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Measurement of ν_e CC π^+ cross
section using the ND280 tracker
and development of optical diffuser
calibration systems for
Hyper-Kamiokande
by

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279		corrected hits as a function of time is shown on bottom right. Plot	
280		modified from [157].	157
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282		at the B5 injection point over the full detector. The time of flight	
283		corrected hits as a function of time is shown on bottom right. Plot	
284		modified from [157].	158
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286		and PTFE diffusers.	159
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288		and PTFE diffusers.	159

¹ Acknowledgments

² Thanks go here.

¹ **Declarations**

² Parts of this thesis have been previously published by the author in the following:

³ [?]

⁴ Research was performed in collaboration during the development of this thesis, but
⁵ does not form part of the thesis:

⁶ [?]

1 Abstract

2

3 Over the last few decades our understand of the physics that governs neutrino
4 oscillations has evolved rapidly through an experimental program designed to measure
5 the key parameters behind neutrino oscillations. This thesis provides an overlook
6 into the Tokai to Kamioka (T2K) long baseline accelerator neutrino experiment, and
7 the next generation water Cherenkov detector Hyper-Kamiokande; both designed
8 to make precise measurements on neutrino oscillation parameters. In the T2K far
9 detector a data excess is seen in the ν_e charged current π^+ sample, a significant
10 channel in electron neutrino appearance studies. An analysis is presented in this
11 thesis to investigate ν_e charged current π^+ production using the off-axis near detector
12 (ND280) tracker of the T2K experiment. A novel selection has been developed and
13 the systematic uncertainties evaluated to measure a flux average cross-section of
14 $\sigma = (2.23 \pm 0.39(\text{stat.}) \pm 0.38(\text{syst.})) \times 10^{-39} \text{ cm}^2$ per nucleon. This result provides
15 the first ever cross-section measurement of ν_e charged current π^+ production on a
16 carbon target. With kinematic constraints applied, analogous to the far detector
17 sample, preliminary studies indicate no data excess in the near detector sample.

18 **Unfinished...**

¹ Sponsorships and Grants

1 **Acronyms**

- 2 **CC** Charged Current.
- 3 **CP** Charge-Parity.
- 4 **HK** Hyper-Kamiokande.
- 5 **INGRID** Interactive Neutrino GRID.
- 6 **MR** Main Ring.
- 7 **NC** Neutral Current.
- 8 **ND280** Near Detector at 280 m.
- 9 **SK** Super-Kamiokande.
- 10 **T2K** Tokai-to-Kamioka.

¹ Symbols

δ_{CP} Charge-Parity violating phase factor

1 Chapter 1

2 Introduction

3 From the postulation of the neutrino to the proposals of next generation detectors,
4 the field of neutrino physics has continuously evolved throughout its 90 year history.
5 Neutrinos are the weak isospin partners of the standard model charged leptons.
6 Existing in three flavour states, neutrinos are electrically neutral, extremely light,
7 and interact with other particles exclusively via the weak interaction. Nevertheless,
8 the neutrino is not feted for its place in the standard model, but rather its role in
9 conclusively confirming the standard model was incomplete. At the turn of the 21st
10 century, a series of discoveries provided experimental proof for neutrino oscillations.
11 The standard model of particle physics predicts neutrinos to be massless [1]. However,
12 the underlying theory for neutrino oscillations requires neutrinos to be massive, which
13 is in direct contradiction to the standard model. Such a discovery provided one of
14 the first experimental indications of physics beyond the standard model.

15 Over the last two decades an experimental program to measure the key
16 parameters that govern neutrino oscillations has been undertaken. Global fits
17 are applied to data, collated across a number of experiments, to give constraints
18 on best fit values for the oscillation parameters. These parameters, which are
19 defined and discussed in detail in Chapter 2, are summarised for the three-flavour
20 neutrino picture in table 1.1. Different experiments have varying sensitivities to
21 different oscillation parameters, often characterised by the source of neutrino (solar,
22 atmospheric, reactor, accelerator). An overview of which types of experiments
23 contribute to the present determination of oscillation parameters is shown in table

Oscillation Parameter	Normal Ordering		Inverted Ordering	
	Best Fit $\pm 1\sigma$	3σ Range	Best Fit $\pm 1\sigma$	3σ Range
$\sin^2 \theta_{12}/10^{-1}$	$3.10^{+0.13}_{-0.12}$	0.275 \rightarrow 0.350	$3.10^{+0.13}_{-0.12}$	0.275 \rightarrow 0.350
$\theta_{12}/^\circ$	$33.82^{+0.78}_{-0.76}$	31.61 \rightarrow 36.27	$33.82^{+0.78}_{-0.75}$	31.61 \rightarrow 36.27
$\sin^2 \theta_{23}/10^{-1}$	$5.82^{+0.15}_{-0.19}$	0.428 \rightarrow 0.624	$5.82^{+0.15}_{-0.18}$	0.433 \rightarrow 0.623
$\theta_{23}/^\circ$	$49.7^{+0.9}_{-1.1}$	40.9 \rightarrow 52.2	$49.7^{+0.9}_{-1.0}$	41.2 \rightarrow 52.1
$\sin^2 \theta_{13}/10^{-1}$	$2.240^{+0.065}_{-0.066}$	2.044 \rightarrow 2.437	$2.263^{+0.065}_{-0.066}$	2.067 \rightarrow 2.461
$\theta_{13}/^\circ$	$8.61^{+0.12}_{-0.13}$	8.22 \rightarrow 8.98	$8.65^{+0.12}_{-0.13}$	8.27 \rightarrow 9.03
$\delta_{CP}/^\circ$	217^{+40}_{-28}	135 \rightarrow 366	280^{+25}_{-28}	196 \rightarrow 351
$\Delta m_{21}^2/10^{-5} \text{ eV}^2$	$7.39^{+0.21}_{-0.20}$	6.79 \rightarrow 8.01	$7.39^{+0.21}_{-0.20}$	6.79 \rightarrow 8.01
$\Delta m_{32}^2/10^{-5} \text{ eV}^2$	$2.525^{+0.033}_{-0.031}$	+2.431 \rightarrow +2.622	$-2.512^{+0.034}_{-0.031}$	-2.606 \rightarrow -2.413

Table 1.1: The best fit 3ν oscillation parameters (from nu-fit [2]) to global data, published in 2019 [3]. Values assuming both normal ordering (NO) and inverted ordering (IO) are shown. All values shown been calculated to include tabulated Super-K atmospheric data measurements [4].

24 1.2. Reactor experiments measuring $\bar{\nu}_e$ disappearances from inverse β -decay provide
25 excellent constraints on θ_{13} , especially with a short-medium baseline on the order
26 of 1 km. Solar experiments have primary sensitivity to θ_{12} and Δm_{21}^2 . Longer
27 baseline reactor experiments, such as KamLAND [5], also have sensitivity to Δm_{21}^2 .
28 Both reactor and Solar experiments measure neutrinos in the few-MeV energy range.
29 With a wide range of oscillation baselines, atmospheric neutrino experiments have
30 sensitivity to most oscillation parameters but focus primarily on Δm_{32}^2 and θ_{23} .
31 Atmospheric experiments measure neutrinos through the decays of π and K mesons
32 created through cosmic ray interactions with the Earth's atmosphere. Long baseline
33 accelerator neutrino experiments use a beam of pure $\nu_\mu(\bar{\nu}_\mu)$ to measure $\nu_\mu(\bar{\nu}_\mu)$
34 disappearances, as well as $\nu_e(\bar{\nu}_e)$ appearances, at far detectors situated on baselines
35 $O(100 \text{ km})$. Measuring neutrinos on the GeV-scale, they have sensitivity to θ_{13} , θ_{23} ,
36 Δm_{31}^2 , Δm_{32}^2 , and δ_{CP} . Long baseline accelerator experimentation is the primary
37 neutrino detection method used within thesis.

38 The 20-30 years have seen a revolution in neutrino physics. Major recent
39 accomplishments include the establishment of non vanishing neutrino masses in

Experiment	Dominant	Important
Solar Experiments	θ_{12}	$\Delta m_{21}^2, \theta_{13}$
Reactor LBL (KamLAND)	Δm_{21}^2	θ_{12}, θ_{13}
Reactor MBL (Daya-Bay, Reno, D-Chooz)	$\theta_{13}, \Delta m_{31,32}^2 $	$\theta_{13}, \theta_{23}, \Delta m_{31,32}^2 , \delta_{CP}$
Atmospheric (SK, IC-DC)		
Accel. LBL $\nu_\mu, \bar{\nu}_\mu$ Dissapp. (K2K, MINOS, T2K, NO ν A)	$\theta_{23}, \Delta m_{31,32}^2 $	
Accel. LBL $\nu_e, \bar{\nu}_e$ App. (MINOS, T2K, NO ν A)	δ_{CP}	θ_{13}, θ_{23}

Table 1.2: The different experiments contributing to the current determination of best fit oscillation parameters. LBL and MBL define long and medium baselines respectively. Reproduced from [1].

40 oscillation experiments, which in turn has led to a solution to missing solar neutrinos.
41 Nevertheless, there are several fundamental questions that remain unanswered.
42 Most notably is the existence, and magnitude, of CP violation in the leptonic sector.
43 CP violation is primarily characterised by the δ_{CP} parameter. Currently T2K
44 and NO ν A have sensitivity to δ_{CP} and can provide hints and constraints on the
45 magnitude, but do not have the sensitivity to confirm CP-violation. The future long
46 baseline neutrino experiments Hyper-Kamiokande and DUNE, with larger detectors
47 and more sophisticated detection techniques, have the measurement of δ_{CP} as a
48 primary goal.

49 The objectives of this thesis can be summarised in two distinct projects:
50 The first is a cross-section measurement to constrain far detector processes using
51 near detector data; the second is a research and development project in optical
52 calibration of water Cherenkov detectors. Both projects are allied towards a common
53 overarching experimental goal of constraining and measuring CP-violation in long
54 baseline neutrino experiments. In detail, chapter 2 will begin with a brief history
55 of experimental neutrino physics, before delving into a discussion of the theoretical
56 models behind neutrino oscillations and neutrino-nucleus interactions. Detailed over-
57 views of the current long baseline water Cherenkov accelerator experiment T2K, and
58 the next generation sister experiment Hyper-Kamiokande, are provided in chapters 3
59 and 5 respectively. The first measurement of the ν_e CC π^+ interaction cross-section
60 on a carbon target is introduced in chapter 4, and provides the preliminary insights
61 into data excesses observed in the T2K far detector. Chapter 6 summarises the

62 research and development into diffuser technology for optical calibration systems
63 currently proposed for Hyper-Kamiokande. This chapter will examine the perform-
64 ance of diffuser systems in the context of both laboratory measurements, and recent
65 deployments in the Super-Kamiokande detector. Finally, chapter 7 will discuss the
66 research and results presented throughout the thesis, closing with a summary of
67 potential avenues for future research.

1 Chapter 2

2 Background

3 2.1 A Brief History of Neutrinos

4 The postulation and eventual discovery of the neutrino arose from interrogating the
5 method in which beta decay occurs. In 1930 the beta decay process was thought of
6 as being the transition of a neutron to proton with the emission of an electron in the
7 form:



8 As an example of a two-body decay process, conservation of energy and momentum
9 requires the energy spectrum for the emitted electron to be theoretically monotonic
10 in shape, appearing likely as a sharp peak. Despite this, empirical data revealed
11 the observed energy spectrum was a wide distribution symptomatic of a 3-body
12 decay. A continuous distribution contradicting the 2-body decay picture thus caused
13 a significant problem for the scientific community at the time. Furthermore, the
14 suggested beta decay process also violated angular momenta conservation when
15 including spin, as a single spin $\frac{1}{2}$ particle cannot produce a final state consisting of
16 exactly two spin $\frac{1}{2}$ particles.

17 Later in that year, Wolfgang Pauli would propose a solution to this problem.
18 Pauli postulated the existence of a third outgoing particle, which he christened the
19 "neutron", thereby generating a three-body decay process. This third particle would
20 take the form of a neutral fermion and was hypothesised to be light and minimally

21 interacting. In 1932 the term "neutron" was given to the newly discovered neutral
22 nucleon, and thus from then on Pauli's particle was known as the neutrino - meaning
23 "little neutral one" in Italian.

24 Given the neutrinos weakly interacting nature, it took a further 20 years
25 for the first experimental evidence of the neutrino's existence. Published in 1952,
26 Rodeback and Allen used the electron capture of ^{37}Ar to measure the recoil energy
27 of the nucleus [6]. But it wasn't until 1956 and the advent of nuclear fission reactors
28 that Reines and Cowan published the discovery of the neutrino [7]. Reines and
29 Cowan used close proximity with the Savannah River nuclear reactor, among the
30 strongest source of (anti)neutrinos at the time, to measure the reaction:

$$\bar{\nu}_e + p \rightarrow e^+ + n \quad (2.2)$$

31 A coincidence of the 511 keV photon associated with the outgoing positron annihila-
32 tion and a neutron capture reaction a few μs later would signal a detection. The
33 experiment consisted of a water tank with dissolved CdCl_2 . Surrounding the tank
34 two liquid scintillators were used to detect both the photons produced from the
35 positron annihilation, as well as from the $^{113}\text{Cd}(n,\gamma)^{114}\text{Cd}$ reaction after neutron
36 capture [8]. The experiment demonstrated that an increased signal was seen when
37 the reactor was running relative to when it was dormant, an observation attributed
38 to the neutrino's discovery. Reines' and Cowan's achievement would be acknowledged
39 with Frederick Reines receiving the 1995 Nobel Prize in Physics, 21 years after the
40 death of Clyde Cowan.

41 Reines and Cowan had successfully discovered the anti-electron neutrino ($\bar{\nu}_e$)
42 yet the story wasn't finished. In 1962 at the Brookhaven National Laboratory, the
43 muon neutrino (ν_μ) was discovered [9]. The experiment used a proton beam to
44 produce pions which subsequently decay to muons and muon (anti)neutrinos¹:

$$\pi^\pm \rightarrow \mu^\pm + \nu_\mu^{(-)} \quad (2.3)$$

¹Note the similarity to the muon neutrino beam approach used by T2K contributing to the work presented throughout this thesis.

45 Brookhaven detected the resulting muon (anti)neutrinos using an aluminium spark
46 chamber. The sole production of only one flavour of neutrino demonstrated that
47 neutrino flavour states are distinct; work that led to Lederman, Schwartz, and
48 Steinberger receiving the 1988 Nobel Prize in Physics.

49 By the late 1970's three different lepton flavours had been discovered; in
50 contrast, despite two more decades passing, only two flavours of neutrinos were
51 known to exist. It therefore came as no surprise when the Large Electron Positron
52 collider (LEP) at CERN hinted at the existence of three light active neutrino flavour
53 states [10]. Over the next decade, searches for the missing neutrino ensued, coming
54 to an end in the new millennium when the DONUT (Direct Observation of NU Tau)
55 experiment discovered the ν_τ [11].

56 **2.1.1 Discovery of Neutrino Flavour Mixing**

57 **Solar Neutrino Problem**

58 With the discovery of the electron neutrino and the new understanding of the
59 Sun's nuclear engine through solar models, Ray Davies was inspired to study solar
60 neutrinos as a means of observing the heart of the Sun [12]. Davies headed the
61 Homestake experiment [13], named after the gold mine in which it was located 1,500
62 m underground. Homestake used a tank filled with pure C_2Cl_4 to observe an inverse
63 beta decay process converting the chlorine to argon via:



64 ^{37}Ar has a half-life of approximately 35 days; radioactive decay results in 2.82 keV
65 X-rays or Auger electrons from K-capture at a ratio of 10:90 [14]. Roughly once a
66 month the argon atoms were extracted by bubbling helium through the tank. The
67 electron neutrino flux was then estimated through the detection of its radioactive
68 decay products. Homestake observed neutrinos at a significantly lower rate than
69 accurate solar models could predict. This observation was further supported by
70 other experiments such as GALLEX [15] and SAGE [16]. Both experiments used

71 inverse-beta decay of gallium into germanium:



72 By using gallium these experiments had access to lower energy higher flux neutrinos
73 from the pp-chain in which Homestake was blind to. The lower energy threshold
74 relative to other experimental targets, such as chlorine (Homestake) and water (Super-
75 K), can be seen in figure 2.1. Interestingly, GALLEX and SAGE observed smaller
76 deficits which would suggest an energy dependence. Nevertheless all experiments
77 saw large discrepancies with the standard solar model, which became known as the
78 "solar neutrino problem".

79 **Atmospheric Neutrino Anomaly**

80 Somewhat ironically the solution for the solar neutrino problem wouldn't begin
81 by probing the Sun as a source, rather it would start through exploring neutrinos
82 from our very own atmosphere. Importantly, for chlorine and gallium experiments
83 the vast majority of solar neutrinos studied were below an energy threshold for
84 ν_μ and ν_τ charged current interactions. This can be seen in figure 2.1. Therefore
85 previous solar neutrino experiments had sensitivity only to (anti)electron neutrinos.
86 Atmospheric experiments however can observe multiple neutrino flavours produced
87 from muon decays in the atmosphere. In particular a double ratio, consisting of the
88 ratio of the ratio of predicted to measured rate of ν_μ to ν_e events, was measured.
89 Super-Kamiokande (SK), described in section 3.4, discovered that the double ratio
90 was lower than expected and the neutrino flux was a function of the zenith angle
91 [18]. Lower ratios was an indication of either ν_μ disappearance or ν_e appearance.
92 Furthermore, changing the zenith angle is equivalent to varying the distance in
93 which the neutrino propagates, thus implying the flux has a dependence on distance
94 travelled. A combination of these two phenomena led to the proposal of neutrino
95 oscillations.

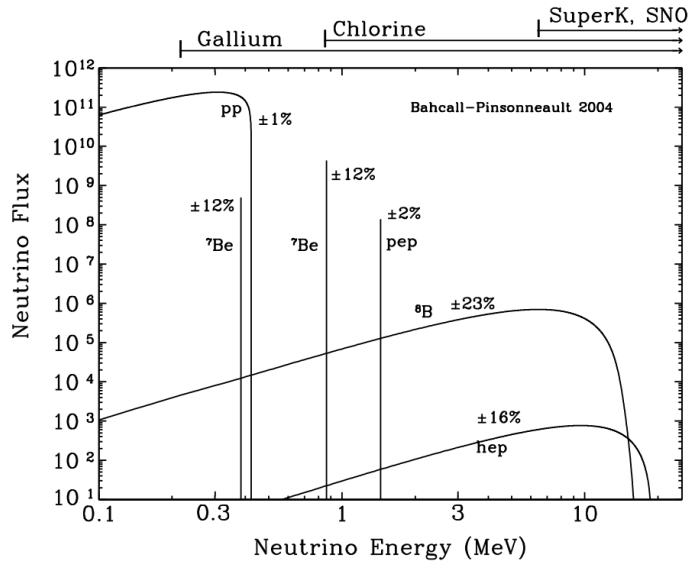


Figure 2.1: The neutrino flux emission as a function of neutrino energy for different fusion processes within the Sun. Regions of neutrino energy space in which experimental detectors are sensitive is also shown. Taken from [17].

96 Evidence of Neutrino Oscillations

97 If previous atmospheric and solar neutrino experiments were not able to measure
 98 ν_μ and ν_τ through charged current interactions, perhaps it would be possible via
 99 neutral current (NC) interactions. The Sudbury Neutrino Observatory (SNO) aimed
 100 to accomplish this through the use of heavy water as a target [19]. SNO aimed
 101 to detect solar neutrinos using Cherenkov radiation much like Super-Kamiokande.
 102 The use of heavy water however allowed SNO to exploit the flavour insensitive NC
 103 interactions on the deuterium:

$$\nu_x + {}^2D \rightarrow \nu_x + p + n \quad (2.6)$$

104 Whereby x can be anyone of the three neutrino flavour states. Furthermore the
 105 neutrons produced can interact with another deuteron, producing tritium and
 106 importantly a 6.3 MeV photon.

$$n + {}^2D \rightarrow {}^3T + \gamma \quad (2.7)$$

107 The interaction's flavour neutrality stems from the fact no charged leptons are
108 produced. Coincidences between the interactions in equations 2.6 and 2.7 could be
109 identified and tagged as NC events. The rate of neutral current interactions seen by
110 SNO matched that predicted by the solar models [19]. This led to the conclusion that
111 unseen neutrinos of previous experiments were not "missing", rather they couldn't be
112 detected as the neutrinos had changed flavour states through oscillations. Moreover,
113 SNO also probed charged current interactions and, much like experiments before,
114 measured a deficiency in neutrino flux. A combination of these two findings led to
115 the discovery of flavour changing neutrino oscillations [20] and the awarding of Nobel
116 Prize in Physics to Takaaki Kajita (Super-K) and Arthur McDonald (SNO) in 2015.

117 **2.1.2 The LSND Anomaly and Sterile Neutrinos**

118 The Liquid Scintillator Neutrino Detector (LSND) [21, 22] was an experiment that
119 took data from 1993-1998. LSND measured $\nu_\mu \rightarrow \nu_e$ oscillations over a short
120 baseline, using a 167 t mineral-oil-based liquid scintillator detector with a cylindrical
121 geometry. An excess on the predicted number of oscillations was observed at low
122 energies, which has subsequently become known as "the LSND anomaly". The
123 result was inconsistent with the atmospheric and solar results in a three-flavour
124 model. Furthermore, confirmation of only three weakly interacting neutrinos, lighter
125 than half of the Z^0 boson mass, existed from the LEP experiment [23]. The LSND
126 anomaly needed a fourth neutrino generation that was unable to couple with the
127 weak force; this became known as sterile neutrino.

128 Experiments have since attempted to test the LSND result. Most notably the
129 MiniBooNE [24, 25] at FermiLab, uses a 0.8 kt mineral oil Cherenkov detector over
130 a short baseline in an attempt to measure the same excess in low energy $\nu_\mu \rightarrow \nu_e$
131 oscillations. Whilst an excess of low energy electron-like CCQE events was found
132 [26], the signal produced by electrons and converted photons is indistinguishable.
133 Some have suggested this could explain the original LSND anomaly, nevertheless
134 searches for the sterile neutrino are ongoing.

135 2.2 Neutrino Oscillation Theory

136 The solution of the solar neutrino problem described in section 2.1.1 provided strong
137 evidence the neutrinos have mass. For non-vanishing rest masses, the weak and mass
138 eigenstates are not necessarily identical. This is a phenomena that has already been
139 studied and observed in great detail in the quark sector whereby the relationship
140 between flavour and mass states is governed by the Cabibbo-Kobayashi-Maskawa
141 (CKM) matrix [27]. Pontecorvo drew analogy to the previously observed $K^0 \rightarrow \bar{K}^0$
142 mixing and suggested that neutrinos could oscillate in a similar manner, if their
143 flavour and mass states were different [28]. The analogous $\nu \rightarrow \bar{\nu}$ process has not
144 yet been observed, but it did lay the foundation to which a full theory of neutrino
145 oscillations was formed².

146 A more general formalism without constraints on the number of flavour
147 states for neutrino oscillations can be found at [29], and sophisticated derivations
148 performed using quantum field theory can be found here [30, 31]. For the purpose of
149 this discussion it is acceptable to simplify the picture to one considering the more
150 experimentally relevant case of three different lepton flavour states ($|\nu_\alpha\rangle$, $\alpha = e, \mu, \tau$).
151 Neutrinos interact solely in these flavour states, and only propagate in the three
152 neutrino mass eigenstates ($|\nu_i\rangle$, $i = 1, 2, 3$). The flavour and mass eigenstates are
153 connected via the unitary mixing matrix, U :

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle \quad (2.8)$$

154 In the case of antineutrinos, $U_{\alpha i}$ has to be replaced by it's complex conjugate such
155 that:

$$|\bar{\nu}_\alpha\rangle = \sum_i U_{\alpha i}^* |\bar{\nu}_i\rangle \quad (2.9)$$

This unitary matrix, U , is known as the PMNS (Pontecorvo-Maki-Nakagawa-Sakata)

²It should be noted that such oscillations through neutrino flavours do not conserve individual lepton flavour numbers, only conserving total lepton number

matrix, and can be written in full generality as:

$$\begin{aligned}
 U &= \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} & (2.10) \\
 &= \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix} \\
 & & (2.11)
 \end{aligned}$$

156 Here the terse notation $c_{\alpha\beta} = \cos(\theta_{\alpha\beta})$ and $s_{\alpha\beta} = \sin(\theta_{\alpha\beta})$ is used for simplicity.
 157 The off-diagonal terms in the PMNS matrix give rise to neutrinos being created in
 158 a superposition of mass states. This mixing of states means that there is a finite
 159 possibility that a neutrino created in one flavour state may be observed sometime
 160 later as a different flavour state. Neutrinos can therefore be considered to change
 161 their flavour state through propagation. This is known by the more common term
 162 "neutrino oscillations".

163 The unitary matrix, U , can be written as the product of four sub-matrices as
 164 demonstrated above. The initial three sub-matrices are separated to contain different
 165 respective mixing angles ($\theta_{12}, \theta_{13}, \theta_{23}$). Distinct types of neutrino experiments exploit
 166 different sensitivities to each of the mixing angles. Solar and atmospheric neutrino
 167 experiments constrain values for θ_{12} and θ_{23} respectively. Reactor neutrino experi-
 168 ments have sensitivity to θ_{13} and θ_{12} . Whereas accelerator neutrino experiments can
 169 measure θ_{13} and θ_{23} . Furthermore the submatrix containing θ_{13} also contains a Dirac
 170 CP-violating phase (δ_{CP}) in which reactor and accelerator neutrino experiments will
 171 also have sensitivity to.

172 The fourth matrix is included only if neutrinos are considered as Majorana
 173 particles. It contains an additional two Majorana CP-violating phases (α_1 and α_2),
 174 but will only have physical consequences if neutrinos are their own antiparticle.
 175 Nevertheless it should be noted that even if neutrinos are Majorana, neutrino
 176 oscillations are unaffected by the Majorana CP-violating phases since oscillation

177 probability only has a dependence on UU^* , where the Majorana phases cancel out.

178 **2.2.1 Flavour Oscillation Probability**

179 When it comes to neutrino oscillation experiments it is important to design the
180 experiment to maximise the probability of oscillations. Using natural units (i.e.
181 $\hbar = c = 1$) the mass eigenstates of the neutrinos $|\nu_i(x, t)\rangle$ are stationary states and
182 can be modelled with a time dependence of

$$|\nu_i(x, t)\rangle = e^{-iE_i t} |\nu_i(x, 0)\rangle \quad (2.12)$$

183 assuming that the neutrinos are emitted by a source from $x = 0$ at time $t = 0$ with
184 momenta p , it is possible to rewrite this equation as

$$|\nu_i(x, 0)\rangle = e^{ipx} |\nu_i(x, t)\rangle \quad (2.13)$$

given the neutrinos are relativistic it is safe to make the assumption that $p \gg m_i$ and
the total neutrino energy $E \approx p$. Therefore the energy of the propagating neutrino
can be written as

$$\begin{aligned} E_i &= \sqrt{m_i^2 + p_i^2} \\ &\simeq p_i + \frac{m_i^2}{2p_i} \\ &\simeq E + \frac{m_i^2}{2E} \end{aligned} \quad (2.14)$$

185 using this and $t \approx L$, whereby L is the distance travelled by the neutrino, equation
186 2.12 can be rewritten as

$$|\nu_i(L)\rangle = e^{-im_i^2 L/2E} |\nu_i(0)\rangle \quad (2.15)$$

187 illustrating that different neutrino masses acquire a different phase factor. Because
188 neutrinos are produced and detected only as flavour states, the neutrino with flavour

189 $|\nu_\alpha\rangle$ emitted by a source at $t = 0$ propagates in time into a state $|\nu_\beta\rangle$ by

$$|\nu(x, t)\rangle = \sum_i U_{\alpha i} e^{-iE_i t} |\nu_i\rangle = \sum_{i, \beta} U_{\beta i}^* U_{\alpha i} e^{ipx} e^{-E_i t} |\nu_\beta\rangle \quad (2.16)$$

190 Combining the two equations previous, the amplitude $A_{\alpha \rightarrow \beta}$ and thus probability
191 $P_{\alpha \rightarrow \beta}$ of neutrino oscillation from state α to state β can be calculated as

$$P_{\alpha \rightarrow \beta} = |A_{\alpha \rightarrow \beta}|^2 = |\langle \nu_\beta(t) | \nu_\alpha \rangle|^2 = \left| \sum_i U_{\beta i}^* U_{\alpha i} e^{-im_i^2 L/2E} \right|^2 \quad (2.17)$$

expanding out, the transition probability becomes

$$P(\alpha \rightarrow \beta) = \sum_i \sum_j U_{\alpha i} U_{\alpha j}^* U_{\beta i}^* U_{\beta j} e^{-i(E_i - E_j)t} \quad (2.18)$$

$$= \sum_i |U_{\alpha i} U_{\beta i}|^2 + 2\text{Re} \sum_{j>i} U_{\alpha i} U_{\alpha j}^* U_{\beta i}^* U_{\beta j} \exp\left(-i \frac{\Delta m_{ij}^2}{2}\right) \frac{L}{E} \quad (2.19)$$

192 whereby $\Delta m_{ij}^2 = m_i^2 - m_j^2$. The first term in equation 2.19 represents the average
193 transition probability; the second term describes the time (or spacial) dependence of
194 the flavour oscillation. Assuming CP invariance and taking only real terms we can
195 simplify equation 2.19 to

$$P(\alpha \rightarrow \beta) = \delta_{\alpha\beta} - 4 \sum_{j>i} U_{\alpha i} U_{\alpha j} U_{\beta i} U_{\beta j} \sin^2\left(\frac{\Delta m_{ij}^2 L}{4E}\right) \quad (2.20)$$

196 This derivation means that the flavour content of the final state differs from that
197 of the initial state. Moreover, it demonstrates the importance of neutrino mass
198 differences in oscillations. Despite the small difference in neutrino masses the effect
199 can still be large at macroscopic distances. Furthermore, it can now be seen from
200 equations 2.19 and 2.20 that three flavour neutrino oscillations can be described by a
201 CP -violating phase term and mixing angles from the PMNS matrix, combined with
202 the neutrino mass state differences. These are the parameters neutrino oscillation
203 experiments aim to measure.

204 **2.2.2 CP Violation**

205 A primary objective of neutrino oscillation experiments is confirmation and meas-
206 urement of CP-violation in the leptonic sector. To explain what CP-violation is, a
207 quantum mechanical charge conjugate operator, \hat{C} , must first be introduced. The
208 operator has the function of replacing particles with their anti-particle counterparts.
209

$$q \rightarrow -q \tag{2.21}$$

210 Here the charge carried by the particle is given as q . With one half of CP represented,
211 it is natural now to define a parity operator, \hat{P} , which reverses the sign of spatial
212 dimensions:

$$t \rightarrow t \quad x \rightarrow -x \quad y \rightarrow -y \quad z \rightarrow -z \tag{2.22}$$

213 The eigenvalues of these operators can hold values of ± 1 . This is because the
214 operators return the original value when applied twice, i.e. $\hat{C}^2 = \hat{P}^2 = 1$. In
215 particular, the eigenvalues of the charge conjugate are obtained through the product of
216 the \hat{C} eigenvalues of its components. The \hat{C} eigenvalue is more commonly known as
217 C-parity; by convention, fermions and anti-fermions are given a parity of $+1$ and -1
218 respectively.

219 The standard model conserves parity and C-parity in EM and strong in-
220 teractions. However, weak interactions have been observed to violate parity [32].
221 The product of C-parity and parity is often referred to as CP. Evidence of CP
222 violation in the quark sector has been observed through Kaon decays via a minority
223 of weak interactions [33, 34]. The search for CP violation in the leptonic sector is
224 a major goal for particle physics, including neutrino oscillation experiments. CP
225 violation, alongside Baryon number and C-symmetry violation, is one of the processes
226 that could help understand the observed matter anti-matter asymmetry seen in the
227 universe.

228 **2.2.3 Matter Effects**

229 When considering experimental searches for neutrino oscillation parameters, one
230 must consider that all neutrino oscillation experiments require neutrinos to pass

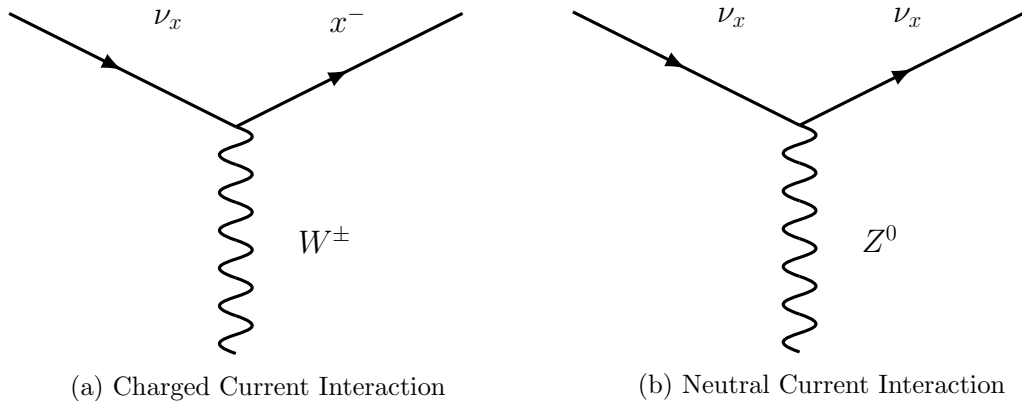


Figure 2.2: Feynman diagrams for the different types of electroweak interactions neutrinos can experience.

231 through matter either at the source, and/or through the Earth. Neutrinos are
 232 susceptible to interactions as they pass through matter, this will modify the vacuum
 233 oscillation probabilities we discussed in section 2.2. In particular, as neutrinos travel
 234 through the Sun or the Earth, ν_e can experience neutral current and charged current
 235 scatterings with leptons because of the existence of electrons in the propagation
 236 medium. Conversely ν_μ and ν_τ can only interact via neutral current scatterings.
 237 These interactions are represented in figure 2.2.

238 As was shown in equation 2.20, neutrino oscillations in a vacuum are only
 239 sensitive to the magnitude square of the neutrino mass splittings, $|\Delta m_{ij}^2|$. In addition,
 240 matter effects also have sensitivity to the signs of the mass splitting. This helps
 241 experiments determine the ordering (i.e. heaviest to lightest) in which the neutrino
 242 mass states occur. Using solar neutrino observations it has been determined that the
 243 mass state ν_2 is larger in value to that of ν_1 [35]. Nevertheless, whilst measurements
 244 of atmospheric mass splitting (Δm_{31}^2) have been made [36], it is not yet known
 245 whether ν_3 is the heaviest or lightest of the three neutrinos. These two occurrences
 246 are more commonly referred to as 'normal neutrino mass ordering' and 'inverted
 247 neutrino mass ordering' respectively, and is demonstrated in figure 2.3. The neutrino
 248 mass ordering problem is currently one of the most active areas of research and
 249 interesting questions to the field of neutrino physics. The next generation of neutrino
 250 oscillation experiments aim to exploit matter effects within the Earth to resolve the
 251 mass ordering picture.

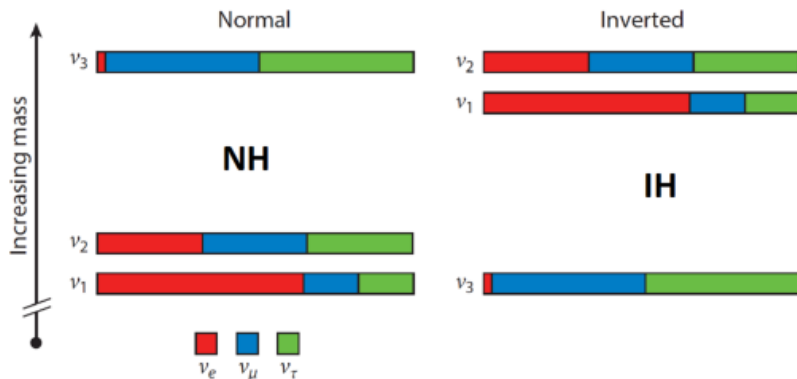


Figure 2.3: An illustration demonstrating the normal hierarchy (NH) and inverted hierarchy neutrino mass orderings. The relative proportions of flavour sharing due to mixing is also shown for each mass eigenstate. Figure is taken from [37].

252 Interestingly, antineutrinos cannot interact via the charged current channel
 253 shown in figure 2.2a. This has big implications on how the matter effect alters any
 254 neutrino oscillation measurement. In particular, this could induce an inequality in
 255 the probability for neutrino and antineutrino oscillations which one could determine
 256 to be a form of CP violation. However, this discrepancy does not inform us of
 257 the fundamental asymmetries in matter and antimatter that neutrino experiments
 258 are trying to observe. Therefore it is paramount that matter effects are taken into
 259 account when constraining values on CP violating phase factors such as δ_{CP} .

260 2.3 Neutrino-Nucleus Interactions

261 It is thought that in the near future the limiting factor in precise neutrino oscillation
 262 parameter measurements will be the systematic uncertainties in neutrino nucleus
 263 interactions. The topic of neutrino interactions is vastly complex, particularly for
 264 intermediate energies between approximately 0.1-20 GeV. In this energy region there
 265 are multiple processes competing against one another, as shown in figure 2.4. Within
 266 the lower and higher energy regions charged current quasi-elastic (CCQE) and deep
 267 inelastic scattering (DIS) processes are dominant; an important contribution to the
 268 total ν -N cross-section within the intermediate range are resonance (RES) processes.
 269 Additionally, a primary neutrino interaction can interact with multiple different
 270 components within the nucleus; a neutrino could interact with a parton, a single

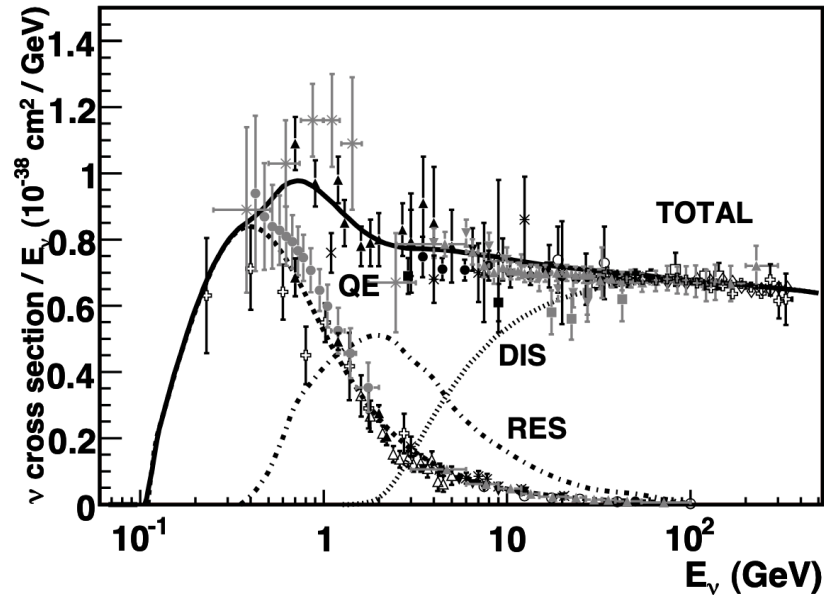


Figure 2.4: The total ν_μ cross section as a function of neutrino energy. The contributions of constituent the interaction modes; quasi-elastic (QE), deep inelastic scattering (DIS), and resonance (RES) are also shown as predicted by NUANCE [38]. Data available up until 2012 is overlaid. Figure taken from [39].

271 nucleon, or even the entire nucleus. Moreover, any component of the nucleus will
 272 also be bound in nuclear potential and have non-zero momentum which needs to be
 273 accounted for. For these reasons, good nuclear models are needed to describe the
 274 behaviour of nucleons inside a nuclear potential.

275 2.3.1 Nuclear Models

276 The simplest nuclear model is the Fermi Gas (FG) model [40] which assumes that
 277 the nucleons are bound in some average nuclear potential and are only co-dependent
 278 on each other through the Pauli exclusion principle. Illustrated in figure 2.5 the
 279 FG picture models the nuclear potential as a rectangular well which is shallower
 280 for protons due to their electromagnetic repulsion. Within the potential, nucleons
 281 occupy discrete energy states up to their respective Fermi energies. In reality this
 282 picture is flawed and can only theoretically exist in temperatures of absolute zero.
 283 Advancements upon the FG model exist and are being tested by current experiments:
 284 Examples of these include the Relativistic Fermi Gas (RFG) [41] model and the

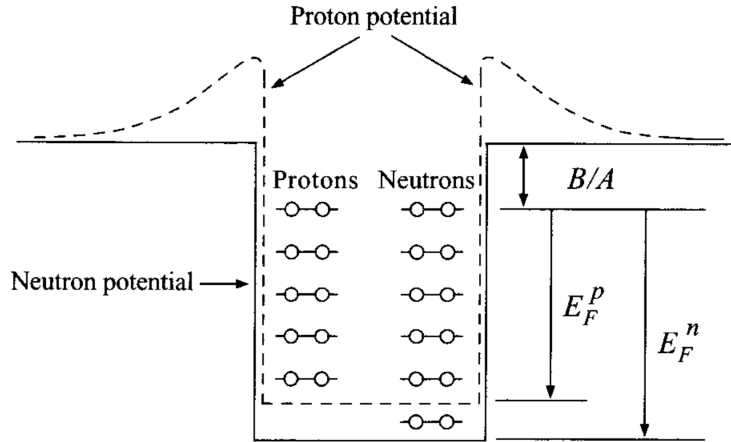


Figure 2.5: The nuclear potential energies for protons and neutrons according to the Fermi Gas Model. The Fermi energies E_F^p and E_F^n are shown for protons and neutrons respectively, as well as the binding energy B/A .

285 Spectral Function (SF) model [42]. The RFG model extends the FG model to include
 286 relativistic kinematics, whereas the SF model takes a new approach by including
 287 nuclear shell structure models to determine nuclear momentum probability densities
 288 [43].

289 2.3.2 Neutrino Interactions in Nuclei

290 The primary interaction is the first action along a chain of events that have to
 291 be accounted for when considering neutrino-nucleus interactions and cross-section
 292 measurements. There are multiple ways in which a neutrino can interact with a
 293 nucleus. This thesis will provide an account of the four main process in which
 294 neutrinos can primarily interact. These are charged current quasi elastic, which is
 295 the predominant interaction limiting the T2K oscillation analysis; deep inelastic
 296 scattering, resonance and coherent interactions which are the dominant sources
 297 within the cross section analysis described in this thesis.

298 Charged Current Quasi-Elastic

Charged Current Quasi-Elastic (CCQE) and Neutral Current Elastic (NCE) scatterings occur when the neutrino scatters off an entire nucleon, usually liberating

it from the nucleus. For NCE the leptonic component of the interaction remains the same through a Z^0 boson exchange. Whereas for CCQE the exchange of a W^\pm boson causes the incoming neutrino to change into its counterpart lepton particle with identical flavour. These are represented by the following interaction modes:

$$\nu + N \rightarrow \nu + N \quad (2.23)$$

$$\nu_l + n \rightarrow l^- + p \quad (2.24)$$

$$\bar{\nu} + p \rightarrow l^+ + n \quad (2.25)$$

299 Elastic and quasi-elastic scattering interactions are of particular importance to
 300 the T2K neutrino oscillation analyses as CCQE scatterings dominate the area of
 301 kinematic space below ~ 1.5 GeV, which corresponds to the neutrino energy region
 302 used to exploit the first oscillation maxima. Calculating the cross-sections of such
 303 processes analytically can be very challenging. Nevertheless, parametrising the
 304 cross-section is possible through the Llewellyn Smith model [44]. The parameters
 305 can then be measured through electron scattering and β -decay measurements.

306 **Resonance**

The largest contributing mechanism for pion production in neutrino interactions, excluding DIS, is that of Resonance production (RES). This is particularly important for the cross-section analysis described in chapter 4 which involves π^+ production in the final state. In resonance interactions the incoming neutrino excites the nucleon to a baryonic resonance. The resonance state then decays back to the ground state, liberating a new final state particle. The most typical resonance state occurs when the neutrino-nucleus interaction centre of mass energy is greater than the mass of a Δ (1232) baryon, which then decays to produce a single pion. For this process the charged current interaction channels are:

$$\nu_l + p \rightarrow l^- + p + \pi^+ \quad (2.26)$$

$$\nu_l + n \rightarrow l^- + n + \pi^+ \quad (2.27)$$

$$\nu_l + n \rightarrow l^- + p + \pi^0 \quad (2.28)$$

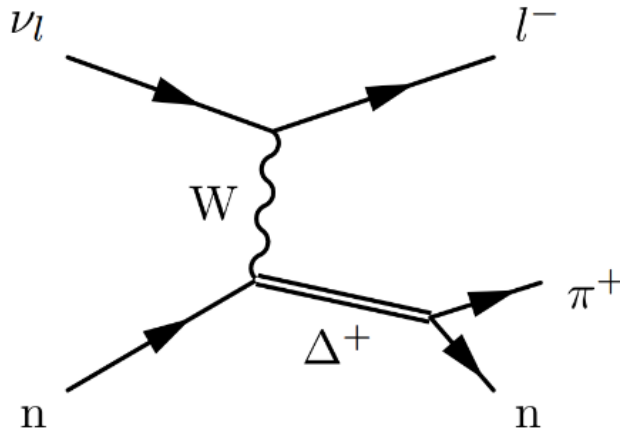


Figure 2.6: An example of a resonance interaction resulting in π^+ production. Figure taken from [45].

