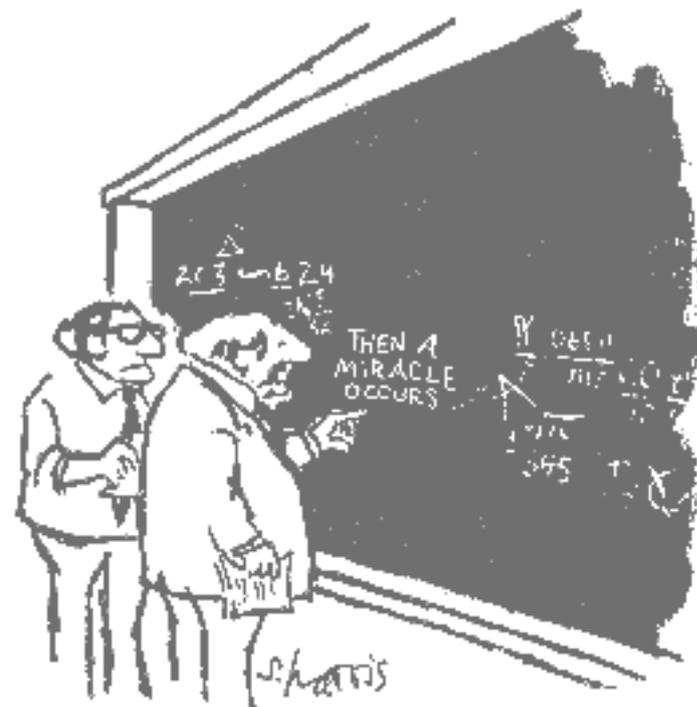
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(iii) constraining systematic errors with neutrino data

- ratio method example: ν_e from μ decay background
- combined oscillation fit to ν_μ and ν_e data

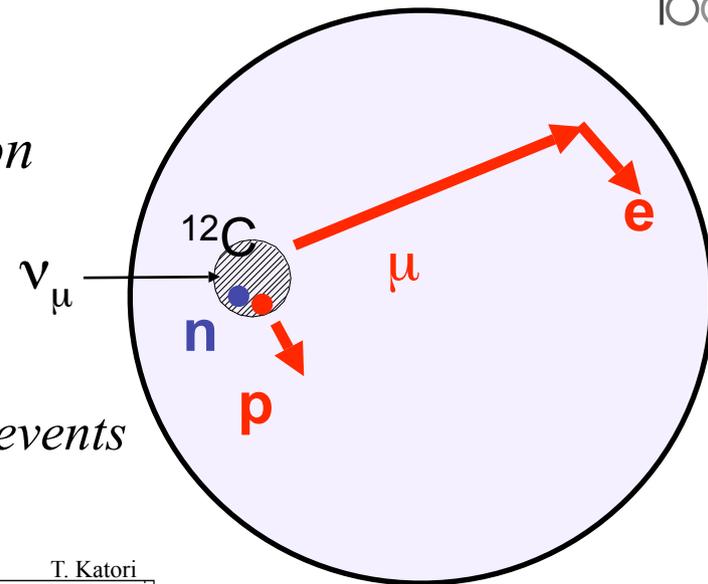


"I think you should be more explicit here in step two."

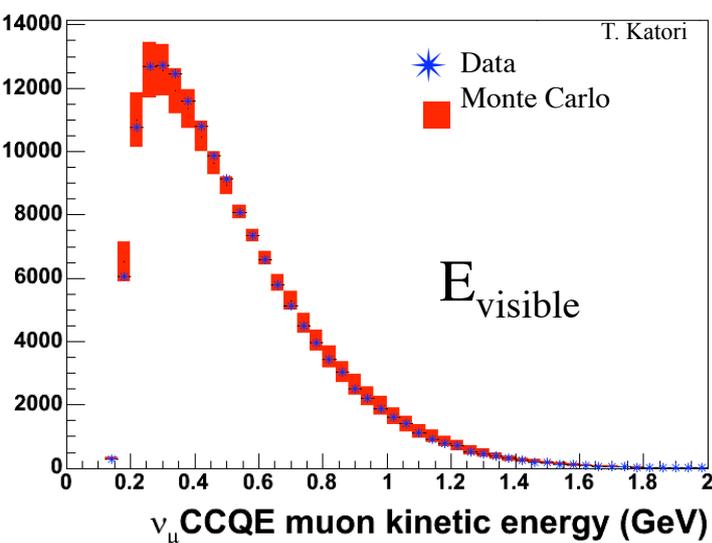
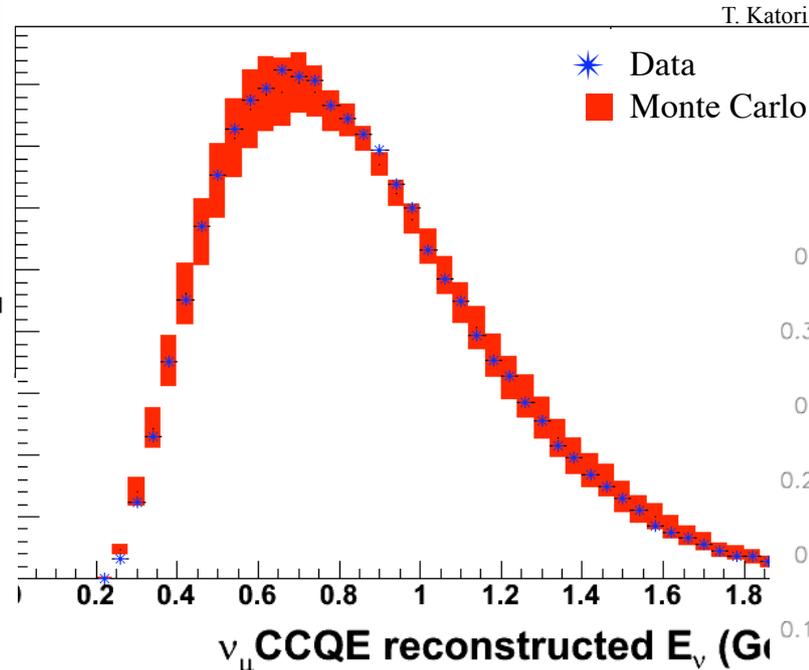
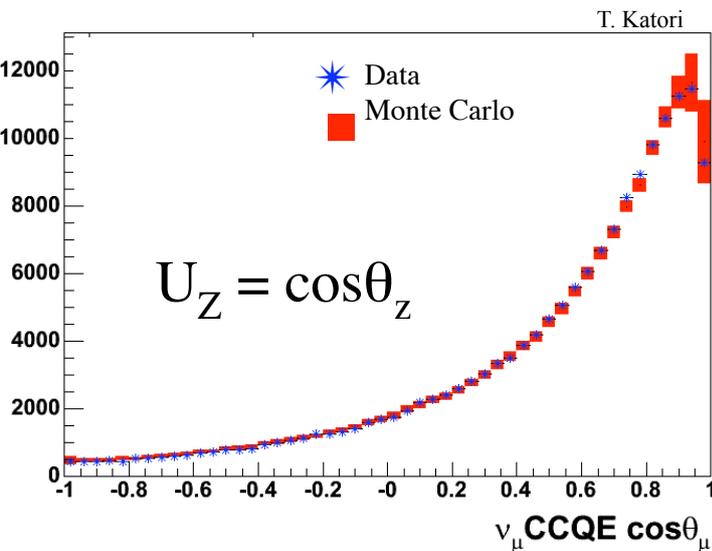
Analysis Strategy: ν_μ CCQE Events

used to measure the ν_μ flux and check E_ν^{QE} reconstruction

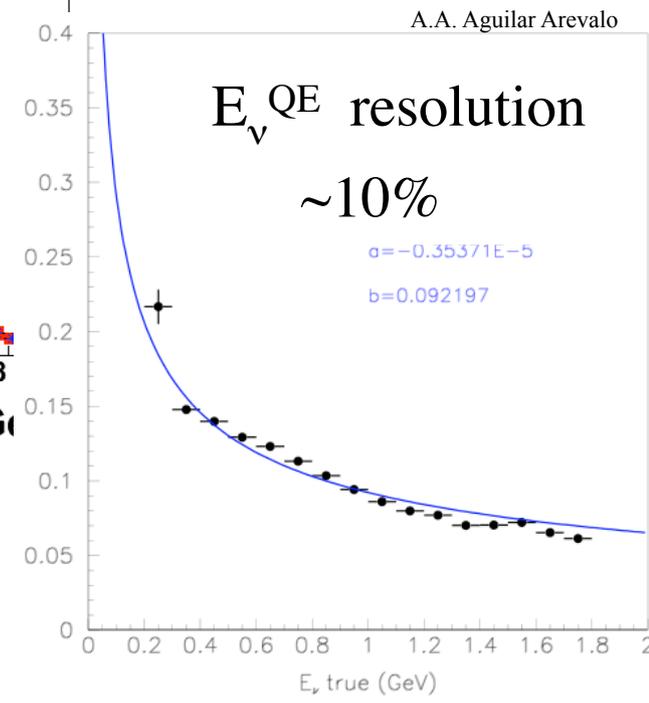
1. tag muons by requiring 2 sub-events in time
2. require reconstructed distance between sub-events $< 1m$



$\sim 74\%$ CCQE purity, $\sim 190k$ events

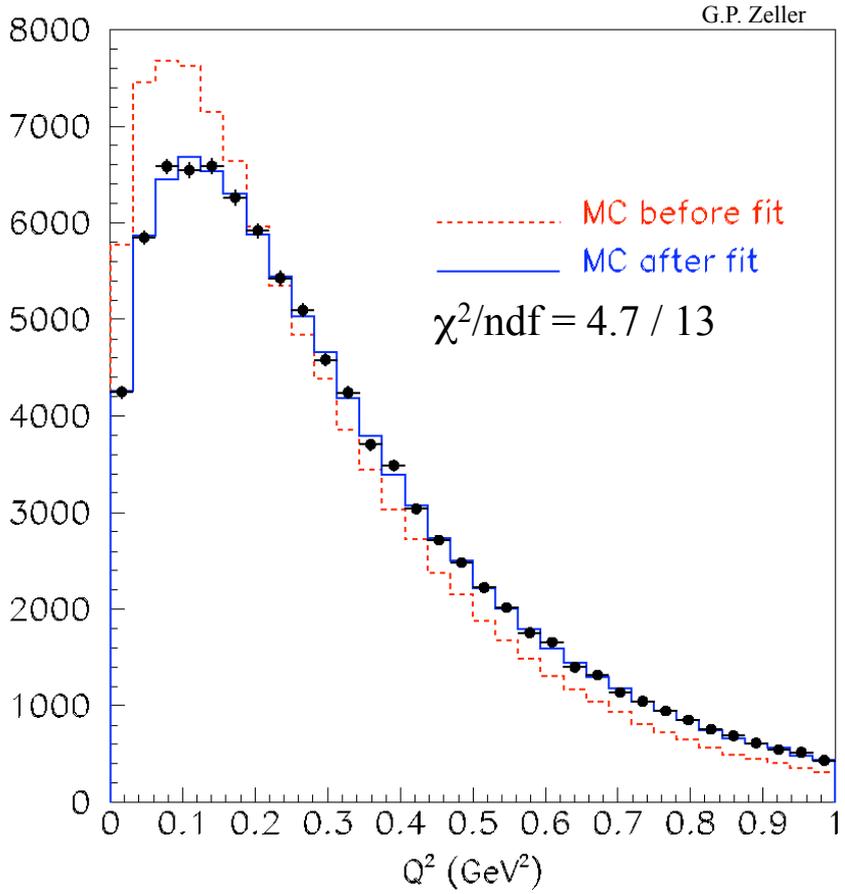


$$E_\nu^{QE} = \frac{1}{2} \frac{2M_p E_\mu - m_\mu^2}{M_p - E_\mu + \sqrt{(E_\mu^2 - m_\mu^2)} \cos\theta_\mu}$$

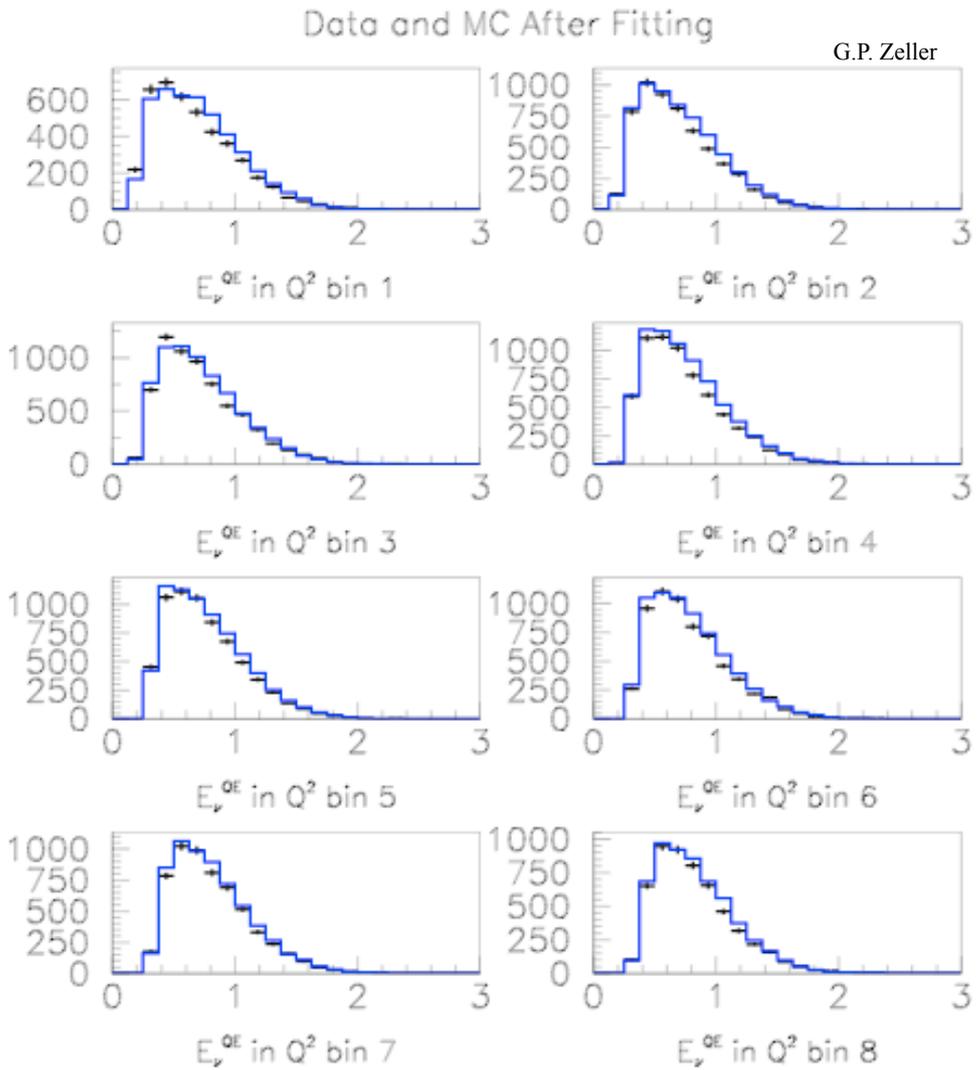


Incorporating ν_μ Data: CCQE Cross Section

The ν_μ CCQE data Q^2 distribution is fit to tune empirical parameters of the nuclear model (^{12}C target)



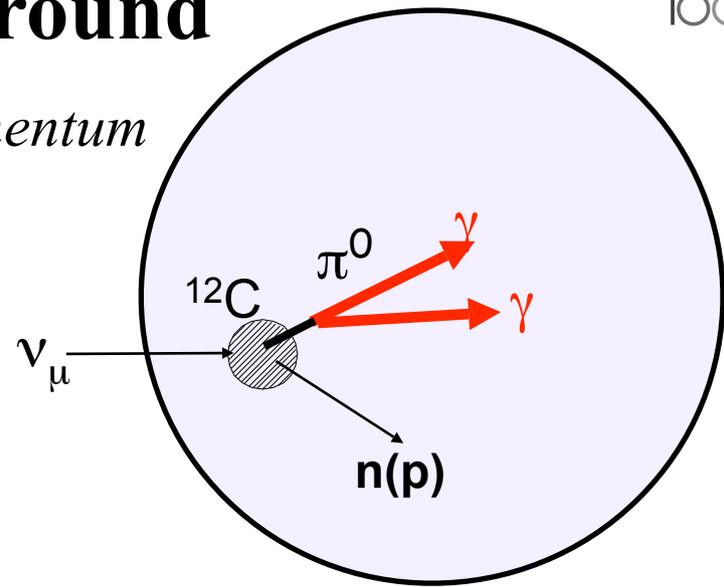
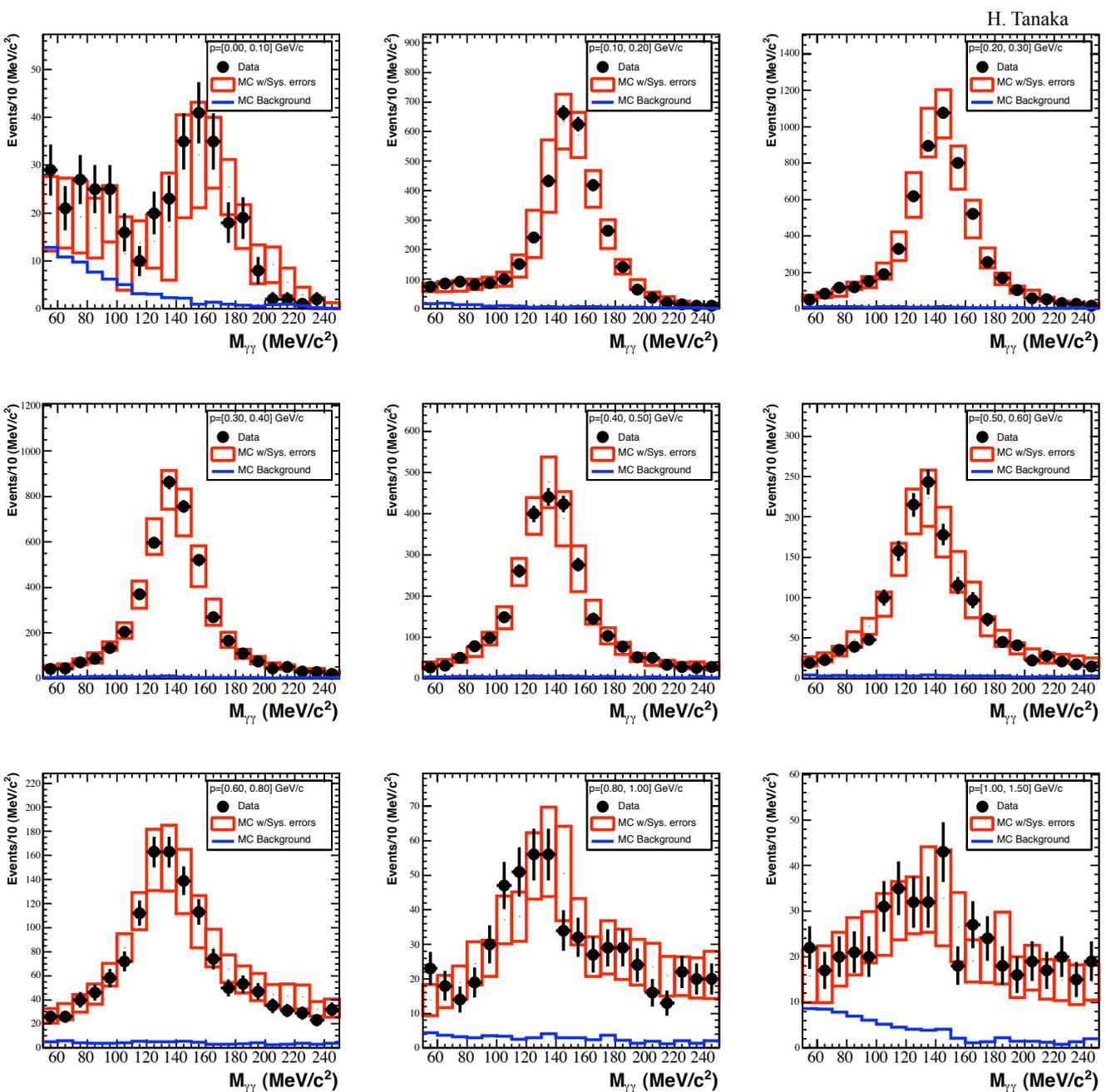
this results in good data-MC agreement for variables **not** used in tuning



the tuned model is used for both ν_μ and ν_e CCQE

Analysis Strategy: π^0 Mis-ID Background

clean π^0 events are used to tune the MC rate vs. π^0 momentum

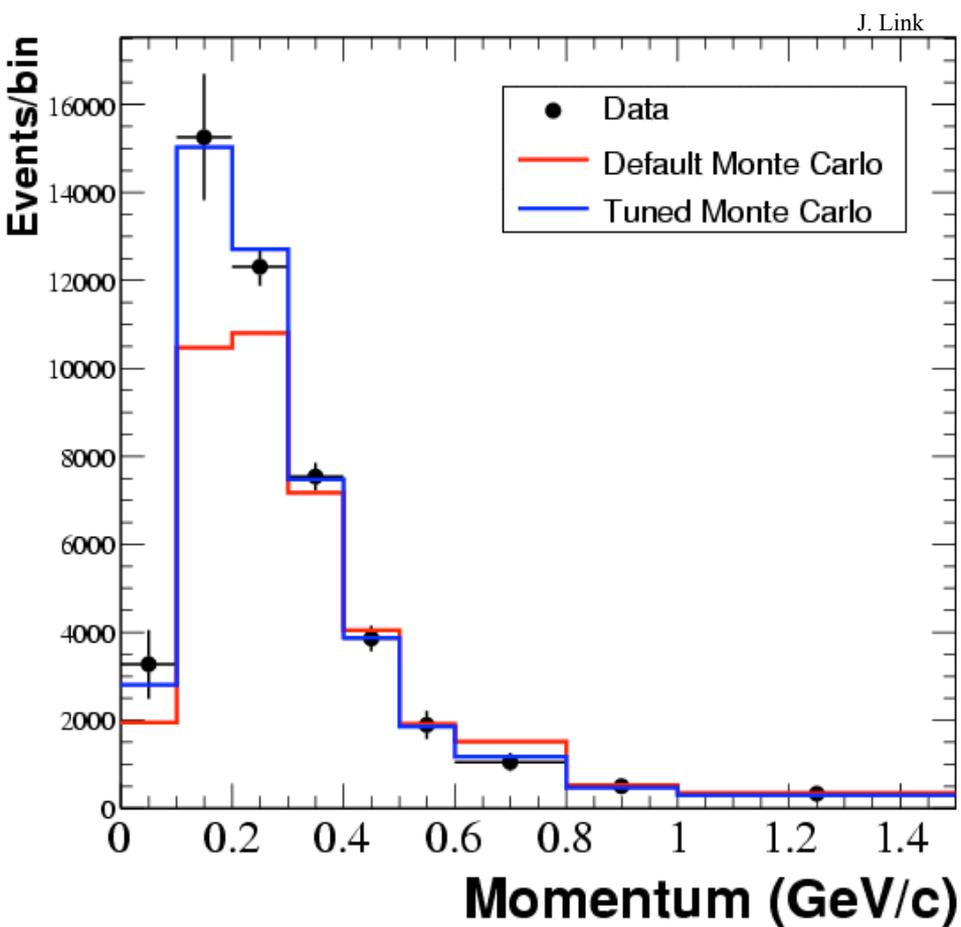


π^0 events can reconstruct outside of the mass peak when:

1. asymmetric decays fake 1 ring
2. 1 of the 2 photons exits the detector
3. high momentum π^0 decays produce overlapping rings

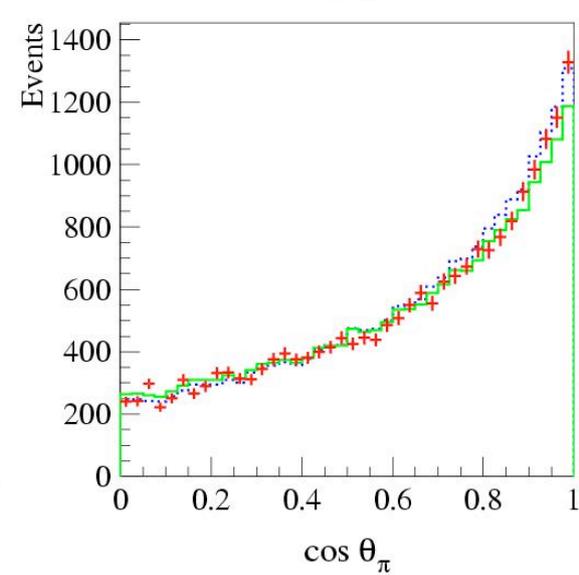
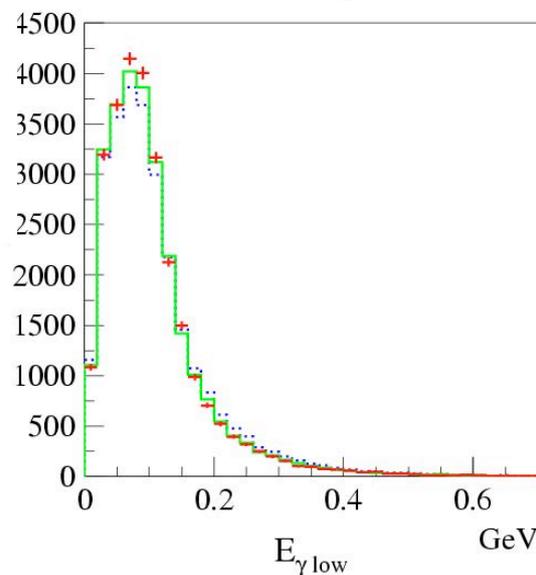
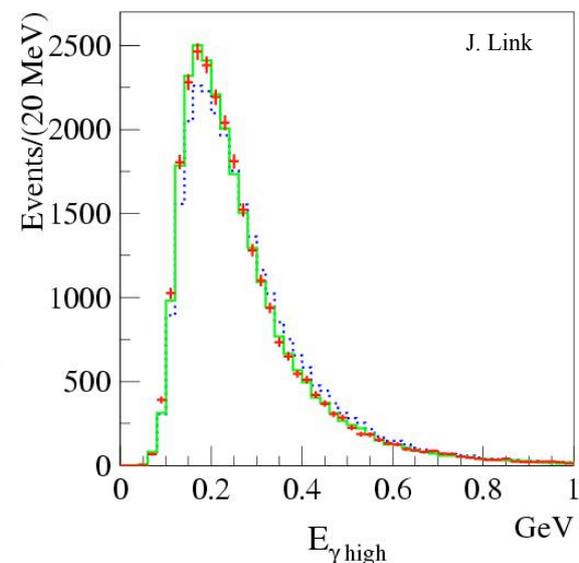
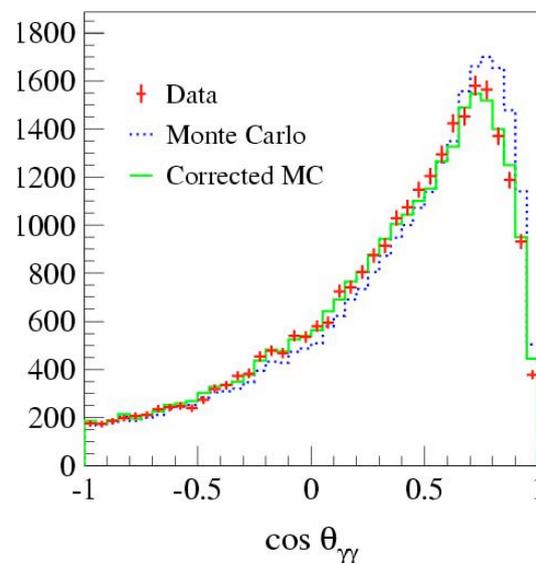
Analysis Strategy: π^0 Mis-ID Background

The MC π^0 rate (flux \times xsec) is re-weighted to match the measurement in p_π bins.



Because this constrains the Δ resonance rate, it also constrains the rate of $\Delta \rightarrow N\gamma$ in MiniBooNE

this procedure results in good data-MC agreement for variables **not** used in tuning



Analysis Overview: Strategy

in-situ data is incorporated wherever possible...

(i) MC tuning with calibration data

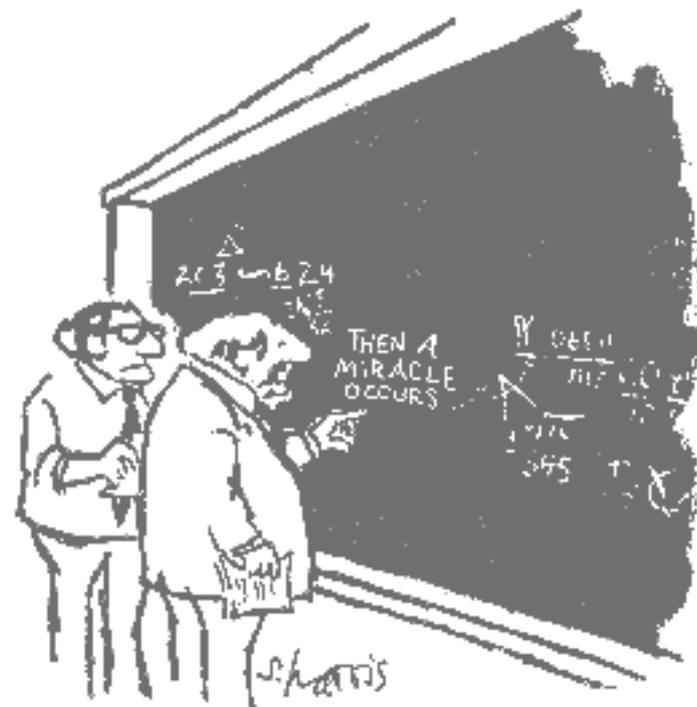
- energy scale
- PMT response
- optical model of light in the detector

(ii) MC fine-tuning with neutrino data

- cross section nuclear model parameters
- π^0 rate constraint

(iii) constraining systematic errors with neutrino data

- ratio method example: ν_e from μ decay background
- combined oscillation fit to ν_μ and ν_e data



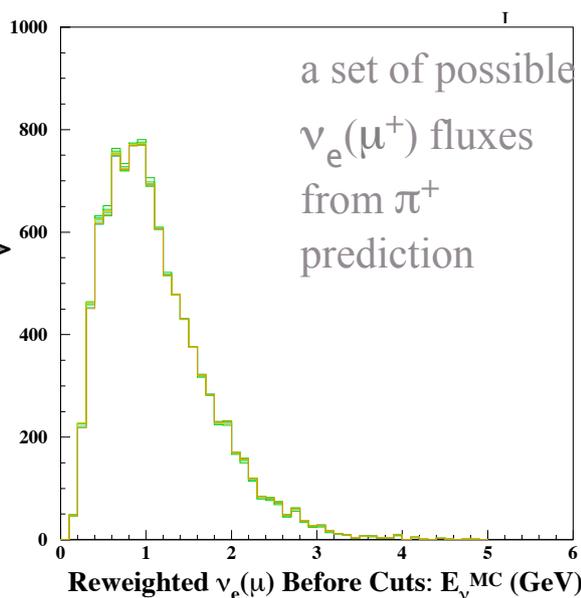
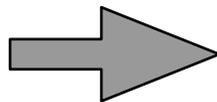
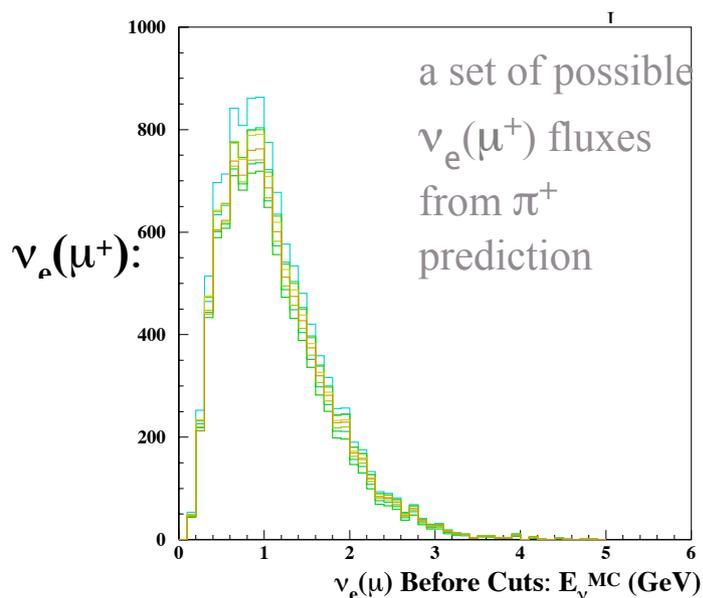
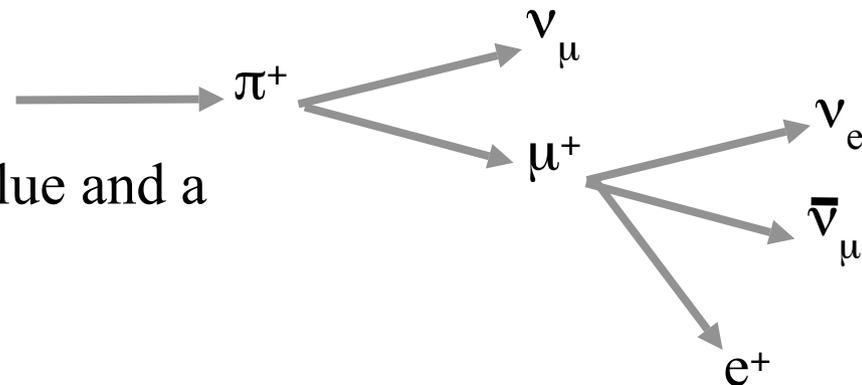
"I think you should be more explicit here in step two."

Analysis Strategy 1: Ratio Method

Example: ν_μ CCQE events measure π^+ spectrum, constrain μ^+ -decay ν_e flux

Ratio Method Constraint:

1. MC based on external data predicts a central value and a range of possible $\nu_\mu(\pi)$ fluxes
2. make Data/MC ratio vs. E_ν^{QE} for ν_μ CCQE data
3. re-weight each possible MC parent- π^+ flux by the ratio (2), including sister μ^+ & niece ν_e



reduction in the spread of possible fluxes translates directly into a reduction in the μ^+ -decay ν_e background uncertainty

Can use ratio method to constrain most BG sources

Analysis Strategy 2: Combined Fit

Fit the E_ν^{QE} distributions of ν_e and ν_μ events for oscillations, together

Raster scan in Δm^2 , and $\sin^2 2\theta_{\mu e}$ ($\sin^2 2\theta_{\mu x} = 0$),
calculate χ^2 value over ν_e **and** ν_μ bins

$$\chi^2 = \sum_{i=1}^{N_{bins}} \sum_{j=1}^{N_{bins}} (m_i - t_i) \mathcal{M}_{ij}^{-1} (m_j - t_j)$$

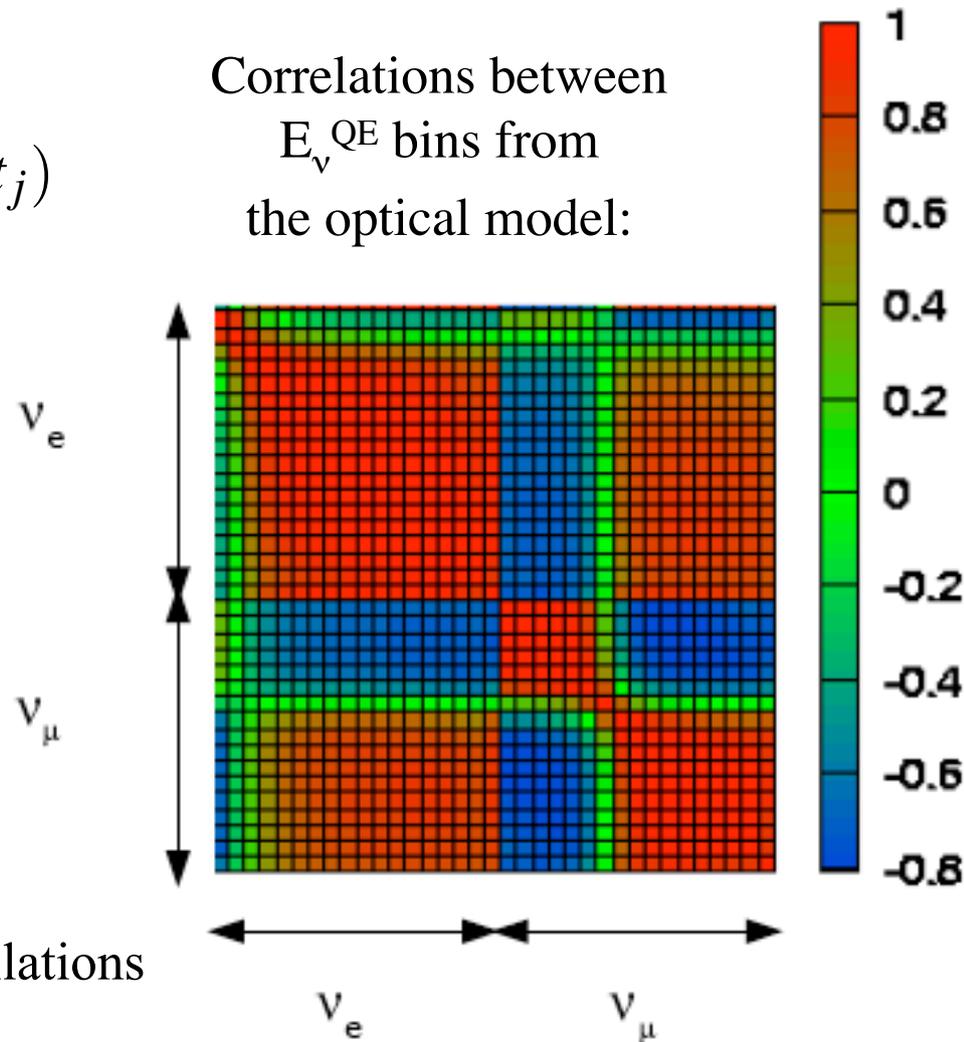
In this case, systematic error matrix \mathcal{M}_{ij} includes predicted uncertainties for ν_e **and** ν_μ bins

$$M_{ij} = \begin{pmatrix} \nu_e & \nu_e \nu_\mu \\ \nu_\mu \nu_e & \nu_\mu \end{pmatrix}$$

Left: example, m_i = "fake data" = MC with no oscillations

a combined fit constrains uncertainties in common

Correlations between E_ν^{QE} bins from the optical model:



Analysis Strategy: Error Matrix

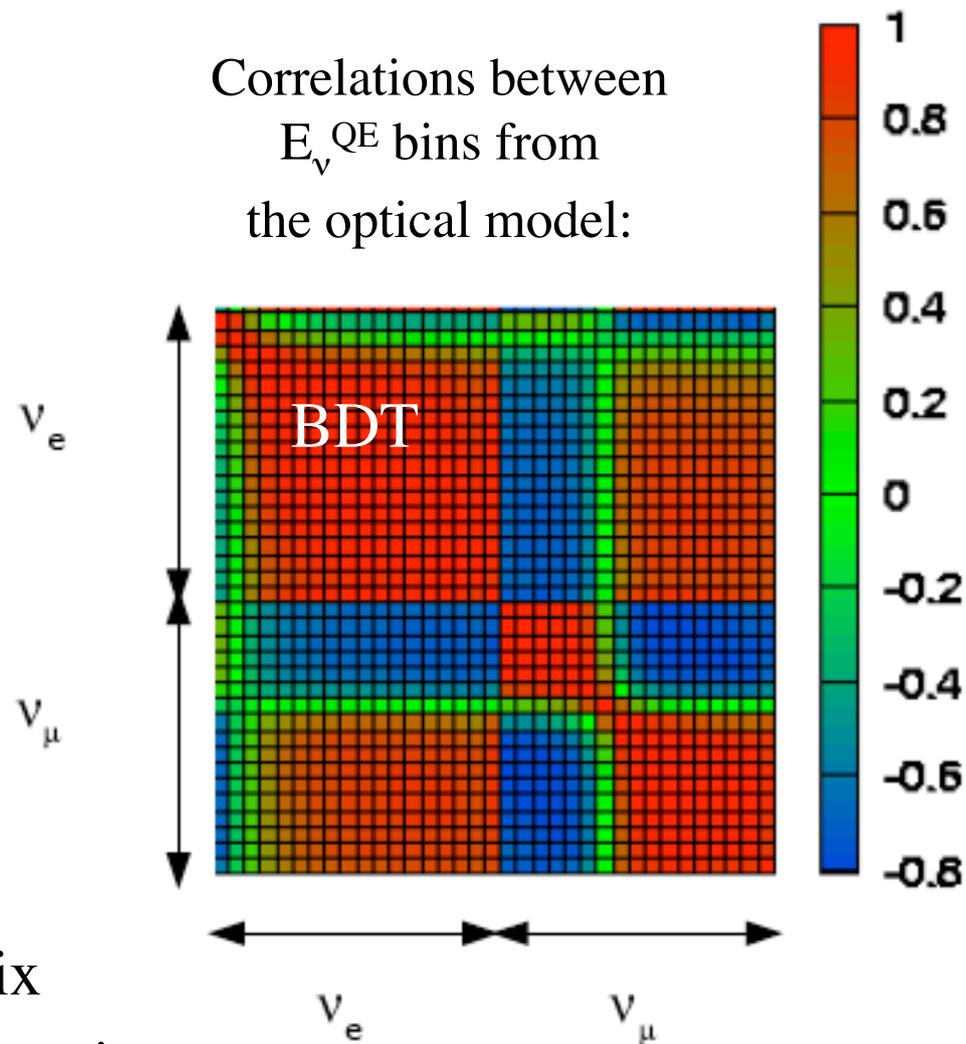
$$E_{ij} \approx \frac{1}{M} \sum_{\alpha=1}^M \left(N_i^\alpha - N_i^{MC} \right) \left(N_j^\alpha - N_j^{MC} \right)$$

- N is number of events passing cuts
- MC is standard Monte Carlo
- α represents a different MC draw
(called a “multisim”)
- M is the total number of MC draws
- i,j are E_ν^{QE} bins

Total error matrix
is sum from each source.

Primary (TB): ν_e -only total error matrix

Cross-check (BDT): ν_μ - ν_e total error matrix

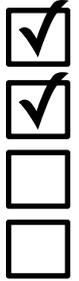


Analysis Overview: Systematic Errors

A long list of systematic uncertainties are estimated using Monte Carlo:

neutrino flux predictions

- π^+ , π^- , K^+ , K^- , K^0 , n, and p total and differential cross sections
- secondary interactions of mesons
- focusing horn current
- target + horn system alignment



neutrino interaction cross section predictions

- nuclear model
- rates and kinematics for relevant exclusive processes
- resonance width and branching fractions



detector modelling

- optical model of light propagation in oil (39 parameters!)
- PMT charge and time response
- electronics response
- neutrino interactions in dirt surrounding detector hall



✓ Most are constrained or checked using in-situ MiniBooNE data.