Neutral current resonance processes are also possible but do not lead to the formation of a charged lepton in the final state. Instead the neutral current resonance production of single pions is described by four processes:

$$\nu_l + p \rightarrow \nu_l + p + \pi^0 \quad (2.29)$$

$$\nu_l + p \rightarrow \nu_l + n + \pi^+ \quad (2.30)$$

$$\nu_l + n \rightarrow \nu_l + p + \pi^- \quad (2.31)$$

$$\nu_l + n \rightarrow \nu_l + n + \pi^0 \quad (2.32)$$

307 Whilst single pion production is most common, higher resonances also have the ability
 308 to produce kaons, photons, other mesons, as well as multiple pions. A representation
 309 of process 2.27 is also shown as a Feynman diagram in figure 2.6

310 As shown in figure 2.4, resonance production is an important interaction
 311 mode for neutrinos of energies between 1.5 GeV and 5 GeV. Neutrino interaction
 312 simulations typically describe resonance production through the Rein-Seghal model
 313 [46].

314 Deep Inelastic Scattering

315 As neutrino energies get higher the neutrino begins to be able to resolve the internal
 316 structure of the nucleon. Interactions with individual quarks via W or Z boson

317 exchanges can break apart the nucleon and produce a jet of hadrons. This process
 318 is known as Deep Inelastic Scattering (DIS) and becomes the dominant neutrino-
 319 nucleus interaction mode above approximately 5-10 GeV (see figure 2.4). DIS is
 320 also a significant process for pion production and is thus an important factor in the
 321 analysis described in chapter 4.

322 Deep Inelastic Scatterings are well understood for high energy neutrinos
 323 given the historical nature of using DIS as a means to validate the standard model
 324 and probe nuclear structure [47–49]. However, there is less understanding of how
 325 RES merges into DIS at the lower energies more relevant to long baseline neutrino
 326 experiments. It is also relatively unclear the accuracy with which current DIS models
 327 can be extrapolated to these lower energies.

328 **Coherent Scattering**

Another method in which pions can be produced is via both neutral and charge
 current Coherent scattering (COH). In coherent pion production the neutrino scatters
 off the whole nucleus, producing a single pion at a small angle relative to the incident
 neutrino:

$$\nu_l + A \rightarrow \nu_l + A + \pi^0 \quad (2.33)$$

$$\nu_l + A \rightarrow l^- + A + \pi^+ \quad (2.34)$$

329 The recoiling nucleus does not fragment and remains in the ground state. This
 330 interaction is only possible at low Q^2 , therefore at neutrino energies relevant to
 331 long baseline neutrino experiments COH scatterings have very small interaction
 332 cross-sections. Coherent interaction simulations are most often modelled with the
 333 Rein-Seghal coherent model [46].

334 **2.3.3 Final State Interactions**

After the primary interactions have occurred the end products then need to propagate
 through the nucleus before they escape into the detector and measurements can
 be made. During this time the hadrons have the possibility to re-interact inside

the nuclear medium. These interactions are known as the Final State Interactions (FSI). At neutrino energies most relevant to long baseline neutrino experiments, the π mesons are the most common form of hadrons produced in primary interactions. Taking the π^+ meson example, the most frequent forms of FSI are elastic scattering (equation 2.35), pion absorption (equation 2.36), and charge exchange reactions (equation 2.37):

$$\pi^+ + N \rightarrow \pi^+ + N \tag{2.35}$$

$$\pi^+ + N \rightarrow N' \tag{2.36}$$

$$\pi^+ + n \leftrightarrow \pi^0 + p \tag{2.37}$$

335 Consequently, the original pion can not only be absorbed, but can also have its
 336 kinematics altered or even stimulate the emission of more hadrons inside the nuclear
 337 medium.

338 Modelling FSI is extremely complex, and imposing constraints on FSI with
 339 experimental data is also very difficult. Nevertheless attempts to model FSI through
 340 cascades has been attempted in neutrino interaction simulations [38]. In such models,
 341 each hadron leaving the interaction vertex is treated independently and a number of
 342 discrete steps are defined on route to the hadrons potential escape. The size of each
 343 step is based on the particles mean free path. At every step each FSI mode has the
 344 potential to occur based on a calculated probability. This process continues until
 345 the hadron either leaves the nucleus or is absorbed. Further details of how neutrino
 346 interaction simulations treat FSI can be found at [50, 51].

Chapter 3

The T2K Experiment

Situated on two sites on opposite sides of Japan, Tokai-to-Kamioka (T2K) [52] is a long-baseline neutrino oscillation experiment to measure parameters of the PMNS matrix that govern neutrino oscillations. On the east coast of Japan in Tokai-mura, the Japan Proton Accelerator Research Complex (J-PARC) [53, 54] provides a high purity ν_μ beam using a 30 GeV proton synchrotron. J-PARC also hosts a number of near detector facilities aimed at observing beam flux and quality before oscillations and characterising neutrino interaction processes useful for oscillation analyses. 295 km west of J-PARC lies the far detector Super-Kamiokande (SK). Stationed under Mount Ikenoyama in the Mozumi mine, Super-Kamiokande is a water Cherenkov detector measuring the status of the neutrino beam post oscillations. Figure 3.1 gives a schematic overview of T2K.

Characterising parameters of the PMNS matrix in T2K is achieved by studying both ν_μ disappearance and ν_e appearance probabilities in the far detector respectively.

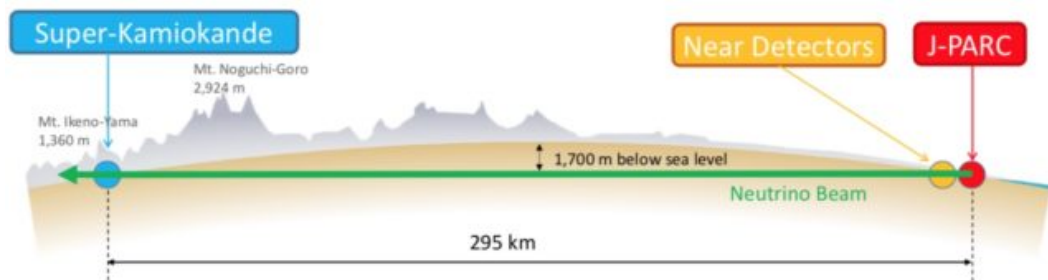


Figure 3.1: An overview of the T2K experiment.

16 In 2014 T2K became the first experiment to successfully measure the $\nu_\mu \rightarrow \nu_e$
 17 appearance channel [55]. A total of 28 events were observed at a significance of
 18 7.3σ . The latest oscillation analysis publication from T2K [56] gives up-to-date
 19 measurements on the oscillation parameters. T2K finds $\sin^2(\theta_{23}) = 0.53_{-0.04}^{+0.03}$ for
 20 both neutrino mass orderings. T2K also found, assuming the normal (inverted)
 21 mass orderings, $\Delta m_{32}^2 = (2.45 \pm 0.07) \times 10^{-3}$ ($\Delta m_{13}^2 = (2.43 \pm 0.07) \times 10^{-3}$) eV^2/c^4
 22 respectively. The best fit values for δ_{CP} and statistically dominated 1σ (68%)
 23 uncertainties, assuming normal (inverted) mass orderings, are $-1.89_{-0.58}^{+0.70}$ ($-1.38_{-0.54}^{+0.48}$).
 24 The T2K results show a preference for values of δ_{CP} that are near maximal CP
 25 Violation. Furthermore CP conserving points, $\delta_{CP} = 0$ and $\delta_{CP} = \pi$, are ruled out
 26 at a 95% confidence level.

27 Whilst not strictly one of T2K's primary goals, the near detector complex has
 28 provided T2K a means of making neutrino cross-section measurements. These meas-
 29 urements are extremely important in understanding the intricate nature of neutrino-
 30 nucleus interactions, and thus are pivotal in constraining systematic uncertainties
 31 [57]. A number of cross-section measurements have been published. Muon neutrino
 32 cross-sections have been studied over a number of target materials and final states.
 33 Published in 2013, the flux-averaged total charged current inclusive ν_μ cross-section
 34 was measured to be $\langle \sigma_{CC} \rangle = (6.91 \pm 0.13(stat) \pm 0.84(syst)) \times 10^{-39}$ cm^2 per nucleon
 35 for a mean neutrino energy of 0.85 GeV [58]. The ν_e charged current inclusive
 36 cross-section has also been published and is outlined in section 4.2.

37 **3.1 Beam**

38 **3.1.1 Proton Beam**

39 T2K's beam is provided by J-PARC's 30 GeV main ring synchrotron. A H^- beam is
 40 linearly accelerated to 400 MeV, before being converted to H^+ at the entry point to
 41 the next acceleration phase - a rapid cycling synchrotron (RCS) injection point. The
 42 protons are accelerated to 3 GeV before supplying approximately 5% to the main
 43 ring, where they are further accelerated up to 30 GeV. Each spill, consisting of eight
 44 bunches, has a spill width of 5 μs and cycles at 0.5 Hz. Fast extraction mode is used

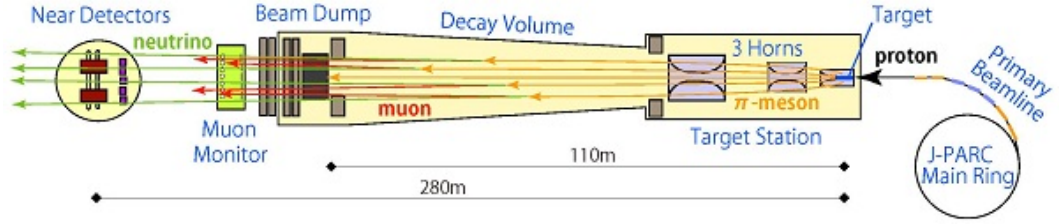


Figure 3.2: A schematic representation of the primary and secondary neutrino beamlines at J-PARC, used for the T2K muon neutrino beam. Reproduced from the J-PARC public website.

45 for the neutrino beam-line, whereby all eight proton bunches are extracted within a
 46 single turn.

47 3.1.2 T2K Neutrino Beamline

48 The T2K neutrino beamline [59] consists primarily of two separate segments shown
 49 in figure 3.2. The primary beamline takes the protons from the MR, steers them
 50 ultimately towards the far detector and collides the protons with a graphite target.
 51 The 2.6 cm diameter 1.8 g/cm³ target core has a thickness of 91.4 cm, corresponding
 52 to 1.9 interaction lengths. The core is surrounded by a 2 mm thick graphite tube
 53 which together are sealed inside 0.3 mm titanium case. Cooling from the pulsed
 54 beam heat load is provided by helium gas flowing through the gaps between the core
 55 and tube, as well as between the tube and case. Upstream of the target an Optical
 56 Transmission Radiation monitor (OTR) is used to monitor the proton beam profile.
 57 The OTR uses titanium-alloy foils placed 45° incident to the beam to produce visible
 58 light in the form of transition radiation. The light is then directed, through iron
 59 and concrete shielding, via four aluminium 90° off-axis parabolic mirrors to a charge
 60 injection device camera, producing an image of the proton beam profile.

61 The secondary beamline collects mesons from the primary beamline, provides
 62 a decay volume and finishes with a beam dump at the far end. Downstream from
 63 the secondary beam-line are three magnetic horns. Running the magnets in forward
 64 horn current mode (FHC) and reverse horn current mode (RHC) will yield beams
 65 dominated by ν_μ and $\bar{\nu}_\mu$ respectively. The reason for this is the magnetic horns
 66 are there to focus mesons with the correct charge for the (anti)neutrino production

T2K Run Number	Start Date	End Date	FHC POT (x10 ¹⁹)	RHC POT (x10 ¹⁹)
Run 1	23 Jan. 2010	26 Jun. 2010	3.288	-
Run 2	18 Nov. 2010	11 Mar. 2011	11.341	-
Run 3	08 Mar. 2012	09 Jun. 2012	16.081	-
Run 4	19 Oct. 2012	08 May. 2013	36.363	-
Run 5	21 May. 2014	24 Jun. 2014	2.465	5.145
Run 6	02 Nov. 2014	01 Jun. 2015	2.149	35.766
Run 7	01 Feb. 2016	27 May. 2016	4.890	35.272
Run 8	27 Oct. 2016	12 Apr. 2017	72.557	-

Table 3.1: T2K data information from runs 1-8 with the recorded POT in both FHC and RHC modes. Information gathered from [61].

of interest and deflect those that do not. All focussed mesons now pass through a decay volume approximately 96 m long. Any particles that do not decay under this volume will hit the beam dump removing any remaining hadrons. It is probably that muons with momenta above 5 GeV/c can pass through the beam dump; these can be measured using muon monitors, such as MUMON. Using an ionisation chamber and a Si pin photodiode, the MUMON monitors can infer the beam intensity to better than 3% and the beam direction to within 0.25 mrad [60].

The beam intensity is measured by five Current Transformers (CTs) which consist of a 50-turn toroidal coil around a ferromagnetic coil. As the protons pass through a current is induced in the coils which can be used to infer the proton flux. The fifth current transformer CT5 is stationed furthest downstream in the primary beam-line, it can therefore be used to count the number of incident protons on the graphite target. Table 3.1 reports the T2K protons-on-target (POT) figures separated by run number; a metric used for data collecting.

3.2 Off-Axis Measurement

This thesis will focus primarily on work performed in FHC mode. There are multiple decay methods in which the neutrinos can be produced in FHC mode. There are three primary meson decays from π^+ , K^+ and K_L^0 ; as well as one meson decay from μ^+ . Figure 3.3 demonstrates the predicted flux of neutrinos from the beam in

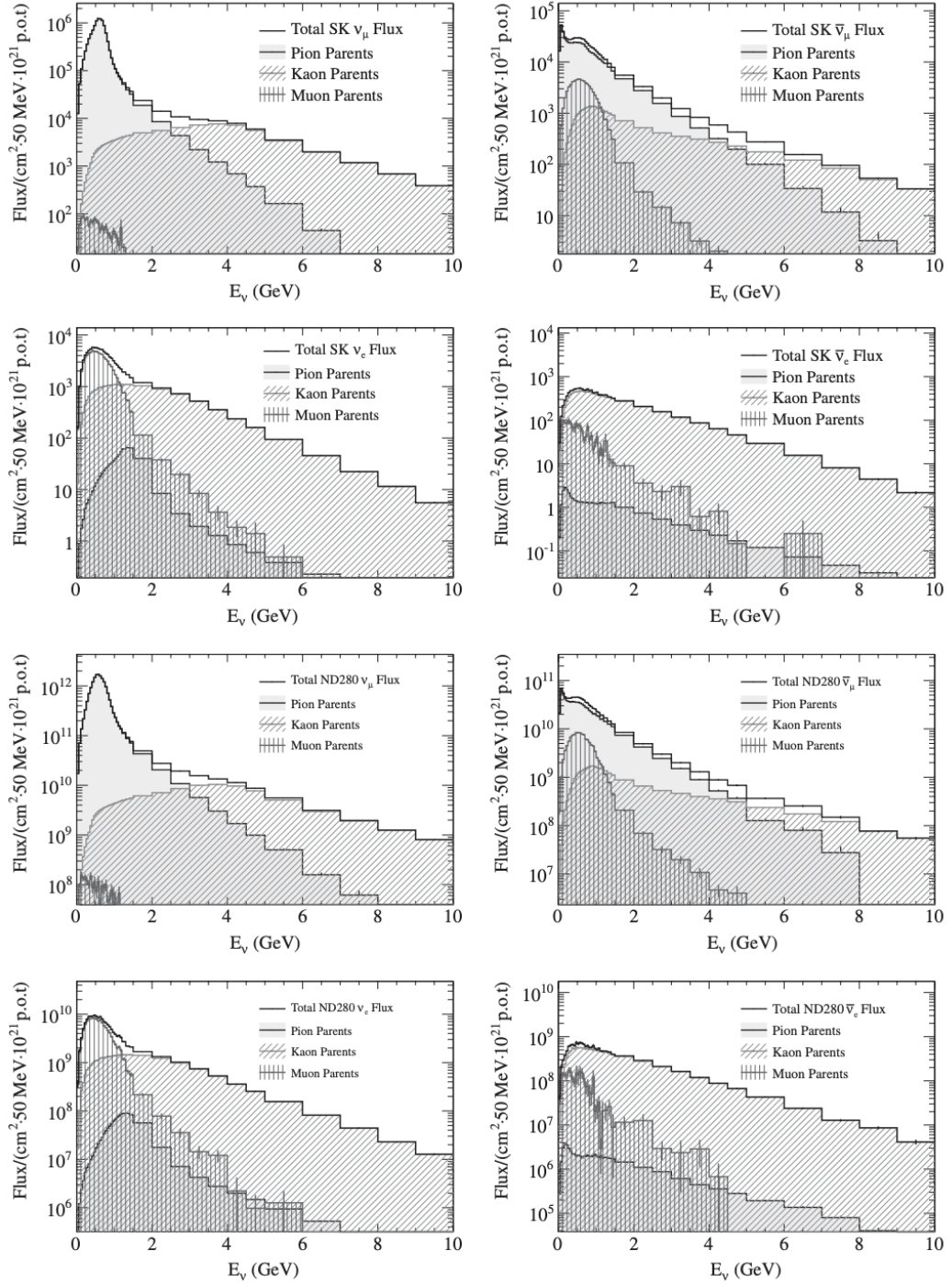


Figure 3.3: The neutrino flux prediction at ND280 and Super-Kamiokande for both ν_μ and ν_e as well as their respective antiparticles. Note due to a large MC statistical error, the error bars in most energy bins are too small to be seen. Figure taken from [59].

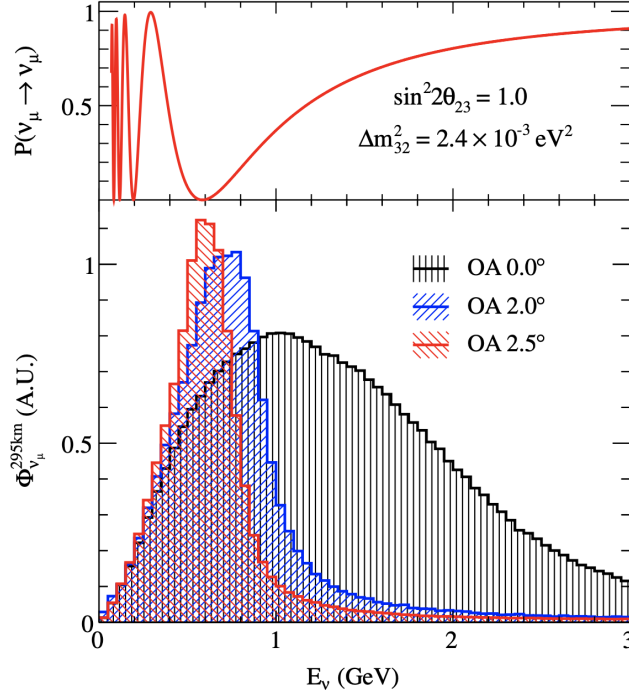


Figure 3.4: The muon neutrino oscillation probability (above) alongside the arbitrarily normalised neutrino flux (below) as a function of neutrino energy over a range of off-axis angles. This figure is used to justify a peak neutrino beam energy of 0.6 GeV and an off-axis angle of 2.5° . Taken from [59].

86 reference to the parent particle. It can be seen in this figure that the dominant parent
 87 of both ν_μ and $\bar{\nu}_\mu$ are π , with a contributions from K above 3 GeV. Pion kinematics
 88 are also integral to the energy spectrum shape of the neutrinos. Figure 3.4 shows
 89 that by moving off-axis the energy spectrum both narrows and shifts favouring lower
 90 energies. T2K was the first experiment to exploit this phenomena with the "off-axis
 91 technique", in which the far and near detector complexes are offset from the neutrino
 92 beam centre [62]. The positive effects of this are two-fold: (1) it focusses and reduces
 93 the neutrino peak energy so that it aligns with the first oscillation maximum for ν_μ
 94 disappearance channel, and (2) it reduces backgrounds present in the high energy
 95 tail, improving sensitivity to both ν_e and ν_μ appearance and disappearance channels
 96 respectively.

97 **3.3 Near Detector Complex**

98 The near detector complex is located 280m downstream of the beam target, it supplies
99 a home for two main near detectors. Located on axis is the Interactive Neutrino
100 GRID (INGRID) and off-axis in line with the far detector is Near Detector at 280m
101 (ND280). The near detector complex equips T2K with beam quality assurance and
102 control, as well as constraints on processes valuable to oscillation analyses. Both
103 detectors are situated 37 metres below ground level in an open air pit lined with
104 concrete surrounded by sand.

105 **3.3.1 INGRID**

106 Using an off-axis technique, it is critical for T2K to understand the neutrino beam
107 properties to a precision of a few percent. The Interactive Neutrino GRID (INGRID)
108 detector [63] is used to measure neutrino flux and beam direction on a spill-by-spill
109 basis to discount any beam discontinuities. INGRID is located in the near detector
110 complex 280 m downstream of the beam target and is centrally aligned to the beam
111 centre axis. The detector is constructed out of 14 identical modules orientated in a
112 cross-shape, as seen in figure 3.5. Each INGRID module consists of a "sandwich" of 9
113 iron plates and 11 scintillator layers, surrounded by veto planes on each side to reject
114 cosmic backgrounds. The modules are arranged such that 7 exist both horizontally
115 and vertically, and 2 are located at the off-diagonal. At 280 m downstream from
116 the beam target the neutrino beam width (1σ) is about 5 m, therefore INGRID was
117 designed to sample the beam in a transverse section of 10 m x 10 m. INGRID has
118 the capability of measuring beam directionality to 0.2 mrad - resolving the neutrino
119 beam centre to 5 cm. In context it has been estimated that an offset of 1 mrad yields
120 an uncertainty of approximately 2-3% on the neutrino energy scale [52]. INGRID
121 can also measure the event rate with an uncertainty of less than 2%. Additionally,
122 an extra module called the Proton Module is used to detect the muons and protons
123 produced by the neutrino beam in INGRID. The module consists of scintillator
124 planes without any iron plates. The goal of this module is constrain the quasi-elastic
125 channel for comparisons of beamline and neutrino interaction simulations. More

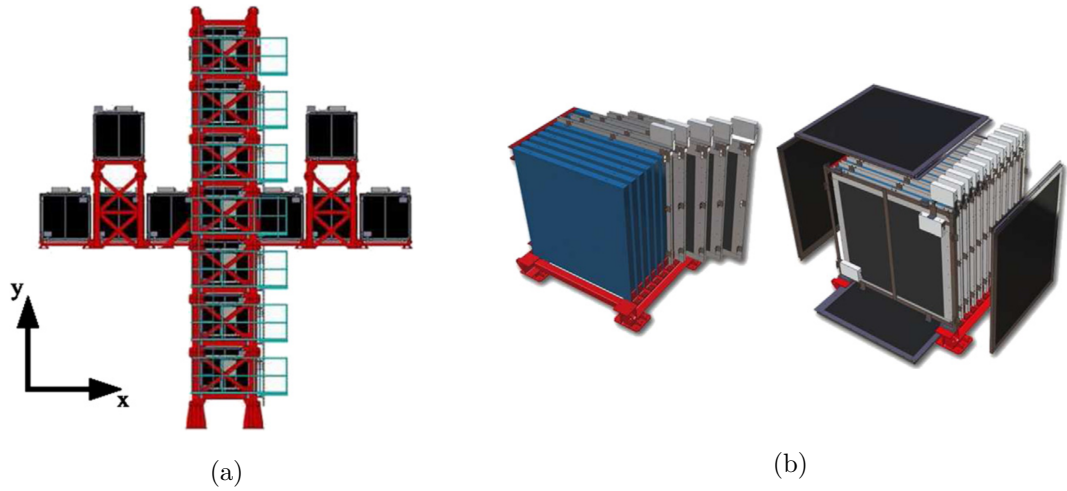


Figure 3.5: Schematic representations of both the INGRID detector (a) and the modules used inside (b). In (b) the left module (blue) shows the tracking planes, the right module (black) shows the veto planes. Taken from [52].

126 recently for a sample of T2K runs, an INGRID water module has been added giving
 127 the capabilities of measuring neutrino interactions on water [64].

128 3.3.2 ND280

129 The near detector complex also houses an off-axis detector called Near Detector
 130 at 280m (ND280). Being off-axis, the role of ND280 is to provide a measurement
 131 of the near neutrino flux to compare with the far neutrino flux observed in Super-
 132 Kamiokande. Primary goals of ND280 were to provide measurements of ν_μ and
 133 ν_e interactions for neutrino oscillation studies with the far detector. Nevertheless,
 134 due to ND280's ability to accurately track and reconstruct particles from primary
 135 scattering in the vertex. An important extra contribution of ND280 is to measure
 136 and study neutrino interaction cross-sections.

137 Figure 3.6 gives a schematic diagram of the components inside of ND280. The
 138 ND280 detector consists of 3 time projection chambers (TPCs) and 2 fine grained
 139 detectors (FGDs) arranged in an alternating pattern. This region is often referred
 140 to as the "tracker". Upstream of the tracker region a π^0 detector (P0D) is located.
 141 Surrounding the inner subdetectors are lead scintillator sampling electromagnetic
 142 calorimeter modules (ECals). The next layer of ND280 is a UA1 magnetic yoke which

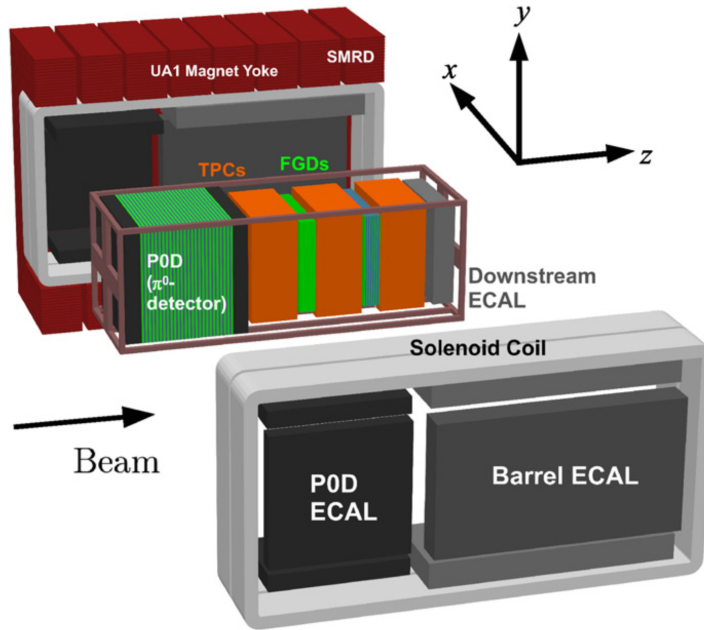


Figure 3.6: An exploded view of the ND280 detector. Taken from [59].

143 provides a 0.2 T magnetic field for accurate sign selection in the TPCs. A Side-Muon
 144 Range Detector (SMRD) made from plastic scintillator strips is interleaved within
 145 the magnetic yoke. The SMRD contributes high angle muon tracking and also acts
 146 as ND280's cosmic trigger.

147 **Pi-zero Detector**

148 The most upstream component of ND280's inner sub detectors is the π^0 -Detector
 149 (P0D) [65]. Surrounded by ECALs the P0D consists of 40 modules containing
 150 scintillator bars for tracking interleaved with lead/brass sheets. Within each module
 151 are 134 vertical bars (2200 mm long) and 126 horizontal bars (2340 mm long),
 152 arranged in perpendicular arrays. Furthermore, the P0D includes pouches that can
 153 be filled with water, giving options to run in water or air mode. The two different
 154 interaction target modes were designed to allow the P0D to measure neutrino
 155 interaction cross-sections on water through subtraction. Given the prominent π^0
 156 background in the ν_e appearance channel at Super-Kamiokande, the P0D was

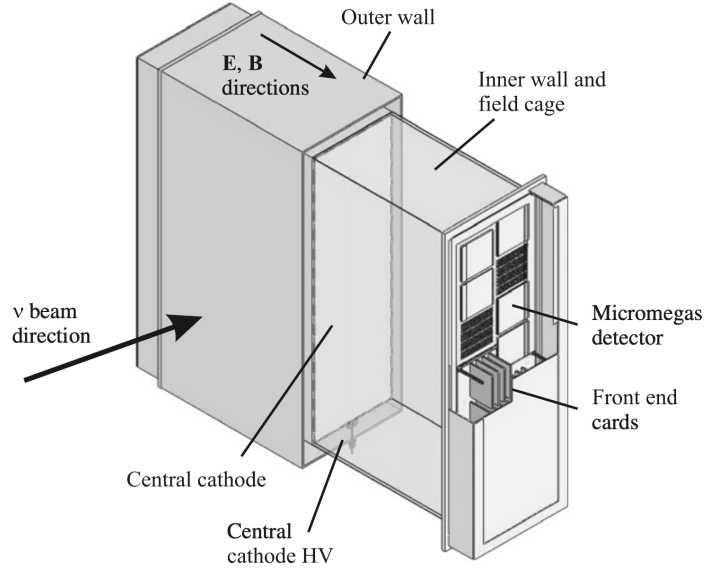


Figure 3.7: A diagram demonstrating the main aspects of the time projection chambers in ND280. Taken from [59].

157 primarily designed to measure the neutral current process:

$$\nu_\mu + N \rightarrow +\pi^0 + X \quad (3.1)$$

158 on a water target [65]. The P0D has also been used to constrain the ν_e contribution to
 159 the beam flux [66], a key intrinsic background for the oscillation analysis. Moreover,
 160 a number of cross-section analyses using the P0D as a water target are still ongoing.

161 **Time Projection Chambers**

162 Moving upstream the next sub detector is the first of three gaseous Time Projection
 163 Chambers (TPC) [67]. The TPCs are situated in an alternating sequence with
 164 the FGDs, as shown in figure 3.6. Each TPC specialises in the high-resolution
 165 tracking of charged particles. Such tracking is important to provide measurements of
 166 particles momentum and identification of particle type. The three TPCs are labelled
 167 numerically in ascending order from 1 for the most upstream TPC, and 3 for the
 168 most downstream. Each of these TPCs consists of two gas-tight boxes, one nested
 169 inside of the other. A schematic of the TPCs is shown in figure 3.7. The inner box

170 contains an argon-based gas doped with small quantities of CF_4 ($\sim 3\%$) and $i\text{C}_4\text{H}_{10}$
171 ($\sim 2\%$). To assemble a drift field in the same orientation as the ND280 magnetic field,
172 a central cathode plane dissects the inner box into two separate halves. At either
173 end wall, readout planes consisting of 12 micromegas modules [68] are placed parallel
174 to the cathode plane. Conducting strips connected by precision resistors are used in
175 the side walls to create a voltage divider and thus produce a uniform electric field in
176 the desired drift direction. The outer box, filled with CO_2 , adds gas contamination
177 protection from the atmosphere, as well as providing electrical insulation between
178 the inner box and ground.

179 As charged particles propagate through the TPCs they ionise the gas. The
180 resulting electrons then drift, away from the central cathode, towards the readout
181 planes. These electrons drift under an electric field of around 280 V/cm, over a
182 maximum distance of approximately 90 cm. Each micromegas module in the readout
183 plane consists of a two stage parallel plate avalanche chamber separated by an
184 amplification region, combined with a conversion-drift space [69]. The advantage of
185 such a design allows for the fast removal of positive ion signal produced during ava-
186 lanche, yielding the potential for sub nanosecond precision signals. Each micromegas
187 modules' objective is to record the charge and arrival time of the drifting particles.
188 These are combined over the 12 modules in each plane to produce 3D reconstructed
189 paths of traversing particles through the TPC.

190 The TPC makes up part of the ND280 tracker region, which is designed to
191 study charged current neutrino interactions. The tracking performance requirements,
192 based at 700 MeV, are to measure the transverse momentum of charged particles with
193 a resolution of 0.1 p_T or less, whereby the transverse momentum is perpendicular
194 to the magnetic field direction. Furthermore to measure ν_e interaction signal, the
195 resolution in ionisation energy loss needs to be at least 10%. This is because the
196 ionisation loss of electrons in 1 atm of argon gas is approximately 45% more relative
197 to muons around the regions of interest in momentum space [67]. To achieve such
198 goals the TPC operates in a magnetic field of 0.2 T with a sampling length of 700
199 mm, and pad segmentation of 70 mm^2 .

200 Particle identification (PID) is also a key goal of the TPCs. Distinguishing

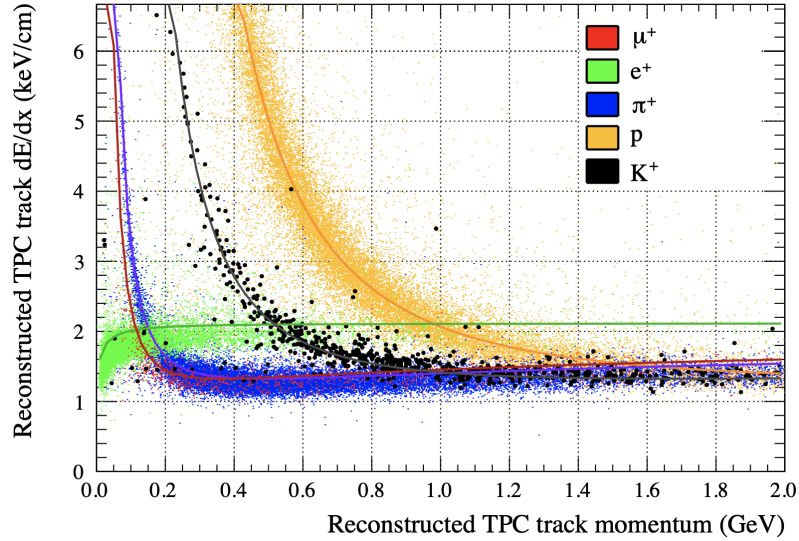


Figure 3.8: The reconstructed energy loss (dE/dx) as a function of reconstructed track momentum in the TPC. The curves show the expected distributions from calibration studies, the scatter points are reconstructed distributions from neutrino interaction simulations in ND280. Taken from [70].

201 between different particle types is achieved by measuring the energy loss (dE/dx) as
 202 a function of momentum. The amount of ionisation gives estimates on the energy
 203 loss and is characterised by the particles velocity, whilst the curvature of a track
 204 yields the particle's momentum which largely depends on both velocity and mass.
 205 These two quantities together can be used to identify particle types - which can be
 206 seen clearly in figure 3.8.

207 **Fine Grained Detectors**

208 Interlaced among the TPC modules reside two Fine Grained Detectors (FGDs)
 209 [71], labelled FGD1 and FGD2 for the most upstream and downstream detectors
 210 respectively. A single functioning unit of each FGD is an extruded polystyrene
 211 scintillator bar oriented perpendicular to the beam axis. Each scintillator bar has
 212 dimensions of 9.6 mm x 9.6 mm x 1864.3 mm, and together they are arranged into
 213 "XY modules". With 192 bars in both the horizontal and vertical direction, each
 214 module is able to achieve the fine granularity and high spatial resolution the name
 215 suggests. FGD1 is comprised of 15 modules, whilst FGD2 has seven; giving them a

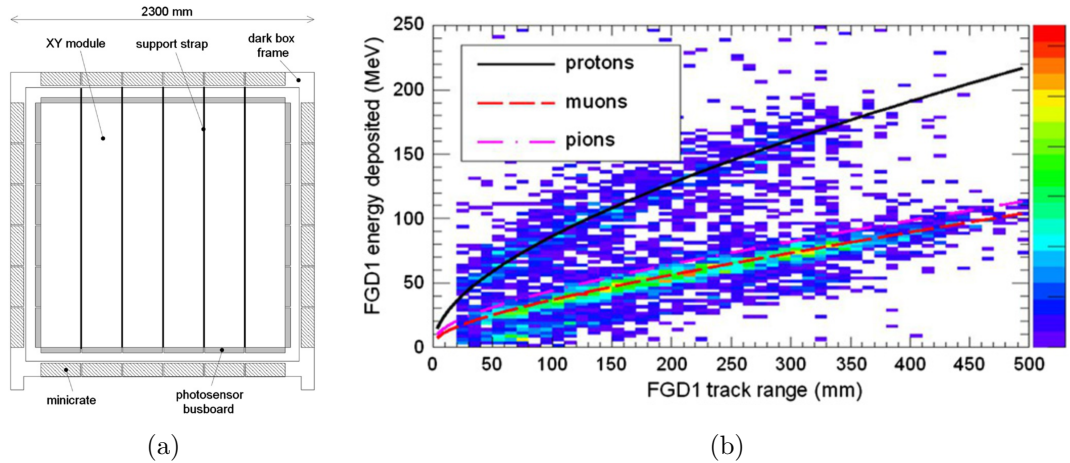


Figure 3.9: (a) A cross section schematic representation of FGD1. (b) The energy deposited in FGD1 as a function of the track range. The scatter-plot is created with stopping particles in neutrino beam data, the curves show the expectation for pions, muons, and protons from MC. Both images sourced from [71].

216 total of 5760 and 2688 scintillator bars respectively. In addition to the bars, FGD2
 217 includes 6 water target modules providing a layer of water as an interaction target.
 218 By comparing interaction rates in both FGD1 and FGD2 it is possible to discriminate
 219 cross-sections between carbon and water targets. A diagram representing the first
 220 FGD module can be seen in figure 3.9a. Each FGD has a total height of 2300 mm,
 221 a width of 2400 mm, and depth 365 mm, corresponding to the beam direction.
 222 FGD1 has a fiducial mass of 919.5 kg, which translates to $(5.54 \pm 0.04) \times 10^{29}$ target
 223 nucleons.

224 Interactions in the FGD are measured via the production of scintillation
 225 light from propagating charged particles through the scintillator bars. Light is
 226 then channelled down a wavelength shifting fibre to a 667 pixel multi-pixel photon
 227 counter (MPPC) [72–74]. The MPPC determines both the charged particle’s time
 228 of arrival, and energy deposited through the light intensity measured in number of
 229 photoelectrons recorded.

230 The primary objective of the FGDs is to provide neutrinos with potential
 231 interaction targets¹, whilst also maintaining a degree of tracking ability required by
 232 T2K for neutrino interaction rates on water. Furthermore, the FGD can also assist

¹The tracker must contain approximately 1 tonne of target mass to yield a sufficient statistical sample of events

233 the TPC in PID through measuring total energy deposited and track length. The
234 main objective is to differentiate protons from pions and muons. The distributions
235 for protons is distinct from the latter and is demonstrated in figure 3.9b.

236 **Electromagnetic Calorimeters**

237 The Electromagnetic Calorimeters (ECal) [75] are lead-scintillator sampling calori-
238 meters organised into three regions that surround ND280: the Barrell-ECal (BrECal)
239 enclosing the tracking region, and the P0D-ECal surrounding the P0D; both consist of
240 six separate modules (2 top, 2 bottom, 2 side)². Additionally, the DownStream-ECal
241 (DS-ECal) is made up of a single module and is located furthest downstream after
242 the final TPC. This equates to a total of 13 ECal modules. For the purpose of this
243 thesis the DS-ECal and BrECal will be referred to collectively as the tracker-ECal, a
244 terminology used commonly given the similar physics motivations of each region.

245 Each module is made up of multiple layers of scintillating polystyrene bars
246 bonded to lead sheets. The polystyrene bars have a cross section of 40 mm x 10
247 mm in all modules, whereas the lead sheets have a thickness of 1.75 mm and 4.00
248 mm in the tracker-ECal and P0D-ECal respectively. The size of such components
249 is constrained by the ECals position between the inner ND280 detectors and the
250 magnet, as demonstrated by figure 3.6.

251 The goal of all ECal modules is to provide a measurement of the energy of
252 particles escaping the inner tracker. Nevertheless, physics aims for the tracker-ECal
253 and P0D-Ecal modules differ from each other. The tracker-ECal is designed as a
254 tracking calorimeter providing detailed reconstruction of electromagnetic showers to
255 complement the charged-particle identification and tracking capabilities of the TPCs
256 [75]. An advantage of this is the ability to measure the energy of neutral particles
257 and assist with particle identification in the ND280 tracker. There are 31 scintillator-
258 lead layers in the BrECal and 34 layers in the DS-ECal. This equates to 10 and 11
259 radiation lengths, X_0 , a quantity that was determined to best contain electromagnetic
260 showers of photons, electrons and positrons of energies up to 3 GeV. At least 10 X_0

²The BrECal and P0D-ECal are attached to the magnet and thus must have two top and bottom modules to allow the magnet to be opened

261 are needed to ensure more than 50% of the energy resulting from photon showers
262 initiated by a π^0 decay is contained within the ECal. 3D reconstruction of tracks
263 and showers is also achieved through rotating alternate layers by 90 degrees. The
264 energy resolution for tracker-ECal modules is approximately $10\%/\sqrt{E}$ [76].

265 The role of the P0D-ECal is to tag escaping energy from the P0D and
266 distinguish between photons and muons. In contrast to the tracker-ECal, shower
267 reconstruction is not needed in the P0D-ECal as it is already performed by the P0D
268 itself. Therefore, the P0D-ECal has only six scintillator layers (approximately 4.3
269 X_0 for reference), but requires thicker lead sheets to promote the higher detection
270 efficiency of photons, the containment of showers, and that photon showers can be
271 recognised from muons.

272 **The UA1/NOMAD Magnet**

273 The magnet installed at ND280 is built around the UA1/NOMAD magnet previously
274 commissioned at CERN [77, 78]. The magnet provides a horizontally orientated
275 dipole magnetic field of 0.2 T. The dipole magnetic field is created by water-cooled
276 aluminium coils. Additionally the magnet also consists of a flux return yoke, split
277 into 2 sections each made of eight C-shaped yokes providing magnetic insulation for
278 the surrounding detector. The external dimensions of the magnet are 7.6 m x 5.6 m
279 x 6.1 m. Nevertheless, it is the internal dimensions at 7.0 m x 3.5 m x 3.6 m that
280 yield the main spatial limitations on ND280s subdetector modules [52].

281 The ND280 magnet has a key role in particle identification through measure-
282 ments of momenta and determination of the signs of charged particles, produced by
283 neutrino interactions within the TPCs.

284 **Side Muon Range Detectors**

285 The Side Muon Range Detector (SMRD) [79] is situated inside the magnetic return
286 yoke described previously. Shown in figure 3.10 the SMRD is placed in the inner-most
287 gaps and surrounds the entire ND280 ECal, P0D and tracker sections. The SMRD
288 consists of 2008 scintillator bars of dimensions 7 mm x 167 mm x 875 mm arranged
289 in 192 horizontal and 248 vertical modules. The purpose of the SMRD is to identify

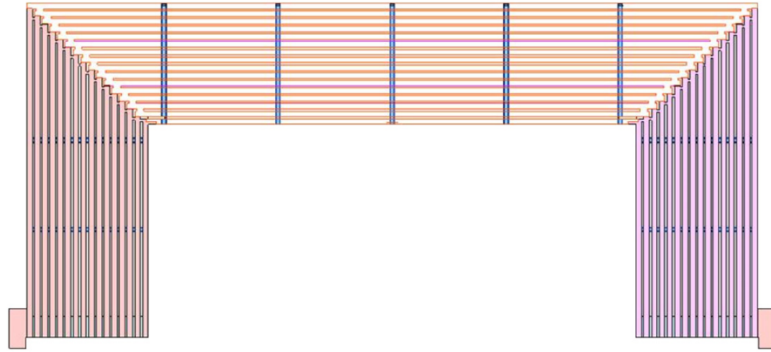


Figure 3.10: An engineers drawing of a single yoke in the UA1 magnet showing the interleaved SMRD. Adapted from [79].

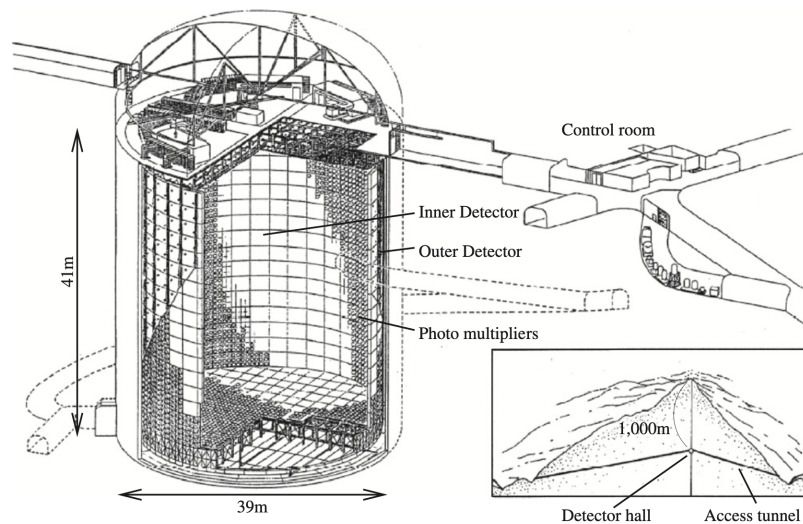


Figure 3.11: A diagram of the Super-Kamiokande detector. Taken from [81].