1. Motivation & Introduction
2. Description of the Experiment
3. Analysis Overview
4. Two Independent Oscillation Searches
 - Reconstruction and Event Selection
 - Systematic Uncertainties
5. First Results

Two Independent Oscillation Searches: Methods

Method 1: Track-Based Analysis

Primary analysis

- *Use careful reconstruction of particle tracks*
- *Identify particle type by likelihood ratio*
- *Use ratio method to constrain backgrounds*

Strengths:

Relatively insensitive to optical model

Simple cut-based approach with likelihoods

Method 2: Boosted Decision Trees

Independent cross-check

- *Classify events using “boosted decision trees”*
- *Apply cuts on output variables to improve separation of event types*
- *Use combined fit to constrain backgrounds*

Strengths:

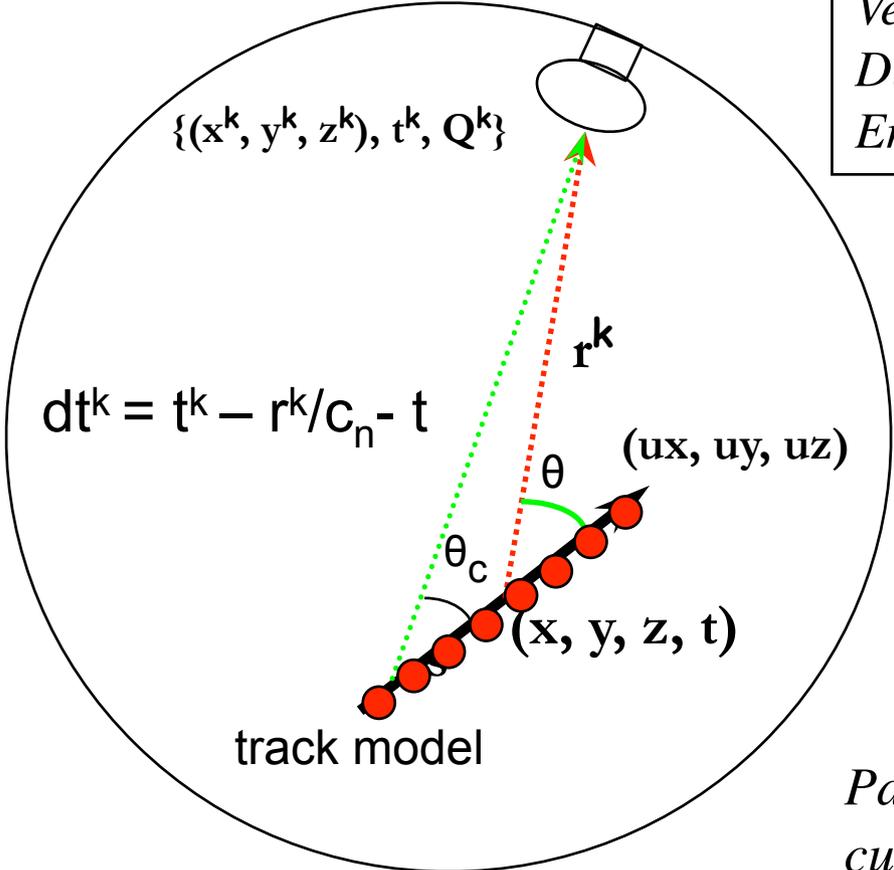
Combination of many weak variables form strong classifier

Better constraints on background events

Method 1: Track-Based Analysis

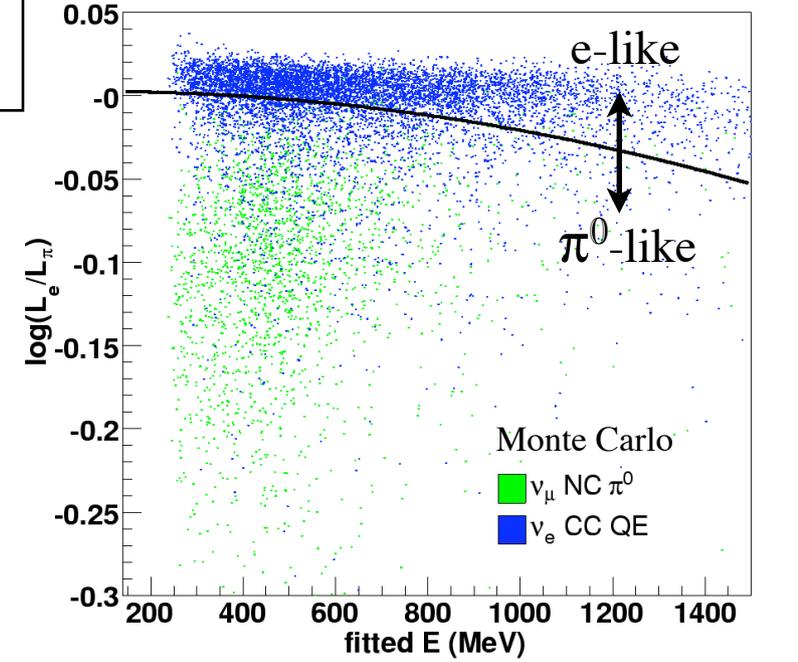
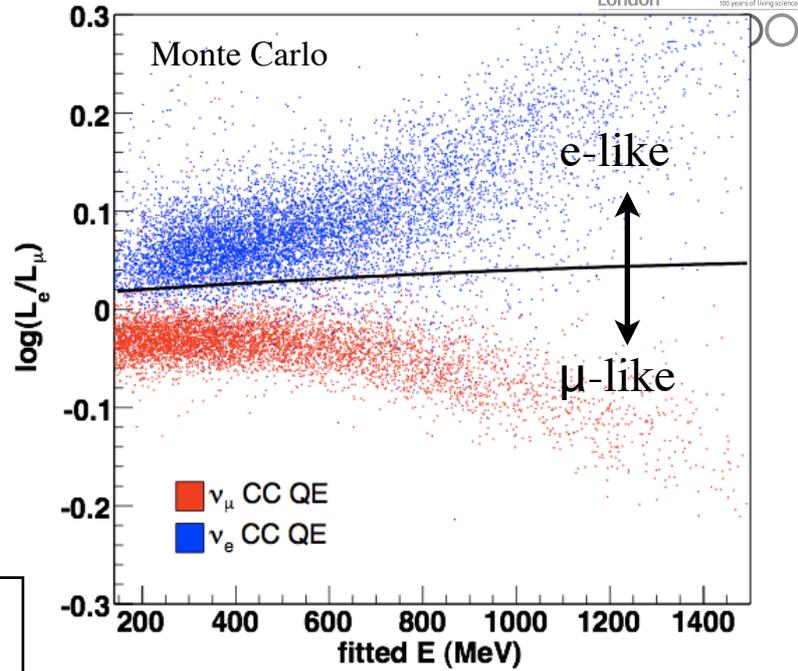
Reconstruction fits an extended light source with 7 parameters: vertex, direction (θ, ϕ), time, energy

Fit events under 3 possible hypotheses: μ -like, e-like, two track (π^0 -like)



Fitter resolution

Vertex:	22 cm
Direction:	2.8°
Energy:	11%



Particle ID relies on likelihood ratio cuts to select ν_e , cuts chosen to maximise sensitivity to $\nu_\mu \rightarrow \nu_e$ oscillation

Track-Based Analysis: e/μ Likelihood

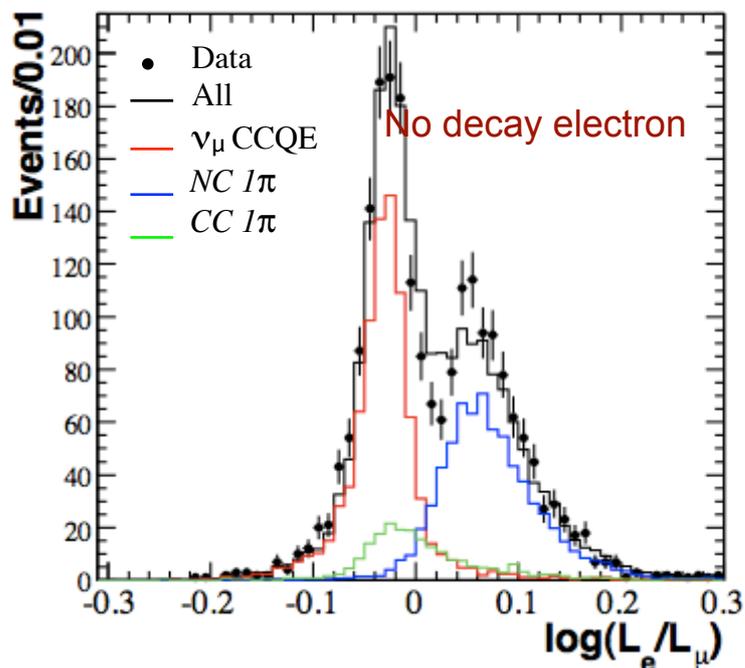
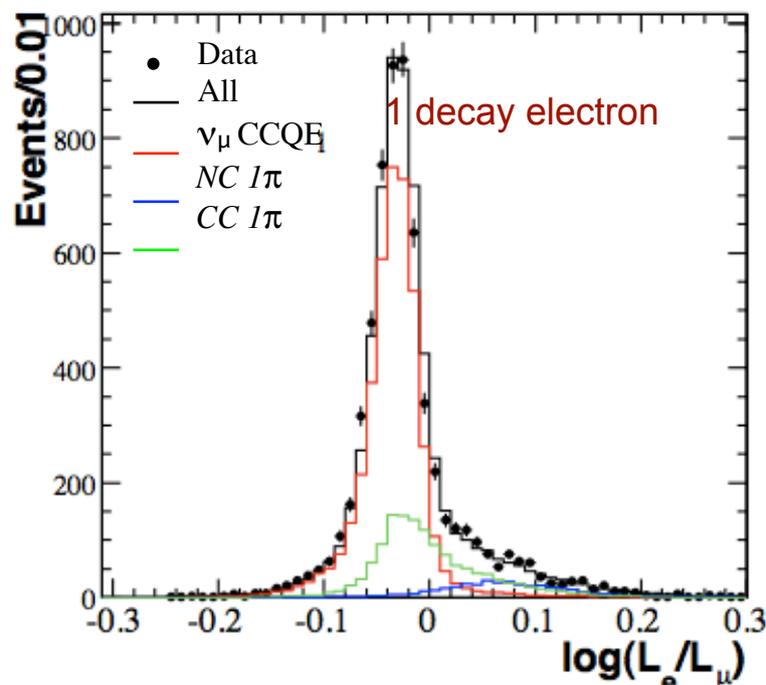
Test μ - e separation on data:

ν_μ CCQE data sample

Pre-selection cuts

Fiducial volume: ($R < 500$ cm)

2 subevents: muon + decay electron



“All-but-signal” data sample

Pre-selection cuts

Fiducial volume: ($R < 500$ cm)

1 subevent: 8% of muons capture on ^{12}C

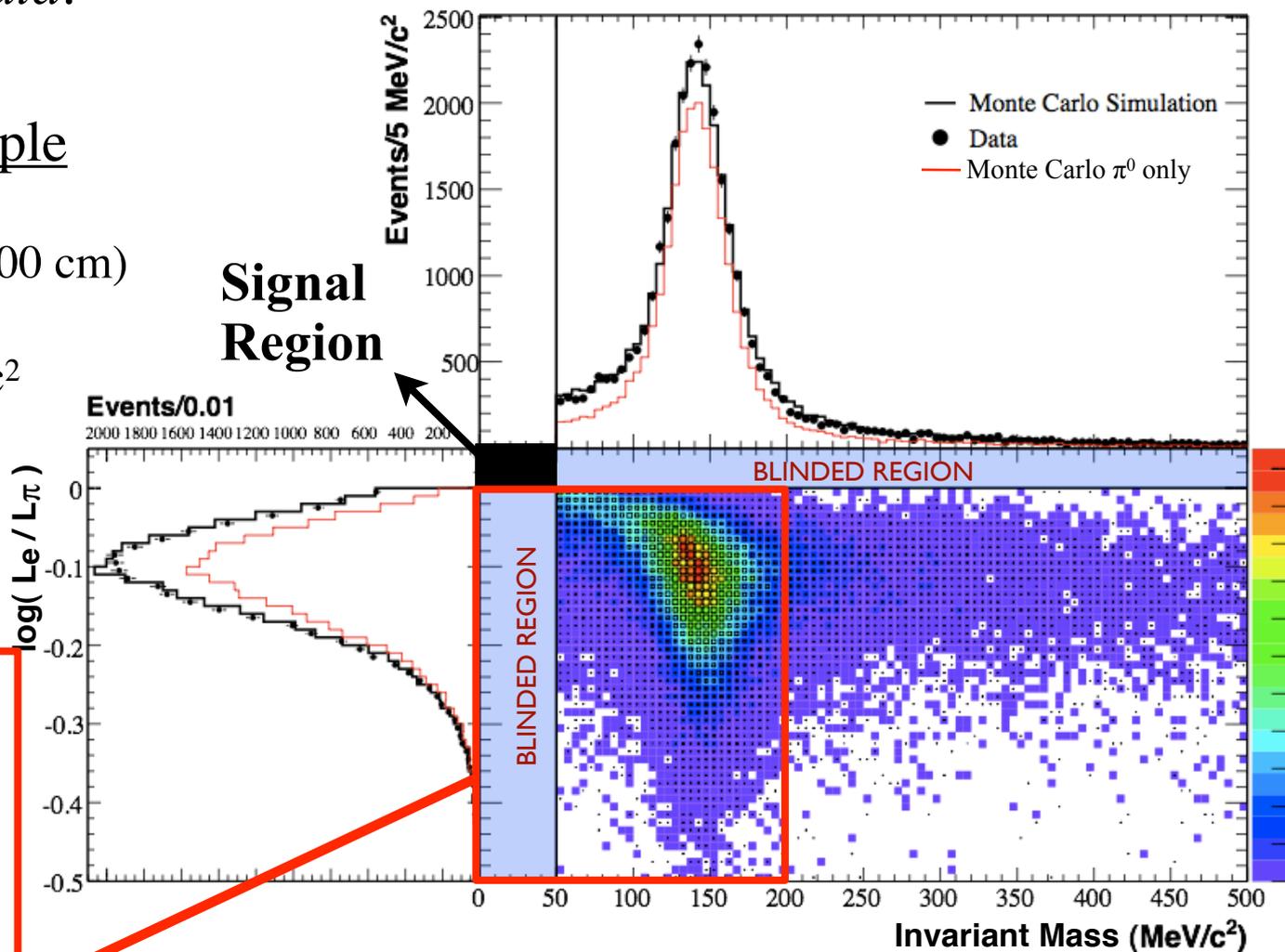
Events with $\log(L_e/L_\mu) > 0$ (e -like) undergo additional fit with two-track hypothesis.

Track-Based Analysis: e/π^0 Likelihood

Test $e-\pi^0$ separation on data:

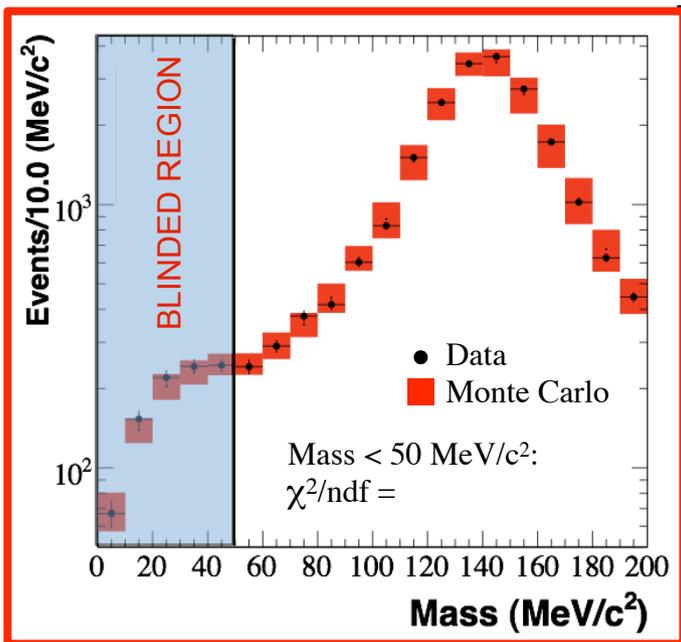
“All-but-signal” data sample

- Pre-selection cuts
- Fiducial volume cut ($R < 500$ cm)
- 1 subevent
- Invariant mass > 50 MeV/c^2
- $\log(L_e / L_\pi) < 0$ (π -like)



Tighter selection cuts:

- Invariant mass < 200 MeV/c^2
- $\log(L_e / L_\mu) > 0$ (e -like)
- $\log(L_e / L_\pi) < 0$ (π -like)



Method 2: Boosted Decision Trees

Decision Trees: A machine-learning technique which tries to recover signal events that would be eliminated in cut-based analyses.

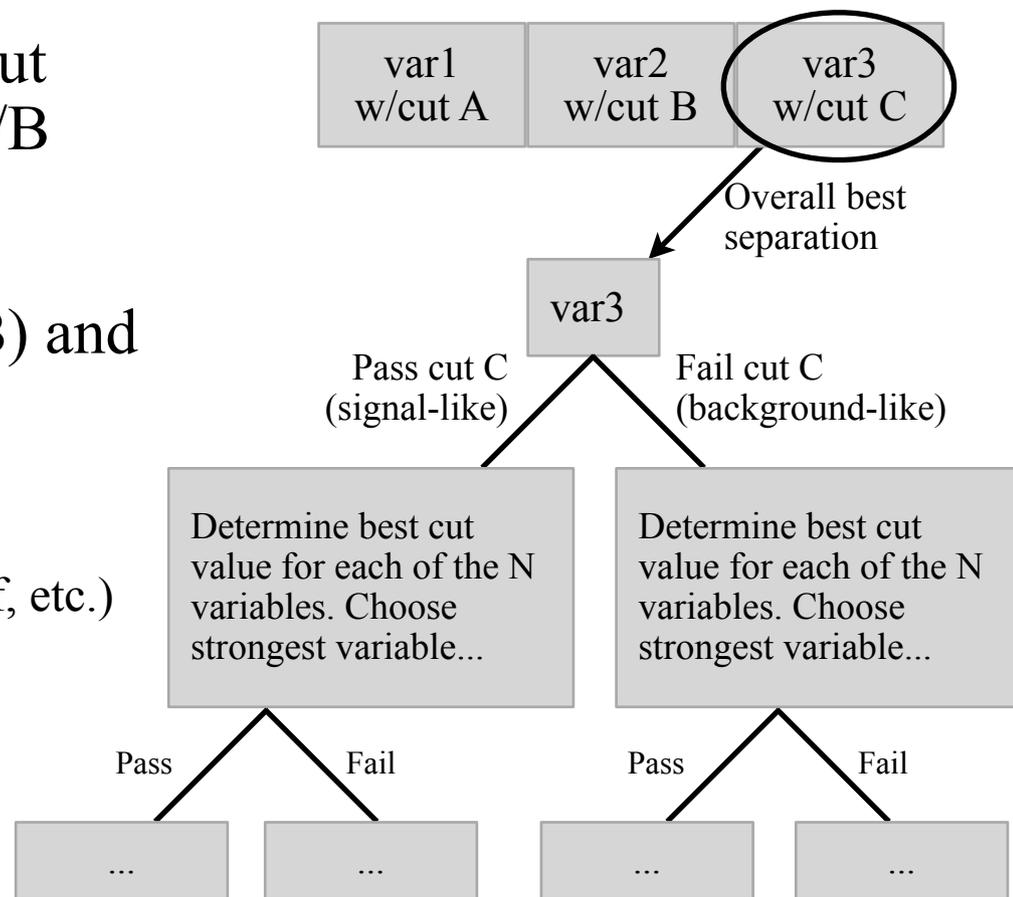
Training a decision tree:

For a set of N variables, determine the cut value for each variable that gives best S/B separation.

Cut on the best variable (i.e. highest S/B) and repeat.

Final score: For each leaf,

- 1 for correct events (signal event on a signal leaf, etc.)
- +1 for incorrect events



Boosting: Increase weight of misclassified events.

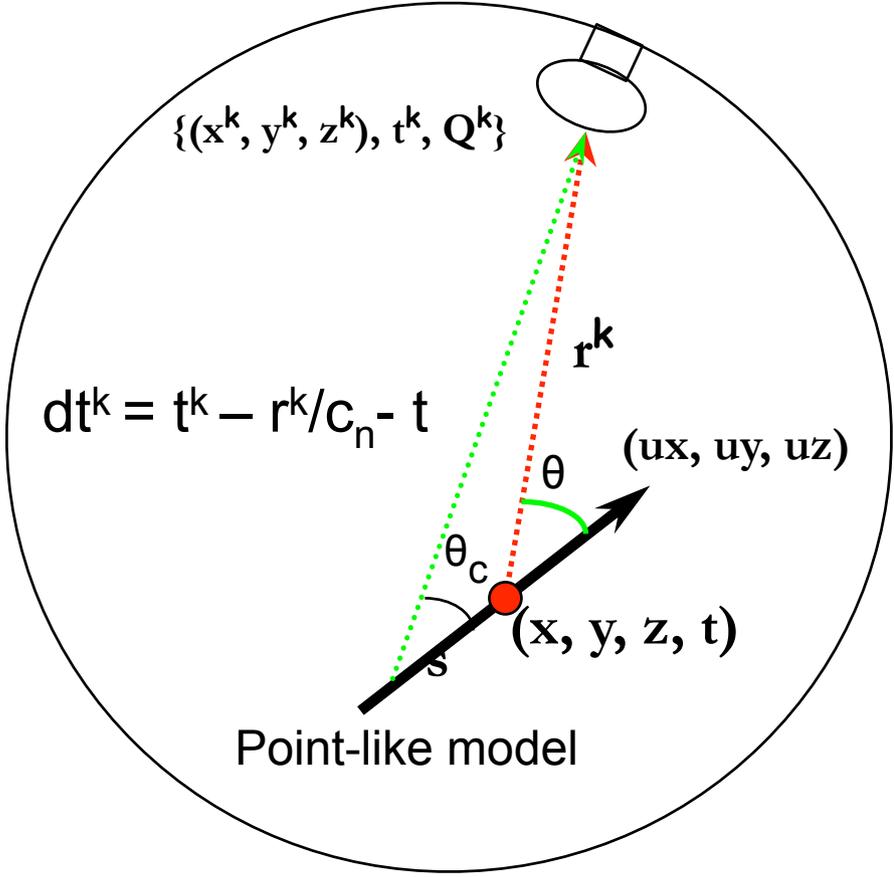
Re-training with newly weighted events improves performance.

Boosted Decision Trees: Reconstruction and Particle ID

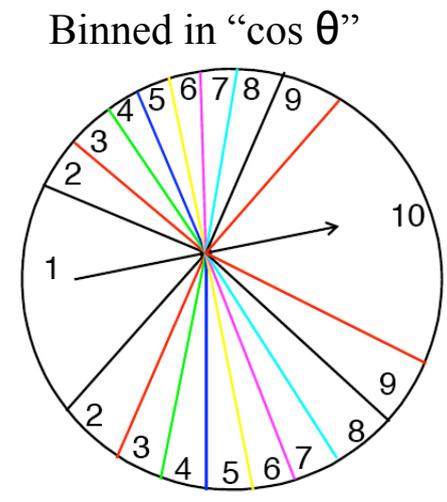
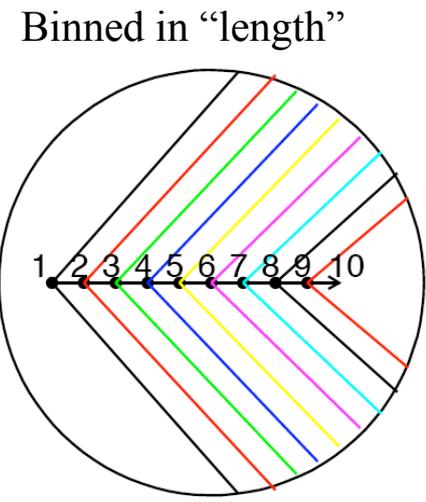
Reconstruction fits a point-like light source:

vertex, direction (θ, ϕ) , time, energy

Fitter resolution	
Vertex:	24 cm
Direction:	3.8°
Energy:	14%



Characterize topology of each event by dividing detector into “bins” relative to track:



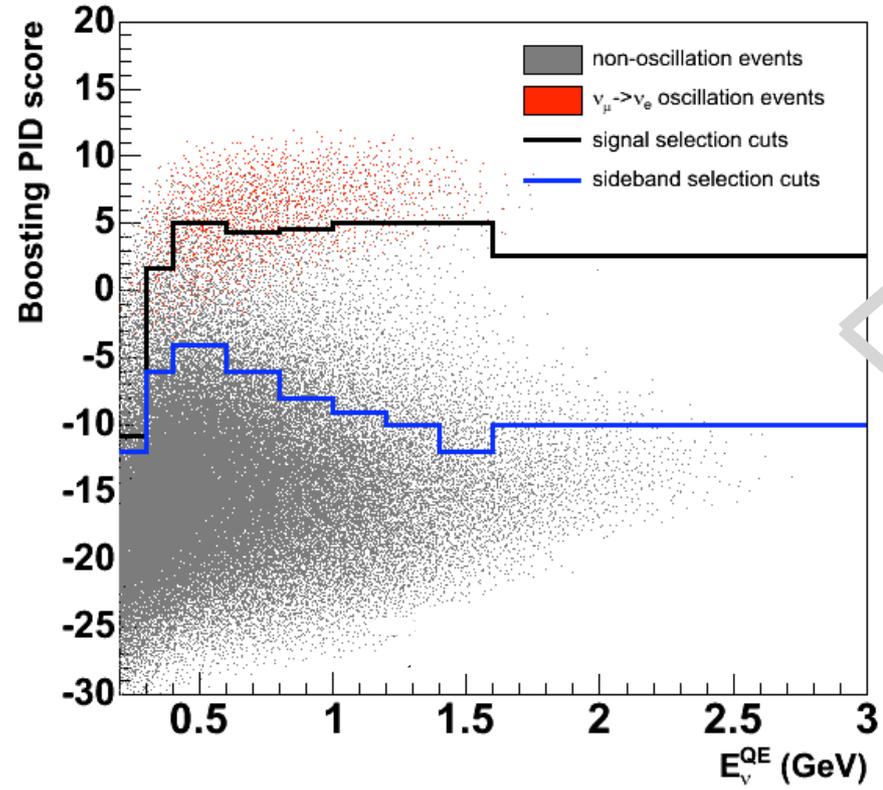
Particle ID “input variables” for the boosted decision trees are created from basic quantities in each bin: *e.g.*, charge, number of hits...

To select events, a particle ID cut is made on the Boosting output score.

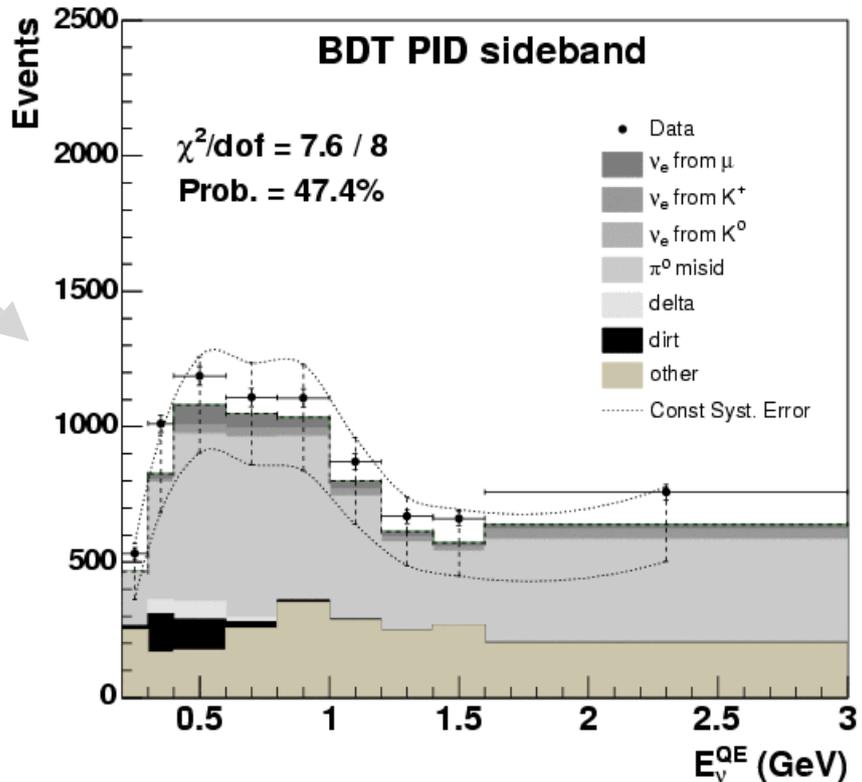
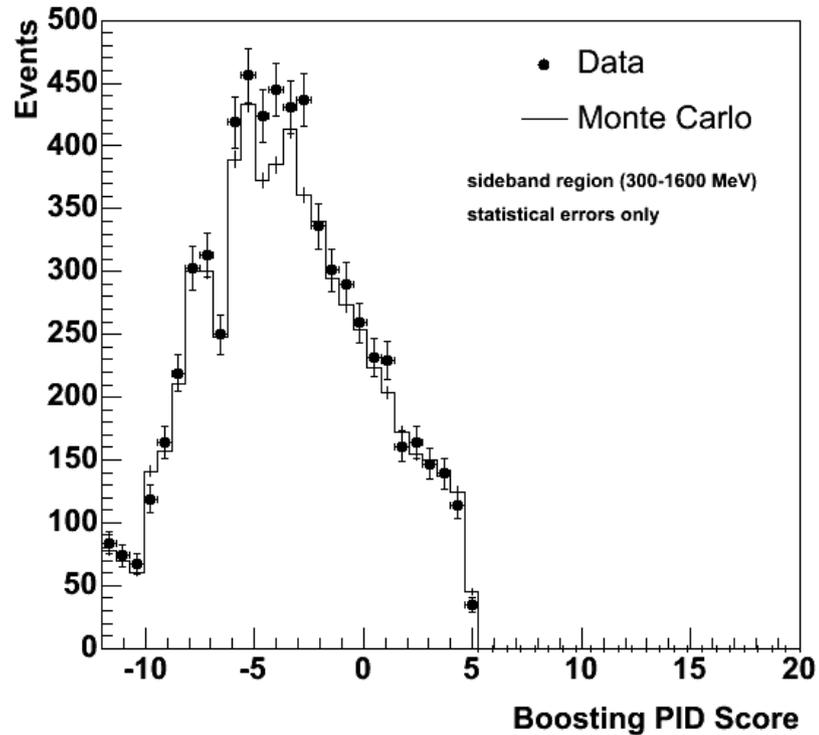
Boosted Decision Trees: Particle ID

A sideband region is selected to validate MC in region near signal.

Sideband contains mostly mis-identified π^0 background events.



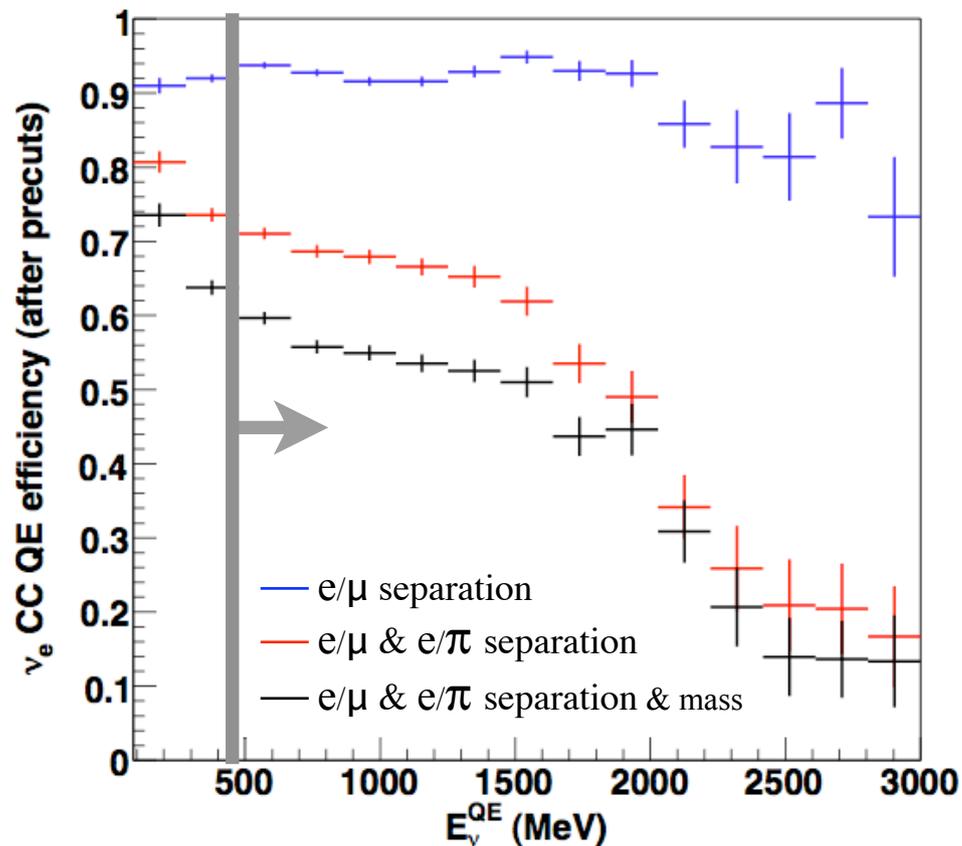
A χ^2 is calculated using the full systematic error matrix, data and MC are consistent.



Comparison: Efficiencies

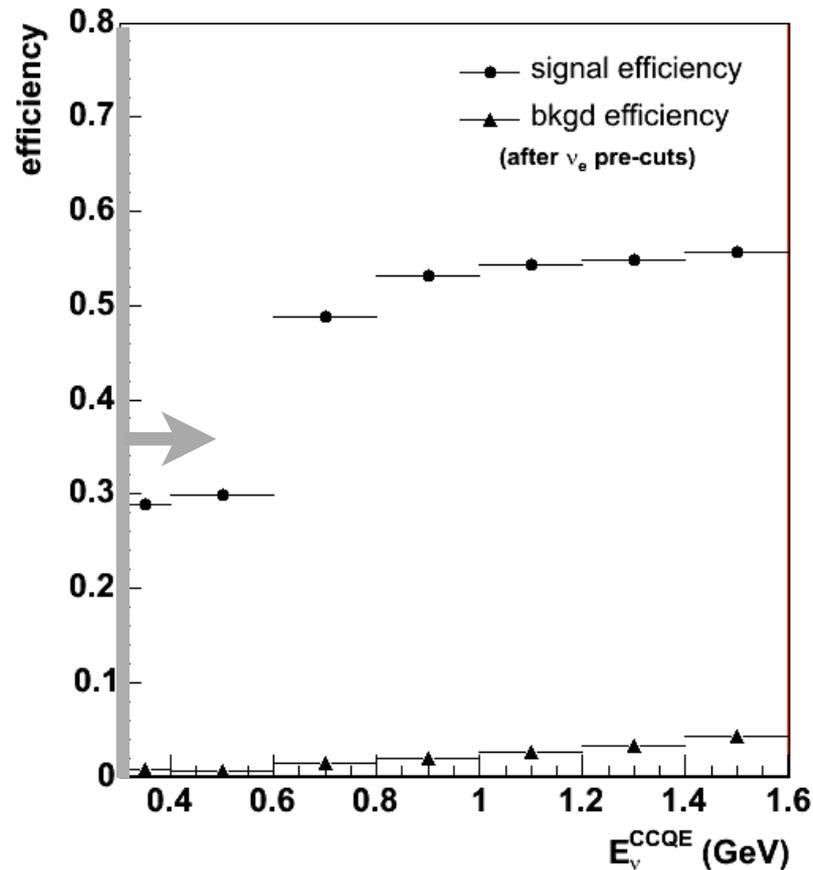
The two analyses have different event selection efficiency vs. energy trends,

Track-Based Analysis



$$E_{\nu}^{QE} > 475 \text{ MeV}$$

Boosting Analysis



$$E_{\nu}^{QE} > 300 \text{ MeV}$$

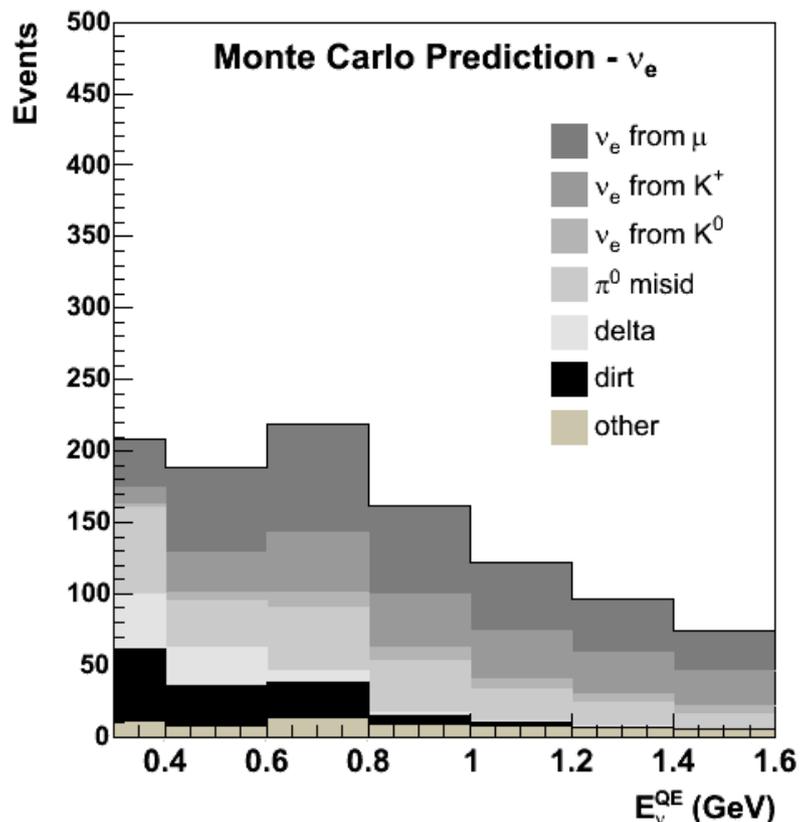
and different reconstructed E_{ν} regions for the oscillation analyses.

Comparison: Backgrounds

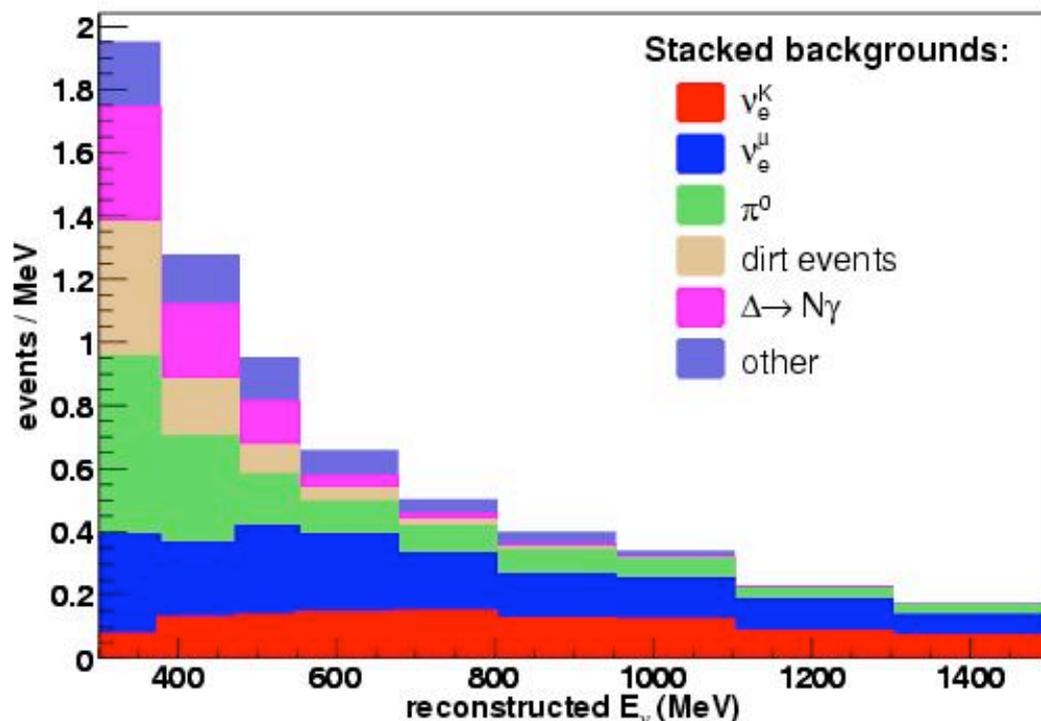
The two analyses have somewhat different background compositions.

Source	T-B	B
ν_e from μ decay	0.37	0.32
ν_e from K decay	0.26	0.24
π^0 mis-ID	0.17	0.21
$\Delta \rightarrow N\gamma$	0.06	0.07
Dirt	0.05	0.11
Other	0.09	0.05

Boosting Analysis



Track-Based Analysis



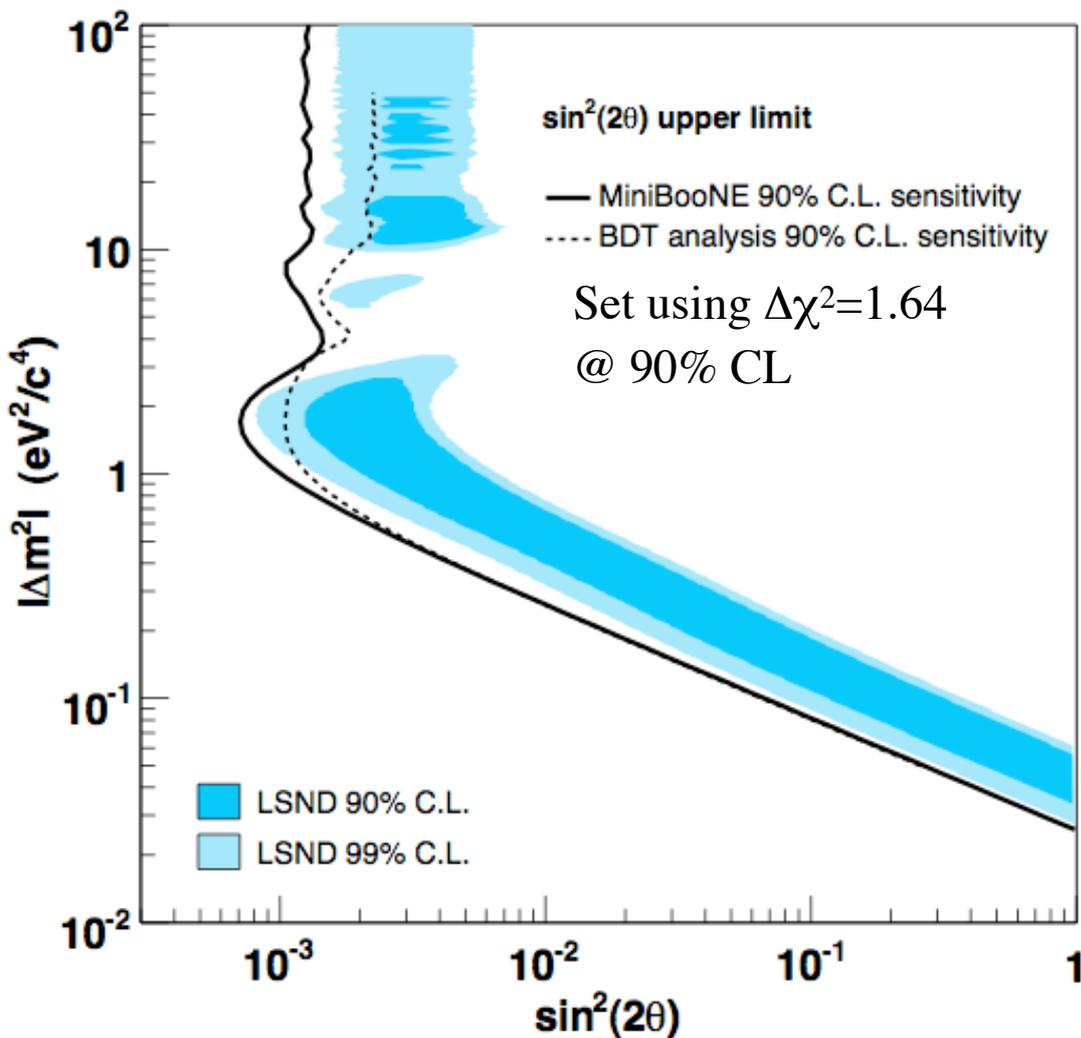
Comparison: Systematic Errors

Both analyses construct error matrices for the oscillation fit, binned in E_ν , to estimate the uncertainty on the expected number of ν_e background events.