290 high angle muons that escape from the inner detector leaving behind little or no
 291 TPC hits. Furthermore, the SMRD acts as both a trigger and a veto for cosmic
 292 muons.

293 3.4 Far Detector

294 Located 285 km away from J-PARC inside the Kamioka mine, Super-Kamiokande
 295 (SK) acts as the far detector for the T2K experiment [80]. The mine is located
 296 1000 m deep under mount Ikenoyama. This is the equivalent of 2700 m.w.e (metre
 297 equivalent water) and thus acts as a natural shield to cosmic rays.

298 A diagram of Super-Kamiokande is shown in figure 3.11. The 41.4 m x 39.3
299 m tank has a cylindrical geometry orientated in the vertical direction. The vessel
300 is made from stainless steel and the detector is split into two coaxial cylinders
301 called the inner (ID) and outer (OD) detectors, with inner dimensions of 36.2 m
302 x 33.8 m. The ID and OD are separated by a black Tyvek sheeting. Tyvek is
303 used for it's high reflectivity, reaching a maximum value of 98.5% at wavelengths
304 of approximately 400 nm [82]. The tank is filled with ultra-pure water providing
305 a 22.5 kton fiducial volume. The ID is surrounded by 11,129 20-inch Hamamatsu
306 R3600 hemispherical photomultiplier tubes (PMTs) directed inwards of the detector,
307 providing approximately 40% photo-coverage. Moreover, the OD has 1,885 8-inch
308 Hamamatsu R1408 PMTs facing outwards - the objective here to provide a veto for
309 the inner detector. Each PMT has single photon detection capabilities and has a
310 combined quantum and collection efficiency of 20%. The working wavelength range
311 of each PMT is 350 nm - 500 nm, with a maximal quantum efficiency reached at
312 approximately 400 nm.

313 The primary method for particle detection in SK is through the production
314 of Cherenkov light from charged particles after neutrino interactions. If a charged
315 particle moves faster than the speed of light with respect to the medium it's propagat-
316 ing through, a respective cone of Cherenkov light will be emitted around its direction
317 of travel. The subsequent ring from this cone seen at any one time has signature
318 properties that can be used for particle identification. The "fuzziness" of the ring, a
319 by-product from the degree of scattering, can be used to differentiate between muons
320 and electrons. Heavier particles, such as muons and pions, will generally without
321 scattering in the medium, whereas electrons being lighter particles will scatter more
322 frequently and produce EM showers when travelling. The contrast of the two will
323 produce clear rings and fuzzier rings for muons and electrons respectively.

324 SK is also capable of detecting delayed signals from Michel electrons as well
325 as detecting charged current interactions with one charged pion in the final state.
326 Furthermore tracks with kinked trajectories are used to discern scattered pions from
327 muons.

328 In the summer of 2018, SK was drained for scheduled maintained. Further-

329 more, gadolinium doping in the water was introduced which will add the capability
330 of neutron tagging [83, 84]. Data taking has since resumed in autumn 2019.

1 Chapter 4

2 Measurement of ν_e CC π^+ with 3 the ND280 Tracker

4 4.1 Motivation

5 For long baseline neutrino oscillation experiments such as T2K and NOVA, as well
6 as the future generation experiments Hyper-Kamiokande and DUNE, ν_e charged
7 current π^+ production provides a significant contribution to the ν_e appearance
8 channel. Despite this, there is currently no measurement of exclusive ν_e charged
9 current π^+ production on a Carbon target in the literature to date.

10 For appearance studies at T2K, the far detector uses two electron neutrino
11 appearance samples in FHC mode: 1-ring ν_e CCQE, and a 1-ring ν_e CC $1\pi^+$ sample.
12 An effort to produce a 2-ring ν_e CC $1\pi^+$, where both e-like and π^+ -like rings sample
13 are reconstructed, is also in development for the oscillation analysis, but is not yet
14 implemented. The T2K far detector data collected to date displays an excess of
15 events over the background prediction in the FHC 1 decay electron sample [85]. It
16 can be seen in figure 4.1 that the expected number of events, assuming maximal CP-
17 violation, is 7; whereas 15 events are observed in data. The probability of observing
18 an excess at least this large in one of T2K's five samples is 6.9% for the best fit value
19 of the oscillation parameters [56]. Currently, T2K has no direct constraint on this
20 process from the near detector.

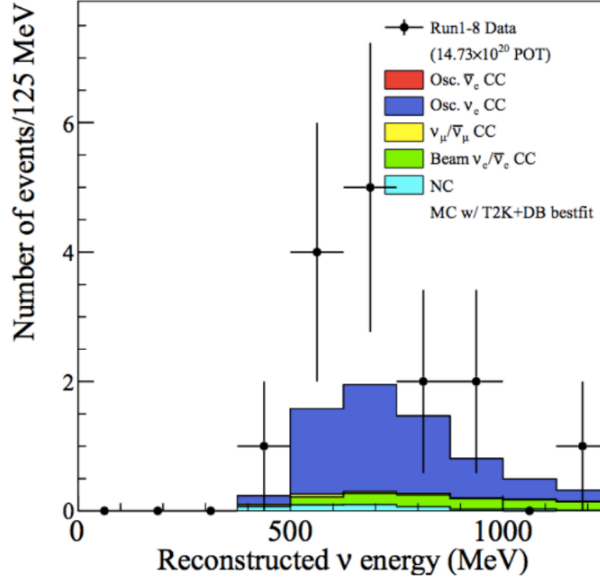


Figure 4.1: The number of events in Super-Kamiokande as a function of reconstructed neutrino energy expected for simulated MC and seen in data.

21 The ν_e CC π^+ analysis within this thesis aims to produce the world's first
 22 exclusive ν_e CC π^+ cross-section measurement on a Carbon target. Furthermore, the
 23 goal of the analysis is to develop a constraint on $\nu_e\pi^+$ background due to intrinsic
 24 ν_e contamination in the T2K beam using ND280 Tracker data. Data and MC
 25 comparisons in the low energy region of phase-space relevant to the far detector can
 26 also provide initial insights into whether an excess, similar to that observed in the
 27 far detector data, is also seen in the near detector dataset.

28 4.2 ν_e Inclusive Cross Section Measurement

29 The analysis outlined in this thesis inherits from a previous study to measure
 30 the ν_e CC inclusive cross-section using the ND280 tracker [86–88]. The primary
 31 motivation of that analysis was to develop a constraint on the intrinsic electron
 32 neutrino contamination in the T2K beam, the single largest background in the
 33 measurement of electron neutrino appearances at the far detector. The inclusive
 34 analysis measured electron neutrino and anti-neutrino cross-sections, in both FHC
 35 and RHC modes. Nevertheless, in this section we will only discuss electron neutrino

MC	Measured σ [10^{-39}cm^2 per nucleon]	Nominal σ [10^{-39}cm^2 per nucleon]	$\langle E \rangle$ [GeV]
NEUT 5.3.2	$6.62 \pm 1.32(\text{stat.}) \pm 1.30(\text{syst.})$	7.18	1.28
GENIE 2.8.0	$6.93 \pm 1.40(\text{stat.}) \pm 1.33(\text{syst.})$	6.87	1.28

Table 4.1: Measurement of the ν_e inclusive cross-section result for two different MC sets, compared against the nominal predicted value. The mean neutrino energy, $\langle E \rangle$, is also shown. Reproduced from [88].

36 measurements in FHC mode, as is the most relevant to the ν_e CC π^+ cross-section
37 analysis provided in this thesis.

38 The ν_e CC inclusive signal was defined as any event that originated from
39 a charged current electron neutrino interaction in the FGD1 fiducial volume, with
40 additional phase space cuts applied to the outgoing electron; the inclusivity of the
41 signal means it is not concerned with the composition of the hadronic final state of the
42 interaction. The selection of electron neutrino candidates followed two distinct paths;
43 the first rejected large muon backgrounds¹, the second reduced a prevalent photon
44 background. The selection required excellent particle identification and was later
45 adapted for the ν_e CC π^+ analysis in section 4.7.1. The dominant background post-
46 selection comes from photon interactions primarily from π^0 decays. Approximately
47 60% of the photon background originated inside the FGD1 fiducial volume; the
48 remaining fraction had interactions occurring in other parts of the ND280 detector,
49 or through sand interactions. A significant amount of photon background was found
50 to populate the low momentum and high angle regions, which was then constrained
51 by an independent photon control selection. A summary of the selection depicted
52 significant data-MC discrepancies in regions dominated by photon backgrounds.
53 These regions are also dominated by large systematic uncertainties. Full details of
54 the ν_e inclusive and photon sideband selections can be found in [86, 87].

55 The total ν_e inclusive cross-section was measured over a limited predefined
56 phase-space ($p > 300$ MeV/c and $\theta \leq 45^\circ$) using NEUT 3.2.0 and GENIE 2.8.0
57 MC; the results have been reproduced in table 4.1. Both results agree within error
58 with the cross-section predictions given by their respective nominal MC. The data

¹Pion and proton backgrounds are also rejected here but these backgrounds are smaller in magnitude

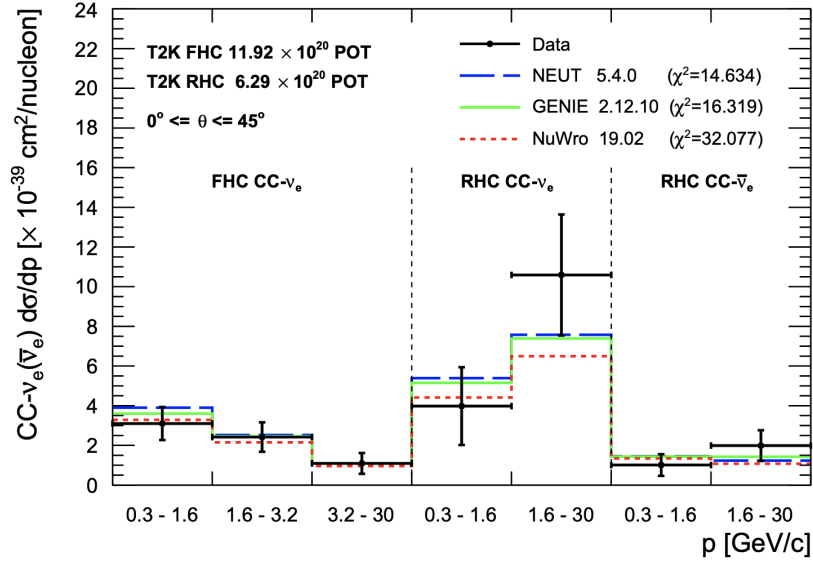


Figure 4.2: Flux integrated ν_e CC inclusive differential cross-section results, in a limited phase-space, as a function of reconstructed lepton momentum. Comparisons to different neutrino event generator models were made. Plot taken from [88].

59 was compared to cross-section predictions from recent neutrino generator models
60 in NEUT 5.4.0, GENIE 2.12.10, and NuWro 19.02. The resulting plot, split into
61 predefined regions of momenta space, is shown in figure 4.2. The best agreement
62 over both FHC and RHC for (anti-)electron neutrinos is observed with NEUT 5.4.0.
63 Nevertheless, all models agree within error for FHC electron neutrino interactions.
64 The ν_e inclusive analysis provided the first $CC-\nu_e$ cross-section measurement using
65 both FHC and RHC fluxes.

66 4.3 T2K Software

67 The T2K software framework used to perform the majority of this analysis was
68 Highland2 (HIGH Level Analysis and the ND280 version 2) [89]. Highland2 provides
69 a framework to analyse Monte-Carlo (MC) simulated and real data on an event-by-
70 event basis. Event selection and cuts are performed on reconstructed objects based
71 on their characteristics. Truth information is parsed throughout, and is used to test
72 the relative performance of the selection.

73 Detector systematic uncertainties and error propagation are calculated and

74 performed by Psyche (Parametrisation of SYstematics and CHaracterisation of
 75 Event) [89], a software package called by Highland2. The cross section and flux
 76 systematic uncertainties are evaluated using a combination of the T2KReWeight
 77 package, and parts of the nueXsLLFitter package which in turn has been used for
 78 the recent ν_e inclusive cross section result [88].

79 4.4 Data and Monte-Carlo Samples

80 The data sample analysed includes T2K runs 2-4, corresponding to 5.87×10^{20} POT
 81 after beam and ND280 data quality cuts are applied. In addition, T2K run 8 data is
 82 used giving an additional 5.73×10^{20} POT. The total exposure of the dataset used
 83 in this analysis is 11.60×10^{20} POT. The full set of ND280 MC produced for runs
 84 2, 3, 4, and 8 was used. The exposure for this MC sample was 7.38×10^{21} POT
 85 for water in and 11.59×10^{21} POT for water out configurations, and a total of
 86 18.96×10^{21} POT for the full MC sample. Only FHC mode data and MC are used
 87 for this analysis, therefore RHC runs 5-7 are not considered. Both MC and data
 88 samples were processed in T2K's software production 6T.

89 4.5 Signal Definition

90 The analysis presented in this chapter aims to measure the cross-section of charged
 91 current ν_e interactions that produce at least one positively-charged pion in the
 92 detector (after FSI). The following signal criterion are imposed on this analysis:

- 93 • The event must include an electron neutrino charged current interaction in
 94 the FGD1 sub-detector fiducial volume (FV). The FGD1 FV cut dimensions
 95 are $|x| < 874.51$ mm, $|y - 55| < 874.51$ mm, and $136.875 < z < 446.955$ mm.
 96 Where the x and y cuts are defined to match the outer boundaries of the
 97 central 182 scintillator bars², and the z cut is placed just after the first XY
 98 module but includes all remaining downstream modules [90].

²The 55 mm accounts for an offset relative to the ND280 coordinate system

- 99 • The interaction must produce an electron and at least one positive pion must
100 exit the nucleus.

101 Any events that pass all of these criteria are defined as signal. For the cross-section
102 measurement, additional phase-space constraints applied to the signal definition;
103 these are outline further in section 4.9.1.

104 It should be noted that the analysis is not designed to select pions stopping
105 in the FGD. Therefore when quantifying selection efficiency performance, a third
106 signal criterion is imposed requiring at least one positively charged pion to pass from
107 the FGD1 to the neighbouring downstream TPC. It should be noted that events in
108 which the positively charged particle is mis-reconstructed as a pion, for example a
109 CC- ν_e event whereby the π^+ is isolated in the FGD1 but the proton escapes, are
110 defined to be signal events.

111 4.6 Significant Background Topologies

112 There are multiple background topologies that impose significant contributions to
113 this analysis. For clarity, these will be grouped and defined as such:

- 114 • The most prominent background topologies in the analysis come from the
115 production of π^0 which consequently decay into photons. There are a number
116 of processes that can produce π^0 which each can mimic signal in different ways.
117 These are:

- 118 – **γ background OOFGD** - Interactions that occur outside of the FGD1,
119 which produce π^0 that decay to photons. A significant contributing
120 background to the ν_e inclusive analysis (40%), these background events
121 can often mimic ν_e interactions as external photons interact in the FGD
122 to produce electron positron pairs.

- 123 – **γ background OOFGDFV** - Analogous to OOFGD photon background,
124 this topology is based on photon production within the FGD, but outside
125 the fiducial volume defined in section 4.5.

- 126 – **γ background ν_μ CC** - This photon background channel consists of
127 events in which a ν_μ CC π^0 interaction that occurs in the FGD. This
128 channel can mimic signal for a number of reasons, the most frequent
129 being poor or no reconstruction of high angle muons in the FGD and/or
130 TPC. Instead the selected lepton track is the final electron from the
131 $\pi^0 \rightarrow 2\gamma \rightarrow e^+e^-$ decay chain. True positive pions are sometimes present
132 in the event, but often positrons and protons are selected as the pion
133 candidate track at low and high track momenta respectively. This occurs
134 as the dE/dx curves used for particle identification intersect one another,
135 as can be seen in figure 4.3. The true particle selected for the pion
136 candidate track for this background can be seen in the appendix, figure
137 A.1a.
- 138 – **γ background NC** - This topology is defined as π^0 production via neutral
139 current interactions. Alongside the aforementioned ν_μ CC photon back-
140 ground, these two topologies make up the dominant photon background in
141 this analysis, and mimic signal in similar manners. The outgoing neutrino
142 remains undetected, whilst the selected lepton remains the electron from
143 pair production. The true pion candidate track for this topology as a
144 function of reconstructed momenta is shown in figure A.1b.
- 145 • The ν_e charged current background is split into two sample topologies. The
146 largest single ν_e CC background contribution comes from ν_e CC 0π interactions.
147 All other charged current topologies are defined as ν_e CC other.
 - 148 • At ND280 the ratio of the total ν_e flux to total ν_μ flux, integrated over all energy
149 space, is approximately 0.012. Naturally, the predominant background before
150 selection cuts is from ν_μ interactions. The initial selection cuts, inheriting
151 from the ν_e inclusive analysis, are designed to inhibit this background and
152 promote selection of electrons as the main lepton track. Nevertheless, muons
153 are occasionally selected as the lepton track, a problem more prominent within
154 this analysis at higher energies.

155 **4.7 ν_e CC π^+ Selection**

156 The selection for this analysis inherits from the event selection used for the ν_e
157 charged current inclusive measurement which is described in reference [86–88]. The
158 principal philosophy of the ν_e CC π^+ selection was to stay as close to the ν_e inclusive
159 selection as possible, to allow for comparisons to be made where necessary. All plots
160 are created with P6T nominal NEUT 5.4.0 MC and, unless otherwise stated, are
161 normalised to data by POT on a run-by-run basis. Plots that do not show data,
162 typically representing truth level information, are normalised by total POT to the
163 full data set.

164 **4.7.1 Selection Cuts**

165 A number of cuts are used during the selection to create a sample of events that
166 maximises both selection efficiency and purity. The efficiency of a sample is the
167 percentage of true signal events that remain in the sample post selection, relative
168 to the number of true signal events pre selection. One can also measure the quality
169 of a sample by the purity, defined as the fraction of signal events in the sample. In
170 general, the addition of selection cuts increases signal purity at a cost to efficiency.

171 Starting from the ν_e inclusive selection, given in [86, 87], a number of cuts
172 designed to reduce out-of-fiducial-volume (OOFV) photon background were removed.
173 This increases signal efficiency, and can be done because the addition of selecting
174 over a pion track naturally has the same effect. The final cuts for the ν_e CC π^+
175 selection are described below in the order in which they are implemented in the
176 analysis. Unless otherwise stated, each plot demonstrating individual selection cuts
177 is taken with $N - 1$ cuts (all cuts but the one in question) applied.

178 **Beam and ND280 Event Quality**

179 The event must pass both T2K’s beam quality and ND280’s data quality cuts.
180 Furthermore, the event time has to be reconstructed within one of the eight distinct
181 beam bunches.

182 **Track Multiplicity**

183 At least two tracks must pass into the TPC downstream of FGD1.

184 **TPC Track Quality**

185 The TPC track quality cut is taken from the ν_e CC inclusive analysis. The most
186 energetic negatively charged track that starts in the FGD fiducial volume is selected
187 as the primary lepton. If this track passes into one of the Tracker-ECal modules, it
188 is required to have at least 18 reconstructed hits within the TPC. Otherwise, it must
189 contain 36 TPC hits. The minimum number of hits required is based on a previous
190 study outlined in [91].

191 **Particle Identification in the TPC and ECal**

192 The particle identification cuts for the electron neutrino beam component, using the
193 TPC and Tracker-ECal sub-detectors, was originally developed in 2013 [91]. The
194 PID cuts are based on the measurement of the truncated mean of the ionisation
195 loss (C_T) by the charged particles as it crosses the TPC gas. The mean value of
196 the charge deposited by charged particles on MicroMegas columns across the TPC
197 is computed. A truncated mean is used to avoid distributions being affected by
198 Landau tails from ionisation processes in the gas. Only the mean value of 70% of
199 the MicroMegas columns recording the least charge are considered. This value has
200 been optimised and TPC performances are outlined in [92, 93]. The distribution of
201 the truncated mean versus the reconstructed TPC momentum is shown for positive
202 and negative tracks in figure 4.3. To perform particle identification a pull variable,
203 δ_i , is defined as:

$$\delta_i = \frac{C_T^{meas} - C_T^{exp}(i)}{\sigma^{exp}(i)} \quad (4.1)$$

204 where i represents different particle species hypotheses, C_T^{meas} and C_T^{exp} are the
205 measured and expected energy losses of particle i respectively, and σ^{exp} is the
206 resolution of the deposited energy measurement.

207 Pull values for cuts have since been defined for the ν_e inclusive analysis, and
208 are outlined in [87]. For this analysis, pull values were once again tuned to maximise

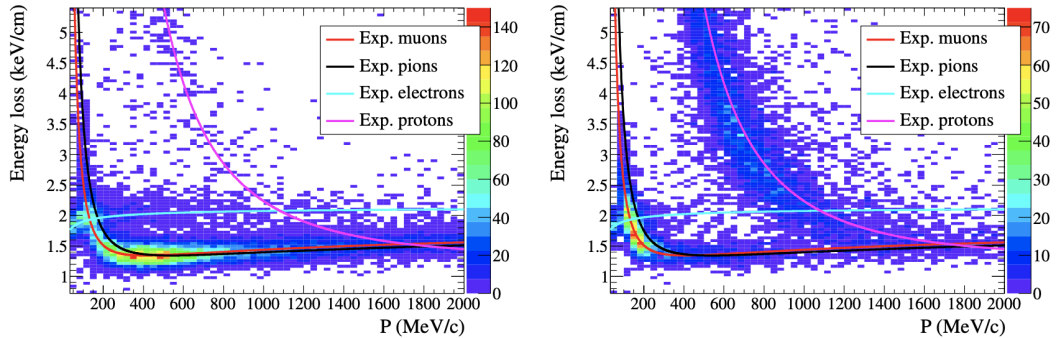


Figure 4.3: Distribution of TPC ionisation loss as a function of reconstructed TPC momentum. The distribution is for the candidate lepton track starting within the FGD fiducial volume. Negative tracks are shown left, positive tracks are shown right. The expected curves for typical particle types are superimposed.

209 signal efficiency and purity, using nominal control MC which is then tested on a
 210 different set of MC.

211 If the selected lepton candidate track does not pass into the ECal, it is rejected
 212 if it fails any of the following cuts.

213 • $-1.5 > \delta_e$, or $\delta_e < 2.5$

214 • $-2.5 < \delta_\mu < 3.0$

215 • $-2.5 < \delta_\pi < 3.0$

216 These cuts are shown in figure 4.4.

217 In addition to the TPC pull cuts, if the momentum of the selected track is
 218 less than 300 MeV/c and enters the ECal, it must pass the ECal MIP-Shower PID
 219 cut, $\text{MIPeM} > 0$. If the selected lepton tracks momentum is above 300 MeV/c, the
 220 TPC and ECal PID criteria for selection are changed. The threshold of 300 MeV/c
 221 is chosen because above this value, the Ecal PID can separate MIP from showers
 222 with good accuracy. A relaxed TPC electron pull criterion of $-2.0 < \delta_e < 2.5$
 223 first used. The ECal PID cut is then dependent on the lepton's momentum. If the
 224 selected track has a momentum larger than 800 MeV/c it and is fully contained in
 225 the ECal, it must have an ECal energy greater than 1000 MeV to pass the PID. This
 226 is shown in figure 4.5a. Otherwise, if $p < 800$ MeV/c or the ECal track is not fully

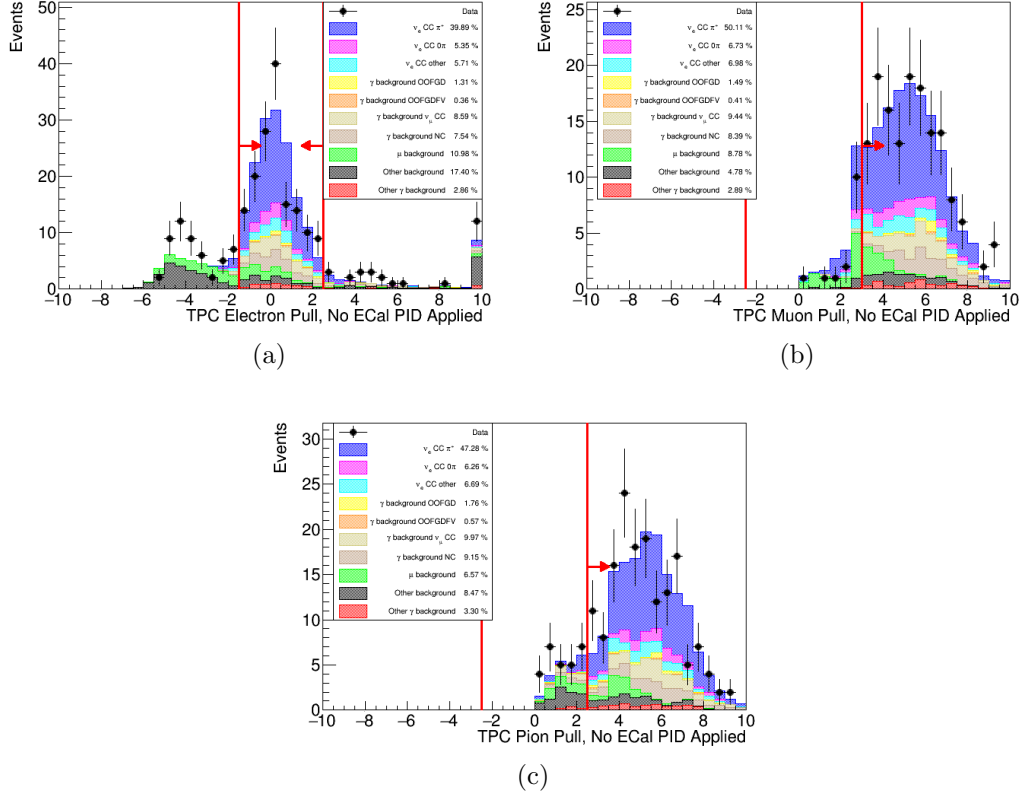


Figure 4.4: The TPC particle identification cuts on (a) the electron pull, (b) the muon pull, and (c) the pion pull.

227 contained, the track must pass an ECal MIP-Shower PID, $MIPEM > 5$, shown in
 228 figure 4.5b. Furthermore if the track does not contain at least 36 hits in the TPC, it
 229 must pass the same pion pull as the TPC PID.

230 **Second TPC PID**

231 Often the main lepton track can propagate into the TPC3 subdetector, the second
 232 TPC downstream of the FGD1. If the lepton track has at least 18 TPC3 hits it is
 233 subject to a second PID cut. The track is rejected if the muon pull falls between
 234 $-2.0 < \delta_\mu < 1.5$. The cuts main objective is to reduce muon background and is
 235 shown in figure 4.6. The cut window was chosen to match the CC ν_e inclusive
 236 analysis [86].

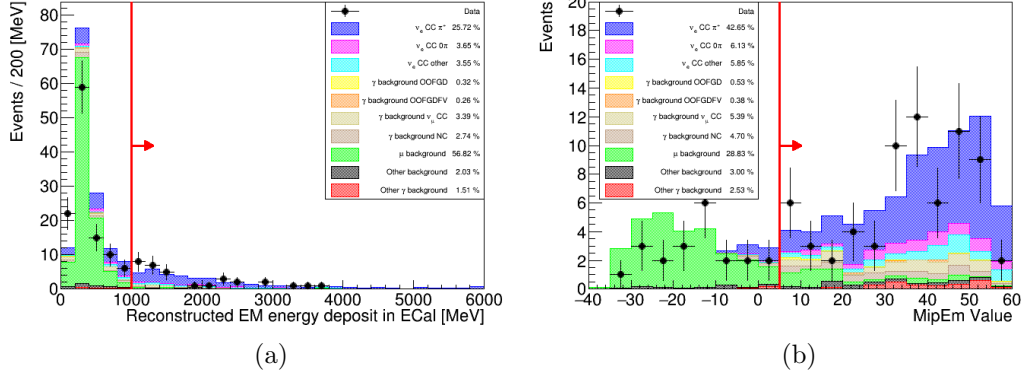


Figure 4.5: (a) The number of events as a function of reconstructed EM energy deposited in the ECal. Cut is used on tracks with momentum above 800 MeV/c. (b) A MIP-Shower cut used on tracks fully contained in the ECal with momentum below 800 MeV/c. A negative value indicates more MIP-like, a positive value indicates more EM shower like.

237 Momentum

238 Inherited from the ν_e inclusive selection, the main lepton tracks that pass the initial
 239 TPC quality and PID cuts are required to pass a lower bound momentum cut of
 240 $p > 200$ MeV/c. The reason for the cut is the observed large background of electrons
 241 from neutrino induced photons, which can be seen in figure 4.7.

242 Pion Selection

243 The previous cuts up to now all had the primary objective of selecting ν_e events. For
 244 this analysis we want to investigate ν_e events that produced at least one π^+ in the
 245 final state through charged current interactions. To find π^+ candidate tracks, all
 246 secondary tracks (that are not the main selected lepton track) that originate from a
 247 vertex in the FGD1 and propagate into the TPC2 are considered. The candidate
 248 track must have a positive charge and originate within 40 mm of the start of the
 249 main lepton track. A cut of 40 mm was established through figure 4.8a. Here, it
 250 can be seen that a cut of 40 mm optimises the selection purity, reducing the overall
 251 photon background levels. Bins of 10 mm were chosen to match the z-dimensions of
 252 each FGD scintillator bar.

253 Once positive pion candidate tracks have been found, the tracks must pass a

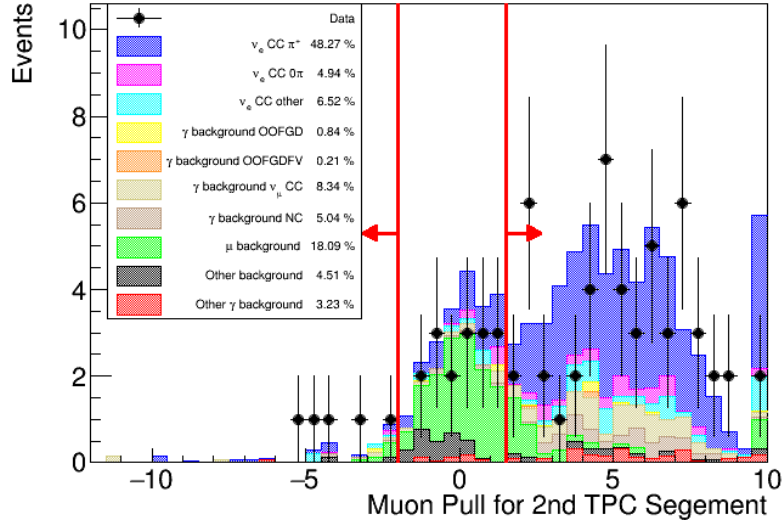


Figure 4.6: The number of events as a function of muon pull in the TPC3. Used for the second TPC PID cut to remove muon background.

254 particle identification cut. The track is rejected if the pion pull does not lie within
 255 $-4.0 < \delta_\pi < 4.0$, as shown in figure 4.8b. These boundaries were found to optimise
 256 selection purity in samples of nominal MC. If multiple tracks pass the cuts described
 257 above, the highest momenta track is selection as the pion candidate track moving
 258 forward in the analysis.

259 Invariant Mass

260 Selecting a π^+ in the final state effectively removes a large fraction of the OOFGD
 261 photon background that is a significant background in the ν_e inclusive analysis.
 262 Nevertheless, an additional invariant mass cut is used to veto the OOFGD and
 263 OOFGDFV background further. The invariant mass cut takes the preselected main
 264 lepton and pion candidate tracks as a pair. The invariant mass for the pair of tracks
 265 is then calculated using the following expression:

$$m_{\text{inv}}^2 = m_i^2 + m_j^2 + 2(E_i E_j - \mathbf{p}_i \cdot \mathbf{p}_j) \quad (4.2)$$

266 Whereby particles i and j are represented by their mass, m , energy, E , and momentum
 267 three-vector, \mathbf{p} . The particle tracks are assumed to be an electron positron pair, and

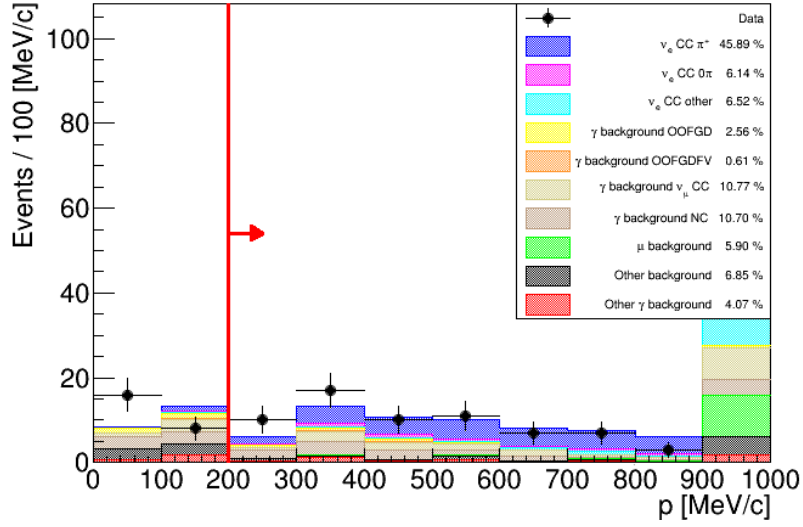


Figure 4.7: The number of events as a function of reconstructed lepton momentum, for the low momentum cut reducing photon background.

268 thus equation 4.2 becomes:

$$m_{inv} = \sqrt{2m_e^2 + (E_1 E_2 - \mathbf{p}_1 \cdot \mathbf{p}_2)} \quad (4.3)$$

269 Here, 1 and 2 represent the main lepton and pion candidate tracks respectively, and
 270 the electron rest mass, m_e , is taken as 0.511 MeV. A cut is placed rejecting all tracks
 271 that do not have $m_{inv} > 110$ MeV, assuming they originated from a photon. A
 272 threshold of 110 MeV was chosen to follow the ν_e inclusive analysis. The cut can be
 273 seen in figure 4.9, and is successful in reducing large amounts of photon backgrounds.

274

275 Momentum Quality

276 The momentum quality cut removes track above 200 MeV/c with negative muon
 277 or pion TPC pulls. It was observed in [86] that tracks with negative muon (pion)
 278 pull fall below the TPC muon (pion) dE/dx curves and as a result are far away
 279 from the TPC electron hypothesis, this is seen in figure 4.3. The cut is named the
 280 'momentum quality' cut as the majority of events attributed to this region are low
 281 momenta events mis-reconstructed to a higher momentum. The performance of the

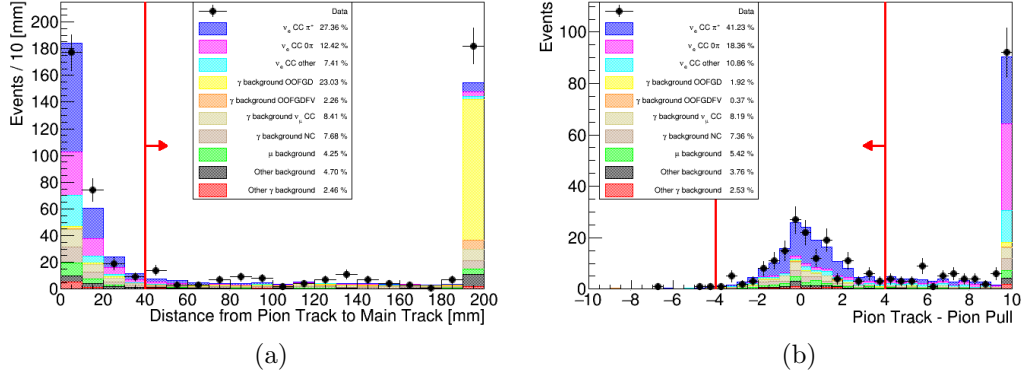


Figure 4.8: The cuts used in positive pion selection. (a) The distance between the pion candidate track and the main lepton track. (b) The pion TPC pull used for PID.

cuts can be seen in figure 4.10.

ECal Veto

Post pion selection and invariant mass cut, the most predominant photon background comes from ν_μ CC, and NC, π^0 production (described in section 4.6). The background usually arises due to mis-reconstructed muons in the FGD or TPC, predominantly caused by high incident angle. It was noticed many of these high angle muons are often reconstructed in tracker ECal modules. To veto this type of event, all reconstructed ECal objects are considered. A vector is then drawn between the start of the main lepton track, and the most upstream segment of the ECal object. This is demonstrated in figure 4.11. The polar angle, θ , with respect to the z-axis is then taken. A cut is made rejecting events below 1 GeV with $\cos(\theta) < 0.6$, removing potential high angle muon events. The cut is shown visually in figure 4.12, and it can be seen that the majority of background events rejected are ν_μ CC photon events.

4.7.2 Full Selection

A summary of the ν_e CC π^+ selection, after all cuts are applied, is outlined in this section. The number of events post selection as a function of reconstructed kinematic phase space for the selected lepton and pion tracks, is shown in figures 4.13 and 4.14 respectively. In general, the selection favours forward going lepton tracks recon-

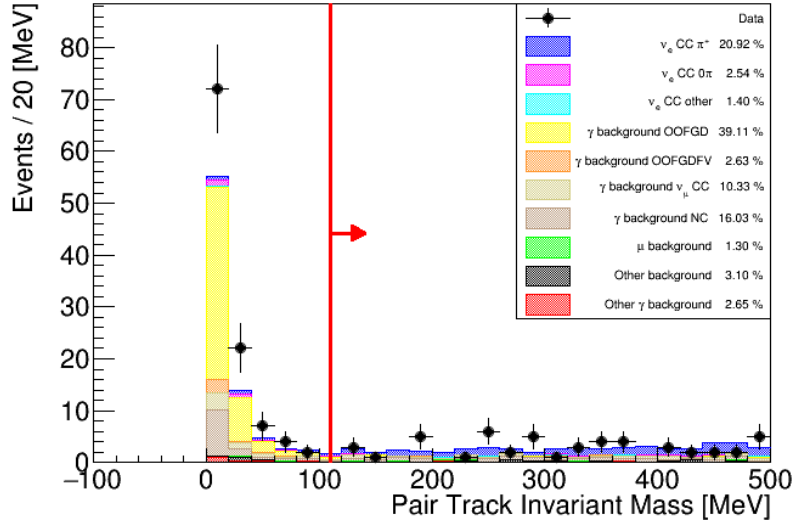


Figure 4.9: The invariant mass of the pion candidate track - main lepton track assuming an e^+e^- pair. The cut aims to removes out of fiducial volume photon interactions.

300 structured in the low momenta region (above $p > 300$ MeV/c). Detector systematic
 301 uncertainties on MC are shown, and errors in data are driven from statistics. The
 302 data-MC agree within error, particularly with respect to track direction. The total
 303 number of data events selected is 152, to be compared with the a total number of
 304 POT normalised selected MC events of 160.6. The predominant background in the
 305 low lepton track momenta region up to 1.5 GeV is from ν_μ charged current and
 306 neutral current π^0 interactions, liberating photons inside the FGD1 fiducial volume.
 307 Above 1.5 GeV, the total background has a larger relative contribution from other
 308 ν_μ CC interactions, and are the dominant background at high momenta. Finally,
 309 backgrounds from other ν_e charged current interactions are most prominent up to
 310 approximately 3 GeV.

311 With a signal selection purity of approximately 51% the total number of signal
 312 MC events predicted is 82. The signal events can be further broken down into reaction
 313 topology to investigate the relative contributions of each interaction mode in ν_e CC π^+
 314 production. The interaction topologies are shown in figure 4.15a as a function of
 315 reconstructed lepton momentum. The vast majority of signal events originate from
 316 deep inelastic scattering ($\sim 56\%$) and resonant ($\sim 40\%$) interactions. Charged current

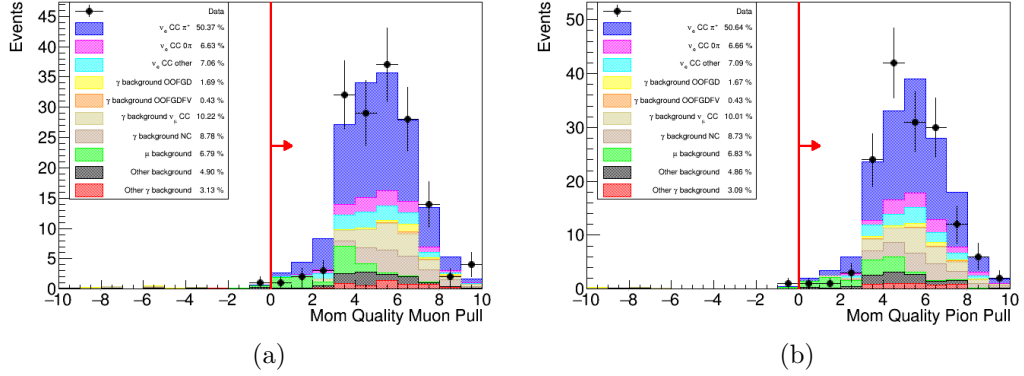


Figure 4.10: The momentum quality cut removing tracks above 200 MeV/c if they have negative (a) muon pull, or (b) pion pull.

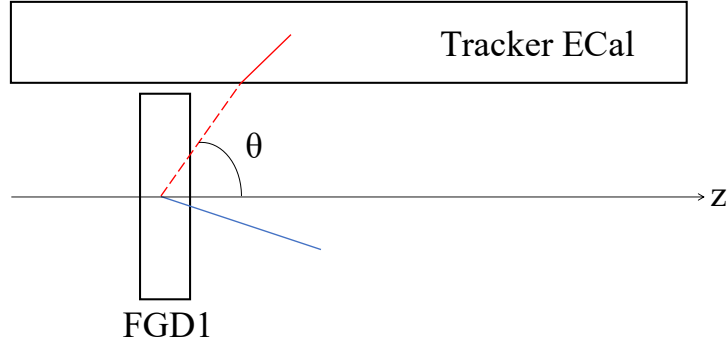


Figure 4.11: Schematic representation of the θ variable used in the ECal Veto cut. The main lepton candidate track is shown in blue, the red solid line represents a reconstructed ECal object, and the red dotted line is the vector joining the two. The FGD1 and Tracker-ECal are shown for context, the other ND280 modules are not shown.

317 coherent scattering interactions also provide a non-negligible contribution to the
 318 total number of signal events. The true number of π^+ particles exiting the same
 319 vertex as the lepton track in signal events is shown in figure 4.15b. Approximately
 320 84% of signal events liberate exactly one positive pion from the true interaction
 321 vertex. The remaining events, with more than one π^+ produced, originate from
 322 DIS interactions only. The true particles selected for both the reconstructed lepton
 323 and pion candidate tracks, as a function of track momentum, is broken down in the
 324 appendix, figure 4.16. It can be seen that in the full selection, an e^- is correctly
 325 selected as the main lepton track roughly 90% of the time. The main sources of
 326 misidentification arise from selecting μ^- (6.6%) and π^- (1.6%) from the muon

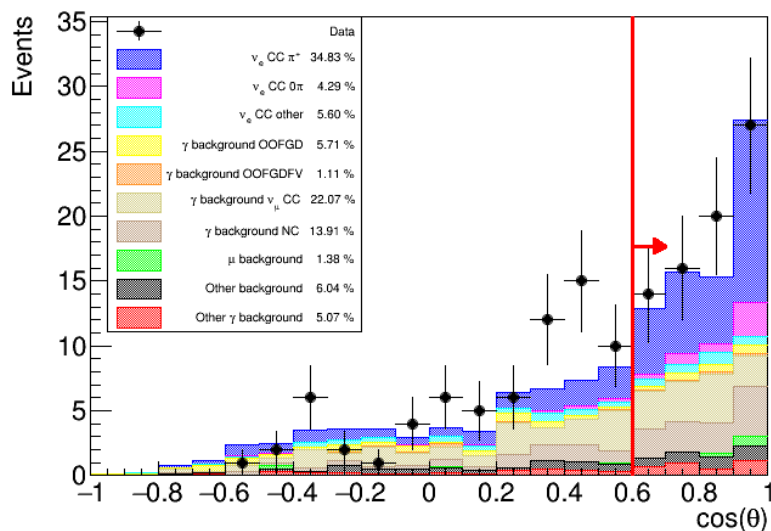


Figure 4.12: The number of events as a function of polar angle used in the ECal veto cut. The angle θ is schematically represented in figure 4.11.

327 background ,and other background topologies, respectively. The relative performance
 328 of pion track selection is divided with respect to momentum space. Below 1.2 GeV,
 329 positive pions are selected roughly 77% of the time, whereas above this threshold π^+
 330 tracks are only selected 28% of the time with protons selected at a rate of 62%. This
 331 is because the pions and protons have similar energy loss curves in gaseous TPCs
 332 above approximately 1.2 GeV (figure 4.3) and thus become indistinguishable through
 333 PID pulls. For the signal topology only, π^+ selection is better at 89% and 41% in
 334 low and high momenta regions respectively. It should also be noted that events in
 335 which a proton is selected as the π^+ candidate track are still signal assuming at least
 336 one true π^+ particle is liberated from the same true vertex as the main lepton track.