	<i>source</i>	<i>track-based (%)</i>	<i>boosting (%)</i>
✓	Flux from π^+/μ^+ decay	6.2	4.3
✓	Flux from K^+ decay	3.3	1.0
✓	Flux from K^0 decay	1.5	0.4
	Target and beam models	2.8	1.3
✓	ν -cross section	12.3	10.5
	NC π^0 yield	1.8	1.5
	External interactions	0.8	3.4
✓	Optical model	6.1	10.5
	DAQ electronics model	7.5	10.8
	<i>constrained total</i>	9.6	14.5

Note:
 “total” is **not**
 the quadrature
 sum-- errors are
 further reduced
 by fitting with
 ν_μ data ✓

Comparison: Sensitivity



Fit the Monte Carlo E_ν^{QE} event distributions for oscillations

Raster scan in Δm^2 , and $\sin^2 2\theta_{\mu e}$
(assume $\sin^2 2\theta_{\mu x} = 0$),
calculate χ^2 value over E_ν bins

$$\chi^2 = \sum_{i=1}^{N_{bins}} \sum_{j=1}^{N_{bins}} (m_i - t_i) \mathcal{M}_{ij}^{-1} (m_j - t_j)$$

m_i = Number of measured data events in bin i

t_i = Number of predicted events in bin i

(t_i events are a function of Δm^2 , $\sin^2 2\theta$,

\mathcal{M}_{ij}^{-1} = Inverse of the covariance matrix

Since the track-based analysis achieved better sensitivity than the boosted decision tree analysis, we decided (before opening the box) that it would be used for the primary result.

1. Motivation & Introduction
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Results: Opening the Box

After applying all analysis cuts:

Step 1: Fit sequestered data to an oscillation hypothesis

Fit does not return fit parameters

Unreported fit parameters applied to MC; diagnostic variables compared to data

Return only the χ^2 of the data/MC comparisons (for diagnostic variables only)

Step 2: Open plots from Step 1 (Monte Carlo has unreported signal)

Plots chosen to be useful diagnostics, without indicating if signal was added
(reconstructed position, direction, visible energy...)

Step 3: Report only the χ^2 for the fit to E_ν^{QE}

No fit parameters returned

Step 4: Compare E_ν^{QE} for data and Monte Carlo,

Fit parameters **are** returned

This step breaks blindness

Step 5: Present results within two weeks

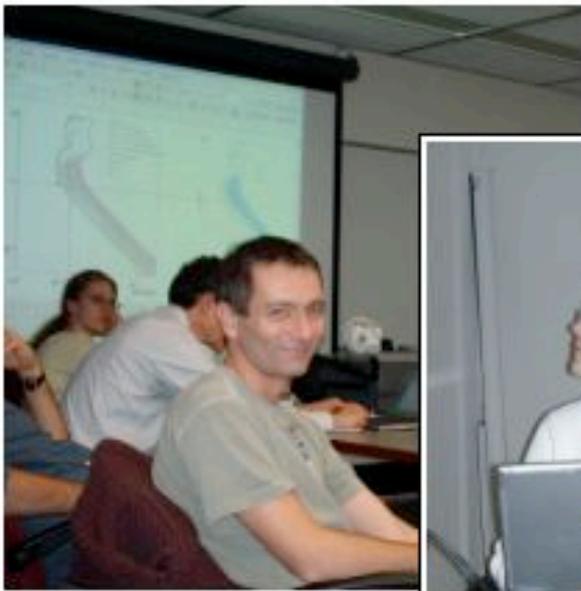
Training for a blind search



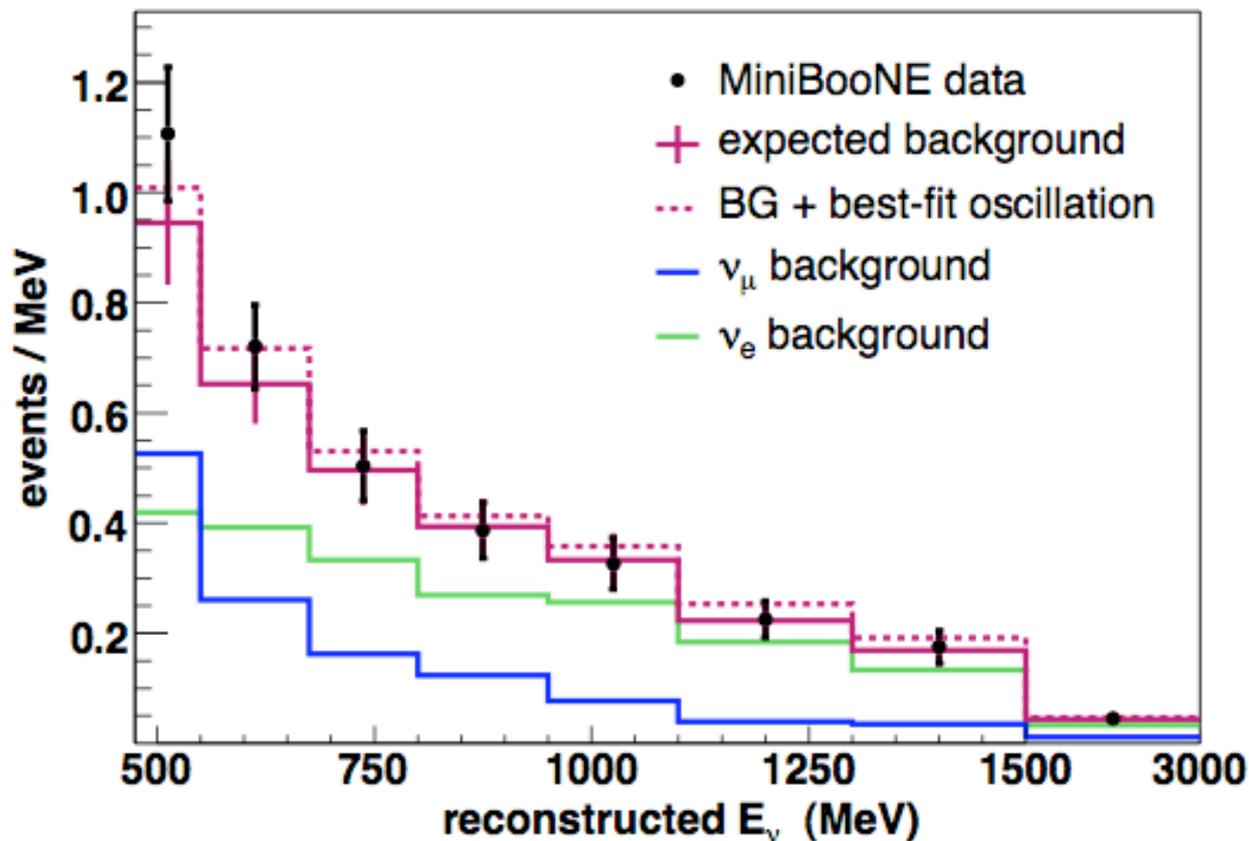
MOW
(blinded)
c.2002

We opened the box on March 26, 2007

The Box Opening



Results: Track Based Analysis



We observe no significant evidence for an excess of ν_e events in the energy range of the analysis.

NB: Errors bars=diagonals of error matrix

Best Fit (dashed):
 $(\sin^2 2\theta, \Delta m^2) = (0.001, 4 \text{ eV}^2)$

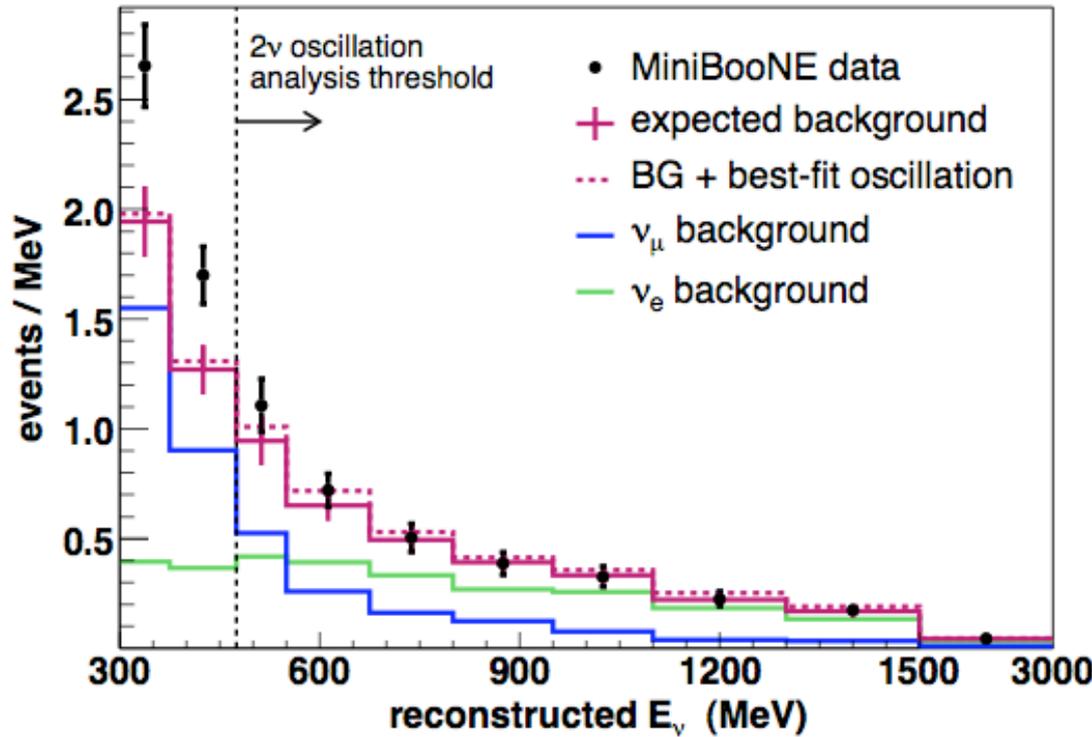
Counting Experiment: $475 < E_\nu^{\text{QE}} < 1250 \text{ MeV}$
 data: **380**
 expectation: **$358 \pm 19 \text{ (stat)} \pm 35 \text{ (sys)}$**

significance:
 0.55σ

χ^2 probability of best-fit point: 99%

χ^2 probability of null hypothesis: 93%

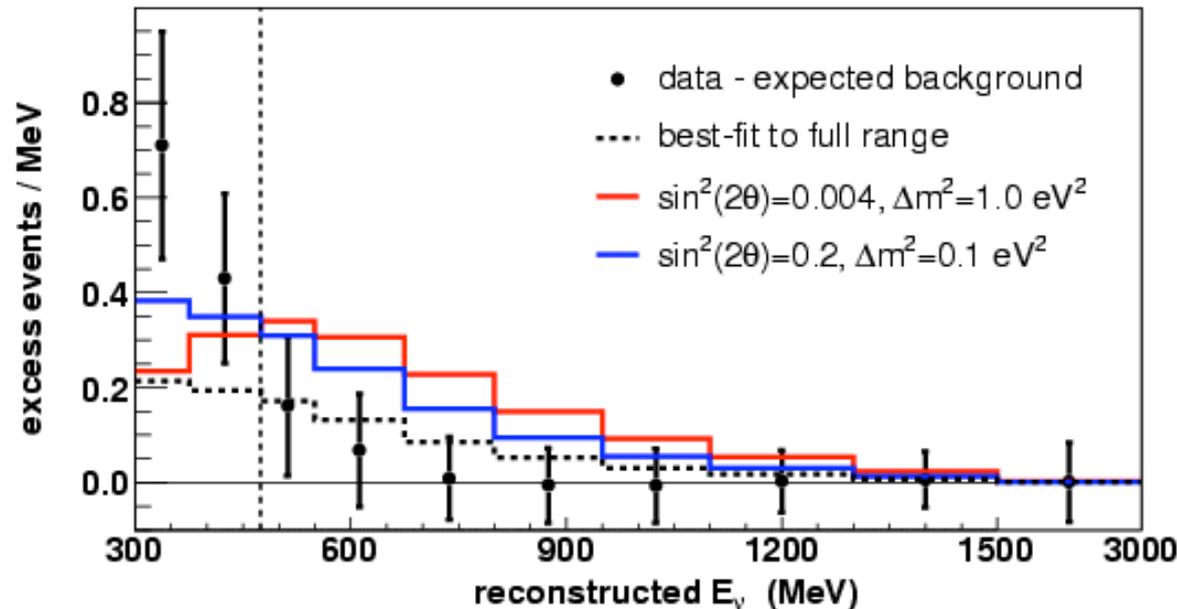
Results: Track Based Analysis, Lower Energy Threshold



Extending down to energies below the analysis range: $E_\nu^{QE} > 300$ MeV (we agreed to report this before box opening)

Data deviation for $300 < E_\nu^{QE} < 475$ MeV: 3.7σ

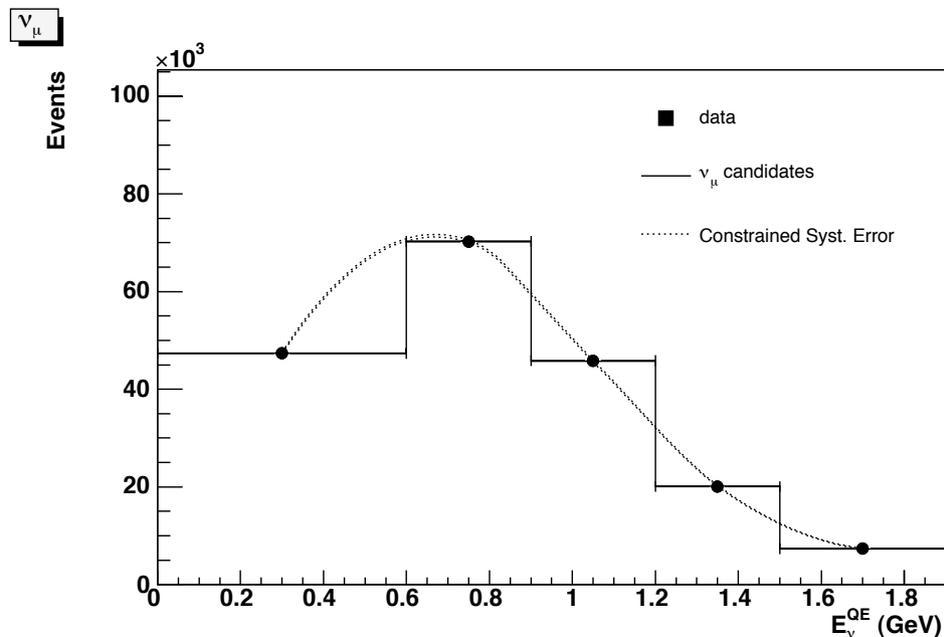
Oscillation fit to $E_\nu^{QE} > 300$ MeV:
 Best Fit ($\sin^2 2\theta, \Delta m^2$) = (1.0, 0.03 eV²)
Ruled out by Bugey



χ^2 prob. at best-fit point: 18%
 No closed contour for 90%CL

Fit is inconsistent with $\nu_\mu \rightarrow \nu_e$ oscillations.

Results: Boosted Decision Tree Analysis



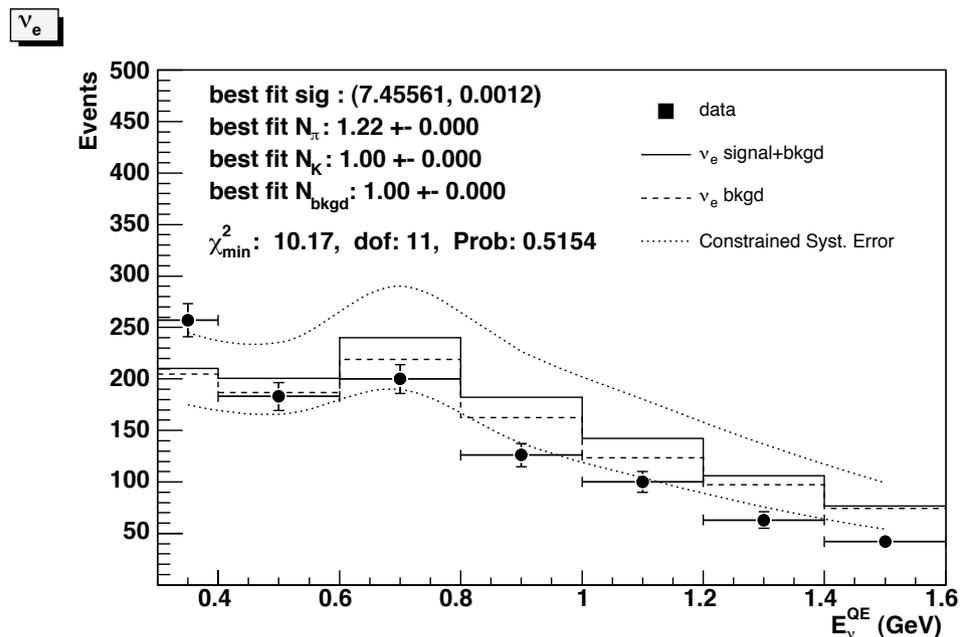
We observe no significant evidence for an excess of ν_e events in the energy range of the analysis.

Counting Experiment:

$$300 < E_{\nu}^{QE} < 1500 \text{ MeV}$$

data: **971**

expectation: **1070 ± 33 (stat) ± 225 (sys)**



Best Fit Point (dashed):

$$(\sin^2 2\theta, \Delta m^2) = (0.001, 7 \text{ eV}^2)$$

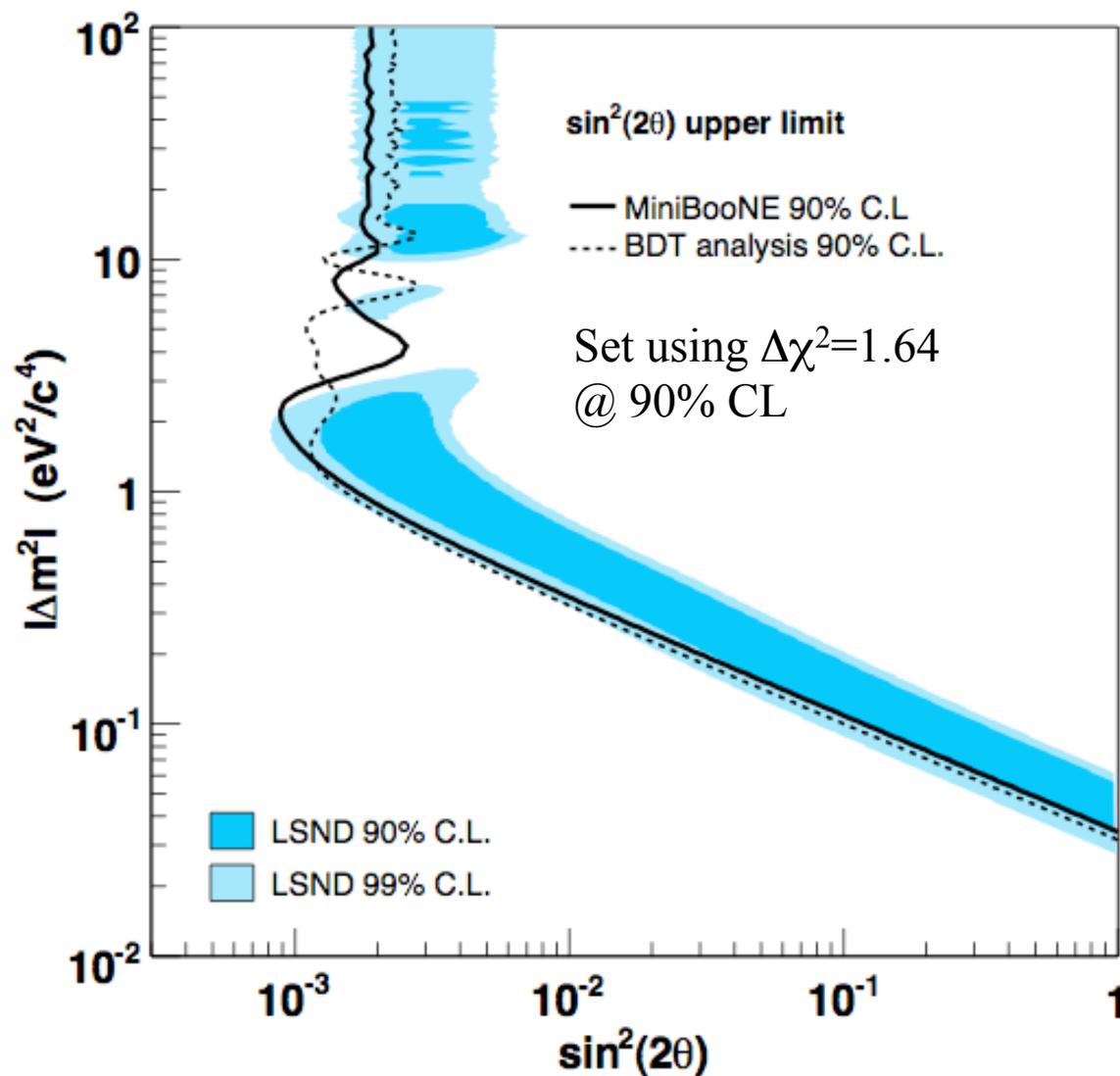
χ^2 probability of best-fit point: 52%

χ^2 probability of null hypothesis: 62%

significance:
-0.38 σ

Results: Comparison

MiniBooNE observes no evidence for $\nu_\mu \rightarrow \nu_e$ appearance-only oscillations.



The two independent oscillation analyses are in agreement.

solid: track-based

$$\Delta\chi^2 = \chi^2_{best\ fit} - \chi^2_{null} = 0.94$$

dashed: boosting

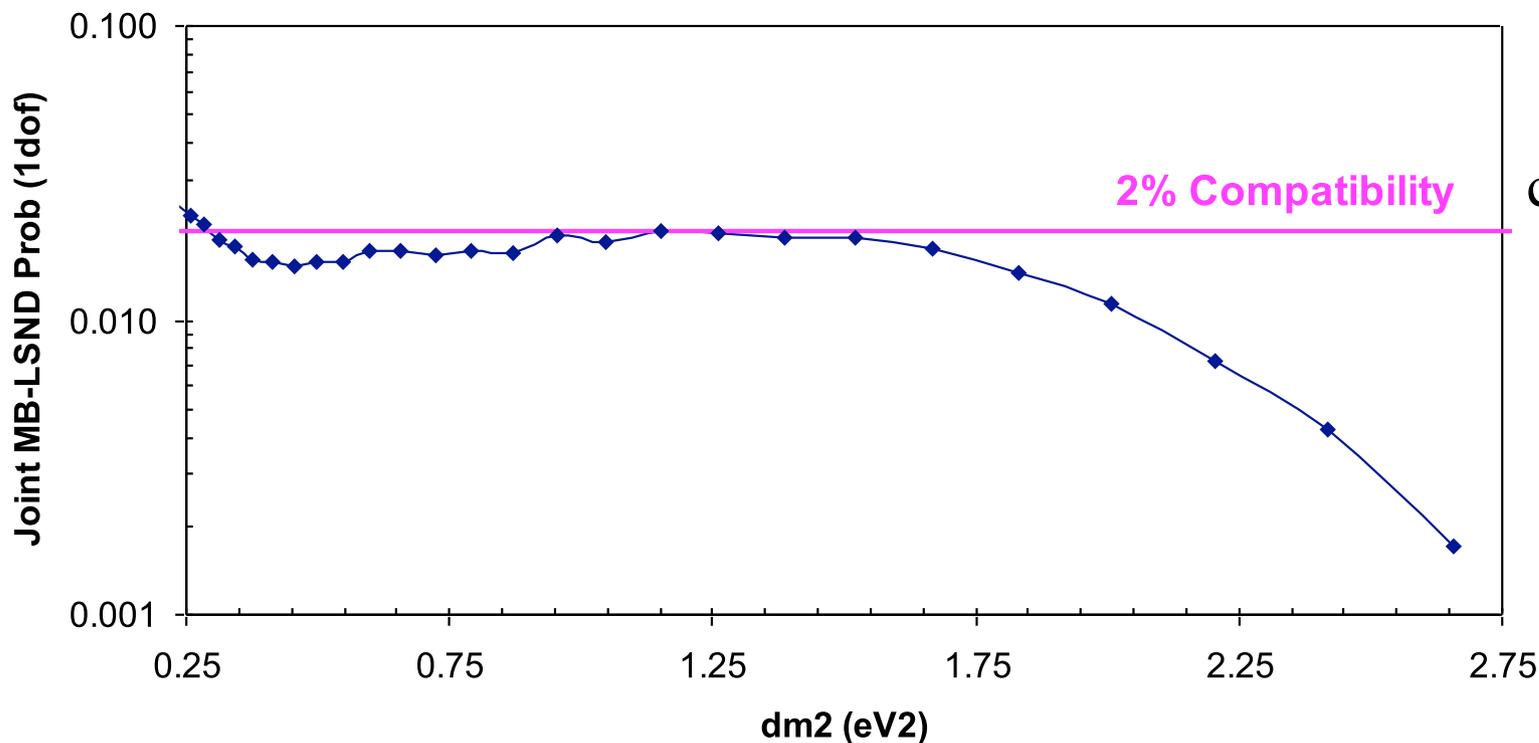
$$\Delta\chi^2 = \chi^2_{best\ fit} - \chi^2_{null} = 0.71$$

Therefore, we set a limit.

Results: Compatibility with LSND

A MiniBooNE-LSND Compatibility Test:
$$\chi_0^2 = \frac{(z_{MB} - z_0)^2}{\sigma_{MB}^2} + \frac{(z_{LSND} - z_0)^2}{\sigma_{LSND}^2}$$

- For each Δm^2 , form χ^2 between MB and LSND measurement
- Find z_0 ($\sin^2 2\theta$) that minimises χ^2 (weighted average of 2 measurements), this gives χ_{\min}^2
- Find probability of χ_{\min}^2 for 1 dof = joint compatibility probability for this Δm^2



cf. LSND-KARMEN:
64% compatibility

*MiniBooNE is incompatible with a $\nu_{\mu} \rightarrow \nu_e$
appearance-only interpretation of LSND at **98% CL***

Results: Plans

A paper on this analysis is posted to the archive.

Many more papers supporting this analysis will follow, in the very near future:

ν_μ CCQE production
 π^0 production

We are pursuing further analyses of the neutrino data, including:
an analysis which combines TB and BDT,
less simplistic models for the LSND effect.

MiniBooNE is presently taking data in antineutrino mode.

SciBooNE will start taking data in June!

Will improve constraints on ν_e backgrounds
(intrinsic ν_e s, improved π^0 kinematics)

Will provide important constraints on “wrong-sign” BGs for
antineutrino oscillation analysis

Conclusions

- 1. Within the energy range of the analysis, MiniBooNE observes no statistically significant excess of ν_e events above background.**
- 2. In two independent oscillation analyses, the observed E_ν distribution is inconsistent with a $\nu_\mu \rightarrow \nu_e$ appearance-only model.**
- 3. Therefore, we set a limit on $\nu_\mu \rightarrow \nu_e$ oscillations at $\Delta m^2 \sim 1 \text{ eV}^2$.
The MiniBooNE - LSND joint probability is 2%.**



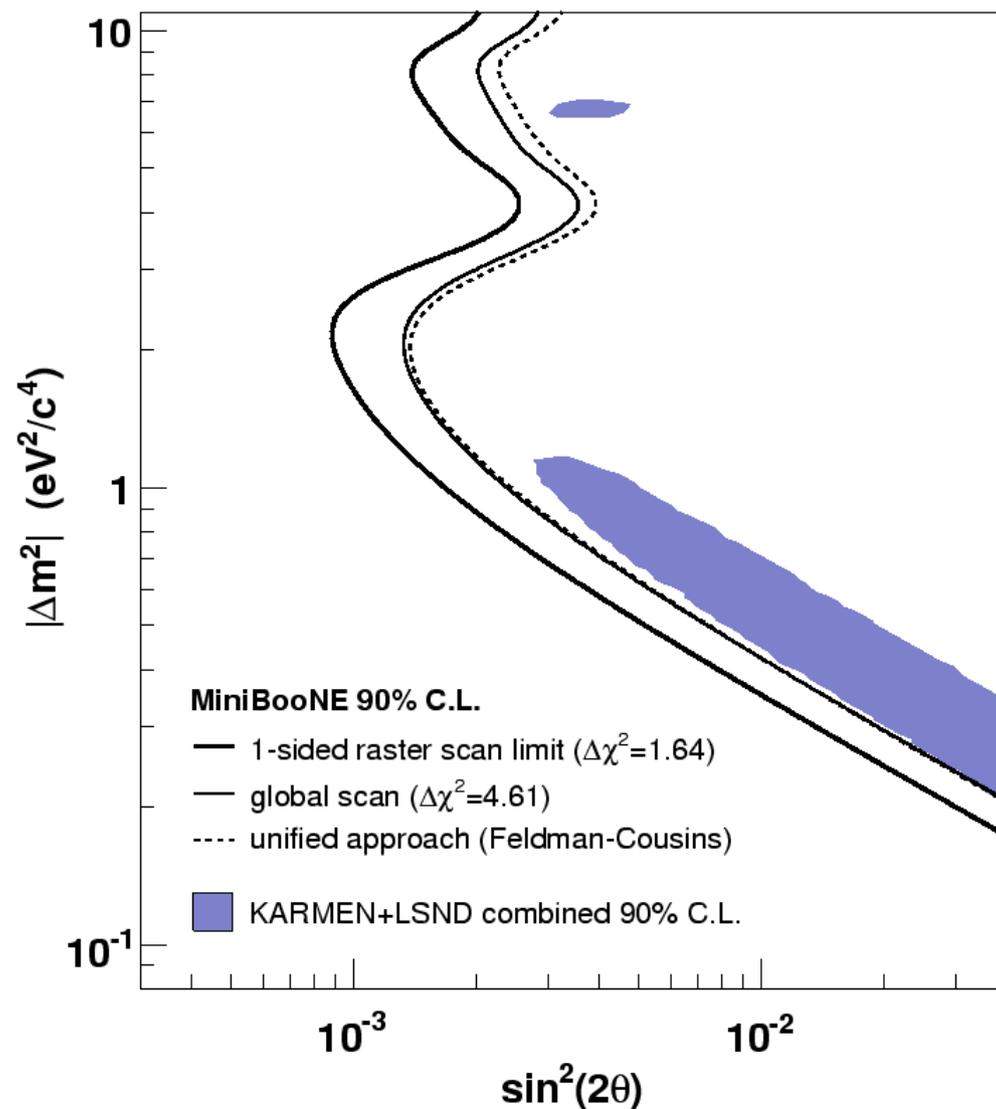
Results: Interpreting Our Limit

There are various ways to present limits:

- Single sided raster scan
(historically used, presented here)
- Global scan
- Unified approach
(most recent method)

This result must be folded into an
LSND-Karmen joint analysis.

Church, et al., PRD 66, 013001



We will present a full joint analysis soon.

Results: Event Overlap

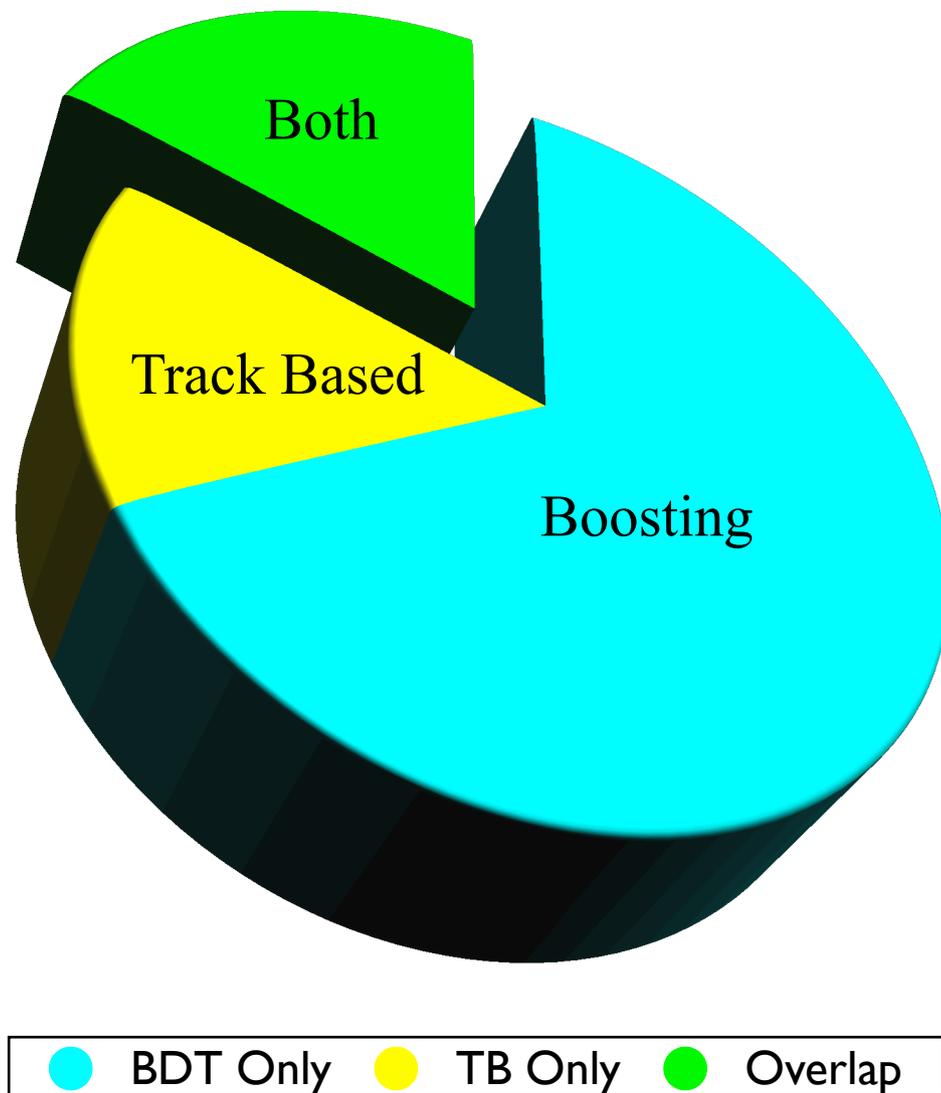
Counting experiment numbers:

Track Based Algorithm finds 380 events

Boosting Algorithm finds 971 events

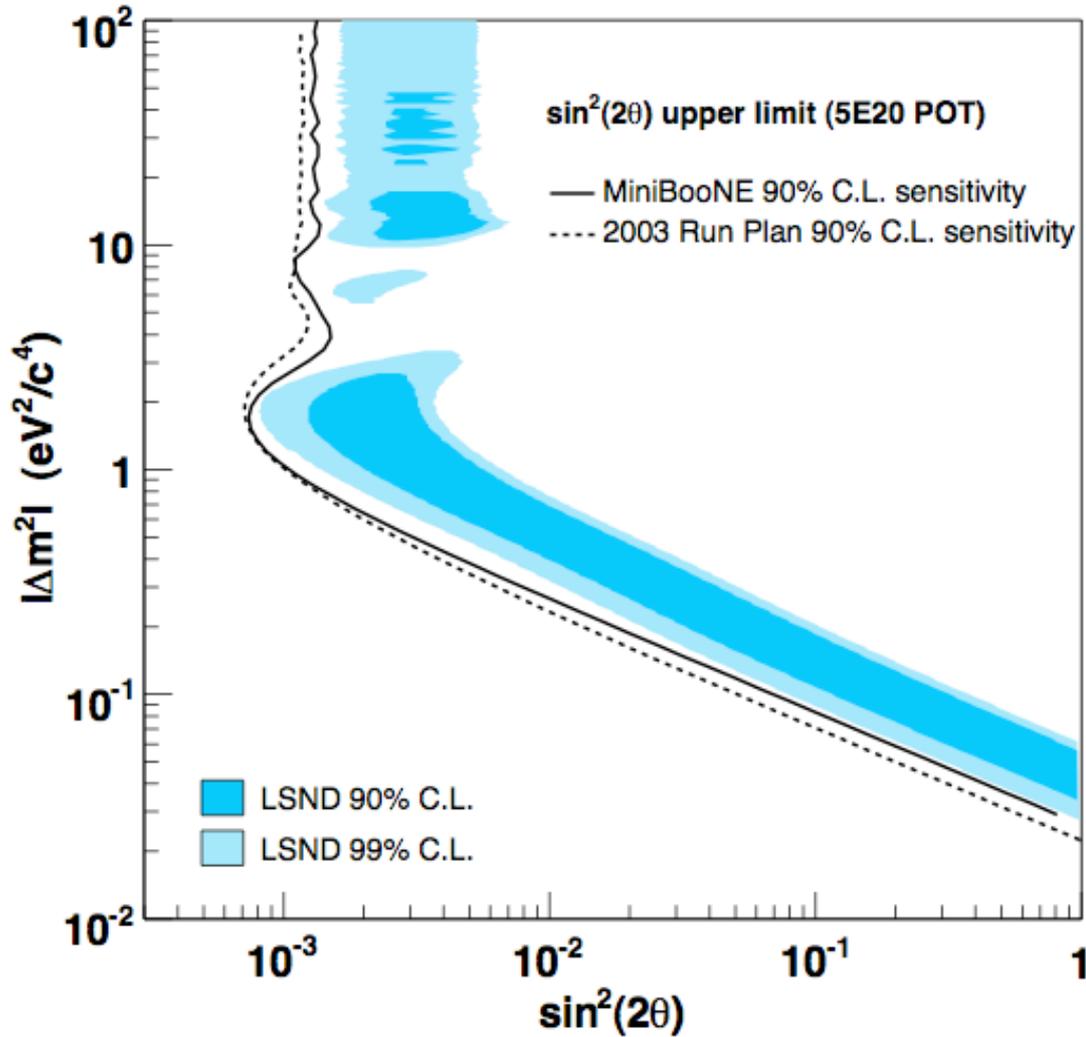
However, only 1131 events total,
because 220 overlap

- chosen by both algorithms!



Results: Sensitivity Goal

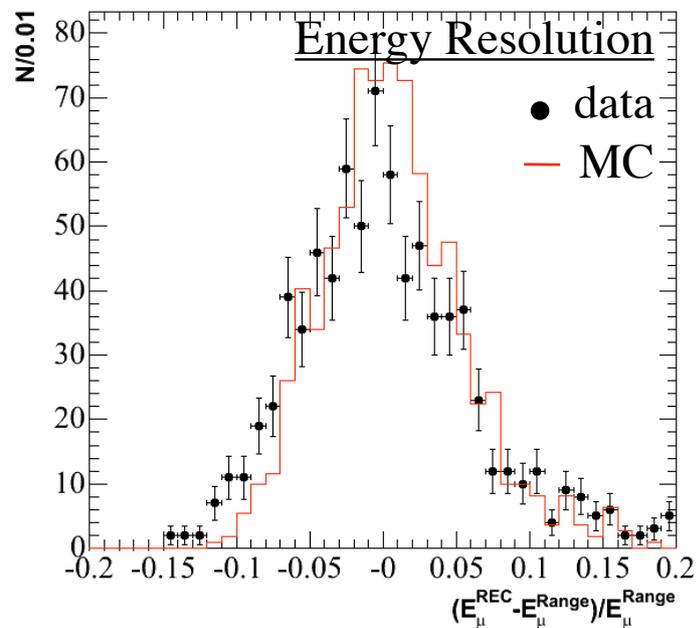
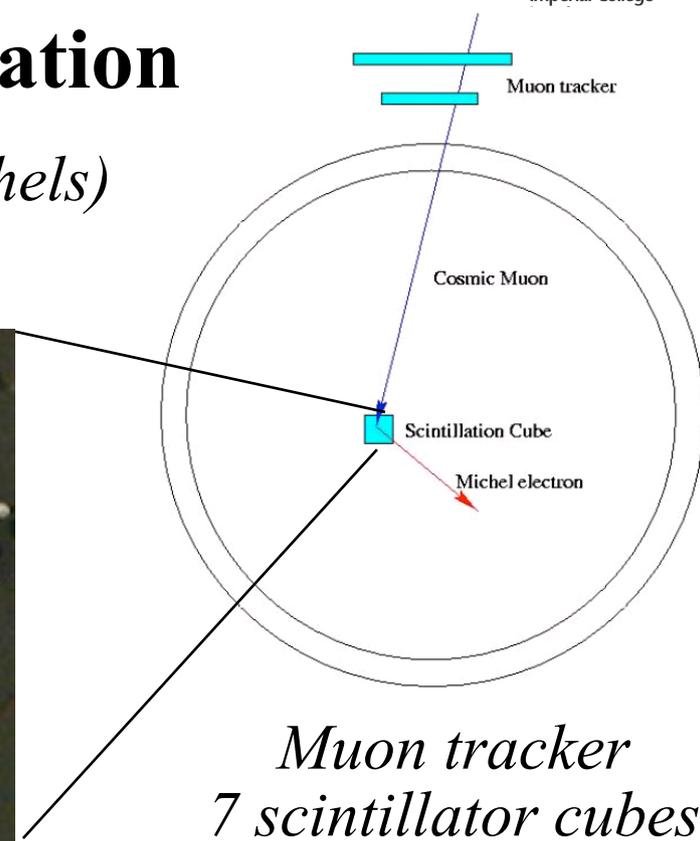
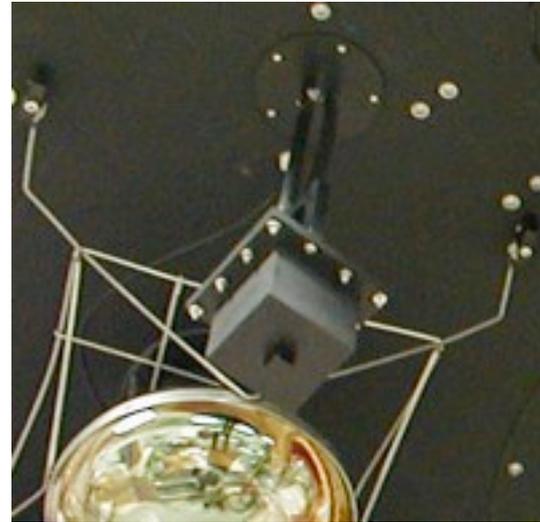
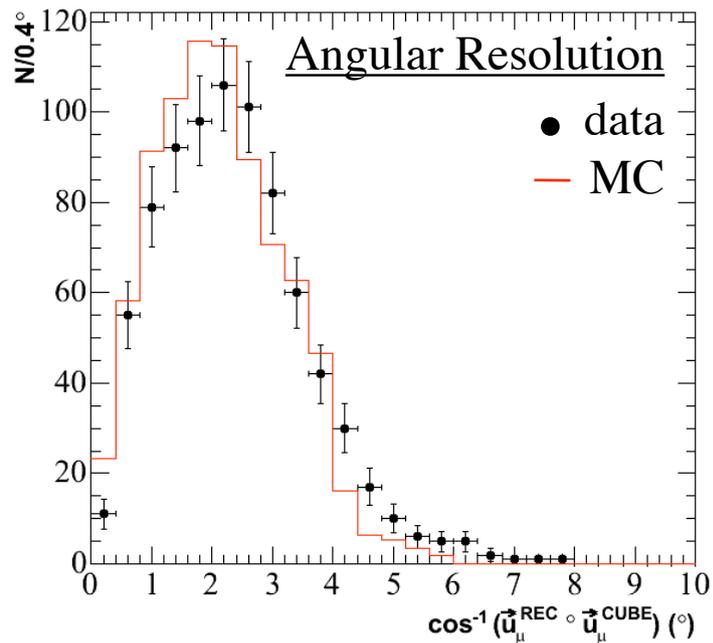
Compared to our sensitivity goal for 5E20 protons on target from 2003 Run Plan