337 4.7.3 Efficiency & Purity Of Selection

The performance of the selection criteria is indicated by the signal efficiency and sample purity. The efficiency defines how many of the signal events pass particular cuts in question (equation 4.4). The purity defines the number of signal topology events as a fraction of the total events (equation 4.5). The efficiency is calculated at

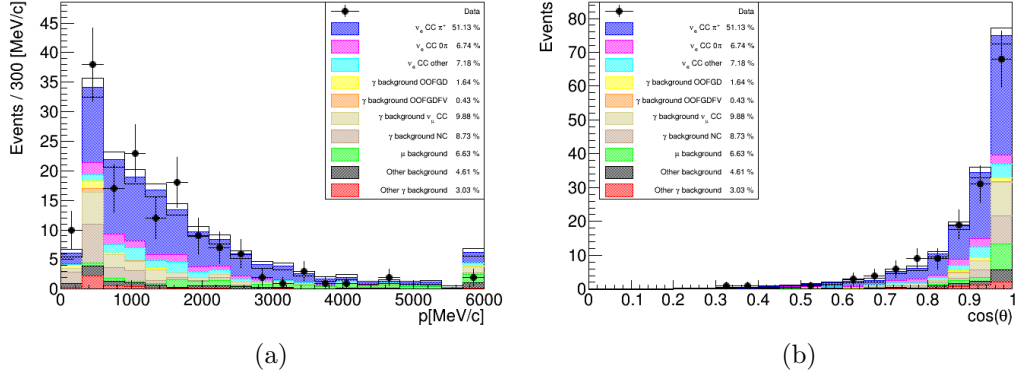


Figure 4.13: The full selection as a function of the lepton track reconstructed kinematic variables.

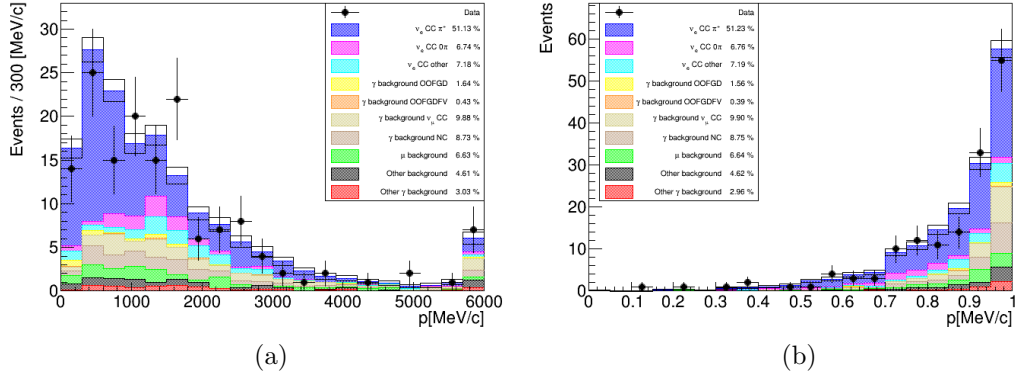


Figure 4.14: The full selection as a function of the pion candidate track reconstructed kinematic variables.

the truth level, whereas the purity is measured at the reconstructed level.

$$\text{Selection Efficiency} = \frac{\text{Number of true signal events selected}}{\text{Total number of true signal events}} \quad (4.4)$$

$$\text{Selection Purity} = \frac{\text{Number of signal events selected}}{\text{Total number of events selected}} \quad (4.5)$$

338 The values for selection efficiency and purity, tracked over the full selection cut-by-cut,
 339 is shown in figure 4.17. This plot provides a good representation of the efficiency
 340 and purity performance as the selection cuts are made. However, one should not
 341 evaluate individual cut performance from this plot as there is a dependence on the
 342 ordering of such cuts. Instead, to evaluate individual cut performance, the $N - 1$

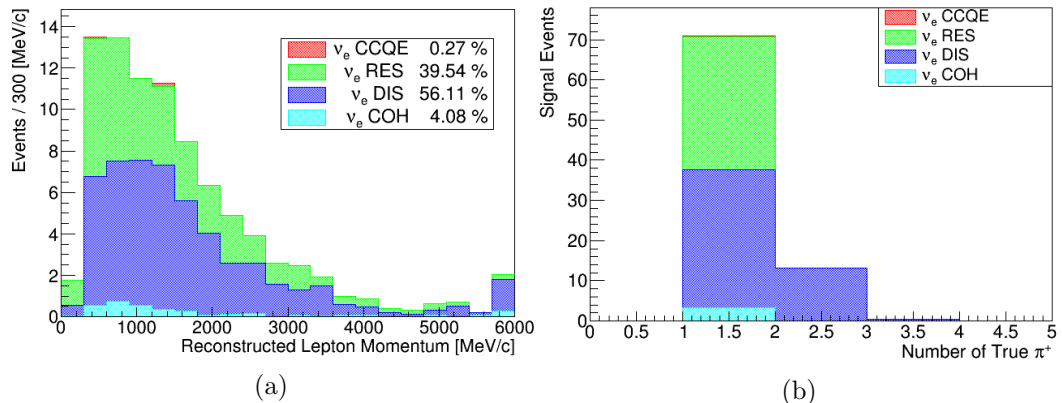


Figure 4.15: (a) The number of signal events as a function of reconstructed lepton momentum broken down by interaction types. (b) The true number of π^+ particles produced in the signal, broken down by interaction type.

343 plots shown in section 4.7.1 are better estimators.

344 Using the signal definition defined in section 4.5, efficiency and purity values
 345 for the full selection can be quantified. The efficiency post-selection, without any
 346 phase-space constraints applied, is calculated to be $(13.8 \pm 0.3)\%$. The signal purity
 347 in the final selection is $(51.1 \pm 0.9)\%$. Imposing the optional signal criterion that the
 348 true pion passes from the FGD1 to the neighbouring downstream TPC, the selection
 349 efficiency is $(19.9 \pm 0.5)\%$. With phase space constraints applied, as defined in section
 350 4.9.1 for the cross-section measurement, the efficiency increases to $(25.4 \pm 0.6)\%$
 351 over the full selection. The uncertainties in efficiency and purity are taken as the
 352 binomial error in calculating them. Calculations of systematic uncertainties affecting
 353 the signal efficiency are outlined in section 4.8.

354 4.8 Systematic Uncertainties

355 The experimental methodology of this analysis at it's most fundamental level is
 356 a measurement of reconstructed event rates both in MC and real data. The MC
 357 predictions on reconstructed event rates at the near detector can be generally
 358 described for this analysis as:

$$R(x) = \int_{E_\nu} \Phi_{\nu_e} \times \sigma(E_\nu) \times T \times M(x_{true}, x_{reco}) \times \epsilon(x_{true}) \quad (4.6)$$

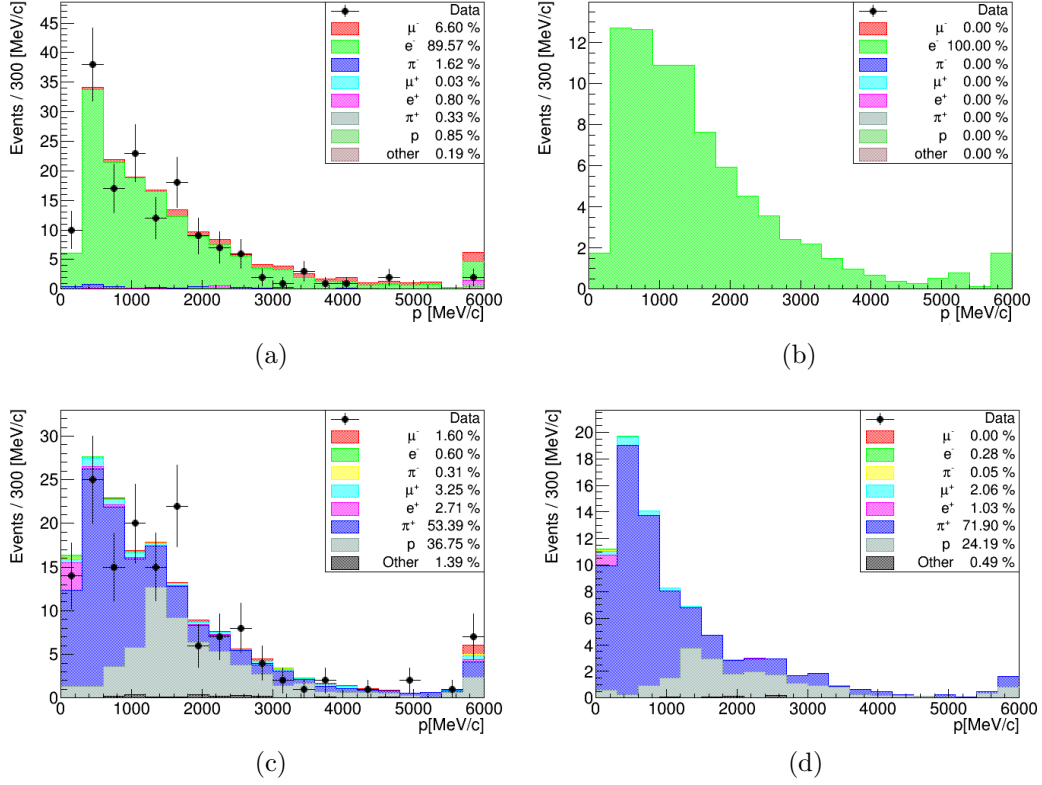


Figure 4.16: The true particle information for the reconstructed lepton track (above) and pion candidate track (below), as a function of track momentum. The full selection is displayed on the left, the signal only is on the right.

359 where R is reconstructed event rate as a function of reconstructed variables, x , for this
360 analysis; Φ_{ν_e} is the electron neutrino flux; σ is the neutrino interaction cross-section
361 as a function of neutrino energy; T is the number of target nucleons; M describes the
362 migration matrix from true to reconstructed variables; and ϵ is the detector
363 efficiency as a function of true variables. Equation 4.6 shows the importance of
364 understanding flux, cross-section model, and detector systematic uncertainties in
365 order to accurately predict the expected number of reconstructed events observed in
366 the near detector. This section will discuss the calculation and propagation of the
367 relevant errors associated to these three sources of systematic uncertainties. Unless
368 stated otherwise, all systematic uncertainties have been evaluated over one global
369 momentum bin from 0 to 30 GeV, and verified using one bin covering the entire
370 angular phase space, $\cos(\theta) = -1$ to $\cos(\theta) = 1$.

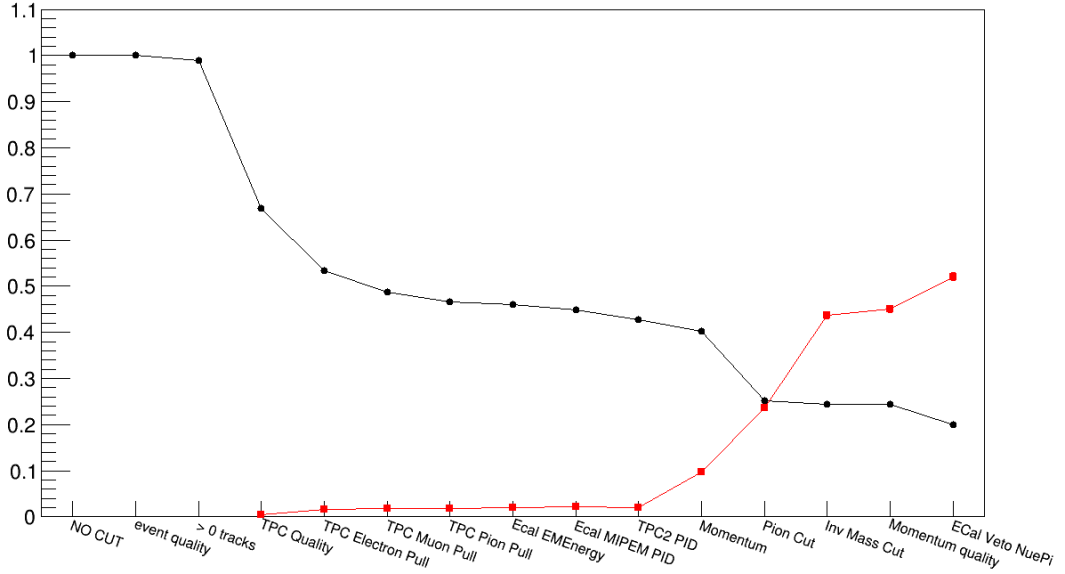


Figure 4.17: Selection efficiency (black) and selection purity (red) as a function of the cuts applied, shown for each stage of the selection. The purity is only shown from the TPC quality cut, but can be assumed to be negligible before this cut.

371 4.8.1 Detector Systematic Uncertainties

372 The detector systematic uncertainties encapsulate the performance of each ND280
 373 sub-detector. The systematic uncertainties are evaluated using the Highland2/Psyche
 374 software packages. The systematic uncertainties are split into two categories: vari-
 375 ation and weight. Variation systematic uncertainties modify the properties of objects
 376 at the event level, whereas weight systematic uncertainties alter the final weight of
 377 the event passing the selection. The decision of whether a systematic uncertainty is
 378 parsed as weight or variation, is defined by the systematic uncertainties affect on
 379 the event. For example, if the uncertainty affects a continuous parameter it must be
 380 implemented as variation; however if the uncertainty affects only event normalisation
 381 it can be treated as a weight. The exception here is if the systematic affects a binary
 382 parameter, for example charge identification. In this case, it may be implemented as
 383 either variation or weight. Efficiency systematic uncertainties are applied as a weight
 384 that depends on more than one variable, and are calculated through comparisons of
 385 data and MC predictions for well known control samples.

386 A full list of the detector systematic uncertainties for both variations and

Systematic Uncertainty	Type	Fractional Error (%)
TPC PID	Variation	1.80
TPC Momentum Resolution	Variation	0.48
TPC Momentum Scale	Variation	0.16
B-Field Distribution	Variation	0.19
ECal EM Energy Scale	Variation	0.31
ECal EM Energy Resolution	Variation	0.07
TPC-ECal Matching Efficiency	Weight	1.91
TPC-FGD Matching Efficiency	Weight	0.17
TPC Track Efficiency	Weight	1.29
TPC Cluster Efficiency	Weight	0.01
FGD Mass	Weight	0.58
Charge ID	Weight	0.68
OOFV	Weight	1.26
Pion Secondary Interactions	Weight	2.01
Proton Secondary Interactions	Weight	2.04
ECal PID	Weight	0.75
ν_e CC π^+ ECal Pile Up	Weight	0.48
FGD Vertexing	Other	1.40
Total	-	4.51

Table 4.2: A full summary of the fractional errors of all detector systematic uncertainties considered for this analysis. The systematic type is also shown. Fractional errors on the number of selected events for the full selection have been calculated over 250 toys.

387 weights can be seen in table 4.2. The fractional error on predicted number of
388 events over the full selection was calculated using 250 toy experiments for each
389 systematic uncertainty. The majority of the ND280 tracker systematic uncertainties
390 are shared with the ν_μ and ν_e inclusive analyses; full descriptions of these systematic
391 uncertainties can be found at [94] and [88] respectively. As a brief overview the
392 systematic uncertainties and their relative affects on the selection are described
393 below:

- 394 • The **TPC PID** systematic for muons, electrons, and pions is estimated using
395 dedicated control samples directly extracted from beam events. The systematic
396 is then estimated from the data-MC differences observed in the pull distribu-

397 tions, which is computed using a Gaussian and considering the correct particle
398 hypothesis.

399 • The **TPC momentum resolution** compares the differences in TPC and
400 global momentum resolutions of data and MC. A smearing factor is then
401 applied as the systematic parameter is propagated through event selection.
402 **TPC momentum scale** uncertainty is taken from the B-field measurement
403 described in [95].

404 • **B-field distribution** systematic performs corrections using two separate
405 methods. The main correction applies a \vec{B} field map, at the reconstruction
406 level, developed using measurements of the magnetic field inside the ND280
407 basket. The second, an empirical correction, is based on a laser system which
408 illuminates aluminium dots on the cathode where expected and measured
409 positions are compared. Magnetic field systematic uncertainties are described
410 fully in [96].

411 • All ECal systematic uncertainties are discussed in full detail in [97]. The
412 ECal EM energy systematic uncertainties are evaluated by comparing first and
413 second moments measured by the TPC. The fractional difference is defined as:

$$\text{Fractional Difference} = \frac{\text{EM Energy} - \text{TPC Momentum}}{\text{TPC Momentum}} \quad (4.7)$$

414 The systematic mean and standard deviation define the **ECal EM Energy**
415 **Scale** and **ECal EM Energy Resolution** systematic uncertainties respect-
416 ively.

417 Control samples for the **ECal PID** systematic use cuts of MIP-EM > 0 and
418 EM-HIP > 0. The efficiency for each particle type is then calculated for both
419 data and MC. Any difference in data and MC samples is interpreted as the
420 systematic error in the modelling of the ECal PID for that particle type.

421 • The **TPC cluster efficiency** is the probability of finding a cluster (group
422 of adjacent single TPC pad hits) that corresponds to one point in the long
423 trace of ionized gas created by charged particle tracks in the TPC. Data-MC

424 discrepancies over both horizontal and vertical directions are used to calculate
425 cluster efficiencies.

426 • The ability in which tracks crossing the TPC are able to be reconstructed is
427 held within the **TPC track efficiency** parameter. Included in the systematic
428 is the evaluation of both TPC pattern recognition and likelihood fit.

429 • **Charge sign identification** systematic assesses global charge identification
430 based on the combination of ND280 subdetectors. Two errors are propagated
431 here: The probability of swapping the local TPC charge identification, and
432 probability of the global tracking to swap the sign of the charge. The systematic
433 is explained in full in [98].

434 • **TPC-FGD and TPC-ECal matching efficiencies** characterise the per-
435 formance of matching reconstructed tracks in the associated subdetectors.
436 Descriptions of the control samples and performances of TPC-FGD and TPC-
437 ECal matching efficiencies can be found in [97] and [99].

438 • The **FGD mass** systematic compares the areal density of an XY FGD module
439 to MC values in the ND280 software. This is then combined in quadrature to
440 the spread in masses over XY modules to give the full systematic uncertainty.
441 Full detailed are given in [100].

442 • **Secondary interaction** systematic uncertainties, for both pions and protons,
443 characterise the uncertainty in the probability for each particle type to undergo
444 interactions outside the nucleus in which it was produced. The pion secondary
445 interactions are modelled using the NEUT cascade method described in [101].
446 Proton secondary interactions are modelled through Geant4 [102]. Proton
447 secondary interaction uncertainties play an important role in this analysis given
448 the difficulty in separating pions and protons identification at high momenta
449 (see figure 3.8), and therefore likelihood of selected protons as pion candidates.
450 These systematic uncertainties account for discrepancies between data and the
451 models used in MC, and are significant in this analysis.

452 • The **OOFV** systematic characterises the case in which the event is recon-
 453 structed as originating within the FGD fiducial volume, yet the true vertex
 454 is outside. A recent study [103] estimated that the OOFV error could be as
 455 large as 100% for analyses with predominant photon backgrounds. If the true
 456 vertex is outside of the FGD a conservative systematic uncertainty of 100% is
 457 assumed, following the ν_e inclusive analysis [87]. If the true vertex originates
 458 inside the FGD but outside the FV, data-MC discrepancies in control samples
 459 are used, following the treatment in [104].

460 ν_e CC π^+ ECal Pile Up Systematic Uncertainty

461 The dominant uncertainty associated to the Ecal veto cut is event pile up in the side
 462 Tracker-Ecals. In the case where a sand muon event is in coincidence with a magnet
 463 event, the activity caused by the sand muon may trigger the veto. This behaviour is
 464 not simulated in the MC, and therefore needs to be characterised in the ν_e CC π^+
 465 ECal pile up systematic.

466 The systematic is evaluated by counting the number of ECal events in a
 467 separate sand muon MC sample with a fixed POT. The data intensity, defined as
 468 $\text{POT}/N_{\text{Spills}}$, (for data, MC and sand MC) is then computed from the respective
 469 data samples and used to calculate the effective number of spills. With eight bunches
 470 per spill, the number of ECal events per spill can be translated to bunches for each
 471 dataset. The number of events should be reduced in the MC since pile up is not
 472 considered. Therefore a re-weight reduction factor is used:

$$w_c = (1 - C_{\text{pileup}}) \quad (4.8)$$

473 where C_{pileup} is the correction to be applied and is defined as the number of sand
 474 ECal events per bunch.

475 There is an intrinsic 10% uncertainty in the total rate of sand muon inter-
 476 actions in neutrino simulations [104]. This value is used for both ν_μ and ν_e T2K
 477 cross-section analyses. Moreover, there are potential differences between data and
 478 MC arising from actual and simulated beam and detector properties. The uncertainty

Sample	ECal/bunch (%)			C_{pileup}	σ_{pileup}
	Data	MC	Sand		
Run 2 - Water Out	16.6265	12.0846	4.23594	0.0423594	0.00423594
Run 2 - Water In	14.3929	10.4291	3.61622	0.0361622	0.00361622
Run 3b - Water Out	15.1766	11.7413	4.16571	0.0416571	0.0073044
Run 3c - Water Out	18.1378	12.9605	4.59826	0.0459826	0.00579119
Run 4 - Water Out	22.4739	16.3985	5.81668	0.0581668	0.00581668
Run 4 - Water In	20.1549	14.5733	5.11788	0.0511788	0.00511788
Run 8 - Water Out	39.7411	30.0092	11.722	0.11722	0.0199007
Run 8 - Water In	35.0465	26.18	10.1318	0.101318	0.0126531

Table 4.3: Table showing the correction, C_{pileup} , and systematic uncertainty, σ_{pileup} , values over each run sample for data, nominal MC, and sand MC. The number of ECal events per bunch is also shown.

479 in data-MC differences is evaluated by

$$\Delta_{\text{data-MC}} = C_{\text{data}} - (C_{\text{MC}} + C_{\text{pileup}}) \quad (4.9)$$

480 where C_{data} and C_{MC} are the number of data and nominal MC ECal events per
481 bunch respectively. Combining these two uncertainties is double counting, and thus
482 the larger uncertainty of the two is taken as the pile up systematic uncertainty,
483 σ_{pileup} . Table 4.3 shows the final values for C_{pileup} and σ_{pileup} for ECal pileup for
484 each data period. A more detailed breakdown of the numbers used to evaluate the
485 correction systematic uncertainties is shown in table A.1.

486 **FGD Vertexing Systematic Uncertainty**

487 Within the pion cut, described in section 4.7.1, a parameter is defined to measure
488 the distance between the start of the lepton track and start of the pion candidate
489 track, as described in figure 4.8. This parameter is most sensitive to the ability to
490 accurately reconstruct the position of vertices within the FGD.

491 A selection has been developed to create a control sample of electron-positron
492 pairs in the FGD1, in which the main lepton track is the electron. The selection uses
493 the electron neutrino selection cuts, described in section 4.7.1, up to and including the

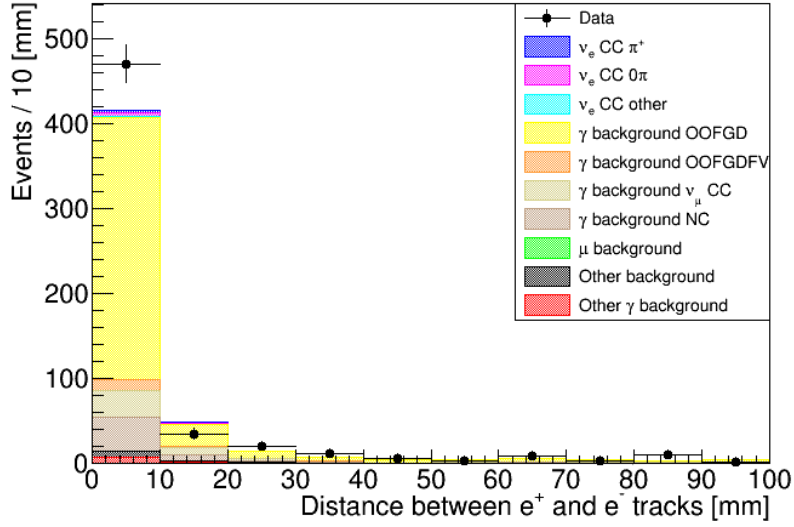


Figure 4.18: The distance between the selected tracks in electron-positron pairs used to calculate the vertexing systematic.

494 momentum cut. An additional cut to veto electron-positron pairs is then applied. All
 495 FGD1 to TPC2 tracks with opposite charge to the main lepton track are considered
 496 as pair candidates. Electron-like particle identification is performed using the longest
 497 segment of the reconstructed track inside the TPC, which in turn must pass TPC
 498 track quality cuts. The track is accepted as a pair candidate if it passes an electron
 499 pull PID of $-3.0 < \delta_e < 3.0$. Next, partnered with the main lepton track, the
 500 pion candidate track must have an invariant mass $< 110 \text{ MeV}/c$, the inverse cut to
 501 that described in section 4.7.1.

Analogous to the separation of pion and lepton tracks in figure 4.8a, the distance from the electron and pair candidate track is shown in figure 4.18. The distribution of data here appears narrower than MC. To attribute a systematic uncertainty, the efficiency of any cut on the given distribution was calculated using:

$$\epsilon_{data} = \frac{N_{data < x}}{N_{data}} \quad (4.10)$$

$$\epsilon_{MC} = \frac{N_{MC < x}}{N_{MC}} \quad (4.11)$$

502 where ϵ is the cut efficiency and x is the chosen value for the cut. To match the pion

503 selection, x is chosen to be 40 mm. The vertexing error is then calculated by:

$$\text{Vertexing Error} = 1.0 - \left(\frac{\epsilon_{\text{data}}}{\epsilon_{\text{MC}}}\right) \quad (4.12)$$

504 This equates to a vertexing error of 1.4% at a cut value of 40 mm. The vertexing
505 error is of course a function of the cut value, and can thus be increased (decreased)
506 by tightening (relaxing) the pion selection cut respectively. A cut value of 40 mm
507 was deemed to be tight enough to optimise selection purity without introducing a
508 dominant systematic error.

509 In verifying the relevance of this systematic uncertainty to the ν_e CC π^+
510 selection, potential kinematic differences between the FGD vertexing sample and the
511 ν_e CC π^+ selection sample have been considered. Different Q^2 values in ν_e CC π^+
512 and e^+e^- interactions could lead to different angular distributions between the two
513 selected tracks. The angle between the two selected tracks in each sample is shown
514 in figure A.2, where no major differences in distribution shape were found.

515 **Discussion of Detector Systematic Uncertainties**

516 The largest detector systematic uncertainties are in the secondary interactions of
517 pions and protons. This is expected given the pion selection within the analysis, and
518 the fact secondary interactions are known to be a large systematic uncertainty in T2K.
519 The interactions pions and protons undergo, outside the nucleus it was produced in,
520 are found to differ greatly between interaction models such as Geant4 and NEUT, and
521 in comparison to external data [104]. The proton secondary interactions contribute
522 similarly to pions as above approximately 1.2 GeV, protons and pions a relatively
523 indistinguishable through energy loss methods in the TPC. In addition TPC PID
524 and TPC-ECal matching efficiencies are significant detector errors to this analysis.
525 All detector systematic uncertainties were cross-checked and validated against the ν_e
526 inclusive analysis found in [86]. The major differences between these analyses is the
527 additional selection of a π^+ candidate track, as well as the subsequent suppression
528 of OOFV photon backgrounds. Errors including TPC PID, TPC track efficiency,
529 Charge ID, and TPC-ECal matching efficiency see comparable increases from the

	γ Bkg (%)	ν_e CC Bkg (%)	All Other Bkg (%)	Signal Efficiency (%)
Systematic Uncertainty	3.12	2.10	2.82	4.13

Table 4.4: A summary of the combined detector systematic uncertainties on background topology event yields. Each background topology uncertainty is normalised relative to its contribution to the total background. The detector systematic fractional error on signal efficiency in a limited phase space is also shown.

Background Topology	Fraction of Total Background (%)
γ Background	47.8
ν_e CC Background	28.6
All Other Background	23.6

Table 4.5: The relative fraction each predefined background topology contributes to the total background event yield.

530 ν_e inclusive analysis due to the identification and tracking of an additional particle.
531 Conversely, the previously dominant OOFV systematic has decreased significantly
532 with the reduction on OOFV photon background yields. Nevertheless, no unexpected
533 differences in detector systematic uncertainties can be seen between the two analyses
534 that cannot be explained by inherent differences in the selection methods.

535 For cross-section measurements it can be more useful to measure the effect
536 of systematic uncertainties on the event yields to different final state topologies, as
537 well as the signal efficiency. Table 4.4 shows the combined effect detector systematic
538 uncertainties have on different background topologies and the signal efficiency. Each
539 background topology uncertainty is normalised to the relevant fraction each topology
540 contributes the total background event yield, as shown in table 4.5. Weighting the
541 uncertainties to topology size provides more context on the detector systematic
542 effects on the full sample. The effect of detector systematic uncertainties upon the
543 total background event yield is 6.08%. The ν_e inclusive analysis sees a background
544 uncertainty from detector effects of approximately 12.7%. The decrease in uncertainty
545 in this analysis is attributed to the significant drop in OOFV photon background.
546 The uncertainty on signal efficiency due to detector effects without phase space
547 constraints applied is 4.22%. In the limited phase space defined in section 4.9.1,
548 the uncertainty on signal efficiency due to detector effects is 4.13%. These are an

549 increase on the ν_e inclusive value of 2.96%, largely because of the effects an additional
550 π^+ selection has on the detector systematic uncertainties stated previously. The
551 magnitudes of the total detector systematic uncertainties are significantly smaller
552 than the data statistical error of approximately 17%.

553 4.8.2 Cross-section Model Systematic Uncertainties

554 A notable source of uncertainty comes from the model choices and parameters used in
555 simulation, and their ability to accurately describe all of the physics undergone in the
556 relevant interactions. Most noteworthy are parameters defining neutrino interactions,
557 nuclear final state interactions and cross-section parameters. T2K estimates prior
558 uncertainties on model parameters using external data constraints. A list of the
559 cross section systematic uncertainties, provided by the T2K NIWG group [105], with
560 their prior values and associated errors is shown in table 4.6. Detailed descriptions
561 of all the cross-section systematic parameters can be found in [106, 107]. A brief
562 overview of each systematic is given below:

- 563 • The axial mass term, M_A , for the axial form factor is implemented for both
564 quasi-elastic and resonance interactions.
- 565 • The Fermi momentum, p_F , is the highest momentum state in Fermi gas models
566 such as RFG. The Fermi momentum parameter has a dependence on the
567 number of nucleons in the nucleus, therefore it is implemented for both Carbon
568 and Oxygen targets.
- 569 • Two-particle two-hole effects, 2p2h, are contributions to the interaction cross
570 section arising from multi-body processes. The contribution to 2p2h interactions
571 can be split into three primary components: meson exchange current, MEC;
572 nucleon-nucleus correlations; and the interference between the two.
- 573 • The binding energy, E_B , is implemented for CCQE interactions on both Carbon
574 and Oxygen targets. There is currently no treatment of binding energy in
575 resonance interactions in the latest oscillation analysis, and thus is not included
576 for this analysis.

Cross Section Parameter	Prior Value and Error	NEUT	Units
M_A^{QE}	1.2 ± 0.41	1.2	GeV/c ²
M_A^{RES}	1.07 ± 0.15	0.95	GeV/c ²
p_F ¹² C	223.0 ± 31.0	217	MeV/c
2p2h ¹² C	1.0 ± 1.0	1.0	None
E_B ¹² C	25.0 ± 9.0	25	MeV
p_F ¹⁶ O	225.0 ± 31.0	225	MeV/c
2p2h ¹⁶ O	1.0 ± 1.0	1.0	None
E_B ¹⁶ O	27.0 ± 9.0	27	MeV
2p2h Other	1.0 ± 1.0	1.0	None
C_A^5 (RES)	1.01 ± 0.12	1.01	None
Isospin = $\frac{1}{2}$ Background	0.96 ± 0.4	1.3	None
ν_e/ν_μ	1.0 ± 0.028	1.0	None
CC Coherent ¹² C	1.0 ± 1.0	1.0	None
CC Coherent ¹⁶ O	1.0 ± 1.0	1.0	None
CC Other Shape	0.0 ± 0.4	0.0	None
NC Coherent	1.0 ± 0.3	1.0	None
NC Other	1.0 ± 0.3	1.0	None
FSI Inelastic Low Energy	0.0 ± 0.41	0.0	None
FSI Inelastic High Energy	0.0 ± 0.34	0.0	None
FSI Pion Production	0.0 ± 0.5	0.0	None
FSI Pion Absorption	0.0 ± 0.41	0.0	None
FSI Charge Exchange Low Energy	0.0 ± 0.57	0.0	None
FSI Charge Exchange High Energy	0.0 ± 0.28	0.0	None

Table 4.6: A list of cross section model systematic uncertainties, their respective prior values with expected range, and their initial values in NEUT nominal MC.

- 577 • Resonance interactions and their associated form factors introduce new para-
578 meters to cross-section models. The first is C_A^5 which affects the scale of the
579 axial form factor at $Q^2 = 0$ GeV/c². The second is an isospin $\frac{1}{2}$ background
580 scaling factor.
- 581 • The difference in ν_e and ν_μ cross-sections is another source of systematic error
582 that is accounted for. The overall effect of this is approximately 3%, and is
583 labelled in table 4.6 as ν_e/ν_μ .
- 584 • Remaining neutrino interactions not covered in fits to external data are char-

585 characterised in the following errors:

- 586 – Charged current coherent pion production carries a 100% error brought
587 forward from similar motivation to the oscillation analysis, namely that
588 the external data is consistent with no coherent production in this region
589 of neutrino energy [106].
 - 590 – The energy shape dependence on other charged current interactions, such
591 as: CC multi- π production, CC DIS, and CC0 π resonant interactions that
592 include γ , K , and η production.
 - 593 – Neutral current coherent pion production which has a 30% normalisation
594 error from [108].
 - 595 – Other neutral current interactions which complements the CC other
596 sample described previously. Poor constraints to external data means a
597 recommended 30% error from [108] is used for this analysis.
- 598 • Final state interaction systematic uncertainties have the effect of migrating
599 events between different observable detector topologies and change pion kin-
600 ematics. For example, pion absorption in the nucleus can move events into
601 CCQE-like samples. FSI uncertainties are broken down into 3 main categories:
602 Inelastic, charge exchange, and pion absorption and production. Uncertainties
603 on FSI parameters are estimated through pion-nuclear scattering data from
604 fits to Carbon (most prominently from the DUET experiment [109]), to the
605 cascade model parameters in NEUT [106].

606 Cross section model systematic uncertainties are evaluated using the T2KReWeight
607 package, which produces splines for each model parameter over a 1σ standard de-
608 viation; each model parameter can affect both the shape and normalisation. The
609 fractional error of each systematic uncertainty on the event yields and signal efficiency
610 is evaluated over 250 toy experiments for each parameter. Either the event yields
611 for a given topology, or signal efficiency are plotted, and the RMS computed as
612 the fractional systematic uncertainty. For simplicity, background topologies are
613 grouped into three categories: Photon background, which covers all background
614 events originating from π^0 decays; ν_e CC background, which is defined as any charged

Systematic	γ Bkg (%)	ν_e CC Bkg (%)	Other Bkg (%)	Total Bkg (%)
M_A^{QE}	0.04	17.24	1.83	5.43
M_A^{RES}	2.87	4.58	3.04	3.39
p_F ^{12}C	0.00	0.13	0.01	0.04
2p2h ^{12}C	0.00	5.38	0.76	1.72
E_B ^{12}C	0.00	0.03	0.01	0.00
p_F ^{16}O	0.00	0.02	0.00	0.00
2p2h ^{16}O	0.00	0.52	0.00	0.15
E_B ^{16}O	0.00	0.00	0.00	0.00
2p2h Other	0.00	0.00	0.00	0.00
C_A^5 (RES)	1.01	2.03	1.58	1.43
Isospin = $\frac{1}{2}$ Background	0.71	1.64	1.07	1.06
ν_e/ν_μ	0.20	2.62	0.07	0.86
$\bar{\nu}_e/\bar{\nu}_\mu$	0.00	0.00	0.01	0.00
CC Coherent ^{12}C	0.14	0.00	2.23	0.59
CC Coherent ^{16}O	0.00	0.00	0.00	0.00
CC Other Shape	4.18	3.95	4.19	4.12
NC Coherent	0.00	0.00	0.00	0.00
NC Other	11.76	0.00	3.30	6.42
FSI Total	2.79	6.04	2.83	3.08
Total Uncertainty	6.31	5.81	1.80	10.76

Table 4.7: The effect of the cross section systematic uncertainties on the background event yields, separated by different topologies. The total systematic uncertainties are also shown taking into account each sample's relative contribution to the total background.

615 current ν_e event that isn't signal; and all other background, to contain any event
616 that isn't included in the first two categories.

617 The fractional systematic uncertainties on background event yield topologies
618 are shown in table 4.7. Each systematic uncertainty has been normalised to the
619 relevant background topology sample size, as shown in table 4.5. Photon background
620 systematic uncertainties are dominated by NC other, which is expected given the
621 large contribution of neutral current interactions (figure 4.19a). The quasi-elastic
622 axial mass term, M_A^{QE} , is the dominating systematic in ν_e charged current back-
623 ground. This is also significantly larger than the two other topology samples, but
624 correlates with the relative CCQE contributions to each sample (figure 4.19b). The

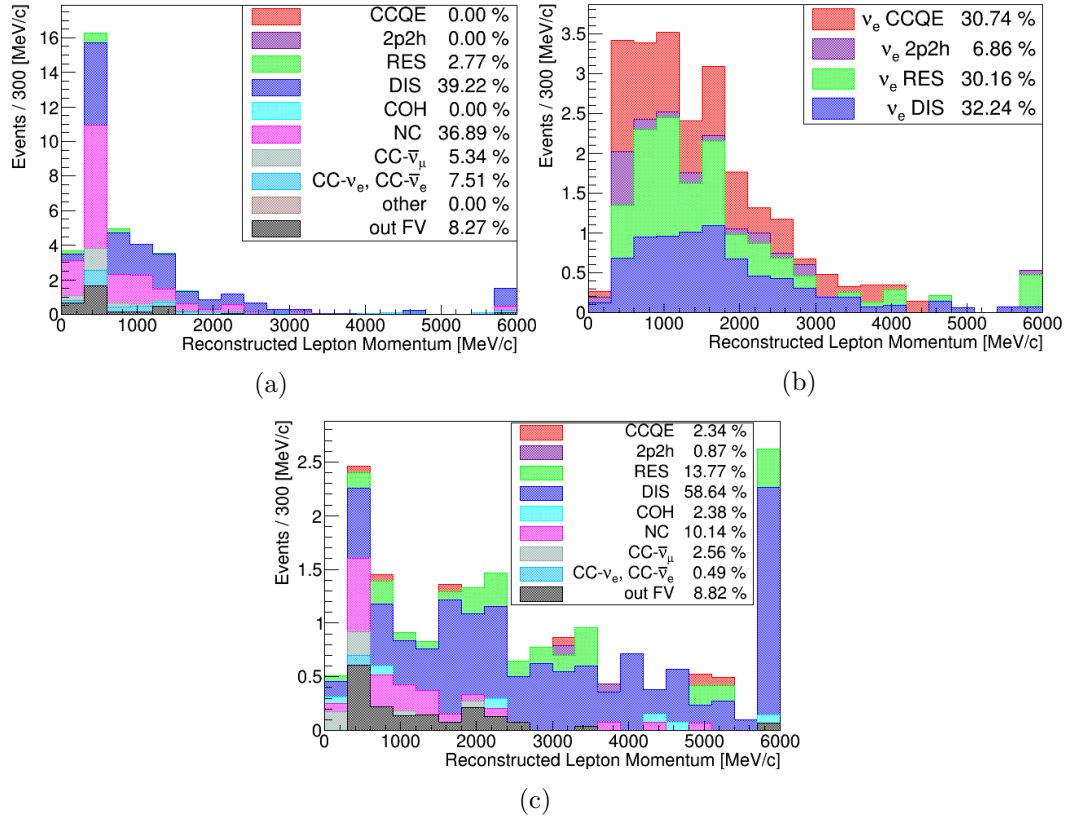


Figure 4.19: Number of events as a function of reconstructed lepton momenta, split into reaction types, for the three background topology samples: (a) Photon background, (b) ν_e CC background, and (c) All other background.

625 largest signal systematic in the ν_e inclusive analysis is M_A^{QE} at 8.52%. Given the
626 ν_e background topology is closely related to the ν_e inclusive signal, except with
627 lower statistics, the systematic values for M_A^{QE} in this analysis makes sense. A
628 systematic of 5.38% is seen for 2p2h interactions on carbon in the ν_e CC background.
629 This largely comes from uncertainties in meson exchange current parameters, and
630 is most prominent in the ν_e CC background sample. Other large systematic con-
631 tributions across all three topologies come from M_A^{RES} , CC other shape, and FSI.
632 The relative resonance interaction contributions to all samples topologies can be
633 seen in figure 4.19. As expected the size of the M_A^{RES} systematic correlates with the
634 fraction of resonance interactions in the sample size. The relative large contribu-
635 tion to resonance interactions in the ν_e CC background originates from charged current
636 $\nu_e \pi^0$ production. The CC other shape systematic is dominated by the energy

Systematic	Fractional Error (%)	
	No Phase Space Cuts	Phase Space Limited
M_A^{QE}	0.37	0.13
M_A^{RES}	0.89	0.39
p_F ^{12}C	0.00	0.00
2p2h ^{12}C	0.03	0.02
E_B ^{12}C	0.00	0.00
p_F ^{16}O	0.00	0.00
2p2h ^{16}O	0.00	0.00
E_B ^{16}O	0.00	0.00
2p2h Other	0.00	0.00
C_A^5 (RES)	0.96	0.30
Isospin = $\frac{1}{2}$ Background	0.45	0.70
ν_e/ν_μ	0.00	0.00
$\bar{\nu}_e/\bar{\nu}_\mu$	0.00	0.00
CC Coherent ^{12}C	1.94	1.39
CC Coherent ^{16}O	0.13	0.09
CC Other Shape	0.73	0.99
NC Coherent	0.00	0.00
NC Other	0.00	0.00
FSI Total	1.33	1.70
Total	2.85	2.56

Table 4.8: The effect of cross section systematic uncertainties on the signal selection efficiency. Fractional errors are quoted before and after phase space constraints are applied.

637 shape dependence of DIS multi-pion interactions. Each systematic uncertainty has
638 been qualitatively compared to the values quoted in the ν_e inclusive analysis in
639 order to test their validity. The largest uncertainty contributions come from photon
640 and ν_e CC background at 6.3% and 5.8% respectively, compared to 1.8% for other
641 backgrounds. The full effect of cross section systematic uncertainties on the total
642 background event yield is 10.76%; with the largest sources of systematic uncertainty
643 coming from NC other and M_A^{QE} .

644 The systematic uncertainties on signal selection efficiency is shown in table 4.8.
645 The different interaction type contributions to signal events are shown in figure 4.15a.
646 A 4% presence of coherent events is large enough to influence the largest cross section

647 systematic uncertainty on signal efficiency. This is likely due to the 100% error on
648 the prior value in NEUT. There are also relatively large contributions from FSI,
649 C_A^5 (RES), M_A^{RES} , and CC other shape which arise from the large presence of both
650 resonance and DIS interactions in the signal sample. The total cross-section models
651 systematic uncertainty on signal efficiency, as a fractional error, is calculated to
652 be 2.85%. With the phase space constraints (defined in section 4.9.1) applied, the
653 systematic uncertainty on signal efficiency is 2.56%.

654 4.8.3 Flux Systematic Uncertainties

655 A major uncertainty to any ND280 cross-section analysis is the modelling of the
656 electron neutrino flux. The secondary beamline is simulated in order to estimate,
657 in the absence of neutrino oscillations, the nominal neutrino flux at ND280. The
658 FLUKA package [110, 111] is used to model primary beam proton interactions and
659 the subsequent hadrons produced in the graphite target. Particles exiting the target
660 are tracked using GEANT3 [112] simulation as they propagate through the magnetic
661 horns and decay volume. GCALOR [113, 114] is used to model any hadron decays.
662 These simulated predictions are bolstered by a significant flux monitoring program in
663 which each beam pulse is measured in the primary neutrino beamline using the suite
664 of detectors described in section 3.1.2. The beam position and width are measured by
665 INGRID (section 3.3.1). Furthermore, data from the NA61/SHINE experiment [115],
666 along with other experiments, taken at 31 GeV is used to improve the modelling of
667 the kinematic distributions of meson from the proton-graphite collisions.

668 The uncertainties in neutrino flux predictions arise from a number of sources
669 including but not limited to: the hadron production model, the proton beam profile,
670 and the currents and alignments of the horns. The underlying parameters of each
671 source of uncertainty are varied to evaluate their affect on neutrino flux binned in
672 both neutrino energy and flavour. Flux tuning files, produced by the T2K beam
673 group, are used to create event weights in Highland2 and propagate these flux
674 systematic uncertainties through the analysis.

675 In this section two systematic errors associated with the flux will be discussed.
676 The first is the error in the total predicted electron neutrino flux at ND280. The

	γ Bkg (%)	ν_e CC Bkg (%)	All Other Bkg (%)	Signal Efficiency (%)
Systematic Uncertainty	4.08	2.43	2.17	0.23

Table 4.9: The flux systematic uncertainty on different background topology event yields. Each background uncertainty is calculated relative to the topologies fractional contribution to the total background yield. The flux uncertainty effect on signal efficiency is also shown for a predefined limited phase space.