Set using $\Delta\chi^2=1.64$ @ 90% CL

MiniBooNE Detector: Cosmic Calibration

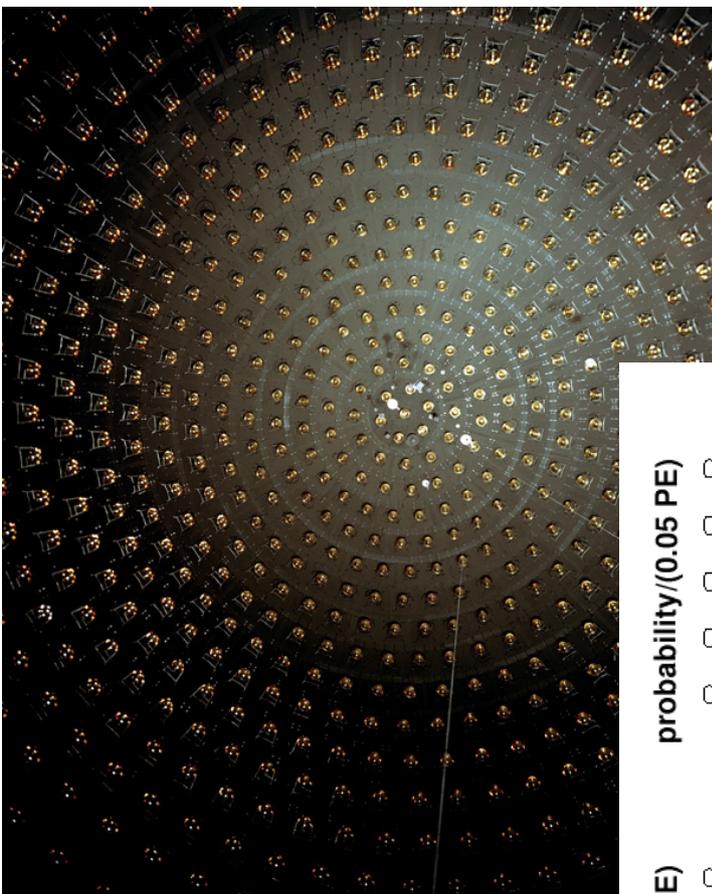
use cosmic muons and their decay electrons (Michels)



- Cosmic muons which stop in cubes:
- test energy scale extrapolation up to 800 MeV
 - measure energy, angle resolution
 - compare data and MC

*Muon tracker + cube calibration
data continuously acquired at 1 Hz*

MiniBooNE Detector: PMT Calibration



10% photo-cathode coverage

PMTs are calibrated with a laser + 4 flask system

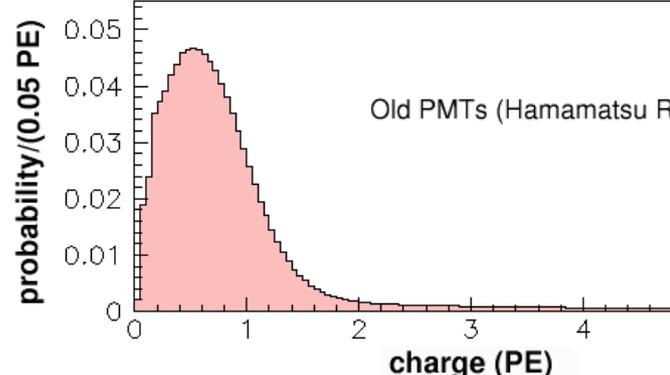
PMT Charge Resolution: 1.4 PE, 0.5 PE

PMT Time Resolution: 1.7 ns, 1.1 ns

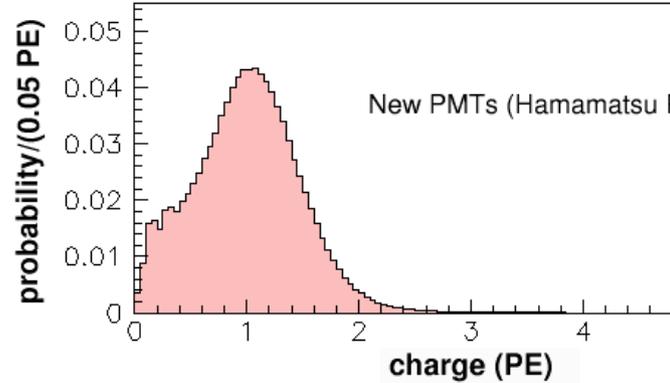
Two types of 8”
Hamamatsu Tubes:
R1408, R5912

*Laser data are acquired at 3.3 Hz to continuously
calibrate PMT gain and timing constants*

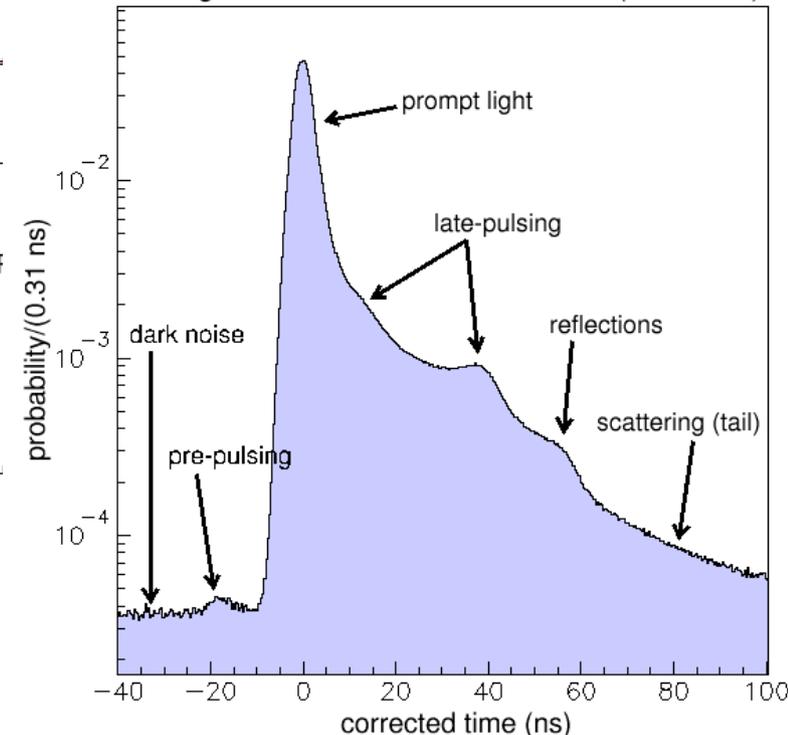
Single Photoelectron Response in MiniBooNE



R. B. Patterson



Timing Distribution for Laser Events (old tubes)

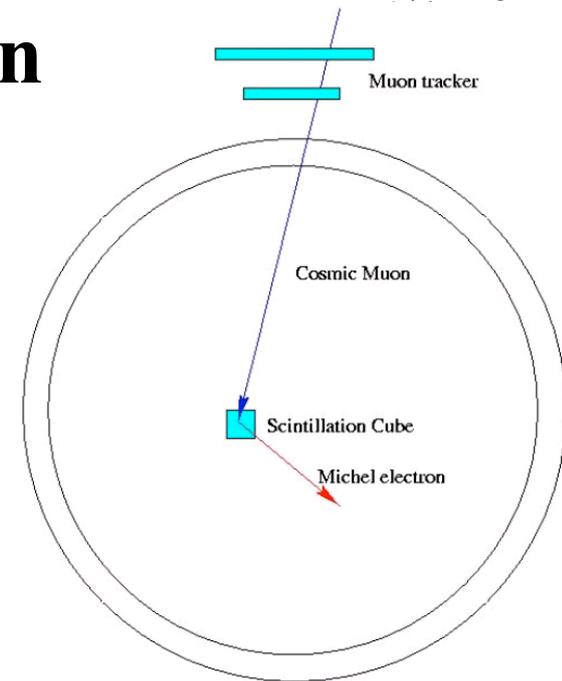
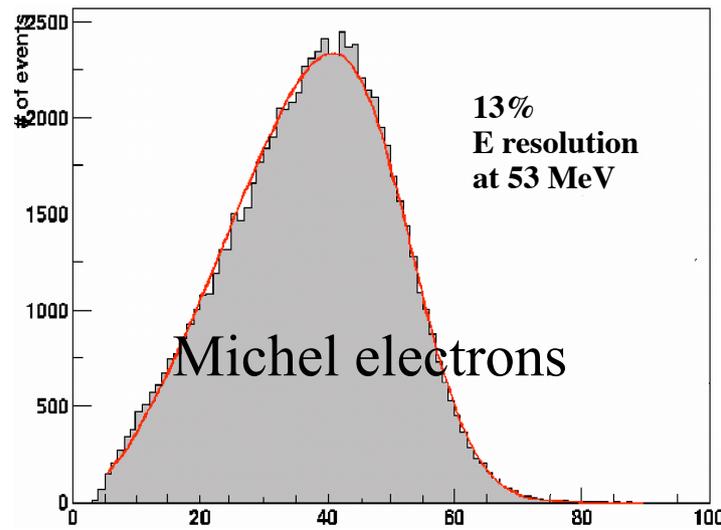


MiniBooNE Detector: Cosmic Calibration

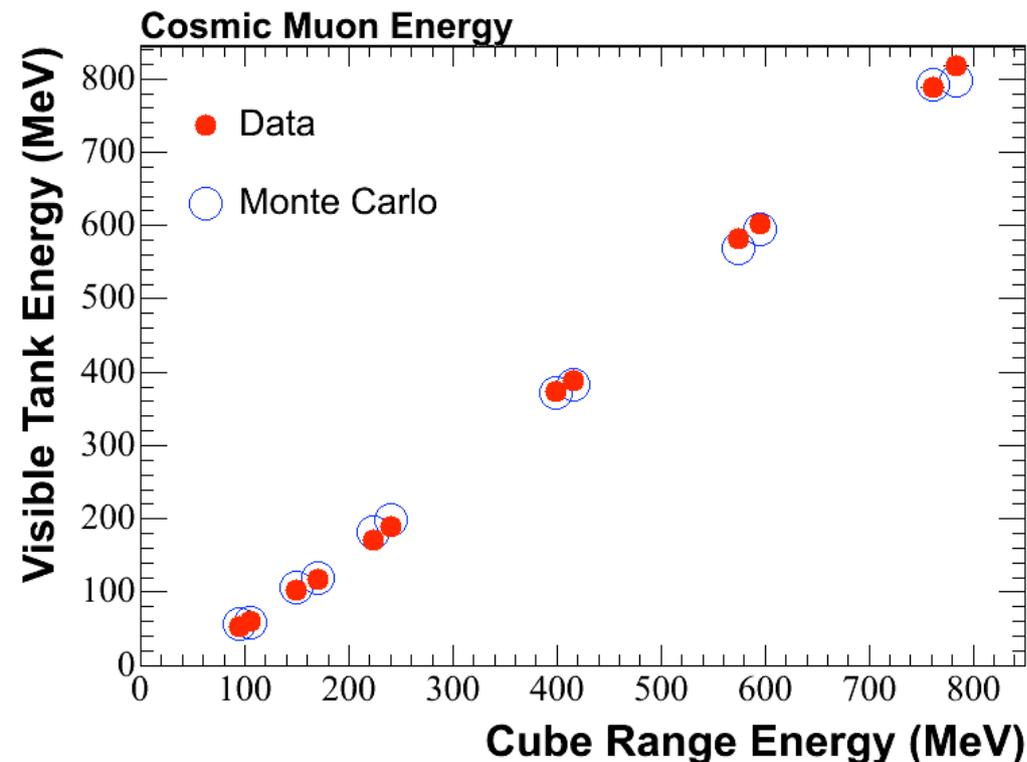
use cosmic muons and their decay electrons (Michels)

Michel electrons:

- set absolute energy scale and resolution at 53 MeV endpoint
- optical model tuning



*Muon tracker
7 scintillator cubes*

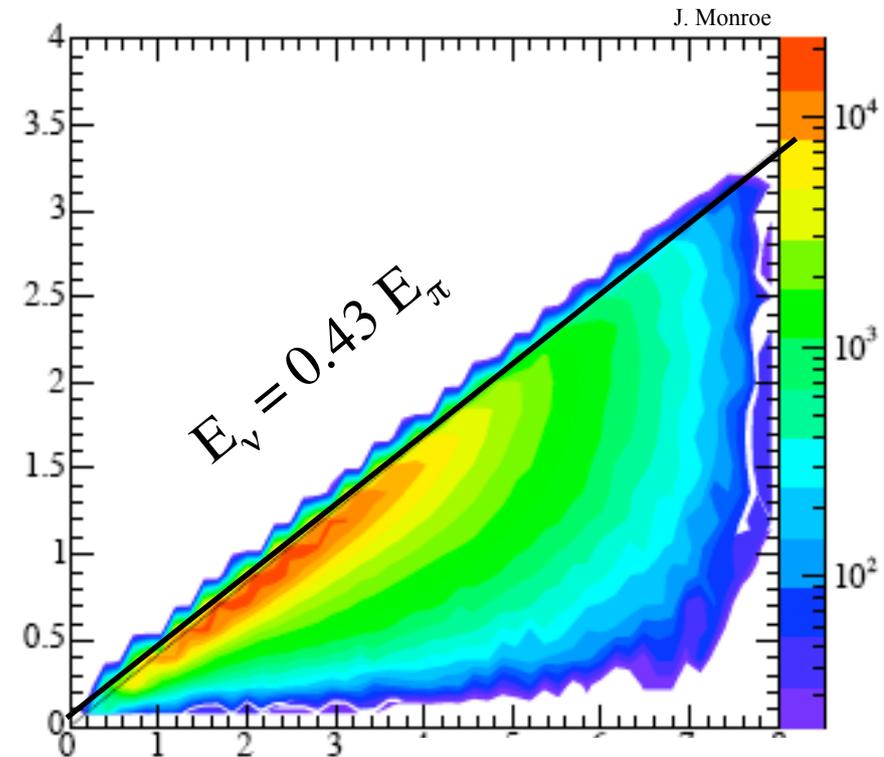
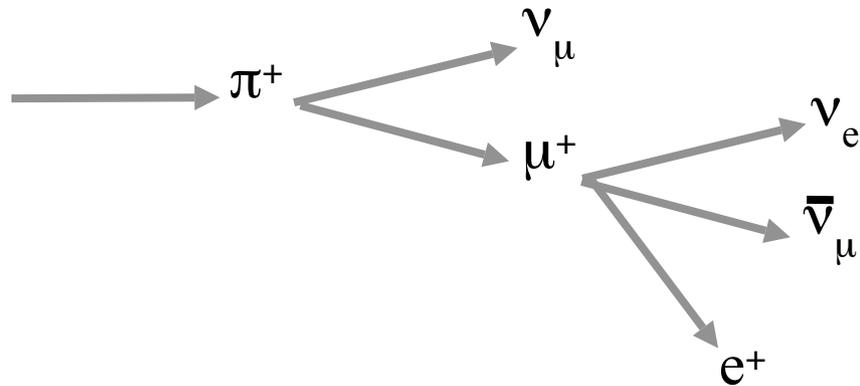


- Cosmic muons which stop in cubes:
- test energy scale extrapolation up to 800 MeV
 - measure energy, angle resolution
 - compare data and MC

*Muon tracker + cube calibration
data continuously acquired at 1 Hz*

Incorporating ν_μ Data: μ^+ -Decay ν_e Background

ν_μ CCQE events measure the π^+ spectrum, this constrains the μ^+ -decay ν_e flux



this works well because the ν_μ energy is highly correlated with the π^+ energy

Ratio Method Constraint:

1. MC based on external data predicts a central value and a range of possible $\nu_\mu(\pi)$ fluxes
2. make Data/MC ratio vs. E_ν^{QE} for ν_μ CCQE data
3. re-weight each possible MC flux by the ratio (2) including the ν_μ , its parent π^+ , sister μ^+ , and niece ν_e

Analysis Strategy: Delta Background

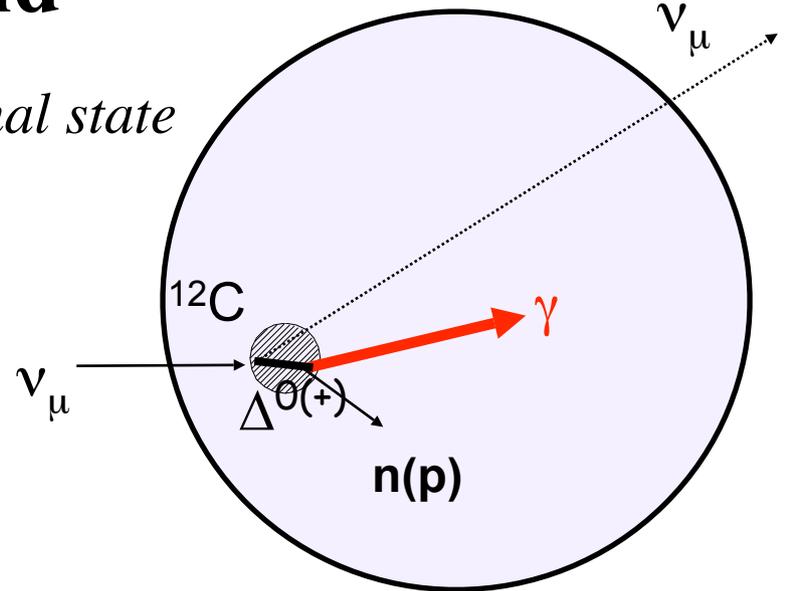
ν induced interactions that produce single Υ s in the final state

Radiative Delta Decay (NC)

- (i) Use π^0 events to measure rate of NC Δ production
- (ii) Use PDG branching ratio for radiative decay
- 15% uncertainty on branching ratio

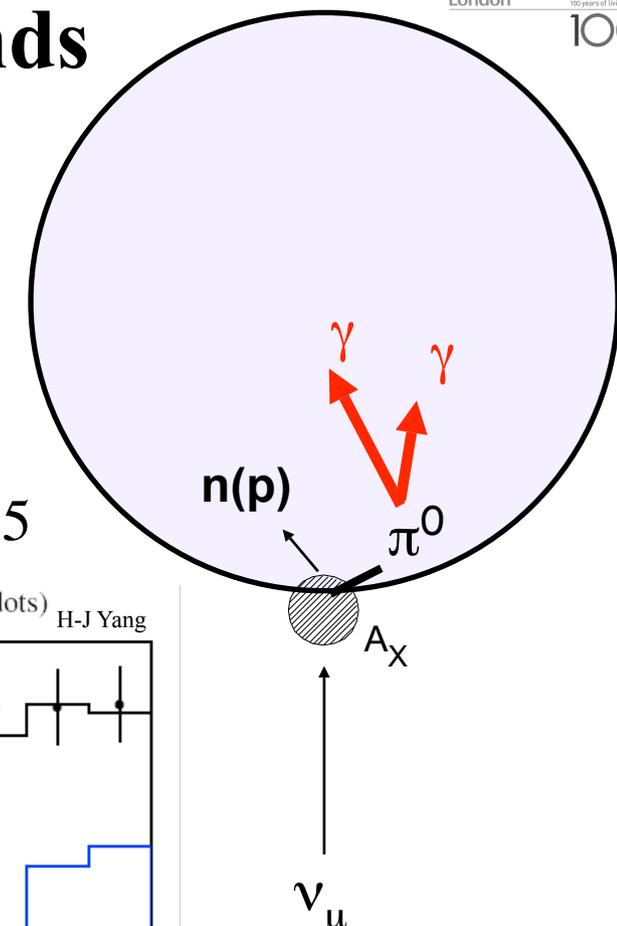
Inner Bremsstrahlung (CC)

- (i) Hard photon released from neutrino interaction vertex
- (ii) Use events where the μ is tagged by the decay e^-
- study misidentification using BDT algorithm.



Analysis Strategy: External Backgrounds

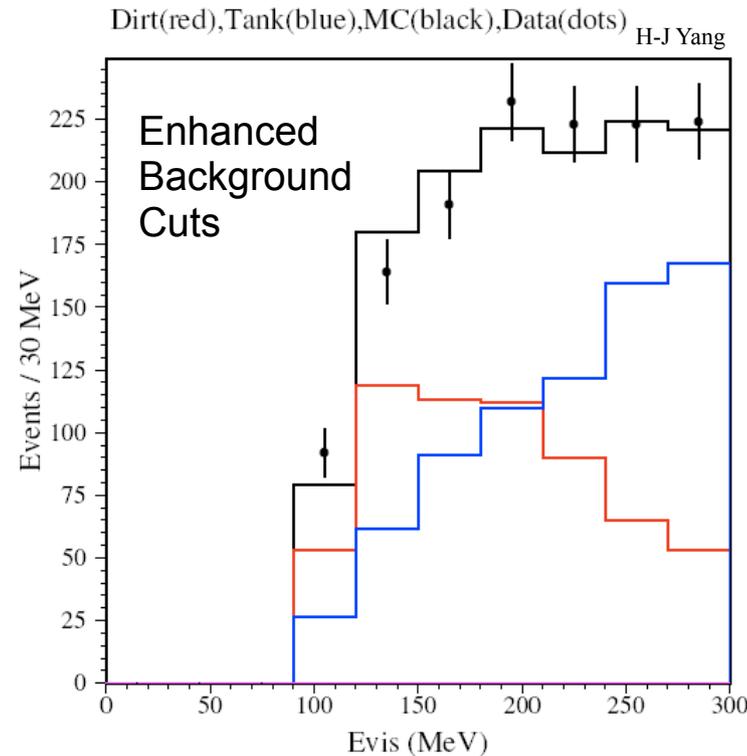
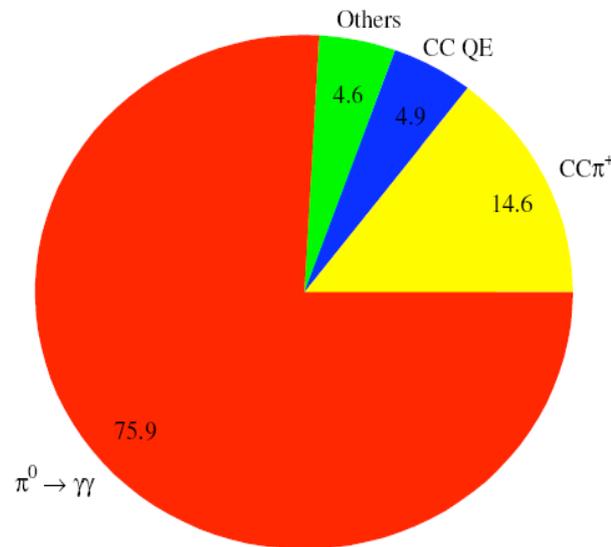
interactions outside the detector that deposit energy in the fiducial volume and pass the veto PMT hits cut



1. "Dirt" Events

ν interactions outside of the detector are measured in the "dirt box:" $N_{\text{data}}/N_{\text{MC}} = 0.99 \pm 0.15$

Event Type of Dirt Events



2. Cosmic Ray Background Events

Measured from 126E6 strobe data triggers: 2.1 ± 0.5 events.