677 second is affect of flux uncertainties on event yield background topologies and signal
678 efficiencies specific to this analysis.

679 **Calculation of Total ν_e Flux at ND280**

680 The total electron neutrino flux at ND280 is calculated using the flux tuning file
681 and covariance matrix used in T2K oscillation analyses [116]. Correlated weights
682 for each flux systematic source are generated using a covariance matrix provided by
683 the T2K flux group. The weights are then applied to the tuned electron neutrino
684 flux, and integrated to calculate the total flux. This procedure is repeated for 10,000
685 pseudo-experiments. Each pseudo-experiment represents the total flux for a given
686 configuration of flux uncertainties. The mean of the distribution of calculated total
687 fluxes is then used to evaluate the mean total flux and the RMS used to estimate the
688 systematic uncertainty. Evaluating these parameters gives a total electron neutrino
689 flux of $\Phi_{\nu_e} = (2.23 \pm 0.14)$ neutrinos/cm²/10²¹ POT with an average electron
690 neutrino energy of $\langle E_{\nu_e} \rangle = 1.31$ GeV.

691 **Effect of flux systematic uncertainties on event yields and signal efficiency**

692 The flux systematic uncertainties are evaluated using the same covariance matrix
693 used in the total flux calculation. The effects of flux systematic uncertainties on
694 background event yields and signal efficiencies over 250 toys are shown in table 4.9.
695 Combining the background topologies, the effect of flux systematic uncertainties on
696 the total background event yield is 8.53%. The total flux systematic uncertainty on
697 signal efficiency as a fractional error is before phase space constraints are applied
698 is 1.11%. In the limited phase space defined in section 4.9.1, the flux systematic
699 uncertainty on signal efficiency is 0.23%. Comparisons can be made to the flux

700 systematic uncertainties calculated for the ν_e inclusive analysis: Here, the flux
701 systematic uncertainties on background event yields and signal efficiencies were
702 calculated to be 5.62% and 0.71% respectively [86].

703 4.9 Cross-Section Measurement

704 A cross-section measurement of the ν_e CC π^+ interaction on a carbon target, over
705 one bin in kinematic space, has been made. This section will outline the choice of
706 binning, phase-space constraints and the calculation made, before discussing the
707 results in the context of nominal interaction model predictions.

708 4.9.1 Phase-Space Constraints

Constraints are applied to the analysis at this stage to define a region of kinematic phase-space in which the cross-section can be best measured. The signal was defined as ν_e CC π^+ events that pass all of the following constraints:

$$\cos \theta_\pi > 0.5$$

$$\cos \theta_e > 0.707$$

$$p_\pi > 100 \text{ MeV}/c$$

$$p_e > 300 \text{ MeV}/c$$

709 The principal factors in choosing regions of phase-space to apply constraints are
710 the selection efficiency and overall number of signal events in any particular bin.
711 Kinematic constraints are applied to both the true lepton and most energetic pion
712 tracks at the truth level. Constraints are applied to high angle events, $\theta > 45^\circ$
713 for electrons and $\theta > 60^\circ$ for pions, since the acceptance due to detector effects is
714 essentially zero. Significant drops in signal efficiency and event yields are seen within
715 these regions, and are shown in figure 4.20. The true lepton angular constraint is
716 taken from the ν_e inclusive analysis and rejects events in the region $\cos\theta < 0.707$.
717 The pion angular constraints are defined by figures 4.20c and 4.20d. Events are
718 rejected in the region of $\cos\theta < 0.5$ due to the significant decrease in both signal

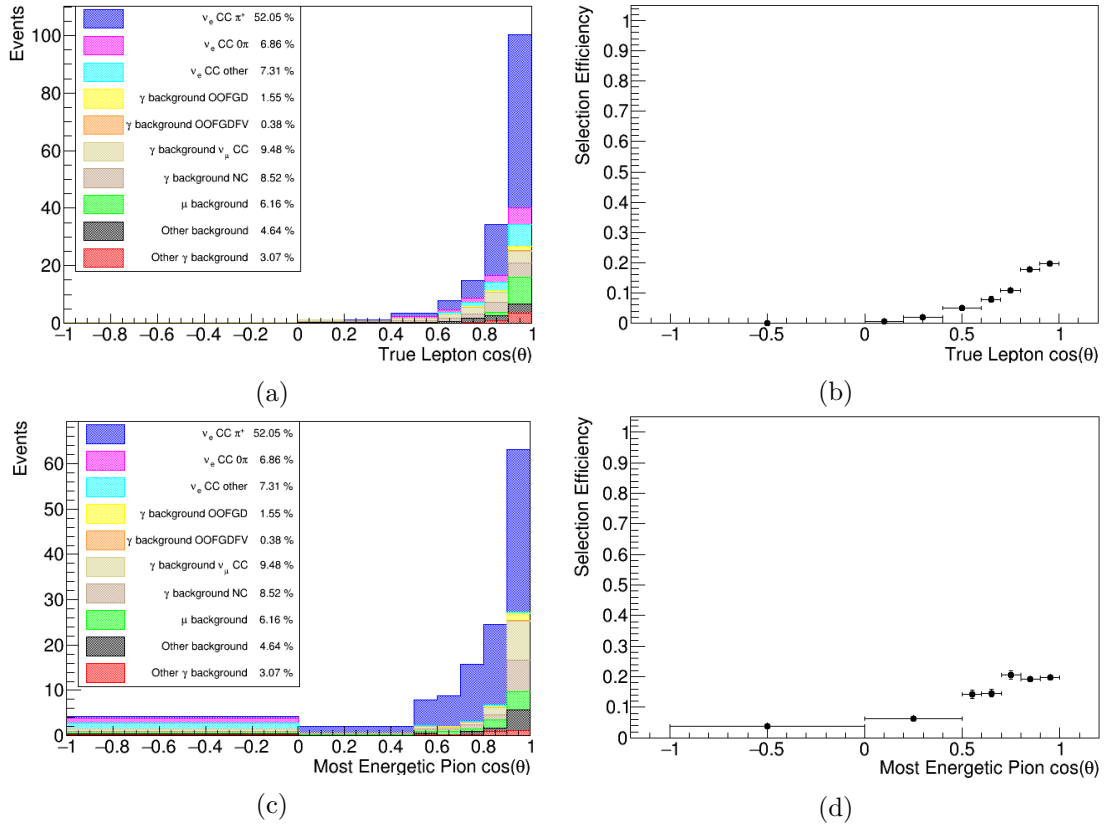


Figure 4.20: The angular kinematic phase-spaces for both the true lepton (top) and most energetic pion (bottom) tracks. The event yields are shown in (a) and (c), and the selection efficiencies are shown in (b) and (d). MC events are normalised to data POT.

719 event yield and efficiency. Backwards going most energetic pion tracks can occur
 720 when the pion candidate track selected at the reconstructed level, is not the most
 721 energetic pion in the event. This region has a negligible amount of signal events
 722 present, and is removed within the constraint.

723 Signal event yields in the background enriched low momenta bins are signifi-
 724 cantly smaller than across the rest of momenta space. Constraints are therefore
 725 applied to these regions also. Once again the post selection signal efficiencies and
 726 event yields, for both the true lepton and most energetic pion tracks, can be seen in
 727 figure 4.21. The signal contribution and selection efficiency at low lepton momenta
 728 is tiny, and thus a constraint is placed at $p > 300$ MeV/c. Moreover, a constraint of
 729 $p > 100$ MeV/c is used is placed on the most energetic pion track.

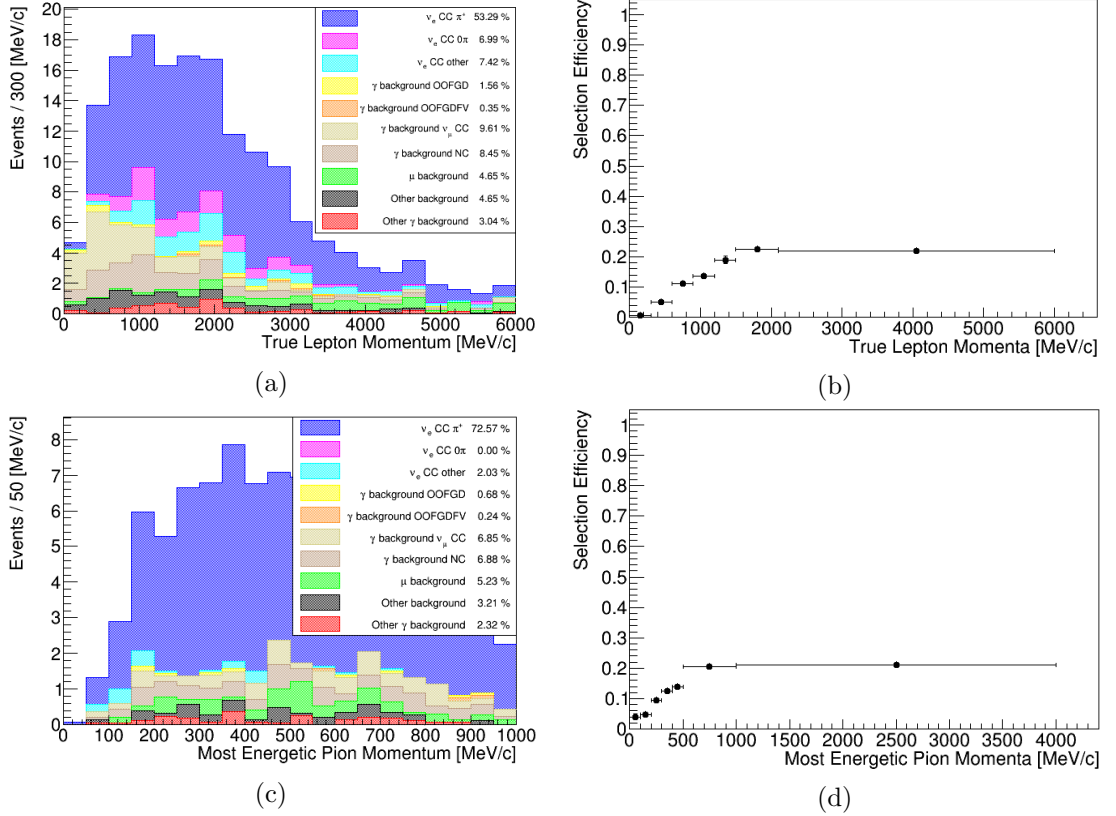


Figure 4.21: The momentum space for both true lepton (top), and most energetic pion (bottom) tracks. The event yields are shown in (a) and (c), and the signal efficiencies post selection are shown in (b) and (d). MC events are normalised to data POT.

730 4.9.2 Nominal NEUT Prediction and Validation

731 The MC used for this analysis is generated using NEUT D 5.4.0 (described further
 732 section 4.4). Cross-section predictions can be made independent of the analysis
 733 framework to contextualise and compare measured values with nominal MC. The
 734 predictions can also be used to validate the cross-section calculation methods to be
 735 used.

736 To obtain a prediction, the total number of predicted events must first be
 737 calculated. One million events were generated using NEUT version 5.4.0, and were
 738 uniformly randomly distributed between 50 MeV and 30 GeV. The simulated events
 739 modelled were all electron neutrino interactions on hydrocarbon targets. The highest
 740 momenta electron and most energetic positive pion are selected from the interaction.

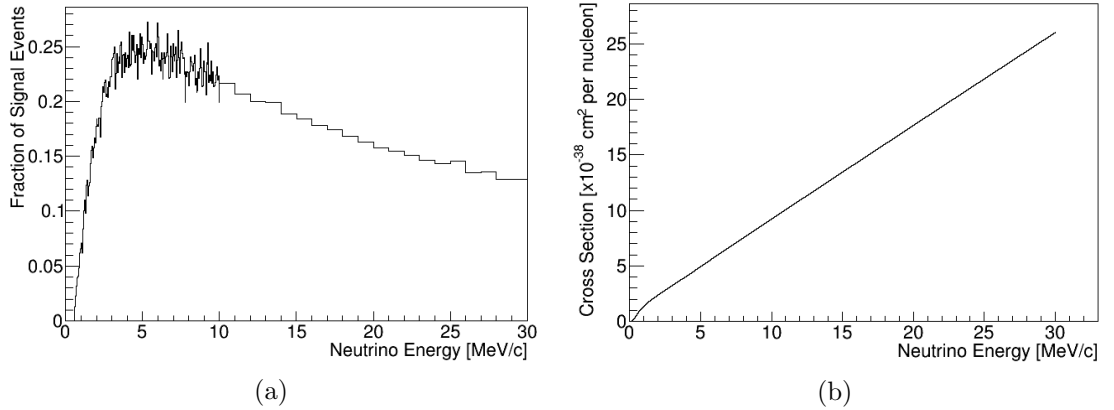


Figure 4.22: (a) The fraction of ν_e CC π^+ events in the NEUT generated sample, and (b) the NEUT ν_e cross-section, both as a function of true incoming neutrino energy.

741 If no post final state interaction electrons or pions are present, the event is defined as
742 background. For the event to be classed as signal, the selected electron and pion must
743 pass further cuts, predefined by the phase-space constraints outlined in section 4.9.1.
744 Events that do not pass the momentum and angular space cuts are put into the
745 background category. The fraction of signal events, and interaction cross-section,
746 both as a function of true neutrino energy are shown in figure 4.22. When multiplied
747 together these plots yield the cross-section of signal events as a function of true
748 neutrino energy. The binning for figure 4.22a is chosen by the fine binned flux MC
749 predictions (shown in figure 4.23) provided by the T2K beam group [116]. A fine
750 binning of 50 MeV is used up to 10 GeV, and then a courser binning of 1 GeV is
751 used from then on. The same flux MC file, "tuned13av7", is used as in the total
752 flux calculation in section 4.8.3. The predicted electron neutrino flux at ND280 as a
753 function of neutrino energy is shown in figure 4.23. The cross section of signal events
754 as a function of neutrino energy is combined with the simulated ν_e flux at ND280, to
755 give a predicted number of signal events. A ratio of the integrated number of signal
756 events against the total ν_e flux, as seen in equation 4.13, provides the predicted flux
757 averaged cross section of signal events over one global bin in neutrino energy space.

$$\sigma = \frac{\int \sigma(E)\Phi(E) dE}{\int \Phi(E) dE} \quad (4.13)$$

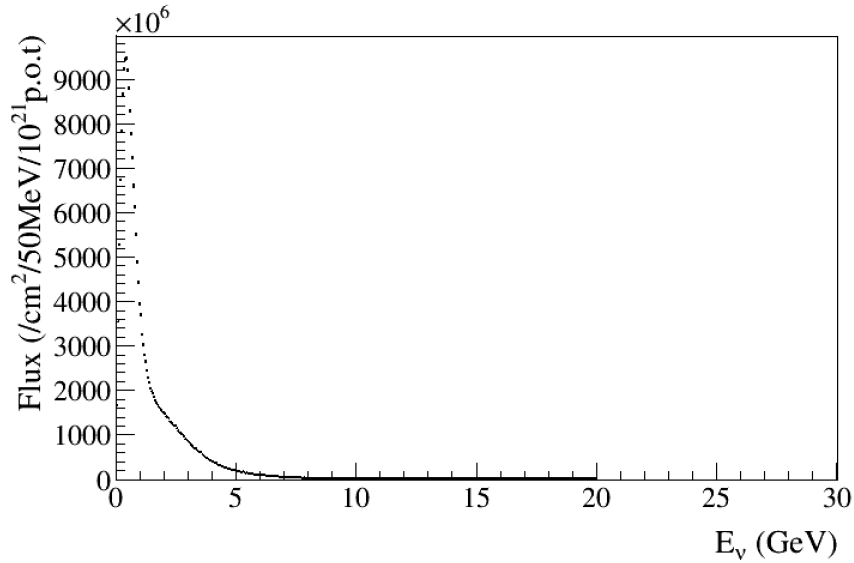


Figure 4.23: The predicted electron neutrino flux at ND280 as a function of neutrino energy.

758 Where σ is the flux averaged cross section prediction for some model, $\sigma(E)$ is the
 759 signal cross-section as a function neutrino energy, and $\Phi(E)$ is the electron neutrino
 760 flux as a function of neutrino energy. Applying equation 4.13 when using NEUT gives
 761 a predicted flux averaged cross section of $\sigma_{\text{NEUT}} = 2.22 \times 10^{-39} \text{ cm}^2$ per nucleon.

762 The same NEUT configuration used to calculate a flux averaged cross sec-
 763 tion prediction is also used to generate production 6T MC simulations. By pars-
 764 ing nominal NEUT ND280 MC in place of data through the calculation frame-
 765 work described in section 4.9.3, cross checks can be performed to validate both
 766 the calculation and the NEUT prediction, as the two methods should result in
 767 the same answer. ND280 MC when processed as data yields a cross section of
 768 $\sigma_{\text{ND280}} = (2.18 \pm 0.05) \times 10^{-39} \text{ cm}^2$ per nucleon. The error is taken from the 2.1%
 769 statistical uncertainty in ND280 MC events passing the phase-space constraints. The
 770 ratio of ND280 MC and NEUT prediction values suggest a difference of approximately
 771 1.8%, and therefore agree within statistical error cross-validating each method.

Parameter	Source of Systematic Uncertainty				Total (%)
	Detector (%)	Cross Section Model (%)	Flux (%)	Other (%)	
S	6.08	10.76	8.53	-	15.02
ϵ	4.13	2.56	0.23	2.43	5.43
Φ_{ν_e}	-	-	6.08	-	6.08
T	-	-	-	0.72	0.72
σ	7.35	11.06	10.47	2.53	17.11

Table 4.10: A summary of each type of systematic uncertainty and its contribution to each parameter in the cross-section calculation (equation 4.14). Other systematic uncertainties originate in the calculation of the relevant parameter and are explained further in the text. All uncertainties are quoted as the fractional error.

772 4.9.3 Cross-Section Calculation

773 A total flux averaged cross-section measurement over one global bin in reconstructed
774 neutrino energy space is calculated using equation 4.14.

$$\sigma = \frac{S}{\epsilon} \times \frac{1}{T\Phi_{\nu_e}} \quad (4.14)$$

775 Where S is number of signal events, ϵ is the signal efficiency, Φ_{ν_e} is the electron
776 neutrino flux, and T is the number of target nucleons. The signal is calculated
777 using $S = N - B_{MC}$, whereby N is the total number of data events and B_{MC} is the
778 number of background events predicted by MC. The electron neutrino flux at ND280
779 is calculated to be $\Phi_{\nu_e} = (2.23 \pm 0.14)$ neutrinos/cm²/10²¹ POT in section 4.8.3,
780 with an average neutrino energy of $\langle E \rangle = 1.31$ GeV. The signal efficiency is taken
781 as the post-selection signal efficiency after phase-space constraints are applied, and
782 is calculated to be $(25.35 \pm 0.61)\%$, whereby the uncertainty is taken as the binomial
783 error. The number of target nucleons, T , is calculated from the FGD1 fiducial mass
784 of 919.5 kg, which corresponds to $(5.54 \pm 0.04) \times 10^{29}$ nucleons [88]. A full breakdown
785 of the systematic uncertainty sources and their contributions to the parameters in
786 the cross-section calculation is shown in table 4.10.

787 The calculation is performed using the 'xsCalculation' package in Highland2,
788 purposely written in C++ for this analysis. The full production 6T FHC dataset
789 outlined in section 4.4 is used, and the MC normalisation to data is performed on a

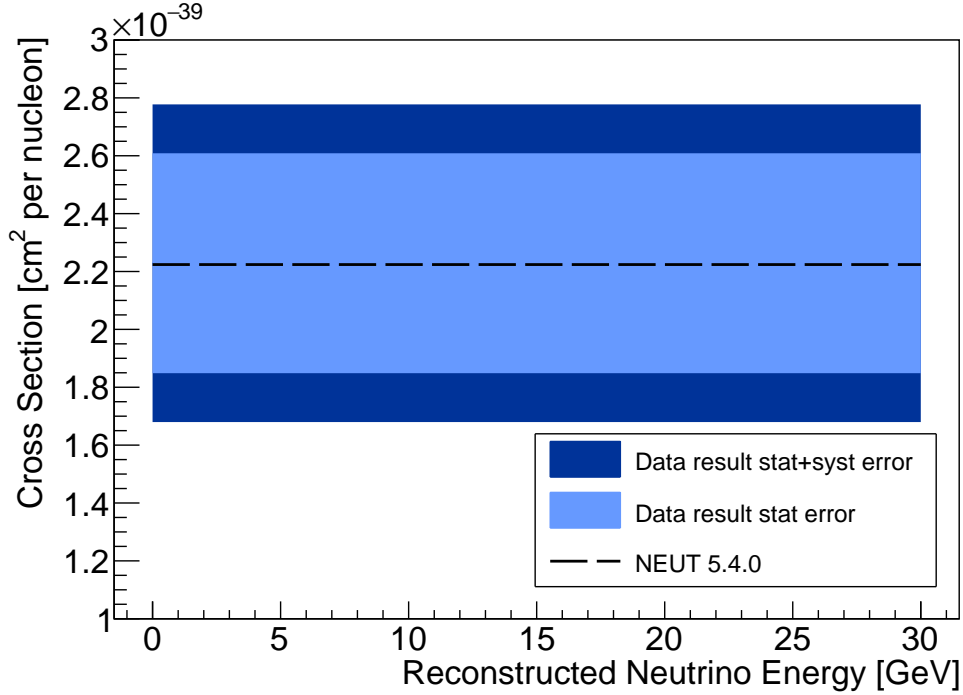


Figure 4.24: The flux integrated cross section prediction for nominal NEUT 5.4.0, compared to the data cross-section measurement in the context of systematic and statistical errors.

790 run-by-run basis. All systematic uncertainties calculated in section 4.8 are added in
 791 quadrature, and the statistical error in data is taken as \sqrt{N}/S . Statistical errors
 792 in MC are deemed to be negligible given the significantly large sample size. The
 793 total systematic and statistical errors are calculated to be approximately 17.11% and
 794 17.70% respectively. The same phase space constraints introduced in section 4.9.1 are
 795 applied to the cross-section calculation. Through equation 4.14, the total ν_e CC π^+
 796 cross-section, over an a reconstructed neutrino energy space of 0 GeV \rightarrow 30 GeV,
 797 is calculated to be $\sigma = (2.23 \pm 0.39(\text{stat.}) \pm 0.38(\text{syst.})) \times 10^{-39}$ cm² per nucleon.
 798 This result provides the first preliminary ν_e CC π^+ cross-section measurement on a
 799 carbon target.

800 A comparison of the ν_e CC π^+ cross-section calculated from ND280 data to the
 801 NEUT 5.4.0 nominal prediction, provided in section 4.9.2, is displayed graphically
 802 in figure 4.24, and numerically in table 4.11. It can be seen that the nominal
 803 MC cross-section prediction lies within both ranges defined by the systematic and

Measured σ [10^{-39}cm^2 per nucleon]	Nominal Predicted σ [10^{-39}cm^2 per nucleon]	$\langle E \rangle$ [GeV]
$2.23 \pm 0.39(\text{stat.}) \pm 0.38(\text{syst.})$	2.22	1.31

Table 4.11: A comparison of the measured ν_e CC π^+ cross section to the nominal prediction from section 4.9.2 using NEUT 5.4.0. The mean neutrino energy $\langle E \rangle$ is also shown.

804 statistical uncertainties on the data measurement. Nevertheless, comparisons against
805 different models including GENIE [117] and NuWro [118], as well as models with
806 more sophisticated resonant pion production treatments, are needed before more
807 complete conclusions on model performances can be made.

808 4.10 Super-Kamiokande Comparisons

809 Section 4.1 introduced an electron neutrino appearance study anomaly in the T2K
810 oscillation analysis. A far detector excess is seen in the $\nu_e\pi^+$ FHC 1 decay electron
811 sample. The probability of observing the 15 events seen in T2K's data samples,
812 assuming maximal CP-violation, relative to a prediction of 7 events, is 6.9% for the
813 best fit oscillation parameters. The analysis in this thesis provides the beginning
814 of a direct constraint on this process using the near detector. A preliminary study
815 of data-MC comparisons in a region of energy phase-space complimentary to the
816 far detector studies, gives initial insights in any potential excess seen in the near
817 detector. Two bins in reconstructed neutrino energy space have been defined. The
818 low energy bin from $0 \rightarrow 1.25$ GeV replicates the region of energy space the Super-K
819 $\nu_e\pi^+$ 1 decay electron sample is sensitive to. The high energy bin contains the
820 remaining phase space ($1.25 \rightarrow 30$ GeV) used in the cross-section measurement
821 above. Figure 4.25 shows the number of events for data and MC as a function of
822 reconstructed neutrino energy for the low and high energy bins. The data-MC ratios
823 are also provided in table 4.12. The Super-K analogous low energy bin shows the
824 data and MC agree within data statistical error. This provides preliminary hints
825 that the Super-K excess in this channel is a result of statistical fluctuation rather
826 than a systematic excess

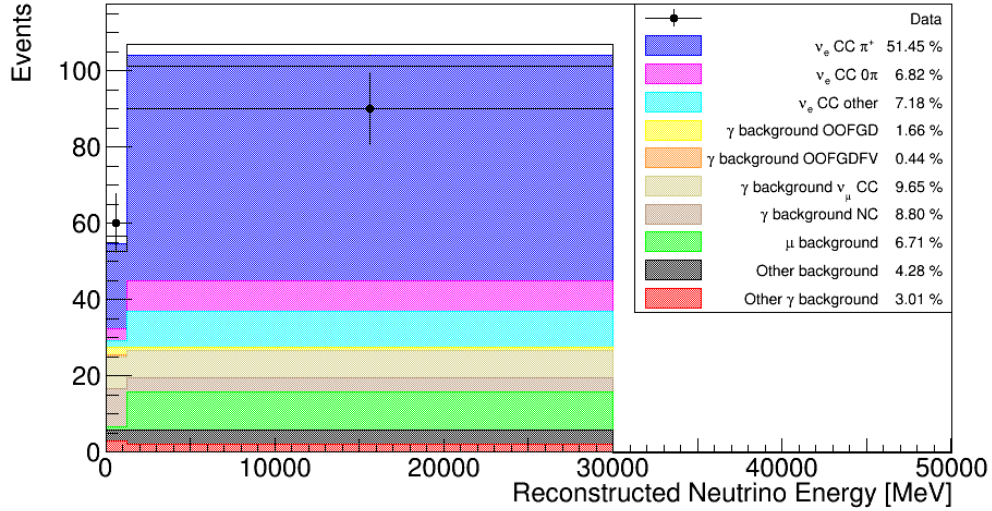


Figure 4.25: A data-MC comparison of the number of events split into two regions of reconstructed neutrino energy space. A threshold of 1.25 GeV is chosen to isolate a region of phase space that is comparable to SK. Detector systematic errors are displayed for the ND280 MC.

Energy Bin	Data	MC	Data-MC Ratio
0.00 → 1.25 GeV	60 ± 8	54.6 ± 2.0	1.10 ± 0.15
1.25 → 30.0 GeV	90 ± 10	104.1 ± 2.4	0.86 ± 0.10

Table 4.12: The number of data and MC events in the low and high bins of reconstructed neutrino energy space, the data-MC ratio is also shown. Error estimates on the data-MC ratios have been provided using the statistical error in data, and the detector systematic uncertainties in MC.

827 The far detector 1 decay electron sample was designed to add an additional
828 ν_e CC $1\pi^+$ channel, increasing the number of signal events in ν_e appearance studies
829 for the oscillation analysis. The selection takes one electron-like ring fully contained
830 in the detector fiducial volume with a visible energy above 100 MeV. Further cuts
831 dictate there must be exactly one decay electron, and the reconstructed neutrino
832 energy, calculated using the same CC Δ picture used in the near detector sample,
833 must be less than 1.25 GeV. The final selection cut is used to reject neutral pions.

834 When building a near detector constraint, comparative similarities and differ-
835 ences between the near and far detector samples should be discussed. An investigation
836 into the number of π^0 particles present in the ND280 signal sample can be seen in

837 the appendix figure A.3. The far detector has the ability to veto π^0 interactions
838 through unique signals. Therefore the likeness of the ND280 signal to the far detector
839 sample, in the region of momenta space comparable to SK, can be tested through the
840 amount of π^0 present. Studies show approximately 85% of signal events are absent
841 of neutral pions; consequently at least one π^0 exists in roughly 15% of events and are
842 topologically different to the 1 decay electron sample in the far detector. Moreover,
843 the far detector sample has 4π angular coverage, whereas the near detector sample
844 is more constrained to the forward going regions of angular phase space.

845 This study provides only a preliminary insight into potential data-MC
846 discrepancies, effects such as event migration across bin thresholds have not yet
847 been considered. A more complete analysis, with further investigations into topology
848 likeness, is needed to constrain any Super-K results using ND280 data.

1 Chapter 5

2 Hyper-Kamiokande

3 Hyper-Kamiokande (HK) [119] is a next generation water Cherenkov neutrino detector
4 that follows on from Kamiokande and Super-Kamiokande. Significantly larger than
5 it's predecessors, Hyper-Kamiokande will be the largest neutrino detector in the
6 world. HK will serve as the far detector in the long baseline neutrino experiment that
7 will eventually supersede T2K. Unless otherwise stated, it may be assumed that the
8 main reference for this chapter is the Hyper-Kamiokande Design Report (2018) [119].

9 5.1 Physics Goals

10 The physics goals of Hyper-Kamiokande are split into three main areas:

- 11 • Neutrino oscillations
- 12 • Nucleon decay searches
- 13 • Astrophysical observations

14 Neutrino oscillations can then further be divided into measuring the magnitude
15 of CP-violation in neutrino oscillations, the determination of normal or inverted
16 mass hierarchy, and precision measurements of known oscillation parameters. Hyper-
17 K aims to measure neutrino oscillation parameters through two neutrino sources.
18 Observing both atmospheric neutrinos and long baseline neutrinos provides com-
19plementary information. Assuming a total of 2.7×10^{22} POT with a beam power

20 of 30 GeV, Hyper-K is expected to be able to determine the leptonic CP violating
 21 phase, δ_{CP} , to better than 23 degrees for all possible values of δ_{CP} . Furthermore
 22 CP-violation could be established with a statistical significance of more than 3σ (5σ)
 23 for 76% (57%) of δ_{CP} parameter space. Currently there has not been an extens-
 24 ive study on the ability for measurements on the sign of Δm_{32}^2 or Δm_{32}^2 for mass
 25 hierarchy determination. It is predicted at the time Hyper-K becomes operational,
 26 the mass hierarchy could be determined to up to $\sim 4\sigma$ thanks to a combination of
 27 data from T2K and NOvA; and future reactor experiments such as RENO-50 [120],
 28 JUNO [121], ICAL [122], PINGU [123], and ORCA [124].

29 The decay of protons and bound nucleons are direct observable consequences
 30 of the violation of baryon number; a process believed to have an important role
 31 during the formation of the early universe. Furthermore baryon number violation
 32 is predicted in many Grand Unified Theories (GUTs) which allow transitions from
 33 quarks to leptons and vice versa. These GUTs predict the lifetime of the proton
 34 to be greater than 10^{30} years, so new experiments must be sensitive to this vast
 35 lifetime. Figure 5.1 demonstrates the future capabilities of Hyper-K in the 90% CL
 36 lifetime limits, in comparison with a number of GUTs. After 10 years of operation
 37 Hyper-K is sensitive to lifetimes that are predicted by a number of GUTs, through
 38 both $p \rightarrow e^+ \pi^0$ as well as channels involving kaons.

39 Hyper-K has the ability to set the energy threshold for detection to as low as
 40 several MeV; this enables event-by-event detection and reconstruction of astrophysical
 41 neutrinos from sources such as the sun, and supernovae. Using solar neutrinos and
 42 higher precision terrestrial matter effect [125, 126] measurements, Hyper-K aims to
 43 better understand neutrino oscillation behaviour in the presence of matter. Terrestrial
 44 matter effects hint at the use of atmospheric and long baseline neutrino experiments
 45 to measure CP-violation and mass hierarchy, as both of these parameters affect
 46 neutrino oscillation probabilities. Moreover, Hyper-K could feasibly help to resolve
 47 a current $\sim 2\sigma$ tension between the best fit values of Δm_{21}^2 in solar and reactor
 48 neutrino experiments; current predictions suggest that the discrepancy is due to
 49 solar neutrino interactions on matter.

50 An important astrophysical source of neutrinos is core collapse supernovae.

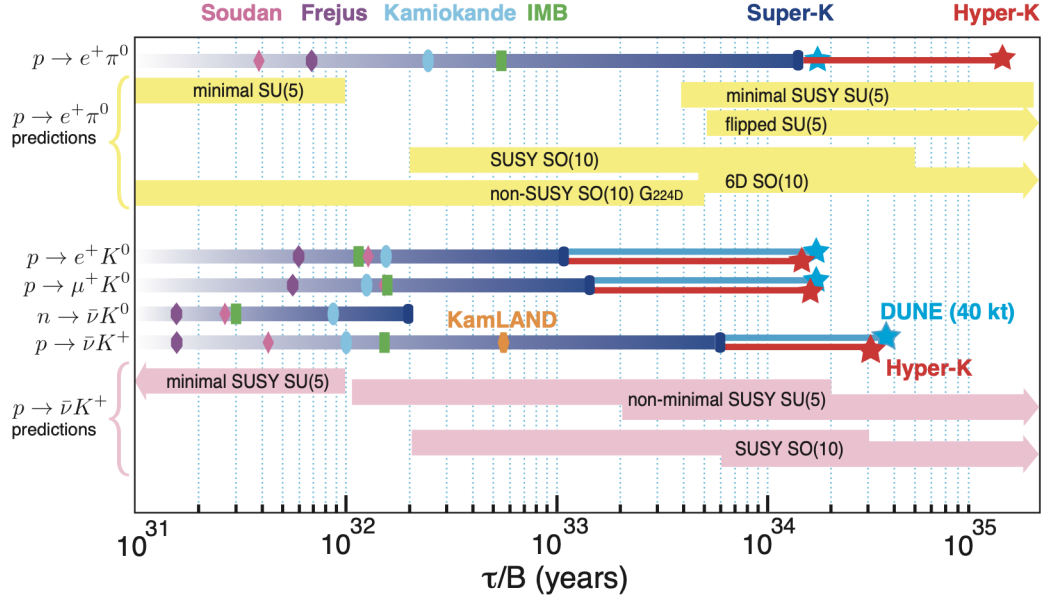


Figure 5.1: A comparison of the theoretically predicted rate of nucleon decay for a number of key modes, and the historical limitations for various experiments. The projected limits for Hyper-K and DUNE are based on 10 years of running. Figure taken from [119].

51 It is anticipated that if a supernova took place near the centre of our galaxy, Hyper-
52 K would observe $O(10^4)$ neutrinos in a time frame of approximately 1 second (a
53 calculation described in detail in [119]). Furthermore, the large volume of Hyper-
54 K increases its sensitivity to distant supernova $O(\text{Mpc})$ away. Hyper-K also has
55 the ability to precisely determine the arrival time of such neutrinos. Analyses
56 from core collapse supernova neutrinos can provide information not only about
57 supernova mechanics, but comparisons of ν_e and $\bar{\nu}_e$ flux during the neutronization
58 burst can yield information on neutrino mass hierarchy. Hyper-K can also study
59 other astrophysical processes, including dark matter and the detection of neutrinos
60 through solar flares [119].

61 5.2 Beam

62 The neutrino beam to be supplied to Hyper-Kamiokande will be an upgraded version
63 of J-PARCs beam [127], currently being used for T2K (see section 3.1.2). As

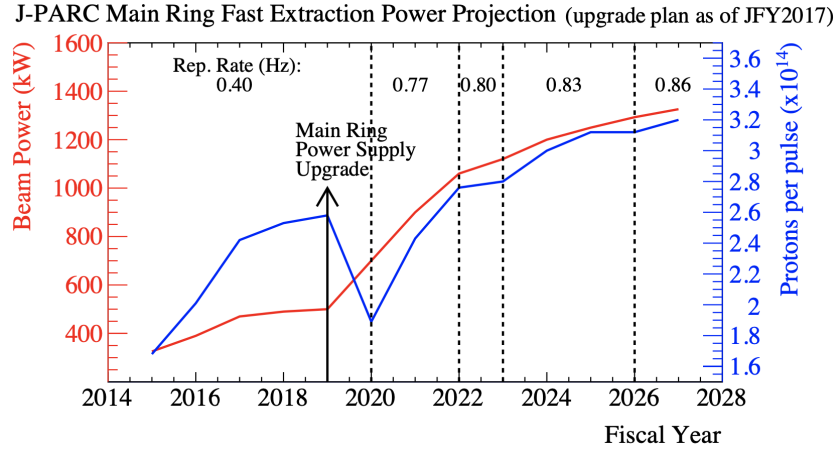


Figure 5.2: The projected main ring performance in fast extraction mode up to the year 2028. The protons-per-pulse, beam power and repetition rate are shown. Figure taken from [119].

64 of 2018 a beam intensity of 2.45×10^{14} protons-per-pulse (ppp), corresponding to
 65 ~ 485 kW of beam power, has been achieved in the main ring fast extraction mode
 66 operation [119]. A number of short term [128], and longer term [129, 130], upgrades
 67 are planned for J-PARCs accelerator chain; starting within the next couple of years
 68 and continuing throughout HKs construction and data-taking periods. The projected
 69 beam performance up to 2028 is shown in figure 5.2. High intensity studies of current
 70 accelerator performance suggest a beam power of 1-1.3 MW can be achieved post
 71 beam upgrades. Conceptual design studies are also in progress for operation at beam
 72 powers greater than 2 MW [131]. The approaches being considered include enlarging
 73 the main ring (MR) aperture, raising the rapid cycling synchrotron (RCS) top energy,
 74 or the insertion of an emittance-damping ring between the MR and RCS.

75 5.3 Near Detector Complex

76 Like T2K, Hyper-K will require a suite of near detectors to measure signal and
 77 background processes relevant for neutrino oscillations. Event rates at Hyper-K will
 78 be predicted through extrapolations from measured event rates at the near detector.
 79 Maximising systematic cancellations when extrapolating is desirable so use of the
 80 same target nuclei as used in the far detector, and enhanced angular acceptance is

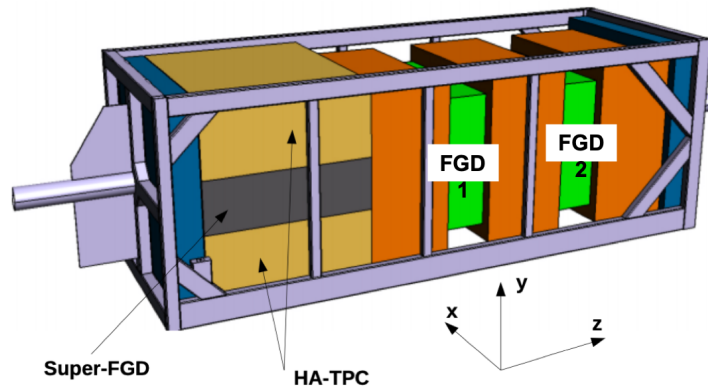


Figure 5.3: CAD model of the proposed ND280 detector post-upgrades. The upstream segment of the detector now consists of two High-Angle TPCs (brown) with a scintillator detector Super-FGD (grey) intersecting them. The beam and magnetic field are orientated approximately parallel to the z and x axis respectively. The two FGD sub detectors present in the current status of ND280 are also labelled for context. Figure edited from [132].

81 required.

82 5.3.1 ND280 Upgrade

83 Hyper-K will use an upgraded version of T2Ks ND280 detector complex previously
 84 described in section 3.3. An official T2K project since 2017, the ND280 hardware
 85 upgrade [57] has a goal of reducing the total systematic uncertainties on neutrino
 86 event rate extrapolation to the far detector to better than 4%. The design aims to
 87 improve the acceptance of high angle or backwards-going particles. This is achieved
 88 through the addition of a new scintillator target detector rotated parallel to the
 89 neutrino beam direction. Sandwiching the target detector with two horizontal High-
 90 Angle TPCs (HA-TPC) achieves almost full 4π angular acceptance. These three
 91 sub-detectors will be situated upstream of the tracker, replacing the current pi-zero
 92 detector. A schematic diagram of ND280 post-upgrade is shown in figure 5.3.

93 The ND280 upgrade keeps the current tracker and surrounding ECal modules.
 94 The POD detector, seen in figure 3.6, is to be replaced but the upstream and POD
 95 ECals will be kept to veto entering muons and photos from interactions in the

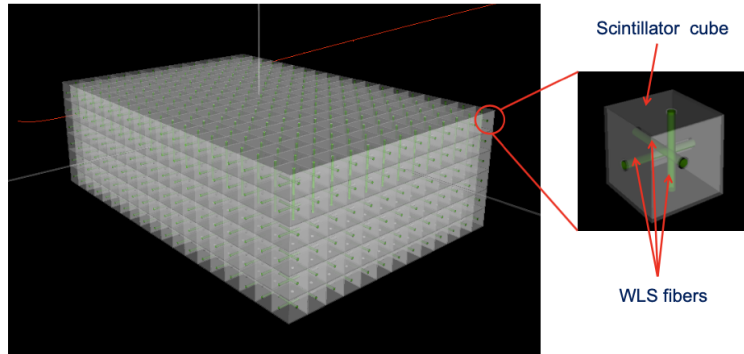


Figure 5.4: A schematic concept of the design of Super-FGD, demonstrating the composition of each scintillator cube and WLS fibres. Taken from [132].

96 sand around the detector. The 2 m long horizontal High-Angle Time Projection
 97 Chambers (HA-TPC) aim to replicate the high performance of the existing TPCs
 98 inside ND280. This requirement ensures key features such as: 3D reconstruction,
 99 particle identification, charge and momentum measurements are retained. Between
 100 the HA-TPCs, Super-FGD will provide a new high resolution 3D scintillator detector.
 101 Conceptual aims of Super-FGD were to provide a sufficiently large target mass, the
 102 acceptance of high-angles charged leptons, and the ability to identify and reconstruct
 103 short tracks of low energy hadrons near the interaction vertex. Super-FGD is
 104 composed of small plastic scintillator cubes read out by three orthogonal wavelength
 105 shifting (WLS) fibres. A concept diagram of Super-FGD can be seen in figure 5.4.
 106 The size of each cube is $1 \times 1 \times 1 \text{ cm}^3$. The total number of cubes in the baseline
 107 design is 2,064,384 arranged in a $192 \times 192 \times 56$ fashion, and 58,368 channels
 108 respectively. Every WLS fibre terminates at a Multi-Pixel Photon Counter (MPPC)
 109 to readout channels for each plane.

110 In addition, surrounding the new horizontal tracker will be six thin Time-
 111 of-Flight (TOF) scintillator layers. The goal with the new TOF system is to
 112 improve reconstruction of backward-going tracks. Studies are currently undergoing
 113 to understand the impact of TOF on particle identification.

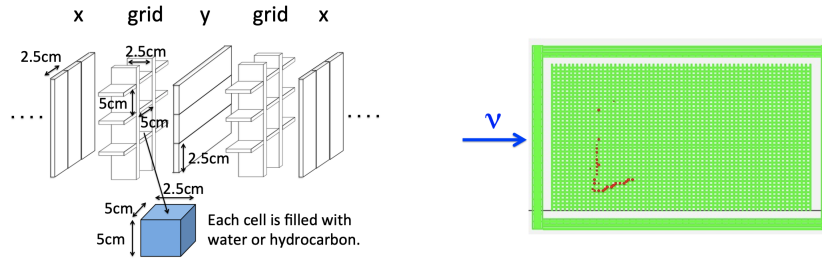


Figure 5.5: Left: A schematic representation of the plastic scintillator bars arrangement inside of WAGASCI. Right: A monte-carlo event display of a charged current neutrino interaction in WAGASCI. Figures taken from [119].

114 5.3.2 WAGASCI

115 Another detector being considered for the ND280 upgrade is WAGASCI (Water
 116 Grid And SCIntillator detector). The concept of WAGASCI is to develop a target
 117 detector, filled primarily with water, to measure neutrino interactions with high
 118 precision and large angular acceptance. Particles can be tracked across the full 4π
 119 solid angle using scintillator bars are arranged into a 3D grid like structure. The
 120 remaining voids created are defined as cells that can be filled with either water
 121 or hydrocarbon, changing the neutrino target medium. Figure 5.5 demonstrates
 122 both the conceptual design, and tracking of charged current interactions through
 123 simulation. WAGASCI modules are first being installed and tested at the J-PARC
 124 near detector hall, surrounded by muon detectors [133]. The test experiment will aim
 125 to measure the cross sections on both water and hydrocarbon targets, and has been
 126 approved by J-PARC PAC as test experiment T59 [134]. The INGRID detector has
 127 previously established the technique of comparing the interaction rates on the two
 128 targets to measure an inclusive water to hydrocarbon charged current cross section
 129 ratio [135]. WAGASCI, as experiment T59, aims to measure the ratio to a precision
 130 of 3% or better.

131 The proposed grid target mass of ND280 is expected to roughly double after the
 132 upgrades. Since the neutrino event rate is proportional to the target mass, it is also
 133 expected that the event rate will approximately double post-upgrade. A simplified
 134 MC study, without full event reconstruction, has been used to predict the number of
 135 neutrino events pre and post ND280 upgrade. The results of this study can be seen

Detector Configuration	Target Mass (ton)	Number of Selected Events		
		CC- ν_μ (ν) beam	CC- $\bar{\nu}_\mu$ ($\bar{\nu}$) beam	CC- ν_μ ($\bar{\nu}$) beam
Current ND280	2.2	95,860	27,433	14,862
ND280 Upgrade	4.3	199,775	54,249	28,370

Table 5.1: A comparison of the predicted number of neutrino events for the current ND280 and ND280 upgrade target mass respectively. The predictions correspond to 1×10^{21} POT. Table adapted from [119]

136 in table 5.1.

137 5.3.3 High Pressure Time Projection Chamber

138 Longer term ND280 upgrades could have the potential of introducing high pressure
139 gas time projection chambers (HPTPC). Advantages of using HPTPC detectors are
140 that they provide detailed vertex resolution, good particle identification, full angular
141 coverage, and sensitivity to low momenta protons. The HPTPCs are proposed to
142 replace the current TPCs. ND280 would be able to contain HPTPCs of size 8 m^3
143 under 10 bar of pressure. HPTPCs, using a gas target, have the strength that the
144 target medium is interchangeable; a wide of successful gas mixtures having already
145 been used in ND280 to test different nuclear model components. Furthermore, a gas
146 HPTPC would yield a relatively pure ν_e sample. This is because of the reduction in
147 photon background prominent in current ND280 ν_e -CC analyses, analogous to that
148 described in chapter 4.

149 5.4 Intermediate Water Cherenkov Detector

150 To better constrain systematic uncertainties, Hyper-K will also house a new inter-
151 mediate water Cherenkov detector (IWCD) [136]. The physics motivation for the
152 detector is to constrain the cross section on water directly, with the same solid angle
153 acceptance as Hyper-K thereby eradicating the need for a subtraction analysis. Water
154 Cherenkov detectors also have the capability of detecting pure ν_μ -CC, ν_e -CC and
155 $\text{NC}\pi^0$ samples due to their excellent particle identification capabilities. Moreover,
156 background rates in nucleon decay searches such as $\text{CC}\pi^0$ and kaon production from

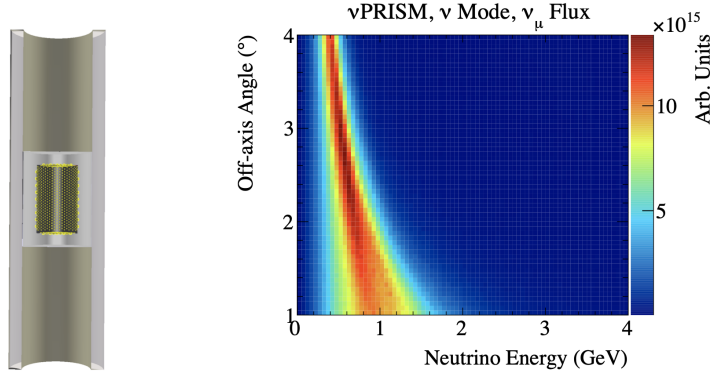


Figure 5.6: (Left) A diagram demonstrating the conceptual design for NuPRISM. (Right) The ν_μ flux energy dependence shown as a function of off-axis angle between $1^\circ - 4^\circ$. Figure taken from [119]

157 neutrino interactions can be measured [119]. The IWCD's measurements, in tandem
 158 with ND280's magnetised tracking abilities, are essential to lowering the systematic
 159 uncertainties needed to achieved Hyper-K's physics goals. To contain enough muons
 160 up to the momentum region of interest to the far detector, the IWCD must be
 161 large enough in size. Furthermore the IWCD must be far enough from the neutrino
 162 production point to minimise pile-up of interactions in the same timing bunch. These
 163 parameters constrain the detector to be of kilotons in size, and approximately 1-2 km
 164 from neutrino beam production point at J-PARC [136].

165 Following on from the conceptual design of NuPRISM [137], one design for
 166 IWCD consists of a detector orientated with it's cylindrical polar axis in the vertical
 167 direction. A feature being considered for IWCD is to span over a range of off-axis
 168 angles to measure the final state leptonic response over numerous neutrino spectra
 169 peaked at different energies. NuPRISM design featured a 10 m tall inner-detector
 170 located 1 km downstream from J-PARC. A crane system enables the detector to be
 171 moved vertically inside of a 50 m pit to yield an off-axis angular range between $1^\circ - 4^\circ$
 172 which the detector can traverse. NuPRISM's inner-detector design holds 3215 8 inch
 173 inward facing PMTs giving a photo-coverage of approximately 40%. A conceptual
 174 drawing of the NuPRISM detector alongside the ν_μ flux dependence on off-axis
 175 angle is shown in figure 5.6. There are three primary reasons why the IWCD would

176 want to probe a range of off-axis angles. The first is to eliminate model dependent
 177 uncertainties in the near to far extrapolations which arise from differing flux at near
 178 and far detectors due to oscillations, and poorly understood nuclear effects in final
 179 state lepton kinematics as a function of neutrino energy [137]. Secondly, the fraction
 180 of intrinsic electron (anti)neutrinos increases as a function of off-axis angle. With an
 181 increase in purity at high off-axis angles, measurements of electron (anti)neutrino
 182 cross sections relative to muon (anti)neutrino cross sections can be achieved with
 183 higher precision. Measurements of $\sigma_{\nu_e}/\sigma_{\nu_\mu}$ have the potential to be made with 3%
 184 precision. This level will decrease for their anti particle counterparts given the larger
 185 NC background in the $\bar{\nu}_e$ sample and more prominent wrong-sign background for
 186 both $\bar{\nu}_e$ -CC and $\bar{\nu}_\mu$ -CC interactions. Thirdly, at 1 km baseline the neutrino spectrum
 187 peak varies from 1.1 km/GeV to 2.5 km/GeV between $1^\circ - 4^\circ$ respectively. Within
 188 this region it is possible to search for sterile neutrino induced oscillations consistent
 189 with the LSND [21, 138] and MiniBooNE [139, 140] ν_e and $\bar{\nu}_e$ appearance anomalies
 190 described in section 2.1.2. The IWCD can search for the oscillation pattern not
 191 only through reconstructed energy but also via reconstructed off-axis angle given the
 192 neutrino spectrum has a dependence on it. Preliminary studies hint that much of
 193 the LSND allowed region can be excluded at 5σ [137, 141].

194 Further design concepts for the IWCD includes the possibility of using gad-
 195 olinium (Gd) doping [142] to exploit neutron tagging. Gd doping in water Cherenkov
 196 detectors is one way of enhancing neutrino and antineutrino separation sensitivity
 197 through the detections of neutrons in the final state, or protons in the case of using
 198 Water-based Liquid Scintillator (WbLS) compounds [143]. Combined with IWCD
 199 4π detector coverage, statistical separation of primary interaction nodes becomes
 200 possible [144]. Neutron tagging allows final state interaction effects within differ-
 201 ent interaction models to be probed in more detail. Moreover, it provides charge
 202 separation information through the enhanced presence of final state neutrons for
 203 $\bar{\nu}$ -CC interactions. This provides a constraint on wrong-sign backgrounds and al-
 204 lows comparisons of neutrino and antineutrino cross-section measurements on water.
 205 These factors reduce critical systematic uncertainties on atmospheric neutrino oscil-
 206 lation and beam δ_{CP} analyses. Simulations of neutron tagging have been performed

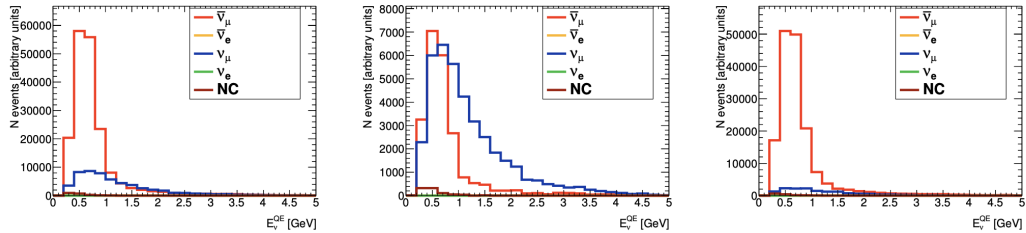


Figure 5.7: The composition of the one muon-like ring sample for the TITUS detector during antineutrino mode running. The effect of different neutron selections is shown. From left to right, before neutron tagging, no tagged neutron, at least one tagged neutron. Figure taken from [145].

207 using the TITUS concept detector. TITUS [145] was originally proposed to be an
 208 intermediate detector for Hyper-K with neutron tagging capabilities. Simulations
 209 demonstrated in figure 5.7 suggested the selection purity of $\bar{\nu}CCQE$ interactions
 210 increases significantly when selection at least one neutron.

211 In principal, it is possible to combine the two techniques described above
 212 using Gd loading inside an off-axis spanning detector. Nevertheless, it should be
 213 noted that a Gd loaded detector must be sufficiently far from the neutrino beam
 214 origin to limit the beam induced entering neutron background. However, the further
 215 downstream, the larger the excavated volume needed for an off-axis spanning detector.
 216 Preliminary studies suggest the entering neutron rate is low enough for the off-axis
 217 spanning detector located 1km downstream from the neutrino production point [119].

218 The IWCD can also be used as a supernova alarm, independent of the far
 219 detector. Additionally with Gd doping capabilities, neutrino type discrimination
 220 would be possible in the event of a supernova in the local galaxy.

221 5.5 Far Detector

222 The Hyper-Kamiokande far detector is the next generation water Cherenkov detector
 223 following Super-Kamiokande. The candidate site for Hyper-K is located 8 km south
 224 of Super-K, in the Tochibora mine near Kamioka town, Japan. Lying 650 m under-
 225 ground, the detector will be situated under the peak of Nijuugo-yama, corresponding
 226 to approximately 1,750 m.w.e. The cavern has been designed to view the same 2.5°

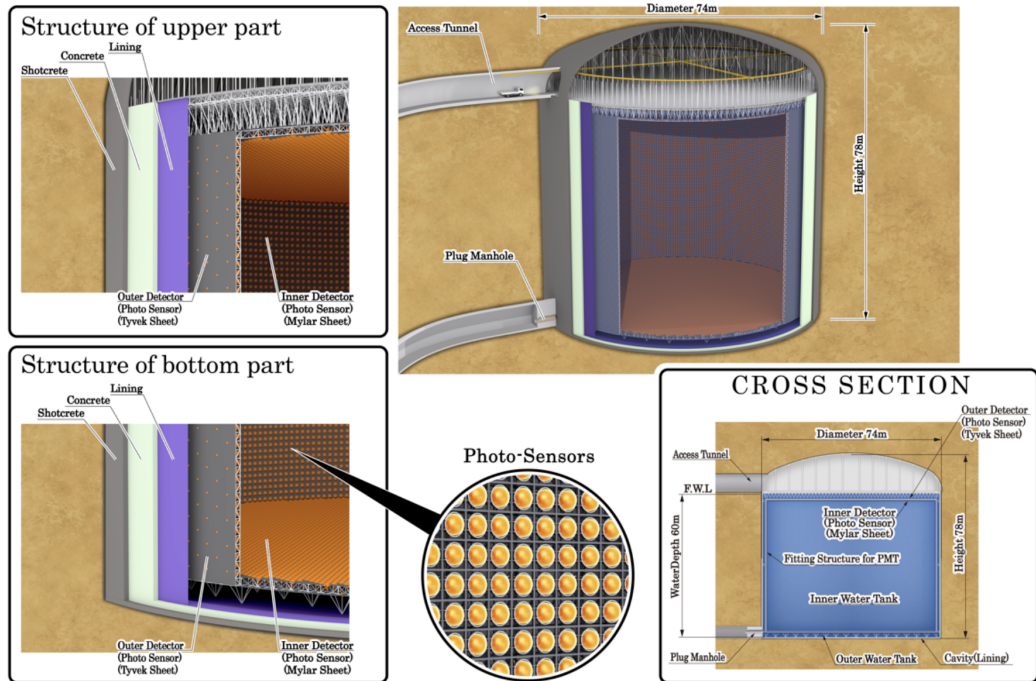


Figure 5.8: Schematic view demonstrating the 1TankHD design for the Hyper-K far detector. The multiple diagrams demonstrate different sections of the detector. Taken from [119].

227 off-axis angle that Super-K currently has with J-PARC.

228 The Hyper-K far detector design is a one vertical cylindrical water tank with
 229 40% photo-coverage. In accordance with the latest Hyper-K design report [119], this
 230 tank design will be referred to as 1TankHD throughout this thesis. A schematic view
 231 of the detector is shown in figure 5.8. The tank design is 60 m in height with a 74 m
 232 diameter, giving it a fiducial volume eight times larger than Super-K. This would
 233 also make Hyper-K the largest water Cherenkov experiment to have ever existed. A
 234 summary of the key parameters of the 1TankHD design relative to it's predecessors,
 235 Super-K and KAM, is shown in table 5.2. The measurement technique of Hyper-
 236 K is analogous to Super-K, employing ring-imaging water Cherenkov techniques
 237 to measure neutrino interactions, and possible nucleon decays. To determine CP-
 238 violation within a few % accuracy, it is estimated that $O(10^3)$ electron neutrino signal
 239 events are needed to be accumulated from the J-PARC beam [119]. This equates to
 240 a fiducial mass of $O(10^2)$ kton. Furthermore, with $O(10^{35})$ nucleons contained in

	KAM	Super-K	Hyper-K
Depth	1,000 m	1,000 m	650 m
Tank Diameter	15.6 m ϕ	39 m ϕ	74 m ϕ
Tank Height	16 m	42 m	60 m
Total Volume	4.5 kton	50 kton	258 kton
Fiducial Volume	0.68 kton	22.5 kton	187 kton
Inner Detector (ID) PMTs	948 (50 cm ϕ)	11,129 (50 cm ϕ)	40,000 (50 cm ϕ)
Outer Detector (OD) PMTs	123 (50 cm ϕ)	1,885 (20 cm ϕ)	6,700 (20 cm ϕ)
Photo-sensor Coverage	20%	40%	40%
OD Thickness	~ 1.5 m	~ 2 m	1 ~ 2 m
Single-photon Detection Efficiency (ID)	unknown	12%	24 %
Single-photon Timing Resolution (ID)	~ 2 nsec	2-3 nsec	1 nsec

Table 5.2: A summary table demonstrating the key parameters of the Hyper-K 1TankHD design with a comparison to its predecessors. Figures for the past KAM [146, 147] and present Super-K [148, 149] experiments have been taken for KAM-II and SK-IV respectively. The single photon detection efficiency is taken as a product of the quantum-efficiency peak at 400 nm, photo-electron efficiency, and threshold efficiency. Table has been adapted from [119].

241 this mass of water a sensitivity to nucleon lifetime of 10^{35} years is possible.

242 The neutrino target medium for tank will be highly transparent ultra-purified
243 water. Hyper-K will adopt the current Super-K water system designs, ensuring
244 the attenuation length for 400 nm-500 nm photons is above 100 m, and a radon
245 concentration level below 1 mBq/m³. For Hyper-K this requires water to be processed
246 at a rate of 310 tons/hour, and 50 Nm³ of radon free air to be generated every hour.
247 Depending on the success of the recent addition of gadolinium sulfate to Super-K, the
248 option to Gd-load Hyper-K is available. Therefore, the design criterion for Hyper-K
249 must be adaptable for this possibility (an important note for chapter 6).

250 Part of the design specification for Hyper-K is the need for single photon
251 sensitivity. This enables reconstruction of the spatial and timing distributions of the
252 Cherenkov photons which are emitted through neutrino interactions and nucleon
253 decay. Hyper-K will use newly developed ultrasensitive PMTs (Hamamatsu R12860),
254 with higher-efficiency and higher-resolution than those currently used in Super-K
255 (Hamamatsu R3600). This upgrade will amplify faint signatures, enhancing the
256 sensitivity to clean proton decay searches via $p \rightarrow e^+ + \pi^0$ and $p \rightarrow \bar{\nu} + K^+$ decay

257 modes. Increased sensitivity to neutron signatures will also benefit the observation
258 of electron anti-neutrinos from supernovae. As seen in table 5.2, the diameter of
259 Hyper-Ks inner-detector and outer-detector are 50 cm and 20 cm respectively; the
260 photo-coverage for the inner-detector and outer detector photomultipliers will be
261 40%.

262 As well as the 1TankHD design, the possibility of an additional second tank
263 (2TankHK-staged) is also being considered. The second detector would be stationed
264 downstream of the first. A possible location being considered is around Mount
265 Hakamagoshi, lying at a baseline of 335 km and an off-axis angle of 2.4° . Due to the
266 magnitude of Mount Hakamagoshi's elevation (1,159 m), the second detector have
267 greater sensitivity to low energy neutrinos such as those from supernovae and the
268 sun.

269 As of February 2020, the Hyper-Kamiokande project has been officially
270 approved [150]. Construction is due to begin in April 2020, and operations will
271 proceed from 2027.

1 Chapter 6

2 Optical Calibration

3 A new generation of neutrino detectors is fast approaching. Water Cherenkov
4 detectors, such as the large scale Hyper-Kamiokande detector with over 40,000
5 photo-sensors, need to include optical calibration systems to monitor the properties
6 of the water, and operation of the photosensors. The Hyper-Kamiokande physics
7 goals dictate that the detector must be understood to the level of a few percent
8 which can only be achieved through careful calibration.

9 This chapter will outline the proposed optical calibration system for the
10 Hyper-Kamiokande experiment with a primary focus on the light injection system
11 developed in the UK. In particular, the research and development of optical diffuser
12 based technologies will be discussed in detail. Diffuser results are shown both in
13 the context of laboratory experiments at the University of Warwick, as well as
14 deployments inside the Super-Kamiokande detector. Later, a discussion of the future
15 of optical diffuser research and development is provided, following the information
16 gained from previous deployments.

17 6.1 Proposed Hyper-Kamiokande Optical Calibration 18 System

19 The Hyper-Kamiokande detector, described in section 5.5, will be the largest water
20 Cherenkov neutrino detector in the world. The technical design of Hyper-K features
21 an inner and outer detector, containing 40,000 50 cm and 6,700 20 cm diameter

22 PMTs respectively. Both inner and outer detectors need optical calibration systems.
23 The calibration system will be an extension of the successful Super-K system, using
24 several calibration techniques established over two decades of operation. This section
25 will outline the proposed optical calibration systems for the Hyper-K detector,
26 with a focus on the light injection system under development by the Hyper-K UK
27 collaboration.

28 **6.1.1 Light Injection System**

29 For water Cherenkov detectors such as Super-K and Hyper-K, it's important to
30 understand the response of the photo-sensors used. The light injection system aims
31 to measure timing, gain and multi-photon responses of the inner detector PMTs.
32 In addition, of particular importance to large scale detectors such as Hyper-K, is
33 to understand properties of the medium, such as absorption lengths and photon
34 scattering probabilities as a function of depth. These are the primary measurements
35 the light injection system aims to cover.

36 As in Super-K, the optical calibration system consists of multiple light injection
37 points around the detector. Pulses of light will be controlled electronically at the
38 top of the tank and fed through optical fibres to the injection points. Two different
39 sources of light pulses are currently being considered. The first is a fibre coupled
40 commercial laser; the second uses short (approximately 1 ns) pulses produced from
41 either LEDs, laser diodes, or solid state optical devices. In order to preserve the fast
42 light pulses, needed for PMT timing calibrations over the dimensions of Hyper-K,
43 graded index fibres will be used. The alternate step-index fibres suffer from modal
44 dispersion reducing the ability to separate pulses over larger distances, and thus
45 inhibiting timing calibration over sub nanosecond pulses. The monitoring of light
46 injection into the detector is essential for some calibrations to be made. A fraction
47 of the light produced will be redirected towards optical sensors, most likely monitor
48 PMTs. This would allow for comparisons to be made on a pulse-by-pulse basis. The
49 light is injected into the tank using two different devices: the narrow-beam collimator
50 and wide-beam diffuser each with their own set of objectives outlined below.

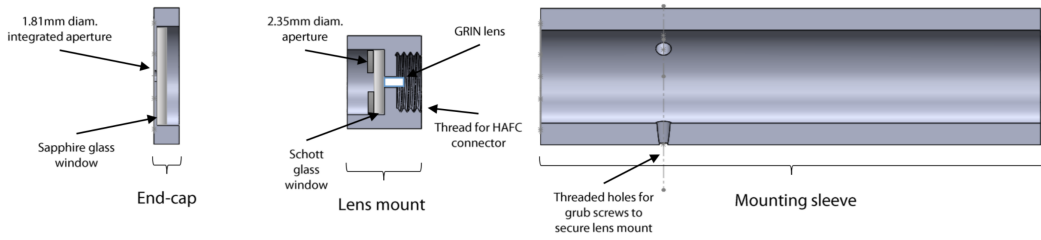


Figure 6.1: A schematic diagram of the collimator design used in the Super-K deployment. Taken from [151].

51 **Narrow-Beam Collimator**

52 In order to measure the optical properties of the water, a narrow beam of collimated
 53 light is needed. The collimator is designed to illuminate a small subset of PMTs.
 54 During the Super-K deployments, an opening angle of 3.5° was used to illuminate
 55 a 5×5 array of PMTs, however this value for Hyper-K is still to be determined.
 56 Illuminating a smaller array of PMTs is ideal for measuring the attenuating and scat-
 57 tering properties of light in water, which will be performed over multiple wavelengths
 58 ranging from 320 nm to 500 nm. Water attenuation is measured by observing light
 59 levels on given PMTs inside the illuminated array from the narrow beam. The
 60 scattering length is measured by monitoring the light detected by PMTs outside the
 61 collimated beam as a function of time and photon path length.

62 At present the final design for the collimator is still in development at
 63 the University of Warwick. The design tested and validated during the Super-K
 64 deployments (described in section 6.5) is shown in figure 6.1. The collimation is
 65 generated by a gradient-index (GRIN) lens, followed by a series of apertures. The
 66 GRIN lens has a pitch of $\frac{1}{4}$ which means it has a length equal to one quarter of a sine
 67 wave, and thus collimates to a point source at the surface of the lens. The optical fibre
 68 was secured in place to ensure correct alignment and the optics were mounted in a
 69 waterproof stainless steel assembly. Other potential designs for collimation, including
 70 the use of an achromatic doublet lens, are currently being considered.

71 **Wide-Beam Diffuser**

72 The motivation of using diffuser technology is to create a wide-angle beam of
73 diffuse light to illuminate as many PMTs as possible. Provided this is done a
74 well-understand beam which is, ideally, uniform in spatial intensity and timing,
75 calibrations of the PMT gain and timing responses can be performed. Measurements
76 of inter-PMT energies can give an indication of "hot" or "cold" PMTs which can
77 then be calibrated. The performance criterion for Hyper-K require the diffuser to
78 produce a well understood light cone over a $\pm 40^\circ$ angular range; this illuminates
79 approximately 1000 PMTs on the far side of the inner detector. The calibration of
80 PMT timing requires a short pulse with known origin and time. The diffuser must
81 therefore also ensure there is no time dependence as a function of angle. Each PMT
82 will be illuminated by at least two diffusers to allow for temporal cross calibration of
83 the fibres.

84 The design of the diffuser is discussed in detail within sections 6.3 and 6.4.
85 The diffuser performance through both laboratory measurements (section 6.2) and
86 deployments in Super-K (section 6.5) are also discussed in the context of optical and
87 temporal measurements.

88 **Inner and Outer Detector Differences**

89 The inner detector (ID) will include injector points on the barrel, directed horizontally
90 into the tank, and the end caps, directed vertically. To measure the difference in
91 optical properties as a function of depth, the proposed system will have seven
92 horizontal injection points at 90° intervals distributed evenly at different depths.
93 Eight vertical line injectors will be deployed, four at the bottom and four at the
94 top of the tank. Each injector point will consist of one diffuser and one collimator
95 system.

96 The geometric nature of the OD requires significantly more injection points
97 to achieve full photo-sensor coverage. Nevertheless, the precision diffusers designed
98 for the ID are not required to achieve the key objectives of the OD. Therefore a
99 simpler diffuser system¹ has been proposed, different to those described in section 6.3.

¹The exact design for the simpler diffuser is not yet known, likely candidates include using bare

100 Extrapolations from Super-K geometries predict that approximately 80 injection
101 points will be needed to cover the full OD PMTs. In addition to these injection
102 points, 12 collimators will be installed in the OD, helping to characterise the region
103 further. Placing collimators in parallel to the PMT wall yields longer travel distances
104 to increase the impact of absorption and scattering effects.

105 **6.2 Laboratory Experiments**

106 An in-house experimental setup at the University of Warwick is used for testing the
107 performance of diffusers during research and development phases. The experimental
108 set up is used to monitor the optical output and temporal signals through the
109 diffusers as a function of angle. This section will describe the experimental setup
110 used in the development of diffusers for the Super-K deployments.

111 **6.2.1 Experimental Setup**

112 The key motivation behind the experimental setup is to provide the light intensity
113 and timing profiles of each test diffuser over a distribution of angles. The setup
114 consists of a system which injects light through a 200 μm core step-index optical fibre
115 into the diffuser being tested. The diffuser is mounted on a 360° rotation stage, and
116 the signal is collected downstream by a photo-sensor. The whole system is enclosed
117 inside an interlocked dark box. A schematic of the experimental setup is shown in
118 figure 6.2, and a photograph in figure 6.3.

119 The light injection system is provided by a single pulsed semiconductor laser,
120 which can be set to emit a wavelength of 450 nm (blue) or 520 nm (green). The laser
121 diode is pulsed at 10 kHz, using the leading edge of a signal generator to trigger
122 an analogue pulser circuit. The light pulse is injected into a 200 μm core 0.5 NA
123 step-index fibre, approximately 1 m in length. The fibre optic cable terminates at
124 a ceramic ferrule which is placed inside the diffuser. Experiments have shown the
125 fibre position within the diffuser can cause differences to the amplitude of the optical
126 output from the diffuser. The magnitude of this effect has a dependence on the

fibres, or diffuse reflectors.

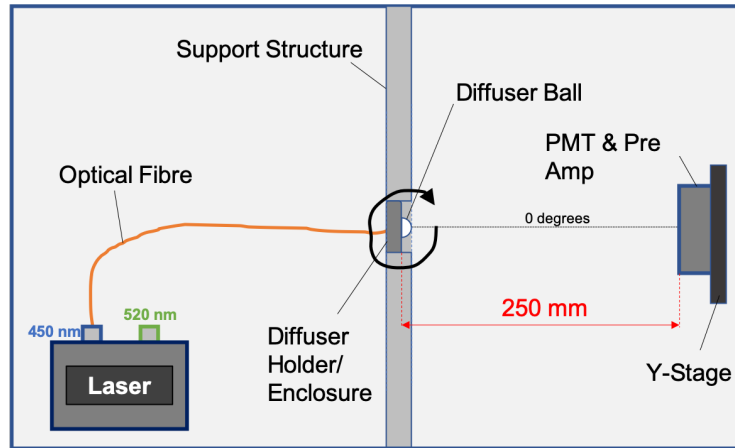


Figure 6.2: A schematic diagram demonstrating the experimental set up from a birds-eye view. *Maybe add coordinates?*

127 diffuser shape. For this reason the fibre is consistently placed touching the back of
 128 the hemispherical diffuser dome, and secured in place with a 3D printed clamp.

129 The diffuser is positioned on a rotation stage which is placed on a 'bridge-like'
 130 support structure. The rotation stage is accurate to ± 0.6 arcmin, and uses a stepping
 131 motor which can be accessed remotely via an Arduino-based control system [152, 153].
 132 The stepping motor is calibrated so that 1000 steps equates to a movement of 1
 133 degree. A metallic tube connected to a ring is attached vertically from the rotation
 134 stage and acts as a grip for either the diffuser, or diffuser enclosure. 3D printed
 135 alignment jigs ensure the diffuser is located at the centre of the grip, and is held
 136 in place using six screws triangulated over three different positions. This set up
 137 has the advantage of the diffuser being effectively suspended, allowing the full 360°
 138 horizontal plane to be scanned. A photo demonstrating the full assembly is provided
 139 in figure 6.3.

140 Downstream from the diffuser a Hamamatsu 6780-02 PMT is used to collect
 141 optical and temporal information for analysis. The PMT provides signal amplification
 142 that is tunable between $0\text{ V} \rightarrow 0.9\text{ V}$. The level of amplification follows a logarithmic
 143 relationship with respect to voltage input. Nevertheless, the PMT alone cannot
 144 provide enough signal amplification. Therefore to achieve a signal with an amplitude
 145 of greater than 1 V, a preamplifier is added inline after the PMT providing baseline
 146 amplification. The Ortec VT120 fast timing preamp is used as it can provide a

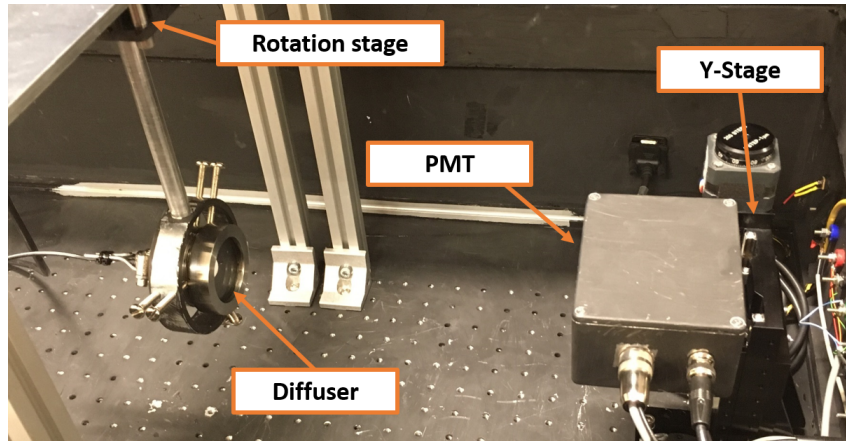


Figure 6.3: A photo showing the assembled experimental set up of the diffuser system (enclosure + diffuser ball) suspended in the grip, as well as the PMT box. Taken from [154].

147 sub-nanosecond rise time. The preamp and PMT are housed inside a box aptly
148 named "PMT box", which is situated 250 mm downstream of the diffuser's centre of
149 rotation. The PMT box is secured to a Y-stage which can be used for alignment, as
150 well as providing an additional dimension in which diffusers can be scanned. At this
151 moment in time, the Y-stage has only been used for alignment purposes, however
152 additional two-dimensional scans are planned for the future. A photo of the PMT
153 box and Y-stage can be seen in figure 6.3.

154 It is important to note that the absolute power of the pulsed laser diode
155 system, and environmental factors such as temperature, are not directly measured.
156 The power was not monitored due to the available meters incompatibility with a
157 pulsed light source; Given this, environmental factors were also not directly measured;
158 instead it was decided to only make comparisons against relative spectra. As a result,
159 quantitative conclusions of absolute light intensity measurements are not best to use.
160 Instead, relative intensity measurements, normalised to the intensity measured at
161 zero degree rotation, are used for comparative conclusions between different scans.
162 Measurements of laser stability in bare diffuser measurements as a function of time
163 have been made, and can be seen in figure 6.4. It can be seen that the laser stabilises
164 at approximately the hour mark; therefore a total of 70 minutes is given between
165 each scan to allow the laser power output to stabilise.

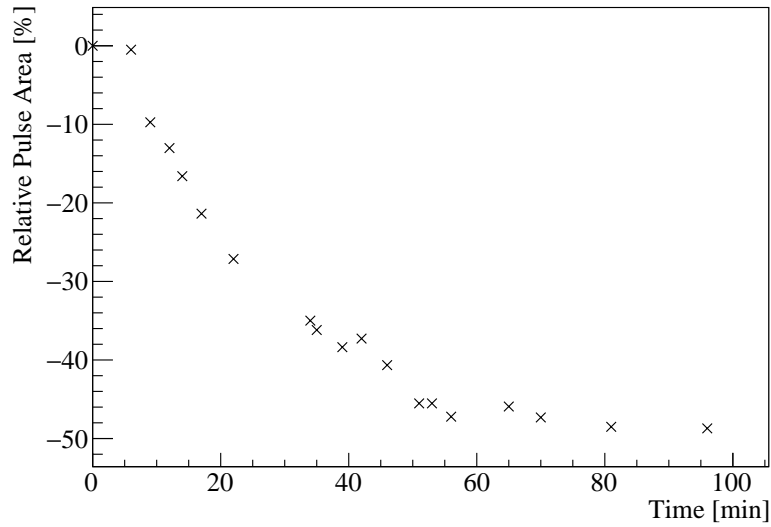


Figure 6.4: The laser power output stabilising as a function of time. The pulse area for a bare diffuser is measured in a zero degree on axis formation. Pulse areas are normalised to an initial time, $T = 0$, defined by the time the laser is switched on.

166 6.2.2 DAQ and Analysis Methods

167 Control of the experiment and DAQ systems are managed by LabView [155]. Scans
 168 are characterized a by series of angular sweeps defined by start and end angular
 169 positions, as well as step intervals. For example, one sweep may be defined from -60°
 170 to 60° at intervals of 2° steps; a scan can consist of multiple sweeps, typically taking
 171 a mean and RMS value over multiple sweeps for each angular position. Amplified
 172 signals from the PMT are fed through to an oscilloscope, where measurements can be
 173 taken. Two distinct methods of measurement are made: The calculated values taken
 174 by the scope itself, and the averaged waveform are both recorded. Both techniques
 175 take the average over 1000 pulses. At each position along a scan, the oscilloscope
 176 is reset and a delay of 10 seconds is applied to allow the scope to settle. This was
 177 done as it was found the sampling speed of the scope was low enough that, without
 178 delay, measurements from the previous position were included in the average for
 179 the current. An example pulse from a bare diffuser ball can be seen in figure 6.5.
 180 Relative light intensity distribution comparisons are taken using the height and area
 181 of the pulse. Timing information can be collected by measuring the pulse full-width
 182 half-maximum, and by measuring the pulse delay - the time between the laser trigger

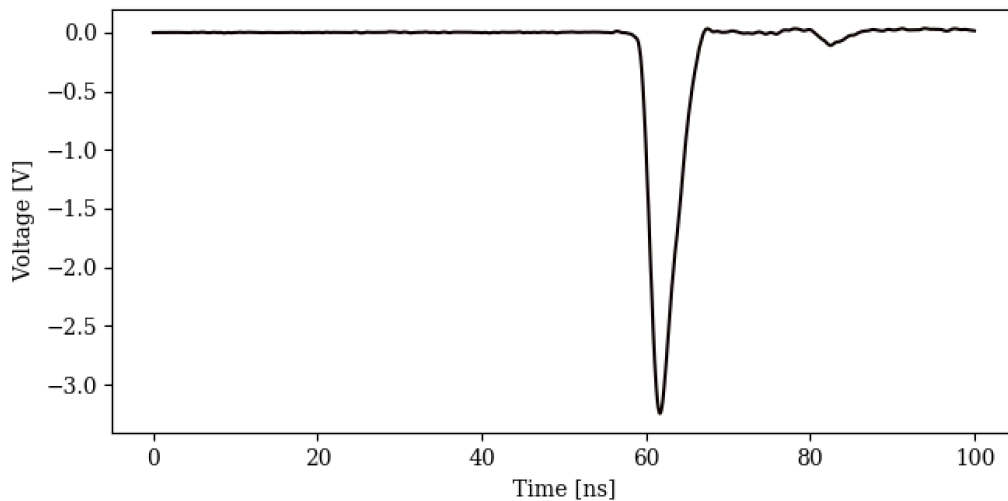


Figure 6.5: An example pulse from a scan of a bare PMMA diffuser.

183 and the rising edge of the pulse, defined by 10% of the peak voltage of the pulse.

184 Diffuser analysis is performed using a python analysis package developed
 185 in-house. The package has the versatility in analysing light intensity and temporal
 186 measurements as a function of angle and time. The majority of results seen within
 187 this chapter, unless otherwise stated, were created using this package.

188 6.3 Diffusers

189 The optical calibration work presented in this thesis will focus primarily on diffuser
 190 based technology. The primary aim of the optical diffusers is to provide a means
 191 of obtaining energy and timing information for calibrating photo-sensors in large
 192 scale projects such as Hyper-Kamiokande. This section and the next will discuss
 193 the research and development of optical diffuser technology and outline the diffuser
 194 calibration systems developed at the University of Warwick.

195 As outlined in section 6.1 the goals for the wide-beam diffuser are to create
 196 a well understood, preferably uniform, distribution over an angular range of $\pm 40^\circ$,
 197 as well as having a uniform timing profile as a function of angle. In addition
 198 the development objectives include practical implications. The diffuser must be
 199 waterproof, passing soak tests, as to not affect either the diffusing optical properties,
 200 or the water it's situated in. Moreover, the diffuser must remain stable in gadolinium

201 doped water. The diffuser must be able to withstand pressures of up to 10 Bar, as
202 well as have a longevity on the order of multiple decades. Furthermore, the diffusing
203 properties must not be affected, as a function of time, by any of the previously
204 described criteria.

205 **6.3.1 Diffusing Material**

206 In 2017, studies were undertaken to find a diffusing material that could achieve 1%
207 uniform light intensity over a $\pm 45^\circ$ light cone ² [156]. The light injection system
208 requires such a material to have good transmission (near 100%) over a range of
209 ultraviolet to visible wavelengths, in particular 350 nm to 550 nm.

210 Generation one diffusers were made out of 50 μm glass beads suspended in
211 a polyester resin. As light enters the resin photons can scatter multiple times off
212 the glass beads due to their relatively small size. By the time photons exit the resin,
213 they are emitted in a diffuse distribution. The light output over the required range
214 was found to be uniform to 4%. However, spectroscopy measurements found that
215 polyester resin absorbs light below 400 nm. Moreover, it was found the addition of
216 glass beads reduced relative intensity by approximately 20% between 500 nm and
217 600 nm wavelengths.

218 For generation two a number of candidate materials were considered. Poly-
219 styrene and poly(methyl methacrylate), PMMA, were two of the options investigated
220 for their optical transmission properties. The relative transmission are between
221 90% - 100% above 400 nm, with a gradual drop to approximately 75% (20%) and
222 55% (0%) at 350 nm (300 nm) for PMMA (polystyrene) respectively. This is shown
223 in figure 6.6. Additionally the 48 μm particle size of PMMA ensures it has good
224 diffusion properties, eliminating the need to add glass beads. For this reason PMMA
225 was chosen as the diffusing material for the Super-K deployed optical calibration
226 system.

227 Soak tests were performed on PMMA samples using both ultra pure water
228 currently used in Super-K, and gadolinium loaded water to ensure future compatibility
229 with the SK-Gd and Hyper-K projects. Traces of PMMA impurities were found to

²Note that the diffusing criteria have since changed

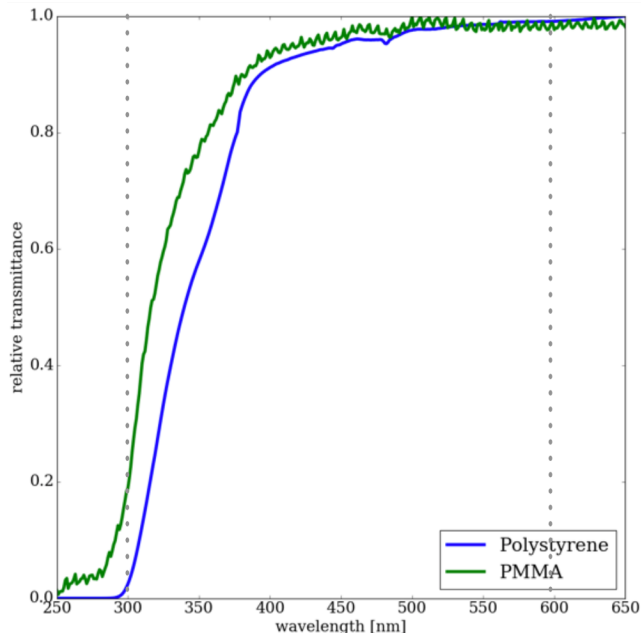


Figure 6.6: Relative transmission properties of PMMA and polystyrene. Taken from [156].

230 have leached into the water, and optical properties of the diffuser ball were adversely
 231 affected [156]. Figure 6.7 shows UV-VIS spectrum analysis of the different water
 232 samples exposed to PMMA. An idea material which does not affect water properties
 233 would follow the absorption and transmittance spectra of the control samples,
 234 indicating an absence of contamination. It can be seen that below 300 nm, the
 235 water absorption and transmittance properties rapidly degrade due to contamination
 236 from the suspended PMMA particles. To mitigate this, a water-tight enclosure was
 237 designed, and is outlined in detail in section 6.4.

238 6.3.2 Diffuser Shape

239 Alongside material, research and development of the diffuser shape was undertaken.
 240 The work primarily focussed on two designs, spherical and hemispherical. In theory
 241 with a perfect diffusing material, the spherical shape would emit a uniform distribution
 242 over 4π sr, whereas the hemispherical shape has geometrical limitations. However in
 243 practice, the hemispherical design is simpler to manufacture.

244 Light intensity distributions for PMMA spherical and hemispherical diffusers

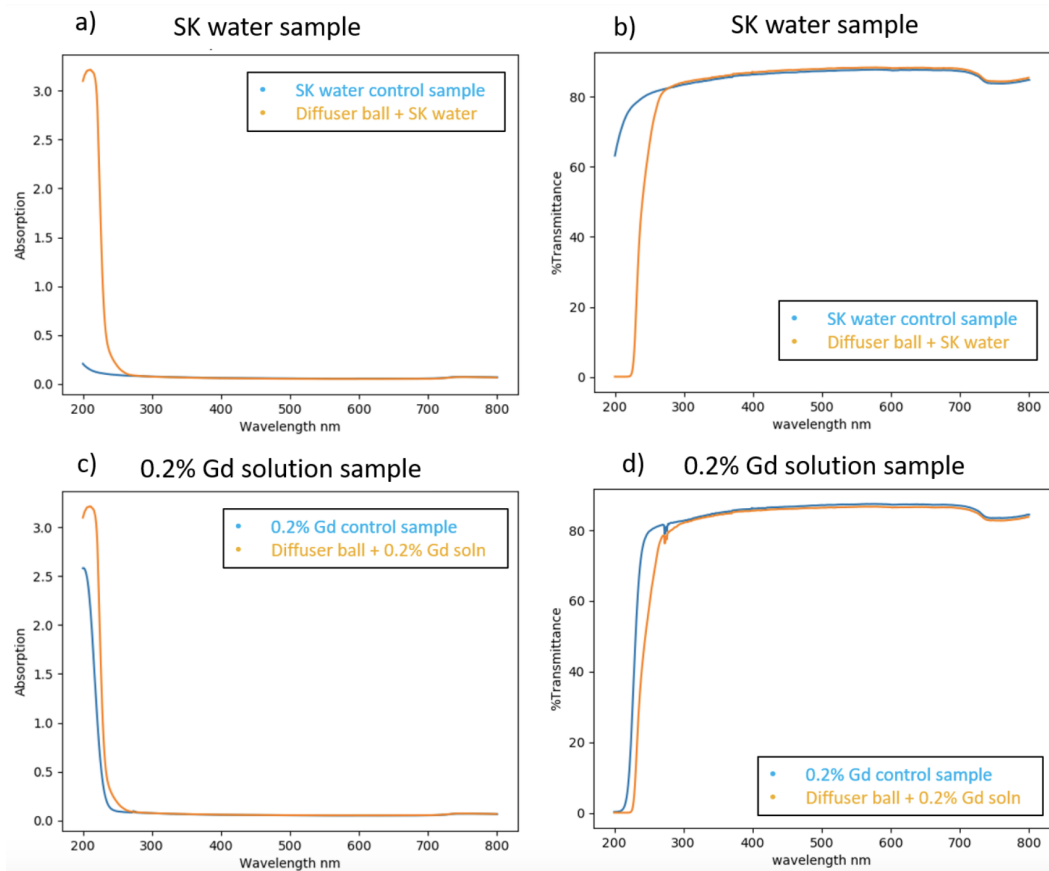


Figure 6.7: Soak test results for the optical absorption(left) and transmission (right) properties of the water over the UV-VIS spectrum, for different water samples. A Perkin Elmer Lambda 850 UV/VIS spectrometer was used. Each sample was measured in 10 mm path length disposable cuvettes and referenced against clean water.

245 measured over a range of $\pm 90^\circ$ is shown in figure 6.8a. A significant problem with
 246 full spherical diffusers was the amount of backscattered light, which had a strong
 247 dependence on fibre position with respect to the centre of the diffuser (figure 6.8b).
 248 Therefore, coupled with manufacturing considerations, a hemispherical diffuser design
 249 was chosen.