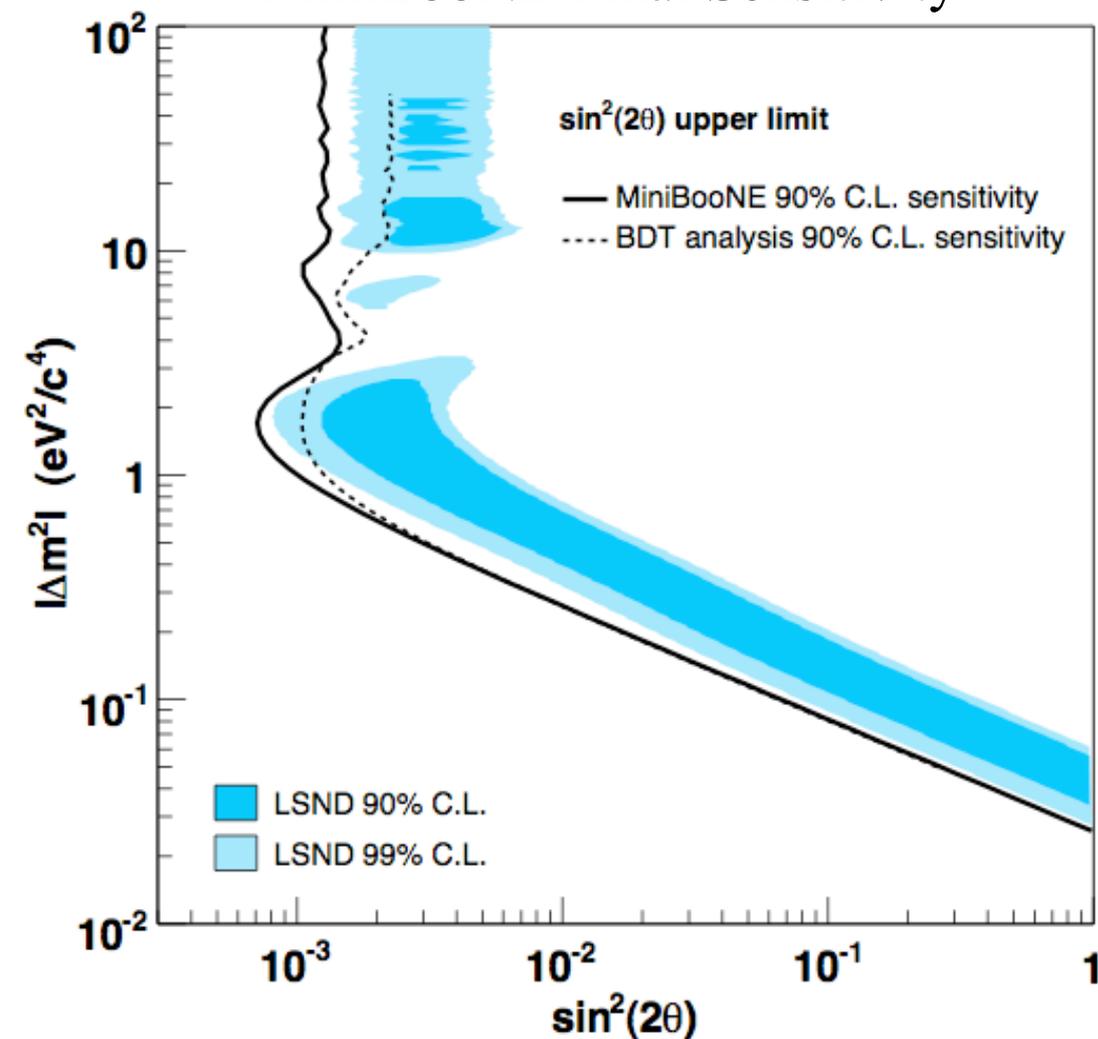
Analysis Overview: Background Summary

*Summary of predicted backgrounds for the primary MiniBooNE result
 (Track-Based Analysis):*

Process	Number of Events
ν_μ CCQE	10
$\nu_\mu e \rightarrow \nu_\mu e$	7
Miscellaneous ν_μ Events	13
NC π^0	62
NC $\Delta \rightarrow N\gamma$	20
NC Coherent & Radiative γ	< 1
Dirt Events	17
ν_e from μ Decay	132
ν_e from K^+ Decay	71
ν_e from K_L^0 Decay	23
ν_e from π Decay	3
Total Background	358
0.26% $\nu_\mu \rightarrow \nu_e$	(example signal) 163

MiniBooNE Summary

MiniBooNE Final Sensitivity



MiniBooNE performed a *blind analysis* for the $\nu_\mu \rightarrow \nu_e$ appearance search

- Did not look at ν_e events while developing reconstruction, particle identification algorithms
- Final cuts made with no knowledge of the number of ν_e events in the box

Final sensitivity to ν_e appearance shown for two independent analyses

- “Primary” analysis chosen based on slightly better sensitivity

Results: Opening the Box

After applying all analysis cuts:

Step 1: Fit sequestered data to an oscillation hypothesis

Fit does not return fit parameters

Unreported fit parameters applied to MC; diagnostic variables compared to data

Return only the χ^2 of the data/MC comparisons (for diagnostic variables only)

Step 2: Open plots from Step 1 (Monte Carlo has unreported signal)

Plots chosen to be useful diagnostics, without indicating if signal was added
(reconstructed position, direction, visible energy...)

Step 3: Report only the χ^2 for the fit to E_ν^{QE}

No fit parameters returned

Step 4: Compare E_ν^{QE} for data and Monte Carlo,

Fit parameters **are** returned

This step breaks blindness

March 26:

Track-Based

χ^2 Probability: 99%

Boosting

χ^2 Probability: 62%

Step 5: Present results within two weeks

Step 1

Return the χ^2 of the data/MC comparison for
a set of diagnostic variables

12 variables are tested for TB
46 variables are tested for BDT

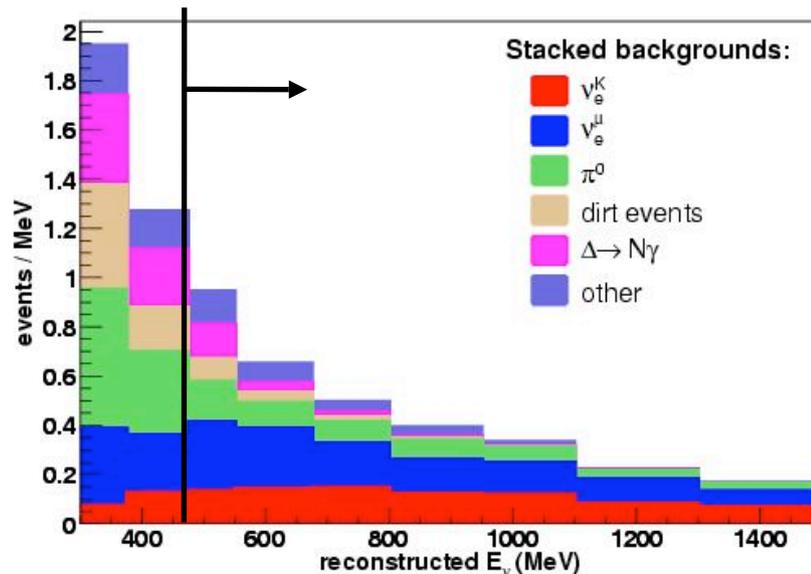
All analysis variables were returned with good
probability except...

TB analysis χ^2 Probability of E_{visible} fit: 1%

**This probability was sufficiently low
to merit further consideration**

In the TB analysis

- We re-examined our background estimates using sideband studies.
 \Rightarrow *We found no evidence of a problem*
- However, knowing that backgrounds rise at low energy,
We tightened the cuts for the oscillation fit:



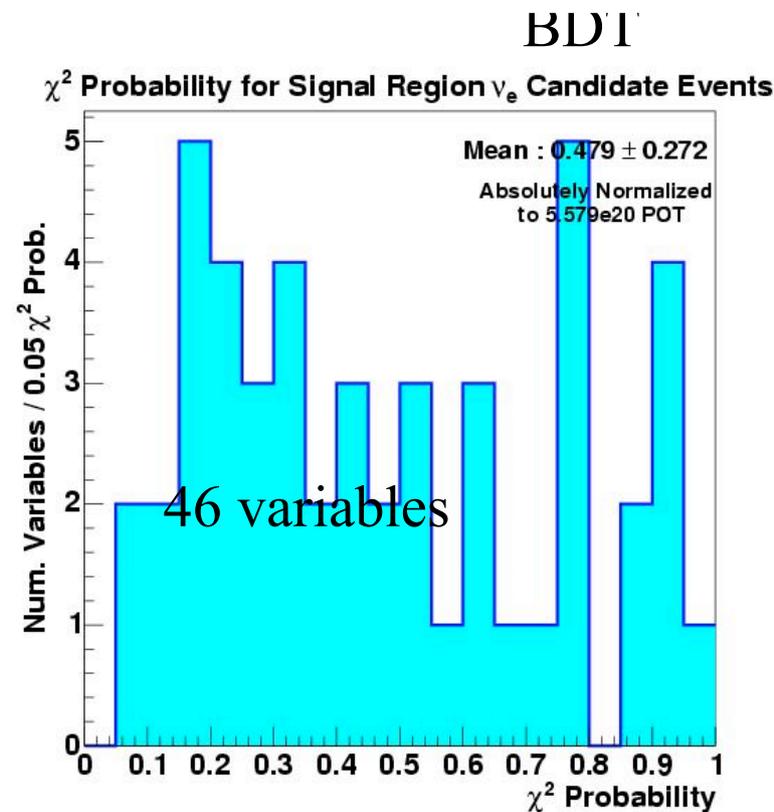
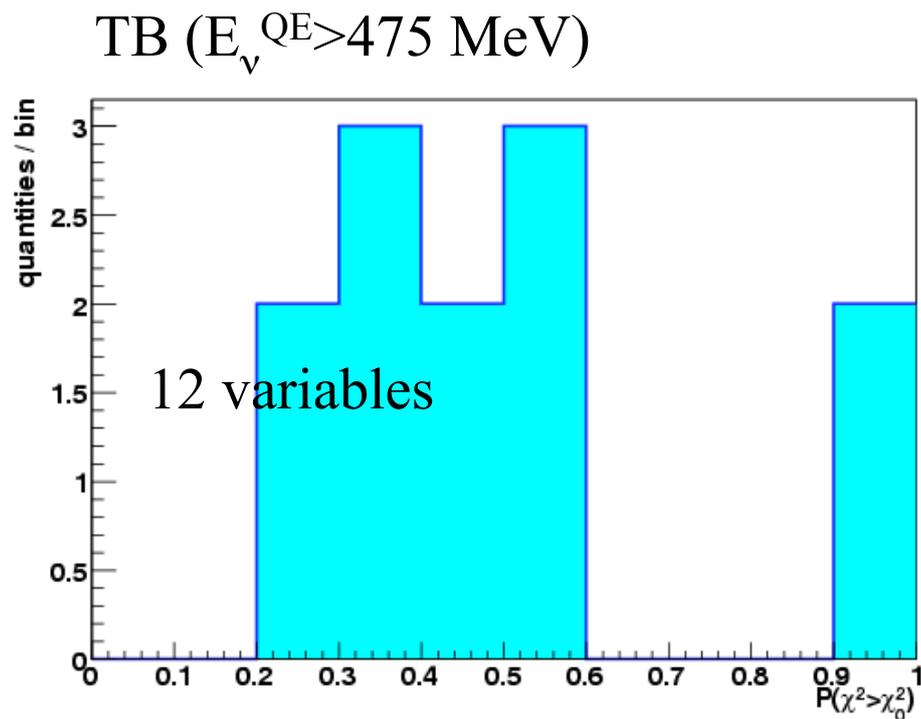
$$E_\nu^{QE} > 475 \text{ MeV}$$

We agreed to report events
over the original full range:
 $E_\nu^{QE} > 300 \text{ MeV}$,

Step 1: again!

Return the χ^2 of the data/MC comparison for a set of diagnostic variables

χ^2 probabilities returned:

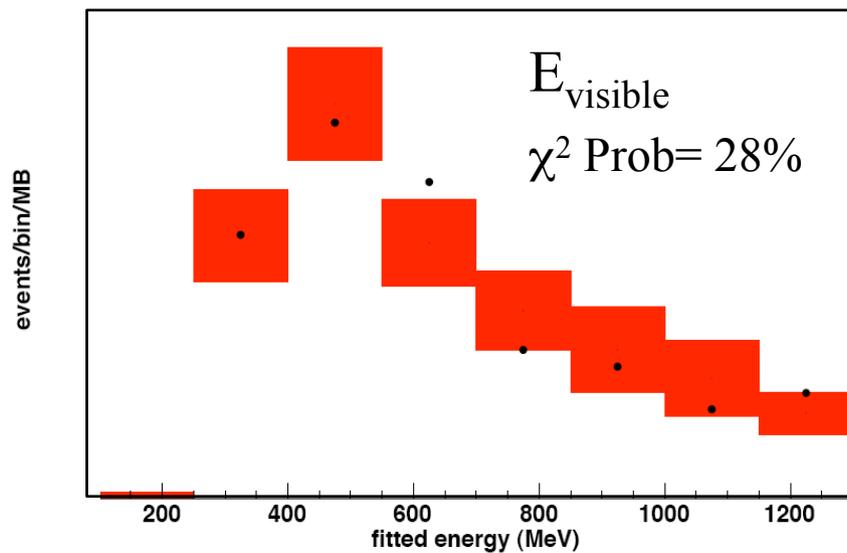


Parameters of the oscillation fit were not returned.

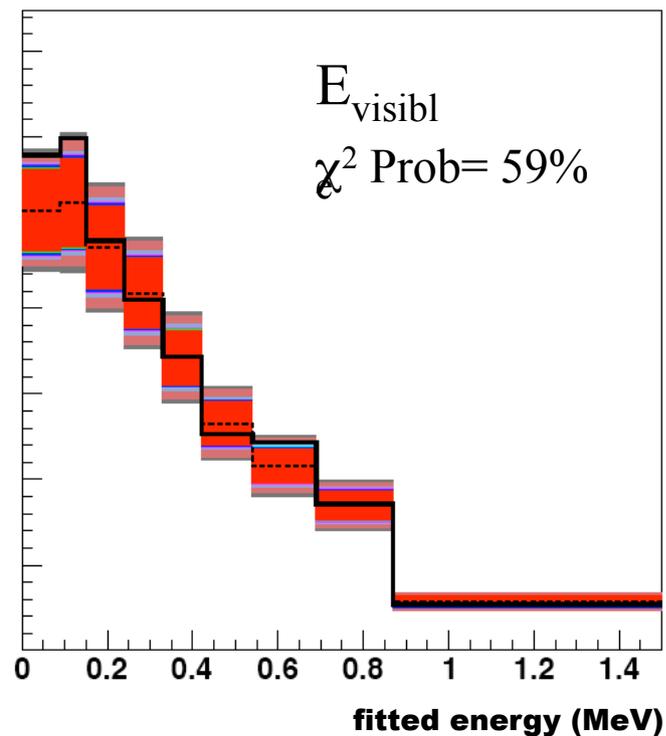
Step 2

Open up the plots from step 1 for approval.

*Examples of
what we saw:*



TB ($E_{\nu}^{\text{QE}} > 475 \text{ MeV}$)



BDT

MC contains fitted signal at unknown level

Step
3

Report the χ^2 for a fit to E_ν^{QE} across full energy range

TB ($E_\nu^{\text{QE}} > 475$ MeV) χ^2 Probability of fit: 99%
 BDT analysis χ^2 Probability of fit: 52%

Leading to...

Step
4
Open the box...

And the answer is...

Primary Analysis

Counting Experiment:

$$475 < E_{\nu}^{\text{QE}} < 1250 \text{ MeV}$$

expectation: **358 ±19 (stat) ± 35 (sys)**

data:

significance:

Cross-check Analysis

Counting Experiment:

$$300 < E_{\nu}^{\text{QE}} < 1500 \text{ MeV}$$

expectation: **1070 ±33 (stat) ± 225(sys)**

data:

significance:

And the answer is...

Primary Analysis

Counting Experiment:

$$475 < E_{\nu}^{\text{QE}} < 1250 \text{ MeV}$$

expectation: **358 ±19 (stat) ± 35 (sys)**

data: **380**

significance: **0.55 σ**

Cross-check Analysis

Counting Experiment:

$$300 < E_{\nu}^{\text{QE}} < 1500 \text{ MeV}$$

expectation: **1070 ±33 (stat) ± 225(sys)**

data:

significance:

And the answer is...

Primary Analysis

Counting Experiment:

$$475 < E_{\nu}^{\text{QE}} < 1250 \text{ MeV}$$

expectation: **358 ±19 (stat) ± 35 (sys)**

data: **380**

significance: **0.55 σ**

Cross-check Analysis

Counting Experiment:

$$300 < E_{\nu}^{\text{QE}} < 1500 \text{ MeV}$$

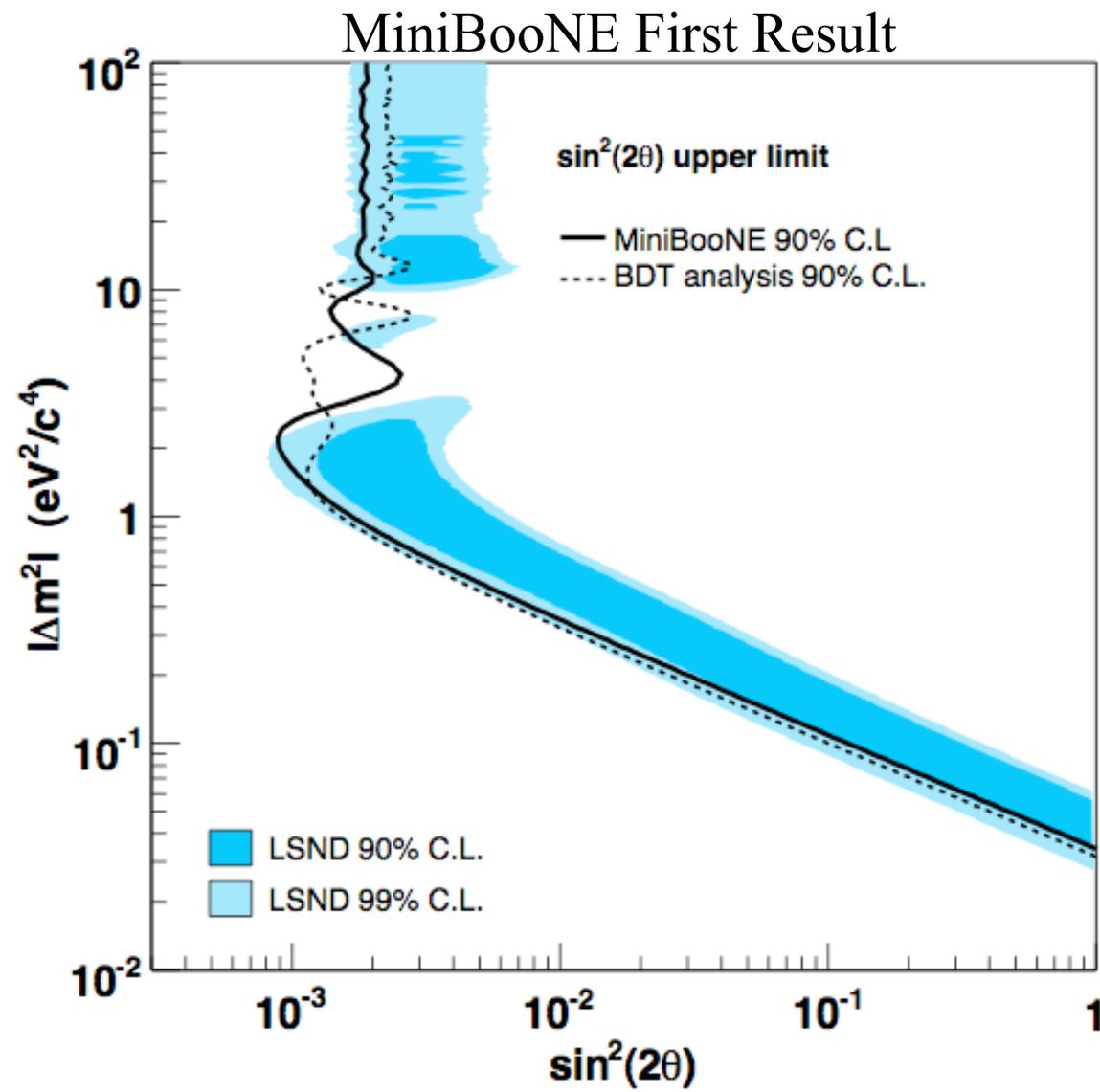
expectation: **1070 ±33 (stat) ± 225(sys)**

data: **971**

significance: **-0.38 σ**

And the answer is...

MiniBooNE observes no evidence for $\nu_\mu \rightarrow \nu_e$ appearance-only oscillations.



The two independent oscillation analyses are in agreement!