250 The performance of hemispherical diffusers was measured with respect to its
 251 dimensions. Diffusers were manufactured at 5 mm, 10 mm, 15 mm radii. The light
 252 output was found to broaden with larger diffusers, and the pulse width increased
 253 by 1.5 ns and 0.5 ns between 5 \rightarrow 10 mm and 10 \rightarrow 15 mm respectively. The final
 254 design featured a hemispherical PMMA diffuser of radius 20 mm with a 10 mm

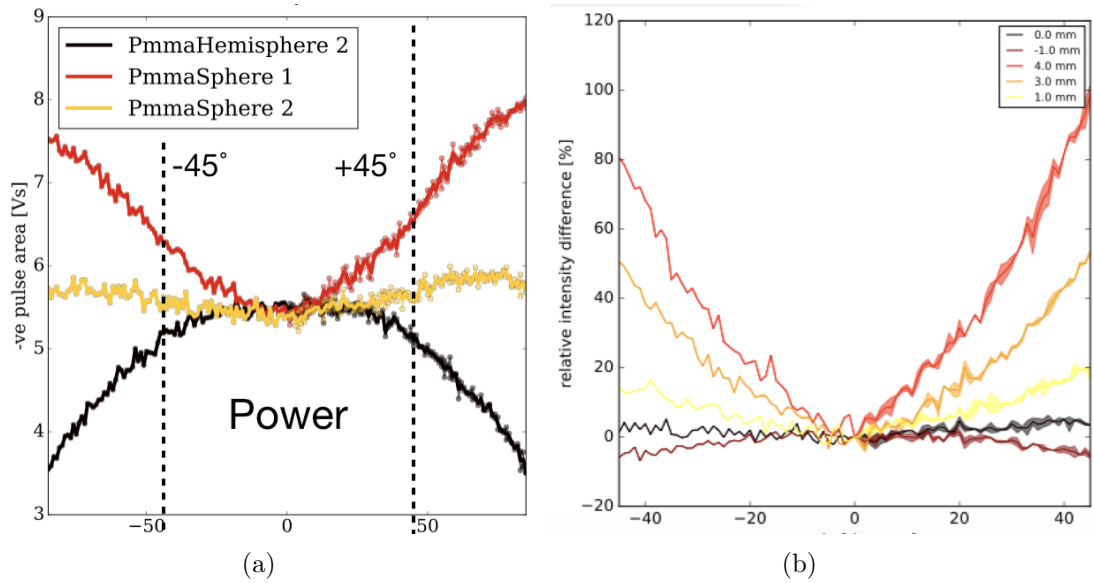


Figure 6.8: (a) The light intensity distributions as a function of angle for spherical and hemispherical diffusers. (b) A plot demonstrating the relative light intensity for various different distances between the fibre and diffuser centre. Both plots have angle in degrees on the x-axis. Taken from [156].

255 thick cylindrical back of radius 30 mm for support. The final design can be seen in
 256 figure 6.9.

257 6.3.3 Manufacturing

258 The manufacturing process of PMMA hemispherical diffusers has a number of steps
 259 that are outlined within this section. PMMA is known to be porous, therefore
 260 all tooling machines has to be scrupulously cleaned of all contaminants as to not

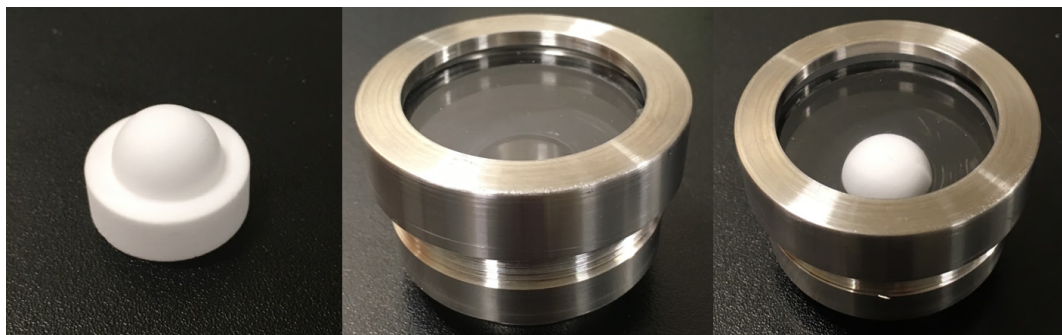


Figure 6.9: Photos showing an example of the diffuser (left), enclosure (middle), and diffuser inside enclosure (right). Taken from [154].

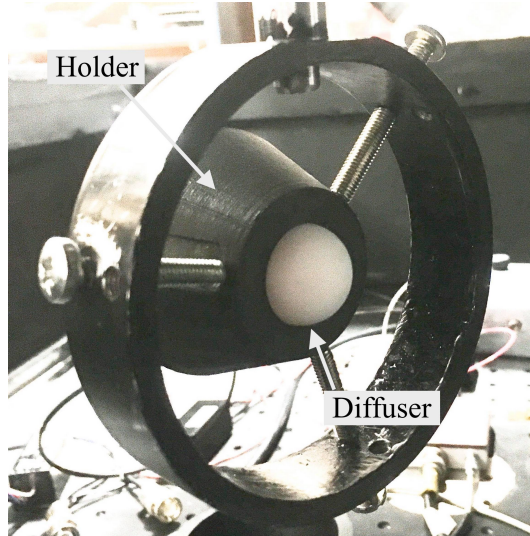


Figure 6.10: A photo of the bare diffuser experimental set up with the diffuser inside the 3D printed holder.

261 risk affecting optical properties. The PMMA is purchased in the form of a powder.
262 A vibration mill is first used to prepare an amalgam, removing any potential air
263 voids and compacting the powder down. The PMMA is then placed into a die and
264 compressed in a hand operated vice at approximately 2 ton of pressure, monitored
265 using a strain gauge. The die is then put into a sash clamp and heated in an oven
266 at 175°C for 3 hours. A PMMA sintered rod is then extracted from the die, and
267 machined into hemispherical diffusers. Each sintered rod of length 100 mm has the
268 capability of producing two to three diffusers.

269 **6.3.4 Bare Diffuser Performance**

270 To measure the optical and temporal properties of bare diffusers, in the experimental
271 set up described in section 6.2, a holder was designed to allow illuminations only
272 from the hemisphere to reach the PMT. An example of a diffuser inside a 3D printed
273 holder is shown in figure 6.10.

274 The light profile from the PMMA diffusers was scanned from $-40^\circ \rightarrow 40^\circ$
275 measuring the area of the light pulse in [Vs] at each step. When comparing the
276 reproducibility over multiple diffuser samples it is more comparable to normalise
277 with respect 0° and plot the relative difference, this is shown in figure 6.11, over

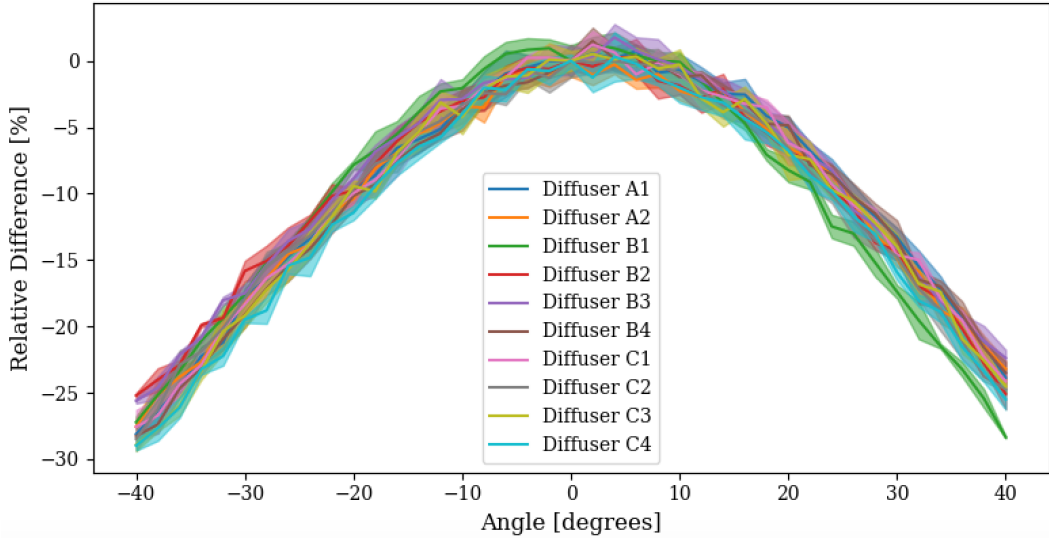


Figure 6.11: The bare diffuser light intensity profile, normalised at 0° , for 10 different diffusers demonstrating a test in reproducibility. The same letter indicates the same diffuser batch. Diffuser pairs 1 & 2 and 3 & 4 are made from the same rod.

278 a sample of 10 bare PMMA diffusers. It can be seen, with the exception of B1
 279 which is believed to have had a minor alignment anomaly, all of the bare diffusers
 280 across the three batches and five rods agree within error. The uncertainty is derived
 281 as the RMS over repeat measurements. It is clear from figure 6.11 that the bare
 282 PMMA diffusers produce a consistent light intensity distribution over a range of
 283 $\pm 40^\circ$. The light profile peaks at 0° and falls linearly with angle from approximately
 284 $\pm 10^\circ$ onwards. It can be noted that figure 6.11 appears to be weakly biased towards
 285 positive angles, which suggests either asymmetry in the diffuser or misalignment in
 286 the set up. Diffuser symmetry was tested by rotating the diffuser through 360° taking
 287 scans at regular intervals. From this test the peak intensity did not change with
 288 respect to diffuser rotation, which suggests the positive bias is due to a systematic
 289 misalignment of approximately 2° .

290 As the diffuser rotates by an angle θ , the total surface area of the diffuser visible
 291 to the PMT changes due to the hemispherical geometry. This can be mathematically
 292 described as:

$$2\pi r^2 - 2\theta r^2 \quad (6.1)$$

293 where r is the radius of the diffuser and θ is an angle of rotation between the limits

294 of $-\frac{\pi}{2}$ and $\frac{\pi}{2}$. The $2\pi r^2$ term represents the surface area of a hemisphere excluding
 295 the base, and $2\theta r^2$ represents the surface area of a wedge from a sphere. The total
 296 surface area visible, relative to the surface area at $\theta = 0$ can then be written as:

$$\frac{2\pi r^2 - 2\theta r^2}{2\pi r^2} \quad (6.2)$$

297 which simplifies to

$$1 - \frac{\theta}{\pi} \quad (6.3)$$

298 thus defining a correction factor for the hemispherical geometry in the context of
 299 this experiment. The intensity profile can then be compared to that of a perfect
 300 hemispherical diffuser by applying this hemisphere correction factor:

$$x' = \frac{x}{1 - (\theta/\pi)} \quad (6.4)$$

301 where x is the variable you wish to correct for (usually light intensity), and θ is
 302 the absolute angle in radians. This derivation assumes perfect alignment, and that
 303 the PMT is at a sufficient distance away as to include the full diffuser inside the
 304 field of view. A plot of the relative light intensity corrected for a hemispherical
 305 geometry, using equation 6.4, can be seen in figure 6.12. Asymmetry effects through
 306 the systematic misalignment are inevitably enhanced through geometry corrections.
 307 However, figure 6.12 indicates that forward going light between approximately $\pm 10^\circ$
 308 is suppressed with more scattered light promoted between $\pm 10^\circ \rightarrow \pm 20^\circ$ region.
 309 Furthermore, accounting for geometric effects, the bare PMMA diffuser is uniform
 310 to 10% over an angular range of $\pm 40^\circ$.

311 The temporal performance of the bare diffuser is measured primarily in the
 312 delay of the signal. The pulse delay is the time between the laser pulse trigger and
 313 the rising edge³ of the pulse arriving at the PMT (figure 6.5). The relative signal
 314 delay as a function of angle is shown in figure 6.13. The signal delay is uniform to
 315 approximately 1% over an angular range of $\pm 40^\circ$. The absolute time delay ranges
 316 from 48.7 ns to 49.3 ns giving a spread of around 1.2% across all manufactured

³This is defined by a threshold of 10% of the peak signal voltage

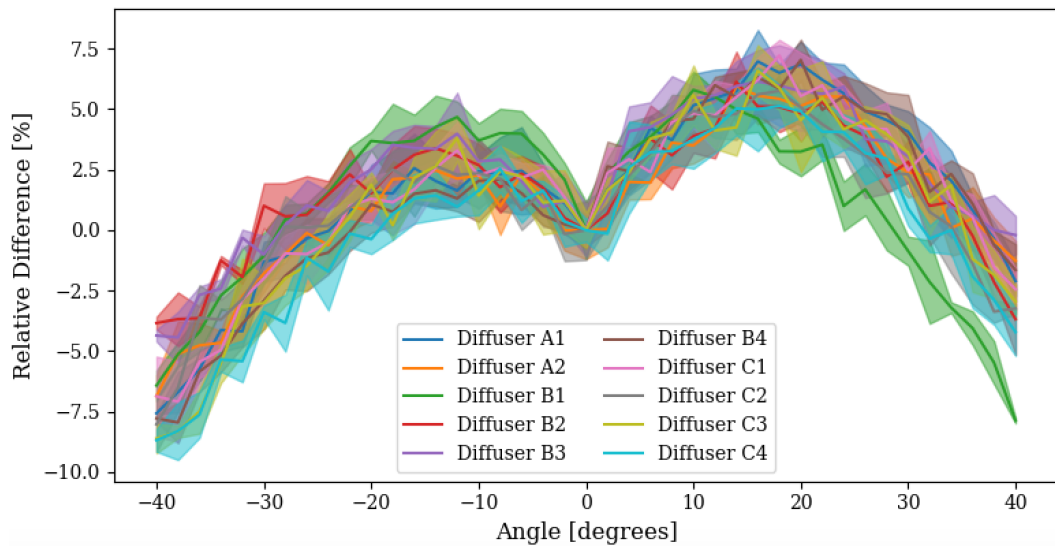


Figure 6.12: The relative light intensity profile of PMMA bare diffusers when corrected for hemispherical geometry effects through equation 6.4.

317 PMMA diffusers. The pulse width was also measured and is shown in figure B.1.
 318 However, a relatively large systematic uncertainty in PMT response means signal
 319 delay information provides a more reliable source of diffuser temporal performance.

320 6.4 Diffuser Enclosures

321 As discussed in section 6.3, soak tests had proven that PMMA diffusers were not
 322 waterproof, and exposure to water changed the light output properties of such
 323 diffusers. It was therefore decided to house the diffusers inside enclosures to protect
 324 both the diffuser and water environment from contamination. Such an enclosure
 325 would need to be watertight to 10 Bar, ideally easy to manufacture, and have no
 326 effect, or positive effect, on achieving the diffuser goals outlined in section 6.1. This
 327 subsection will discuss the enclosure research and development, and introduce results
 328 of the diffuser and enclosure system as a whole.

329 6.4.1 Base Enclosure Design

330 A number of key concepts were drawn up and gave rise to what will be referred to
 331 within this thesis as a base enclosure design. The concepts included fibre injection,

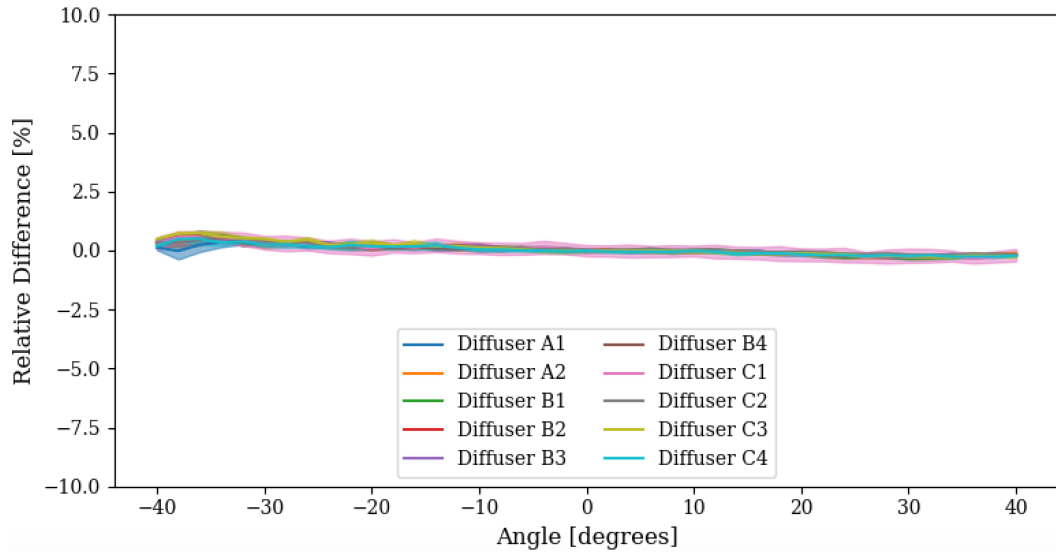


Figure 6.13: The relative signal time delay as a function of angle, normalised to 0° . The same ten PMMA diffusers from figure 6.11 are used.

332 materials, and water proofing measures. The base design consists of a stainless steel
 333 enclosure, with a glass window on the front, and a hole for fibre injection at the back.
 334 Water-tightness is achieved through a mixture of o-rings and epoxy resin.

335 6.4.2 Materials

336 The main body of the diffuser enclosure during the early stages of development
 337 was made out of 304 stainless steel. This was changed to 316 stainless steel for
 338 the deployment in Super-Kamiokande (section 6.5) because of its better chemical
 339 resistant properties. Two different glass materials, sapphire and Schott, were tested
 340 for the window and were found to have little to no difference in optical transmission
 341 and profile properties. A 6 mm thick Schott glass with a 50 mm diameter was chosen
 342 for the window. Water-tightness was ensured using a combination of Viton o-ring
 343 gaskets, as well as a water and chemical resistant epoxy resin. All materials were
 344 subject to soak tests in ultra pure and gadolinium loaded water; each material also
 345 underwent pressure tests up to 10 bar pressure for at least 12 hours underwater.

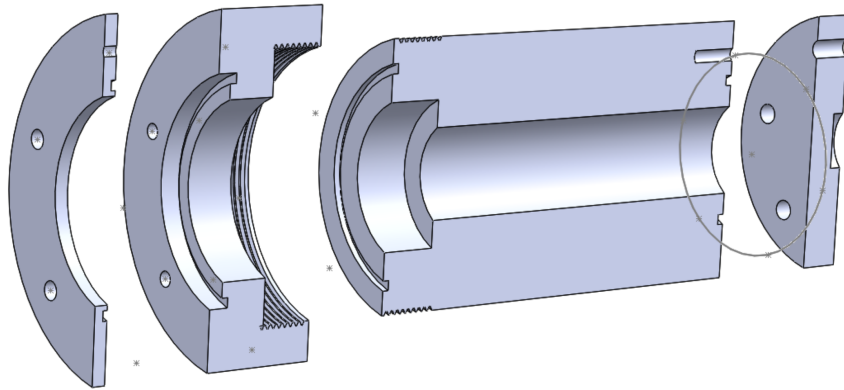


Figure 6.14: A cross-section of the version 1 (V1) diffuser enclosure design.

346 **6.4.3 Enclosure Development**

347 Major developments of the diffuser enclosure revolved around two key base designs,
 348 known as version 1 (V1) and version 3 (V3). Intermediate phases between these are
 349 referred to as version 2 (V2) designs. However, since these acted only as conceptual
 350 stepping stones between V1 and V3, and no measurements were ever taken with
 351 these designs, they will not be discussed in detail within this thesis. Schematic CAD
 352 drawings of concept V2 designs are provided in the appendix (figure B.2).

353 **V1 Design**

354 A cross section for the initial enclosure design, labelled V1, is shown in figure 6.14.
 355 The design consists of a 75 mm long main body, with a 27 mm front screw cap, and
 356 3 mm window cap which holds the window in place whilst screwing onto the main
 357 body. A 10 mm thick solid disc with a hole for fibre feed-through makes up the
 358 back end piece, known as the base lid, which is attached via 6 threaded screws in
 359 hexagonal formation. Moreover, the design also allows the enclosure to be directly
 360 attached to a pressure vessel in place of the base lid. Fully assembled, the V1 design
 361 measures 100 mm in length, with a 60 mm main body diameter, which rises to
 362 75 mm at the front. The diffuser base sits 30 mm away from the front end of the
 363 enclosure. The origin of diffusion, defined by the point at which the light is injected
 364 from the fibre into the diffuser, is located 17 mm from the front of the glass window.
 365 The diffuser shoulder is also exposed, so to obtain hemispherical diffuser results

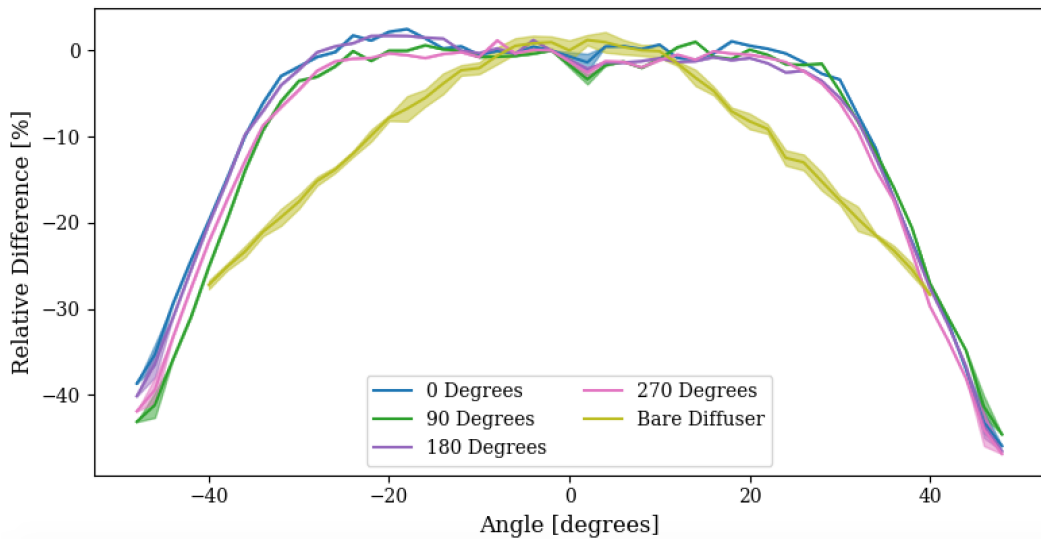


Figure 6.15: The optical profile using the V1 diffuser enclosure system. The full system was rotated through 360° over 90° steps, at one sweep for each scan. An example bare diffuser profile is also shown for comparison.

366 improvised shielding from black electrical tape was often used. O-ring gaskets at
 367 the front and back of the enclosure provide water-tightness. The V1 design was too
 368 large and heavy to match any practical considerations for deployments into Super-K
 369 or Hyper-K, but did allow for important preliminary enclosure studies to be made.

370 **V1 Optical Performance**

371 The light profile from a 20 mm hemispherical diffuser installed inside enclosure V1
 372 shown in figure 6.15. The diffuser and enclosure were rotated as a full system through
 373 360° to compare the optical profiles over two different planes. The measurements
 374 suggested the V1 enclosure promoted high angle photons, within it's respective field
 375 of view, relative to the bare diffuser profile in figure 6.11. The resultant profile is
 376 flat within approximately 2% over an angular range of approximately $\pm 25^\circ$ to $\pm 30^\circ$.
 377 The light intensity drops linearly as the enclosure gradually eclipses the diffuser
 378 towards higher angles. The change in profile is largely because of two factors. The
 379 first is that the often shoulder of the diffuser is not covered in V1, meaning more
 380 than the intended hemispherical diffuser is visible. The second is because of specular
 381 reflections from the enclosure which promotes otherwise lost light at higher angles.

382 It was standard practice in the early stages of research and development to

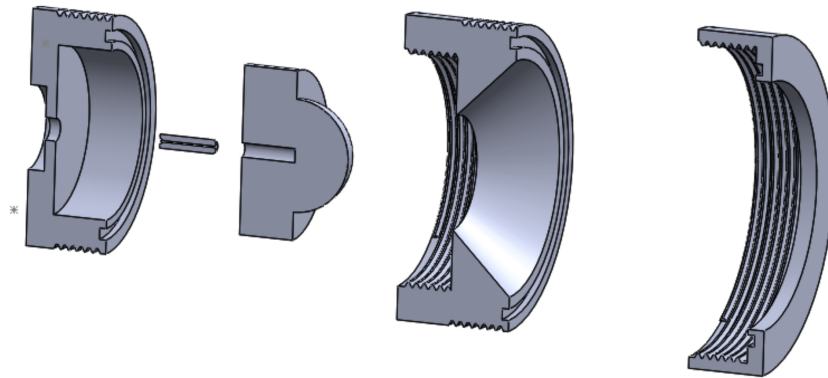


Figure 6.16: A cross-section schematic of the V3 diffuser and enclosure with a torch design.

383 perform measurements for only one sweep at a time. Combined with an uncertainty
 384 in unmonitored laser power between scans, an error in repeated measurements is not
 385 able to be taken. Therefore the results shown in figure 6.15, must be taken with a
 386 relatively large qualitative uncertainty, and were used as a guide for developmental
 387 paths only.

388 **V3 Design**

389 The premise for version 3 was to design an enclosure fit for the Super-Kamiokande
 390 deployments outlined in section 6.5. Figure 6.16 shows a schematic drawing of the
 391 cross-section of both the diffuser and the enclosure V3. The V3 design consists of
 392 three segmented parts which screw together, named the base, main body, and end
 393 cap. The base segment, 15 mm long with a diameter of 42 mm, holds the diffuser
 394 and includes a hole for fibre injection. The role of the main body is to provide
 395 shielding against the diffuser shoulder as well as facilitate light output objectives
 396 using internal enclosure reflections. During development phases there were two
 397 designs for the main body: The first, shown in figure 6.16, was a "torch" with a
 398 96° field of view designed to promote forward going light using diffuse reflections of
 399 high angle photons. The second was a "bucket" design with a flat face to the outer
 400 edge and then a perpendicular wall. Both face designs are painted matte black in
 401 order to remove specular reflections off the stainless steel surface, encouraging any
 402 remaining reflections to be diffuse. The main body is 21 mm long with a diameter of

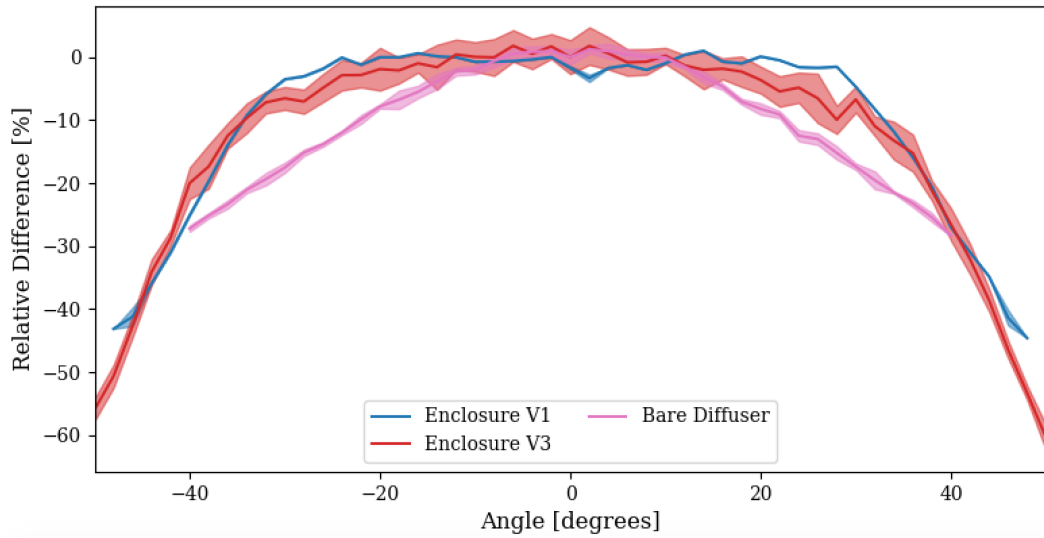


Figure 6.17: The relative optical profile of the V3 enclosure, with comparative profiles for the V1 enclosure and bare diffuser also shown. The optical profiles are normalised to 0°.

403 54 mm. As with V1, the front end cap holds the glass window; the cap is 14 mm
 404 in length and has a diameter of 60 mm. When fully assembled the enclosure is
 405 approximately 35 mm in length and is designed to house the "final design" diffusers
 406 outlined in section 6.3.2. Pictures of the fully assembled V3 enclosure design are
 407 shown in figure 6.9.

408 **V3 Optical Performance**

409 The optical profile for the V3 enclosure, with visual comparisons to the V1 enclosure
 410 and bare diffuser, is shown in figure 6.17. The profile is flat to within 10% over
 411 an angular range of approximately $\pm 35^\circ$ to $\pm 40^\circ$. Despite the loss of the diffuser
 412 shoulder, the diffuser reflections from the matte black torch design promote angle
 413 photons in the forward direction to help flatten the distribution. The field of view
 414 also increased between V1 and V3 by an estimated $5^\circ - 10^\circ$ thanks to the shallower
 415 end cap design. The enclosure V3 designs optical profiles had good reproducibility
 416 which was seen for the Super-K deployment in figure 6.27.

417 Temporal performance of example V1 and V3 enclosures, and the bare diffuser
 418 are shown in figure 6.18 and figure 6.19. The pulse signal delay for enclosure V3

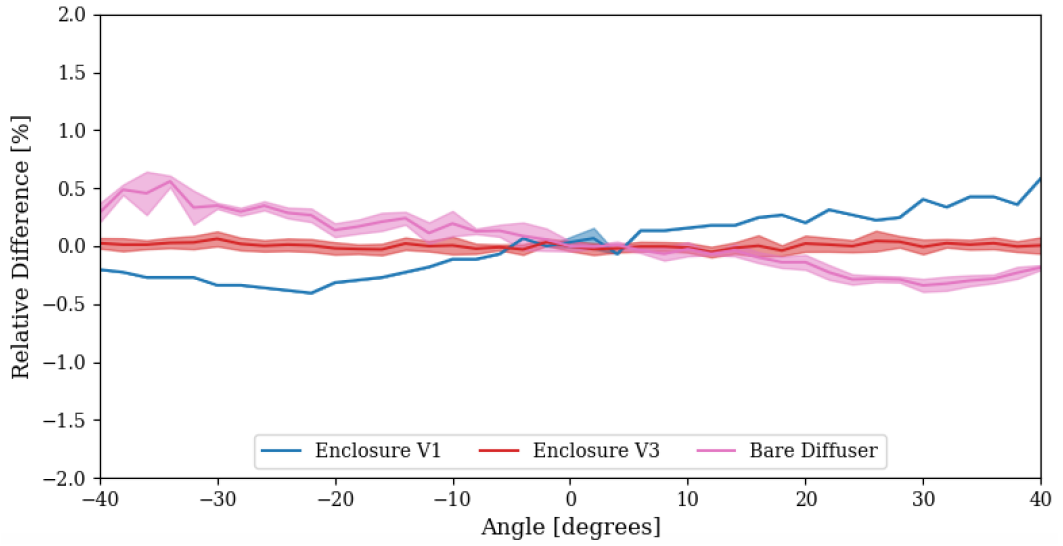


Figure 6.18: The relative signal delay, normalised to 0°, as a function of angle. Measurements for the V1 and V3 enclosures, and the bare diffuser are shown.

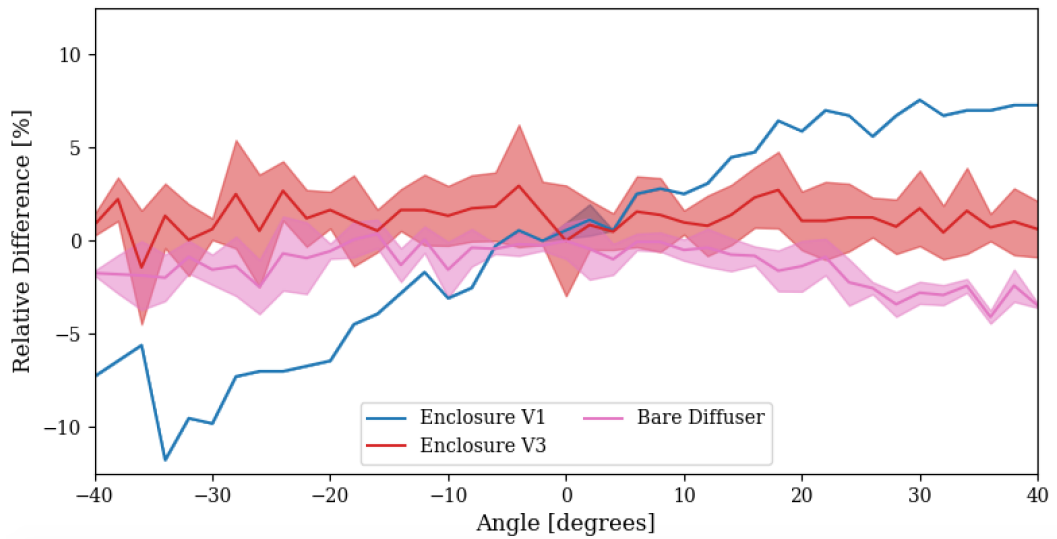


Figure 6.19: The relative pulse width as a function of angle for the V1 and V3 enclosures, as well as the bare diffuser. Each plot is normalised to 0°.

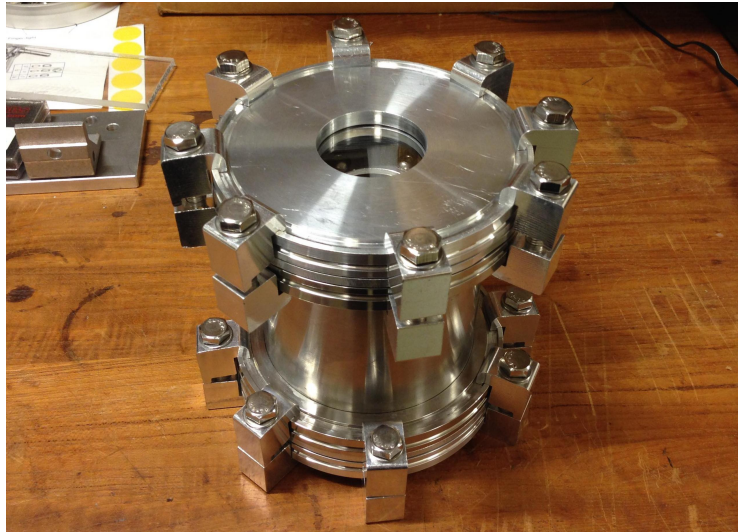


Figure 6.20: Pressure vessel used for diffuser and enclosure pressure tests at the University of Warwick.

419 is uniform as a function of angle. Moreover there is circumstantial evidence that
420 V3 performs better than V1 or even the bare diffuser, however the likelihood is
421 the sloped features for V1 and the bare diffuser in the signal delay is from minor
422 misalignments in the experimental set up. Absolute measurements indicate pulse
423 delay of 44.5 ns for V1 enclosure, (49.0 ± 0.3) ns for the bare diffuser⁴, and 52.5 ns
424 for enclosure V3. The increased delay from bare diffuser to inside the V3 enclosure
425 could be attributed to addition of propagating through a glass window. The shorter
426 delay seen in the V1 enclosure is likely due to changes in the experimental setup
427 between measurements. The pulse width, measured by the full-width half-maximum,
428 also seems to favour the V3 enclosure in terms of uniformity over angular space. The
429 V1 enclosure exhibits unusual behaviour for the pulse width that could be attributed
430 to asymmetries in specular reflections, however this is not known. In summary, the
431 V3 enclosure has good temporal performance, yielding uniform distributions in both
432 signal delay and pulse width over and angular range of $\pm 40^\circ$.

433 6.4.4 Pressure Testing

434 Water pressure tests are performed using a vessel shown in figure 6.20. The vessel is
435 made out of stainless steel and consists of a main container with two lids. The front
436 lid contains a glass window from which observations can be made during tests. The
437 back lid contains a feed-through injection point for the fibre optic cable. Sixteen
438 clamps, eight top and eight bottom, are tightened using a torque wrench to squash
439 two large o-rings gaskets and ensure a water-tight seal up to and beyond 5 bar of
440 pressure.

441 Originally the V1 enclosure main body attached directly to the back plate of
442 the pressure vessel. This had the effect of exposing the inside of the enclosure to
443 external atmospheric pressures through the fibre feed-through. Furthermore, because
444 the main body was directly attached to the vessel, the full V1 enclosure was never
445 pressure tested in its fully assembled state. Nevertheless, V1 enclosure pressure
446 tests provided performance validation in the materials, gaskets, and front assembly
447 mechanisms.

448 Failures in the ability to test fully assembled V1 enclosures, set about changes
449 in the pressure testing methodology for future enclosures. Instead of being attached
450 directly, the V3 enclosure would instead be fully submersed in the water. This
451 brought with it logistical challenges in sealing the fibre feed-through during pressure
452 tests. The initial solution was to use a silicone gel to plug the gaps between the fibre
453 and vessel, a system which worked with relative success throughout the Super-K
454 deployment phase. More recently, a high-pressure fibre feed-through has been used
455 to couple the fibre into the vessel. This was not, however, implemented until after
456 the Super-K deployments.

457 The definition of successful and failed pressure tests comes down to a number
458 of factors. Firstly, quantitative measures of the pressure inside the vessel are made
459 as a function of time through the test. Consistent drops in pressure can signify
460 a vessel failure, whereas a small singular drop may indicate an enclosure failure.
461 Secondly, qualitative observations of the enclosure and diffuser are made once they
462 have been removed from the vessel. Evidence of water ingress into the enclosure,

⁴Where the error is taken from repeat measurements in figure 6.13

463 condensation, and diffuser damage are looked for. Finally the optical profiles are
464 retaken and compared against relevant profiles pre-tests. It is known that water can
465 affect the optical properties of PMMA, and thus changes in the optical profile may
466 indicate water contaminations and thus a failed pressure test.

467 The V3 enclosure waterproofing consisted of two systems: internal o-rings,
468 and external epoxy resin. The internal gasket system had already been proven to
469 work during V1 enclosure testing, but did not provide any waterproofing through
470 the fibre-enclosure coupling. A chemical resistant epoxy resin was liberally applied
471 both at the fibre injection point at the base of the enclosure, and at the screw
472 connection points between each enclosure segment to allow the epoxy to seal any
473 gaps via capillary action. At this stage a number of enclosures failed pressure
474 tests. The caused was deemed to be poor bonding between the epoxy resin and the
475 fibre furcation tubing ⁵, which under stress formed minute cracks and gaps which
476 water could penetrate under pressure. Strain relief measures were manufactured and
477 added to the V3 design to counteract stress on the convex fibre-epoxy bond. These
478 included an epoxy filled stainless steel "top-hat" and a lateral support attached to
479 the fibre with polypropylene tie wraps. The strain relief measures can be seen in
480 figure 6.21. Once strain relief measures had been put in place, the V3 enclosure
481 designs successfully passed all pressure testing criteria outlined previously. The
482 optical profiles of an example diffuser enclosure before and after successful pressure
483 tests are shown in figure 6.22. It can be seen that the profiles are unchanged over
484 the pressure tests within the RMS error from repeated measurements.

485 **6.4.5 Condensation Testing**

486 The water inside Super-Kamiokande has an ambient temperature of approximately
487 13°C. A relatively cold temperature, a potential concern was the build up of con-
488 densation inside the diffuser enclosure. To address this concern an assembled diffuser
489 was placed in a cold box at 5°C and then 0°C for three and two consecutive days
490 respectively. No visual condensation was found over the 5 days of testing. The dew

⁵The furcation tubing material was predominantly PVC, however the exact makeup was not disclosed by the manufacturer.

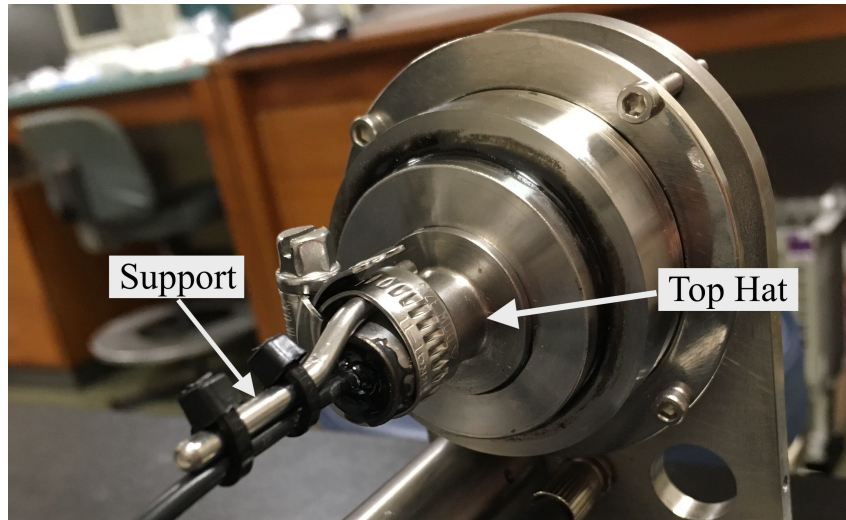


Figure 6.21: A photo demonstrating the strain relief measures used to protect the epoxy resin bonding to fibre furcation tubing in the Super-K deployed V3 enclosures.

491 point changes as a function of both ambient temperature and humidity. To achieve
492 a dew point of below 10°C , the laboratory temperature and humidity during the
493 diffuser enclosure assembly needed to be considered. Efforts were made to lower
494 the environmental humidity during diffuser enclosure assembly, in an attempt to
495 suppress the internal dew point to below the 10°C threshold.

496 **6.5 Super-Kamiokande Deployment**

497 As the Hyper-Kamiokande experiment is under construction, the proposed optical
498 calibration light injection system was deployed into the sister experiment Super-
499 Kamiokande. The objective was to both test the calibration optics as well as provide
500 direct physics measurements for the Super-K experiment. This section will outline
501 both the temporary winter 2018 test and more permanent summer 2018 deployments,
502 as well as provide a preliminary discussion on the diffuser performance inside Super-K.

503 **6.5.1 Test Deployment**

504 In January 2018 a test deployment was scheduled for the light injection system.
505 The purpose was to trial the proposed light injection system and gain experience in
506 preparation for the summer deployment. The optical calibration devices, consisting of

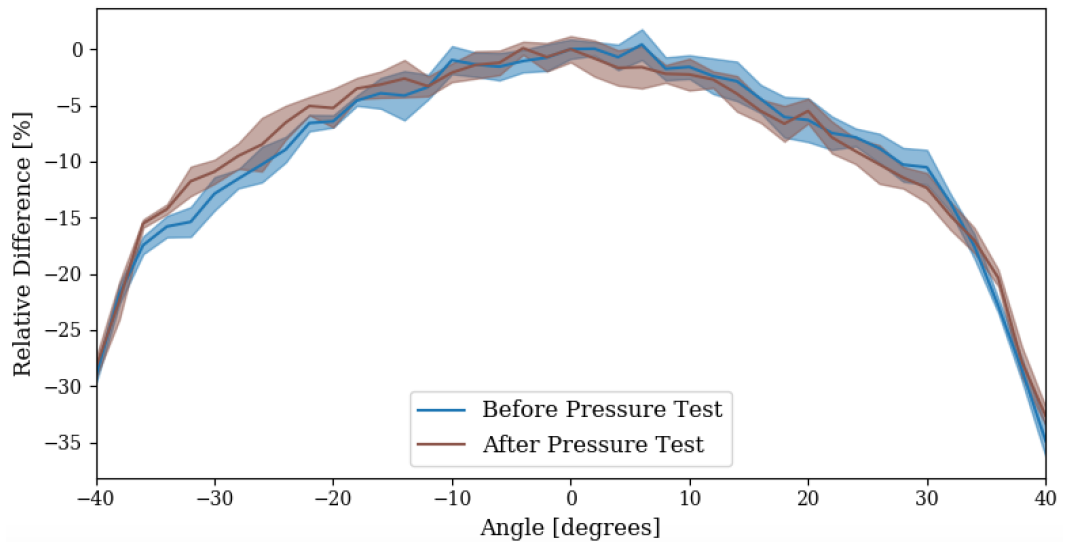


Figure 6.22: The optical profiles of a diffuser inside enclosure D1 before and after pressure tests. Each optical profile is normalised to the pulse area at zero degrees. The solid line shows the mean, the shaded part is the RMS, over repeat measurements.

507 a diffuser, collimator, and bare fibre, were attached to a mounting plate and lowered
 508 into the top of the Super-K tank via a calibration port. A simplified representation
 509 of the optical equipment installation methodology during the test deployment is
 510 shown in figure 6.23a.

511 A 20 mm PMMA hemispherical diffuser was used inside a V3 enclosure similar
 512 to that shown in figure 6.9. An early design collimator was also attached, and a
 513 bare fibre was used as a control. The test deployment used 200 μm core 0.22 NA
 514 multimode step-index fibres, notably different to the 0.5 NA fibres used for the
 515 summer deployment. Threaded screws held the optical elements in place, and the
 516 mounting disc was lowered using a mechanical winch with stainless steel chains
 517 attached to three triangulated shackles.

518 Light injection was provided through a 1 mm core fibre by an electronically
 519 pulsed LED system provided by the University of Liverpool. The light from the
 520 1 mm core fibre was then separated into three 200 μm core step-index fibres coupled
 521 in a triangular formation. The first injected light into the tank, the second fed
 522 through to a monitor PMT, and the last was attached to an oscilloscope for in-house
 523 monitoring. Light pulse durations were varied to effectively increase the magnitude of
 524 light emitted from the optical devices. The duration of the pulses were not measured

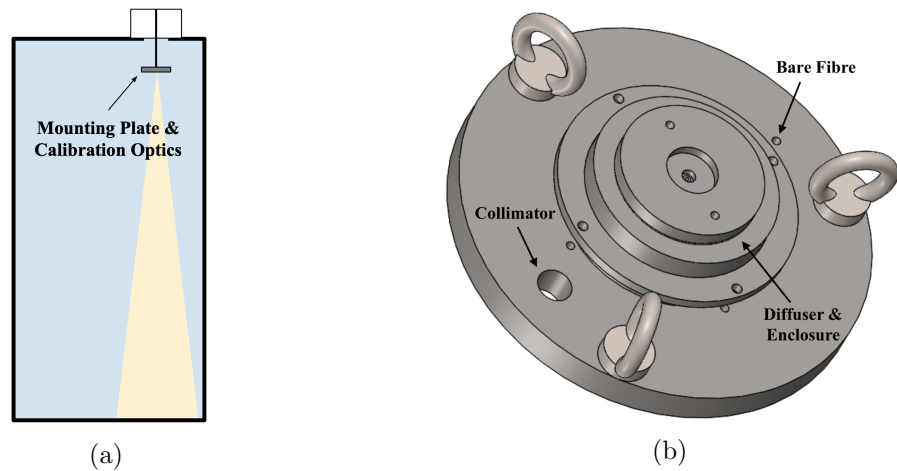


Figure 6.23: (a) A simplified drawing demonstrating the installation of calibration optics in Super-K during the test deployment. (b) The mounting plate used during the Super-K test deployment, with mounting positions for each optical system labelled.

525 directly, but intensity values corresponding to the pulse duration were arbitrarily
 526 defined through the input register on the FPGA board. The nature of the diffuser
 527 devices required higher relative intensity compared to the bare fibre and collimator,
 528 which was achieved through longer pulse lengths. A combination of long undefined
 529 pulse lengths and step-index fibres meant information regarding potential timing
 530 calibration performance was not possible.

531 The PMT hit response was recorded for the top, bottom, and walls of the tanks
 532 to obtain the average number of hits per event over each individual PMT. From this
 533 diffuser and bare fibre event displays can be made to visually characterise the diffuser
 534 performance, these are shown in figure 6.24. A trigger was taken using a 473 nm
 535 laser from the bottom tank injection point. Each run consisted of approximately
 536 100,000 events triggered at a rate of 100 Hz. Temporal cuts over the laser trigger
 537 and pulse width were applied to isolate the relevant events. Example event displays
 538 for Super-K PMT hit occupancy using the diffuser and bare fibre assemblies are
 539 shown in figure 6.24. Drawing quantitative conclusions between the two plots is
 540 difficult given the different, also arbitrary, intensities in light injection. Nevertheless
 541 qualitative comparisons can be made. Firstly the emitted light from the diffuser
 542 assembly can be seen in the Super-K PMTs approximately 40 m away from the top

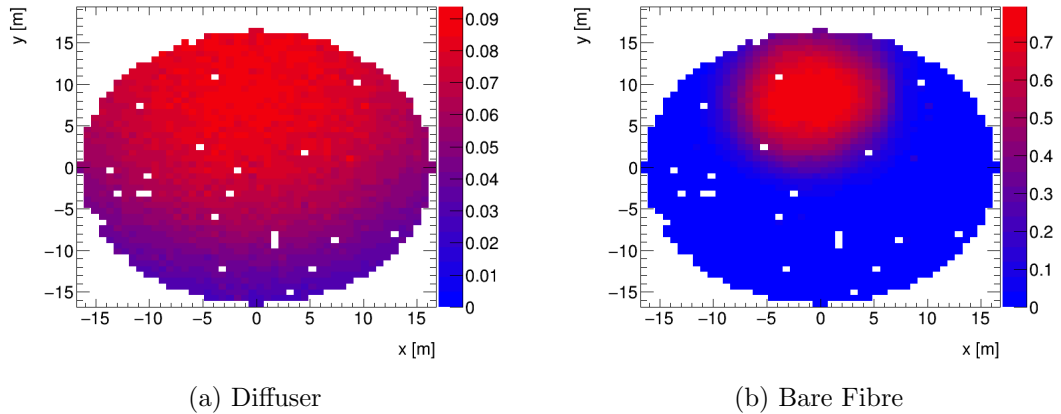


Figure 6.24: The Super-K PMT hit occupancy over the bottom of the tank for (a) the diffuser and (b) the bare fibre control. The units of hit occupancy are number of hit per event/ns.

543 to bottom of the tank. The spot size for the diffuser is significantly larger than
 544 the bare fibre control, which indicates the diffuser is performing it's intended role
 545 in illuminating a wider spread of PMTs. The hit occupancy for each individual
 546 illuminated PMT is an order of magnitude lower than the bare diffuser, despite the
 547 longer pulse length used for the diffuser. Future development considerations were
 548 therefore made to maximising light input into the diffuser and minimising losses
 549 through attenuation in the fibre and coupling. In theory this would also allow short
 550 enough pulse widths to enable timing calibration.

551 Zeroth order comparisons can also be made to the experimental laboratory
 552 results shown outlined in figure 6.17. The geometric field of view from the top to the
 553 bottom of the tank spans approximately $\pm 20^\circ$. It can be seen from the V3 enclosure
 554 profiles in figures 6.17 and 6.27, the relative light intensity in air varies no more
 555 than around 10% over this field of view. A projection of the x-axis in the diffuser
 556 event display (figure 6.24a) at the y-axis injection point can be seen in figure 6.25.
 557 Roughly a 20% decrease in intensity is seen over approximately $\pm 20^\circ$, more than
 558 the laboratory results. The reason for such a discrepancy can be explain through
 559 unaccounted factors such as, the greater refractive index of water, attenuation length,
 560 as well as geometric affects such as a $1/r^2$ relation and PMT solid angle. The short
 561 time between the test and summer deployments meant that only basic data analysis
 562 studies could be performed before production for the summer deployment had to

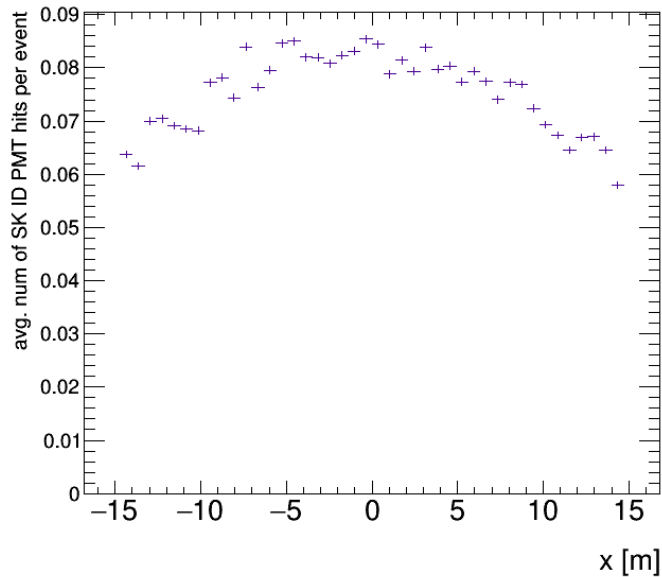


Figure 6.25: A projection of the diffuser light profile in the x-axis, taken at the point of injection in the y-axis.

563 begin. The optics were left in the tank for approximately six months. Afterwards
 564 the optics were taken out with no obvious mechanical failures found in the diffuser
 565 assembly.

566 6.5.2 Summer Deployment

567 In the summer of 2018, the Super-Kamiokande detector was drained for scheduled
 568 maintenance, and an updated optical calibration system from the test deployment
 569 was installed. Similar to the proposed Hyper-K inner detector configuration outlined
 570 in section 6.1, five injectors were installed at regular vertical intervals (B1 to B5) on
 571 the Super-K tank (figure 6.26a). Each light injection system consists of an updated
 572 collimator, bare fibre control, and an amended V3 diffuser assembly which included
 573 the additional strain-relief waterproofing measures outlined in section 6.4.4. The
 574 calibration equipment was mounted on the inner detector PMT frame using the
 575 bracket shown in figure 6.26b. Alignment over 3 degrees of freedom was controlled
 576 through adjusting triangulated screws at the bottom of the bracket. Additionally
 577 an optional laser pointer could be used to illuminate opposing PMTs acting as an
 578 alignment aid. Tyvek sheeting surrounded the mounting bracket in an attempt to

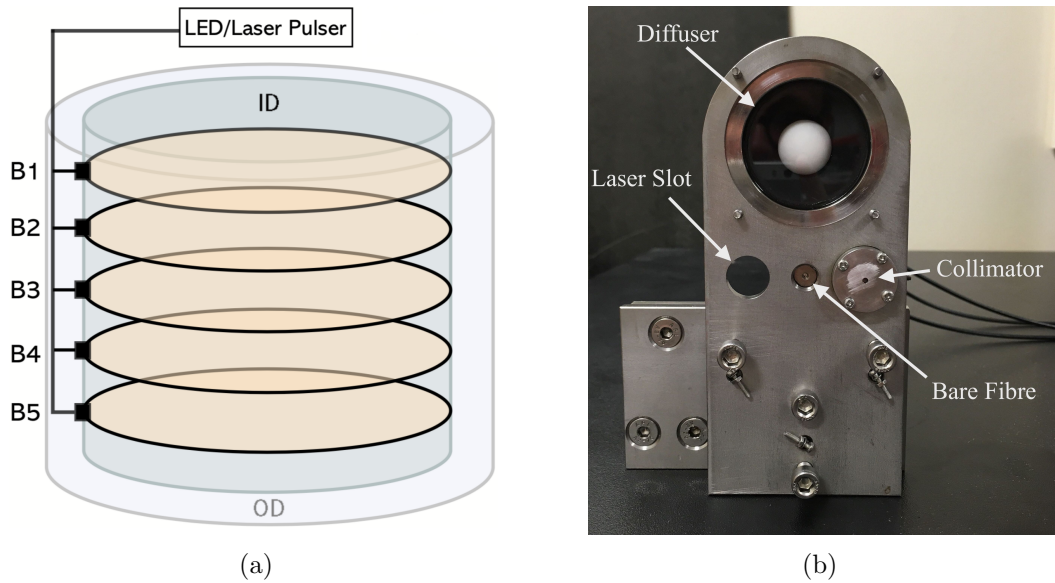


Figure 6.26: (a) A representation of the five light injection points (black squares) used for the summer deployment, taken from [151]. (b) The redesigned mounting bracket for the summer deployment.

579 reduce backscattered light from the optical devices entering the inner detector tank.
 580 The first light injector was installed on the 29th June 2018; the final injector was
 581 installed on the 13th August 2018.

582 Ten bare diffusers were manufactured for the summer deployment, with each
 583 individual light profile shown in figure 6.11. Seven fully assembled diffuser assemblies
 584 were made, five originally designed for deployment and two acting as spares. A
 585 selection process based on qualitatively selecting the most consistent relative light
 586 profiles dictated which bare diffusers would be used in assembly. The relative light
 587 profiles for the resulting fully assembled diffuser systems, labelled D1 to D7, are
 588 shown in figure 6.27. Table 6.1 outlines which diffuser assembly was installed at each
 589 light injection point.

590 Photon injection into the calibration optics was provided by the same set up
 591 used in the test deployment described previously. A pulsed LED provided a light
 592 source which is then partitioned into three outputs: A designated monitor PMT,
 593 on-site monitoring system for validity checks, and the calibration optics. The light is
 594 propagated through a 200 μm core step-index fibre optic cable which changed from
 595 0.22 NA from the test deployment to 0.5 NA to maximise the light yield through the

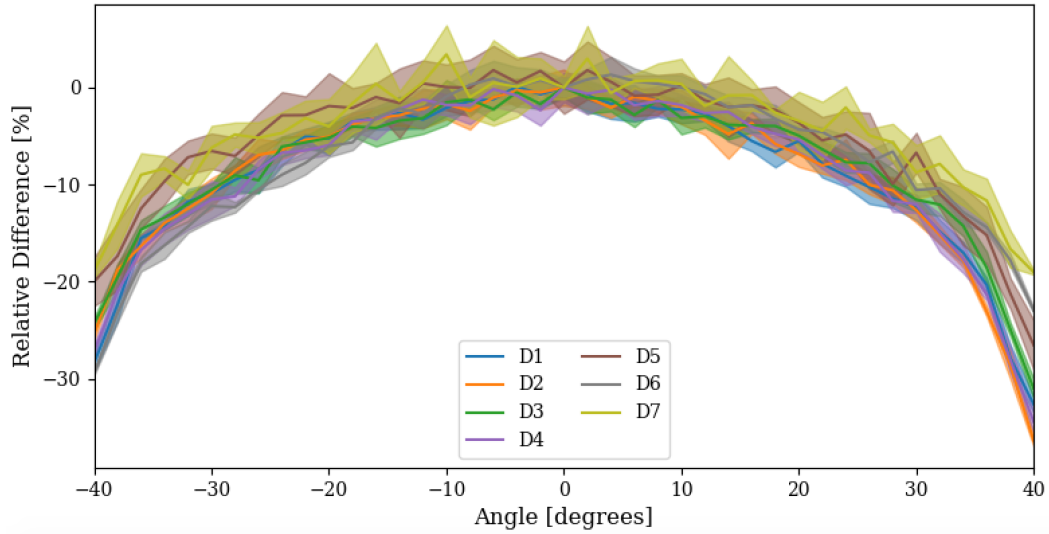


Figure 6.27: The seven full diffuser enclosure assemblies for the Super-K deployment, labelled from D1 to D7 during the assembly phase for clarity during measurements.

Full Diffuser Assembly	Injection Point Installed	Comments
D1	B1	Spare at Super-K Kept at Warwick for testing
D2	Spare	
D3	B3	
D4	B2	
D5	Spare	
D6	B5	
D7	B4	

Table 6.1: A summary of which diffuser assemblies were installed at each of the injection points for the summer deployment.

596 fibre. Underwater fibre optic connections were used close to the mounting bracket
597 to limit the strain on long fibres during deployment, particularly with injection
598 points near the bottom of the tank. Water-tightness of the connections was ensured
599 by submerging the connections in boxes filled with Mineguard™, a viscous epoxy
600 material developed in the mining industry previously used in waterproofing Super-K
601 PMTs [80]. The optical calibration system was designed with a longevity on the
602 order of approximately 20 years, and is expected to remain in the tank collecting
603 calibration data for the foreseeable future.

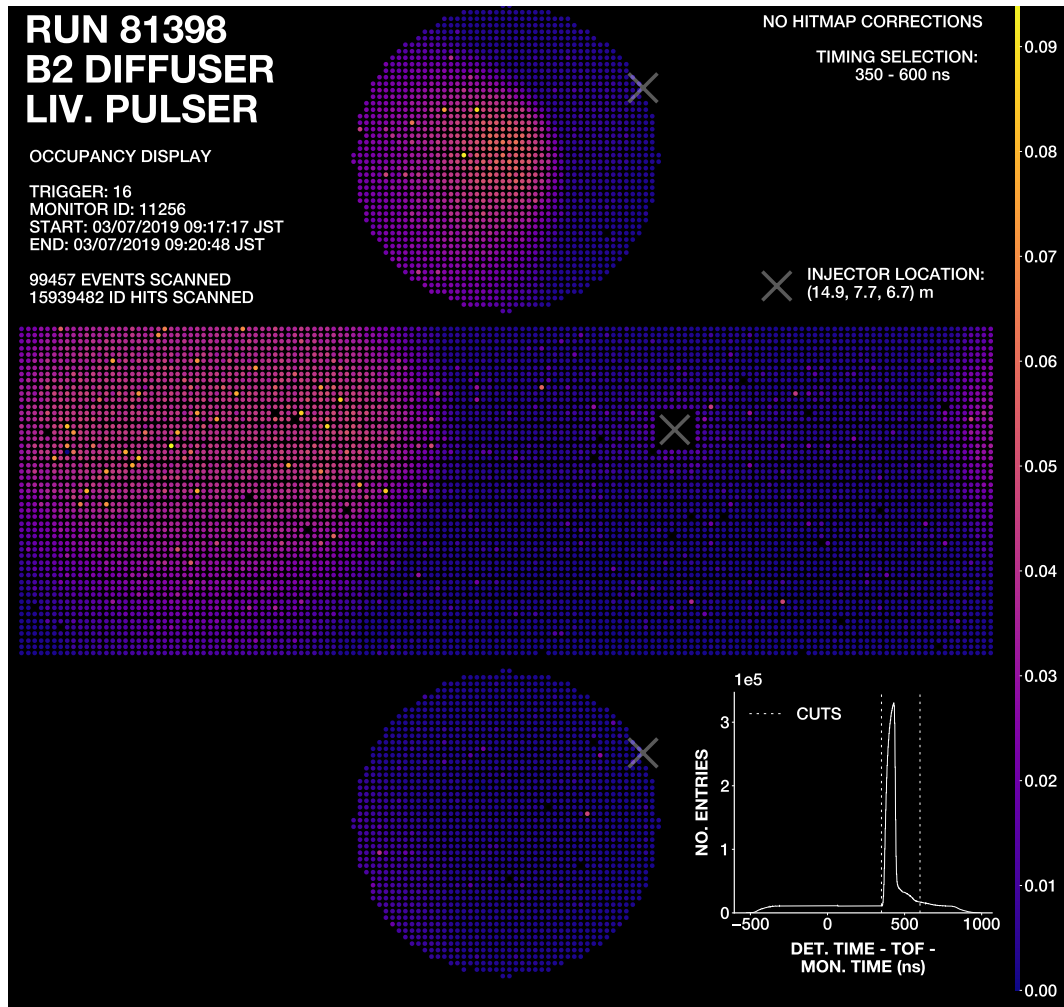


Figure 6.28: An example PMT hit occupancy event display for diffuser installed at the B2 injection point over the full detector. The number of SK PMT hits as a function of time is also shown in the bottom right, from which cuts are applied. Plot modified from [157].

604 6.5.3 Results

605 Analogous to the test deployment the PMT hit response for all the calibration optics
 606 was measured around the tank; firstly using the pre-existing Korean laser system,
 607 secondly using the UK light injection system described above. An example diffuser
 608 event display, located at the B2 injection point, is shown in figure 6.28. The same
 609 473 nm laser trigger was used as the test deployment. The number of Super-K PMT
 610 hits is recorded as a function of time, a monitor PMT time pedestal is subtracted
 611 and a correction is also applied accounting for time-of-flight. Temporal cuts can be

612 applied to the resulting plot, shown in figure 6.28, to isolate the light injection pulse.
613 From the duration of the temporal cut one can estimate the length of a typical light
614 injection event to be on the order of a 200-300 ns. The lower cut threshold is strict to
615 veto any hits before the pulse arrives; the upper threshold can be relaxed (tightened)
616 to include (exclude) internal detector reflections. Super-K event displays for all other
617 installed diffuser systems can be found in the appendix, figures B.3 to B.6. A shadow
618 can be seen in some of the diffuser displays, this is believed to be caused by the
619 collimator assembly protruding too far outwards in the mounting bracket (figure
620 6.26b). The shadowing effect is most prominent in B4 and B5 whereby the bottom
621 of the tank is illuminated and the collimator blocks the line-of-sight.

622 Preliminary qualitative conclusions suggest that the diffuser is working as
623 expected. The deployment appears to have been successful and the event displays
624 are promising. A full analysis is in progress which aims to quantify the calibration
625 optics performance, make comparisons with laboratory profiles, and outline the
626 systems potential to perform PMT calibration and water property measurements
627 inside Super-K. The analysis will also be extended to make performance predictions
628 for other water Cherenkov detectors, most notably Hyper-Kamiokande.

629 **6.6 Future Development**

630 Diffuser research and development has continued since the Super-K deployment, with
631 the intention to develop a final system for mass production for the use of Hyper-K,
632 and potentially other large scale water Cherenkov detectors. This section will discuss
633 the recent investigations into PTFE as a new diffusing material as well as enclosure
634 design research and development moving forward.

635 **6.6.1 PTFE**

636 A discussion into diffusing materials is provided in section 6.3.1. Poly(methyl
637 methacrylate), otherwise referred to as PMMA, was chosen as the diffusing material
638 for the Super-K deployments. Whilst PMMA is known to have good diffusing
639 properties and produces well understood optical profiles, the notable disadvantages

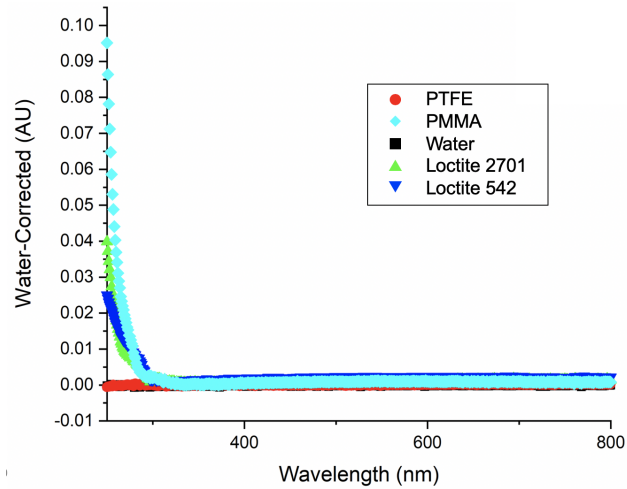


Figure 6.29: Optical transmission measurements as a function of wavelength across the UV-VIS spectrum, for different diffusing and sealant materials.

640 are in its difficulty in manufacturing and porous nature, which in turn require the
 641 use of a water-tight enclosure. There has since been a push to find new potential
 642 diffusing materials; one candidate is poly-tetrafluoroethylene, otherwise known as
 643 PTFE or Teflon.

644 Virgin PTFE is renowned for its excellent chemical and water resistant
 645 properties [158]. Soak tests in ultra-pure and gadolinium loaded water sample are
 646 in progress with preliminary results indicating no visible leeching into the water
 647 solutions. If successful, and the transmission properties of water exposed PTFE
 648 diffuser are unchanged, the need for a water-tight diffuser enclosure is put into
 649 question. Nevertheless, studies have demonstrated the enhancements to optical
 650 profiles that enclosures may potentially provide (section 6.6.3). Moreover, a water-
 651 tight enclosure may be used to ensure longevity on the scale of multiple decades.

652 The proposed optical calibration system for Hyper-K requires the diffusing
 653 material to have good transmission properties over the UV-VIS spectrum from
 654 approximately 300 nm to 500 nm. The optical transmission of PMMA and PTFE
 655 using a Shimadzu UV-2600 spectrometer is shown in figure 6.29. Each measurement
 656 is corrected to a water control to eliminate any water band features. A value of
 657 zero indicates the same transmittance as water, which is ideal for the calibration
 658 optics. It can be seen that both PMMA and PTFE perform well at wavelengths

659 across the UV-VIS spectrum. However below approximately 350 nm, PTFE retains
660 its transmittance, whilst PMMA begins to absorb more light, indicating that PTFE
661 performs as good, if not better, than PMMA over the full UV-VIS spectrum.

662 The manufacturing process of PMMA, as described in section 6.3.3, is not
663 ideal and has proved problematic in scaling up to mass production. PTFE rods can be
664 purchased directly from industrial manufacturers. The only in-house manufacturing
665 needed after this stage is crafting the hemispherical diffusers. Each 1 m PTFE rod
666 can be manufactured into an estimated 30 to 40 diffuser balls, compared to 2 to 3
667 diffusers from each 100 mm rod of PMMA. This results in a reduction in material
668 costs of 65%, and an even larger saving in labour costs as the sintered rods are no
669 longer manufactured in-house. Furthermore, PMMA is extremely porous and requires
670 machining tools to be scrupulously cleaned of oils, suds, and other containments
671 beforehand. PTFE by nature is more impermeable, meaning less cleaning is needed
672 before machining.

673 **6.6.2 PTFE Optical Performance**

674 In a review process, diffusers made out of PTFE must pass all of the optical and
675 temporal performance tests that their PMMA counterparts had passed previously.
676 Any major issues can then be highlighted, advantages and disadvantages discussed,
677 before any decisions are made.

678 The optical light profiles of bare PMMA and PTFE diffusers are shown in
679 figure 6.30. It can be seen that the PTFE relative light intensity profile is comparable
680 to PMMA with a marginally narrower distribution. The magnitude of light emitted
681 is larger for the PTFE diffusers with approximately 15% more light emitted in the
682 forward going region (figure B.7). A comparison of the pulse delay, shown in figure
683 B.8, demonstrates PTFE has the same uniform timing profile as PMMA. Together
684 the temporal and light intensity profiles suggest the performance of PTFE as a
685 diffuse calibration device is similar to PMMA. Further investigations into PTFE
686 batch reproducibility and pressure testing are currently in progress before any final
687 decisions are made about the diffusing material moving forward.

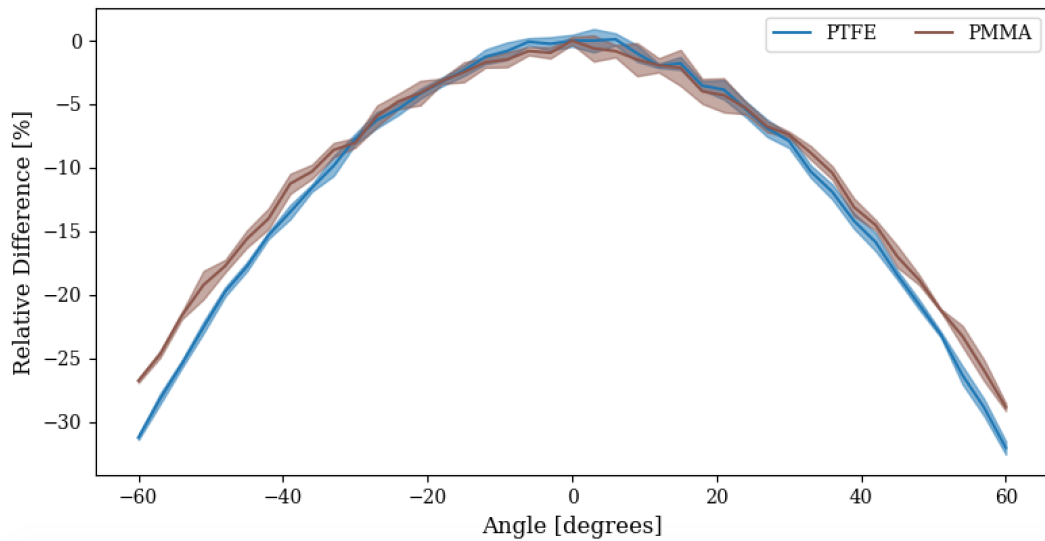


Figure 6.30: A comparison of the relative light intensity profiles, normalised to 0° , for bare PMMA and PTFE diffusers.

6.6.3 Enclosure Development

Post Super-K deployment reviews highlighted particular flaws in the V3 enclosure design. The most significant problem, as section 6.4.4 has already alluded to, is the measures used to waterproof the enclosure. Studies had shown that the Vitron o-ring gaskets had performed well in enclosures V1 and V3. However, sealing the fibre feed-through point in the back of the enclosure was extremely problematic. Epoxy resin did not bond well with the PVC fibre furcation tubing, and had to be applied liberally around the entire enclosure. Application of epoxy resin made assembly of enclosure V3 intricate and not feasible on the large scale mass production needed for Hyper-K. The philosophy was to turn to mechanical waterproofing and create a new enclosure design aimed towards large scale mass production.

V4 Enclosure

A schematic diagram of enclosure V4 is shown in figure 6.31, along with a photograph of the fully assembled front of the enclosure. Similar to the previous designs, enclosure V4 consists of three cylindrical segments. The individually threaded segments used for assembly has been replaced by six long bolts that feed-through the entire design. Three o-ring gaskets, one either side of the window, and one between the main

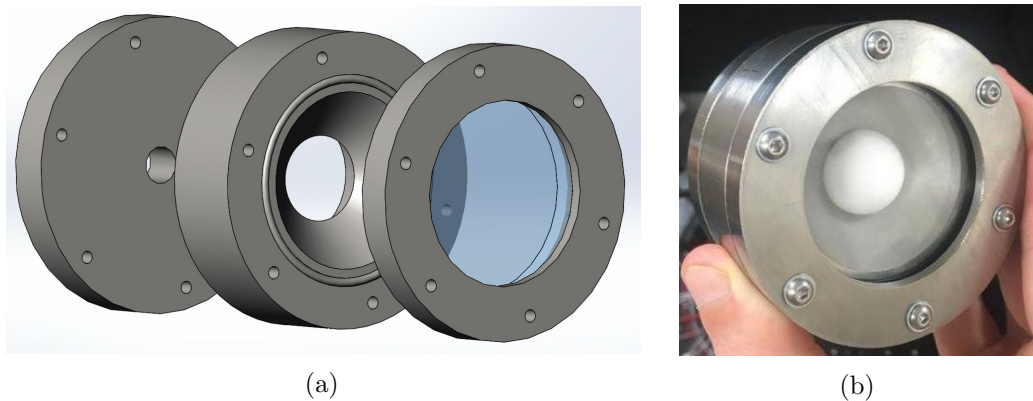


Figure 6.31: (a) A schematic CAD drawing of the V4 enclosure. (b) A front facing photo of enclosure V4, fully assembled with a sand-blasted stainless steel torch surface.

705 body and base, are quashed when the bolts are tightened providing mechanical
706 watertight seals. Replacing the epoxy resin, preliminary concepts for fibre feed-
707 through waterproofing use screw in fibre ports with a thread sealant. Studies are
708 in progress to pressure test the enclosure up to 10 bar, with further development
709 expected for future water Cherenkov detector experiments such as Hyper-K.

710 **Surface Treatments**

711 When considering internal enclosure reflections, a distinction should be made between
712 specular and diffuse reflections. Specular reflections are often unwanted as they
713 strongly bias the light output in a particular direction. However, diffuse reflections
714 scatter incident light rays at many different angles during reflection. Different surface
715 treatments to the internal design of the enclosure have been tested to compare the
716 various effects on the outgoing light profile. All tests were performed using a V4
717 enclosure design, which is shown in figure 6.31. Each enclosure was made out of
718 304 stainless steel, except for the 3D printed enclosure which was made out of a
719 carbon fibre based composite material. A measurement of untreated stainless steel
720 was used as a control, and then the internal torch base was painted matte black and
721 white. The torch was then sandblasted to finely roughen the surface, the previous
722 treatments were then applied, and measurement retaken. The resultant light profiles
723 using a standard PMMA diffuser are shown in figure 6.32. The results indicate a

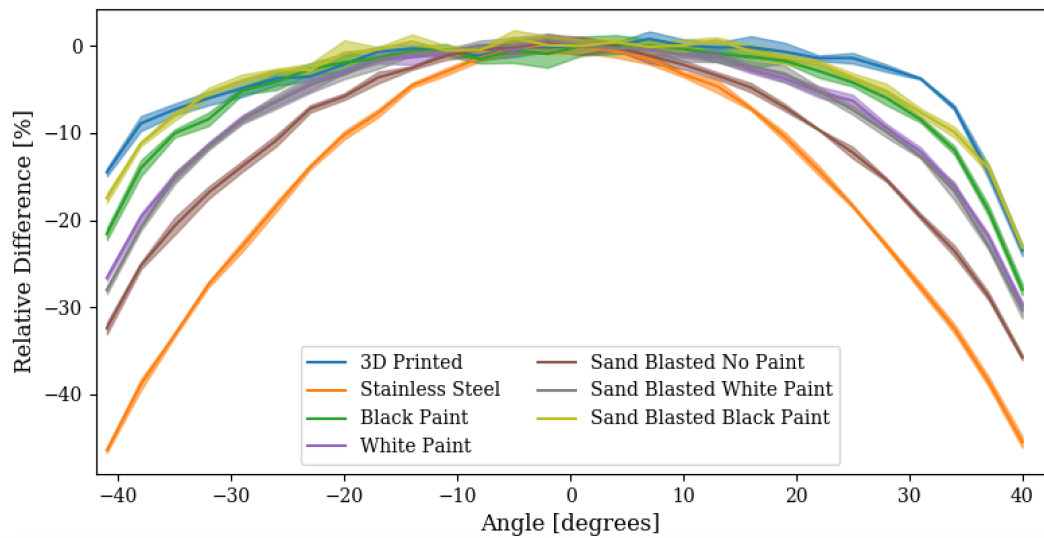


Figure 6.32: The relative light profiles, normalised to 0° , for the PMMA diffuser inside enclosure V4 for different surface treatments of the torch.

724 correlation between the surface treatments and the uniformity of light profiles within
 725 the enclosure field of view. In particular, smoother metallic treatments were found
 726 to encourage specular reflections and increase the forward going light output. Sand
 727 blasting has the affect of roughening the surface and replacing potential specular
 728 reflections with their diffuse counterparts. Painting the surface with a matte black
 729 paint effectively removes most reflections. Perhaps interestingly, painting with matte
 730 white paint acts as a mid ground between sand blasting and painting black. The
 731 absolute intensities are also shown in figure 6.33, here it is more trivial to conclude
 732 which surfaces are promoting and inhibiting internal enclosure reflections. Matte
 733 black paint is found to reduce the pulse intensity, whilst painting white appears to
 734 increase the overall pulse intensity integrated over all angular space.

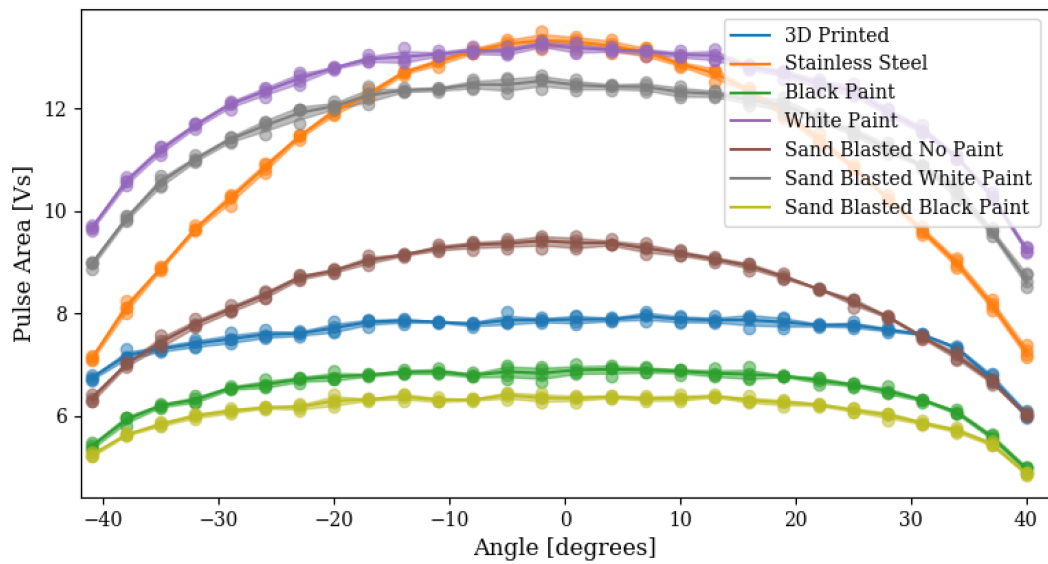


Figure 6.33: The pulse area as a function of angle for the PMMA diffuser inside enclosure V4 for different surface treatments of the torch.

1 Chapter 7

2 Summary and Closing Remarks

3 In this thesis the topic of neutrino physics has been introduced, with an emphasis
4 on neutrino-nucleus interactions, and how cross-section measurements can help
5 with an overarching goal of measuring key neutrino oscillation parameters. A brief
6 history of neutrinos has been explored, from their discovery through to modern
7 day neutrino oscillation experiments. Chapter 3 outlined a detailed description
8 of the T2K experiment, including a discussion of the ND280 near detector which
9 has subsequently been used to measure the ν_e CC π^+ cross-section in chapter
10 4. Proposals for the Hyper-Kamiokande experiment, the next generation water
11 Cherenkov detector from Super-Kamiokande, were introduced in chapter 5. Finally,
12 the optical calibration system, as well as diffuser research and development for the
13 Hyper-K detector were examined. This closing chapter will summarise both the T2K
14 cross-section analysis, and Hyper-K diffuser calibration studies, reviewing the results,
15 and proposing potential avenues for future research.

16 7.1 ν_e CC π^+ Cross Section Analysis Summary

17 A selection has been developed to analyse post-FSI π^+ production from charged
18 current electron neutrino interactions in ND280. The lepton selection inherits from
19 the CC- ν_e inclusive analysis, and includes an additional new π^+ selection. Out of
20 fiducial volume photon background, prevalent in the CC- ν_e inclusive analysis, is
21 significantly reduced through pion selection. Photon backgrounds from ν_μ CC π^0

22 and neutral current interactions are the predominant source of background in the
 23 forward going low lepton momenta regions; backgrounds from muon interactions
 24 dominate in higher lepton momentum regions. A signal purity of 51.1% is selected
 25 at an efficiency of 25.4% over the full selection with phase space constraints applied.
 26 Systematic uncertainties on detector effects, cross-section model parameters, and flux
 27 have been calculated for their relative effects on both background event yields and
 28 signal efficiencies. The flux integrated cross-section, over one bin in momenta space,
 29 was measured to be $(2.23 \pm 0.18 \text{ (stat.)} \pm 0.31 \text{ (syst.)}) \times 10^{-39} \text{ cm}^2$ per nucleon.
 30 This is the first measurement of the ν_e CC π^+ cross-section on a carbon target ever
 31 made. The result agrees with nominal NEUT 5.4.0 MC, within both statistical
 32 and systematic errors. Finally, a region of low momenta phase space was defined
 33 analogous to the Super-K FHC 1 decay electron sample that observes an excess in π^+
 34 production. Data-MC comparisons within this region give preliminary indications
 35 that no excess is seen at the near detector, and is the beginning to providing a
 36 constraint on the far detector process.

37 Whilst achieving preliminary results, further investigations could be made to
 38 help understand the interaction process at the near detector. Firstly, time constraints
 39 limited model comparisons to nominal NEUT predictions only. A more complete
 40 study should compare results to other neutrino event generator predictions, such
 41 as GENIE and NuWro. Resonant pion production in NEUT is described, at an
 42 invariant mass $W \leq 2 \text{ GeV}/c^2$, using the Rein-Seghal model [46], with a resonant
 43 axial mass set to $0.95 \text{ GeV}/c^2$. Deep inelastic scattering is modelled using the GRV98
 44 parton distribution function [159], including the Bodak and Yang corrections [160],
 45 for $W \geq 1.3 \text{ GeV}/c^2$. GENIE has very similar treatments to NEUT for resonant
 46 pion production and DIS processes, but uses a resonant axial mass of $1.12 \text{ GeV}/c^2$
 47 and a slightly different Bodak and Yang correction respectively. Resonances are also
 48 switched off in GENIE above $W > 1.7 \text{ GeV}/c^2$ to avoid double counting with DIS.
 49 A comparison against NuWro would be interesting given it's different treatment of
 50 resonant pion production; A single Δ -model by Adler-Rarita-Schwinger [161] is used
 51 with an axial mass term of $0.94 \text{ GeV}/c^2$ at $W < 1.6 \text{ GeV}/c^2$. A smooth transition
 52 from resonance to DIS processes then takes place from hadronic masses of $1.3 \text{ GeV}/c^2$

53 to $1.6 \text{ GeV}/c^2$. Comparisons of data against multiple neutrino event generators
54 provides a measure of testing the performance of these different interaction models.
55 These studies were not able to be performed within the thesis time scale, but provide
56 suggestions for the analysis moving forward.

57 Multiple improvements to the analysis could be made in the future. One
58 suggestion would be to perform a multivariate analysis on the systematic uncertainties
59 to study the interplay and correlations between individual systematics, which is not
60 yet considered. Furthermore, the cross-section measurement was performed using a
61 zeroth-order calculation. A likelihood fit package, such as `nueXsLLFitter` used in the
62 $\text{CC-}\nu_e$ inclusive measurement, allows for more sophisticated error propagation and
63 cross-section extraction. Multi-bin measurements of the cross-section over different
64 areas of phase space would be possible with a likelihood fitter. The cross-sections
65 of ν_μ and ν_e should be similar at higher energies, and so it is more interesting to
66 investigate the low Q^2 regions of phase space. The analysis could also be expanded
67 to include RHC data runs, FHC runs 9-10, and FGD2 interactions thereby increasing
68 statistics. Limitations of the analysis include its preference to forward going events,
69 and criterion for tracks to leave the FGD and enter the TPC. Finally, multiple efforts
70 have seen recent T2K ν_μ analyses attempt to measure over a 4π angular coverage
71 [162], and include isolated FGD pions [163]. One could in theory extend these ideas
72 to ν_e analyses in the future.

73 **7.2 Diffuser Systems For Optical Calibration**

74 Proposals for a light injection system to optically calibrate water Cherenkov detectors
75 such as Hyper-Kamiokande are underway. The proposed system uses two optical
76 calibration devices: A wide beam diffuser for PMT energy and timing calibration,
77 and a narrow beam collimator for monitoring water properties. This thesis focussed
78 on the research and development of diffuser technology. A principle bare diffuser
79 has been designed with a well understood light and uniform timing, profile over a
80 wide angular coverage, from $-60^\circ \rightarrow 60^\circ$. The diffusers are made out of PMMA
81 and are held in a water-tight stainless steel enclosure. Not only have the principal

82 designs have been demonstrated to work under laboratory conditions, but have also
83 twice been successfully deployed in the Super-Kamiokande detector. Preliminary
84 qualitative analyses have indicated that the diffusers are working as expected.

85 A full quantitative analysis on the Super-K deployment, based on the diffuser
86 performance and the subsequent calibration potential has begun, and will heavily
87 influence future research and development. A post-deployment external review
88 identified potential weaknesses in the diffuser design. Two key areas identified
89 were the waterproofing measures, and scalability for mass production. Bonding
90 between epoxy resin and fibre furcation tubing was poor, with ad-hoc strain relief
91 accessories needed to mitigate against failures under pressure. Furthermore, the
92 liberal application of epoxy resin and intricate assembly of the diffuser enclosures
93 are problematic when scaling towards mass production. Development of diffuser
94 enclosures V4 and above need to ensure systematic water-tightness under pressure; the
95 design philosophy has moved to mechanical seals, which are known to work through
96 previous enclosure body pressure tests. Screw in fibre ports with a thread sealant,
97 will provide a mechanical watertight seal at the fibre feedthrough point. The porous
98 nature of PMMA has also led to a search for alternative diffuser materials; PTFE
99 has been highlighted as a potential candidate due to it's water and chemical resistant
100 properties. Investigations into PTFE as a diffusing material, and it's comparative
101 performance against PMMA, are underway with preliminary results suggesting
102 similar diffusing characteristics. PTFE also provides an easier means to machining,
103 and scaling to mass production. Upgrades are planned for the experimental setup
104 outlined in section 6.2. The light injection system is to be upgraded with a pulsed
105 laser, which will allow for accurate laser power monitoring. The upgraded test system
106 will also allow for 2-dimensional diffuser scans, providing a more complete mapping
107 of the diffuser light profile.

108 The optical calibration diffuser work presented in this thesis demonstrates a
109 successful diffuse light injection system for PMT energy and timing calibrations. With
110 further research and development planned the diffuser system has been proposed
111 for installation in Hyper-Kamiokande, and has the potential to be adapted for
112 installations in other water Cherenkov detectors.

¹ **Appendix A**

² **T2K Analysis Appendix**

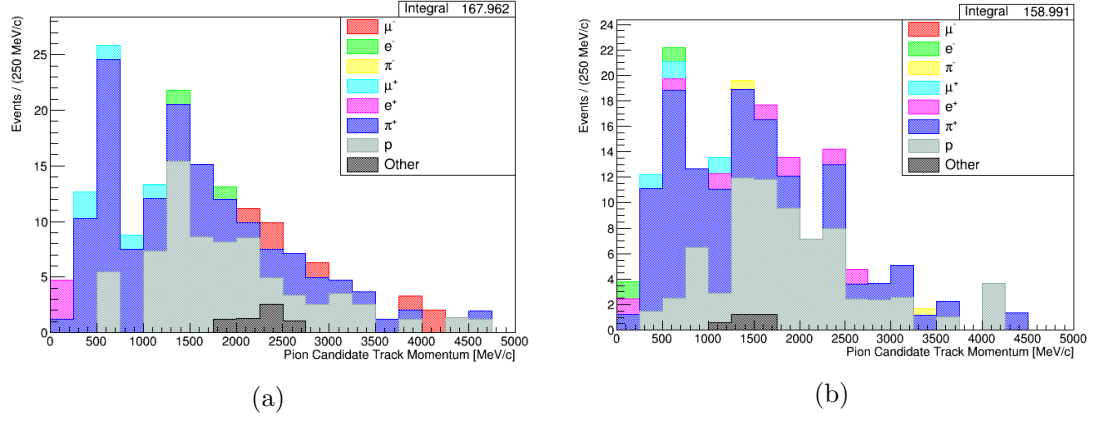


Figure A.1: A histogram demonstrating the true particle selected for the pion candidate track, as a function of the track's reconstructed momentum. The ν_μ CC photon background topology is isolated on the left, the NC photon background topology on the right.

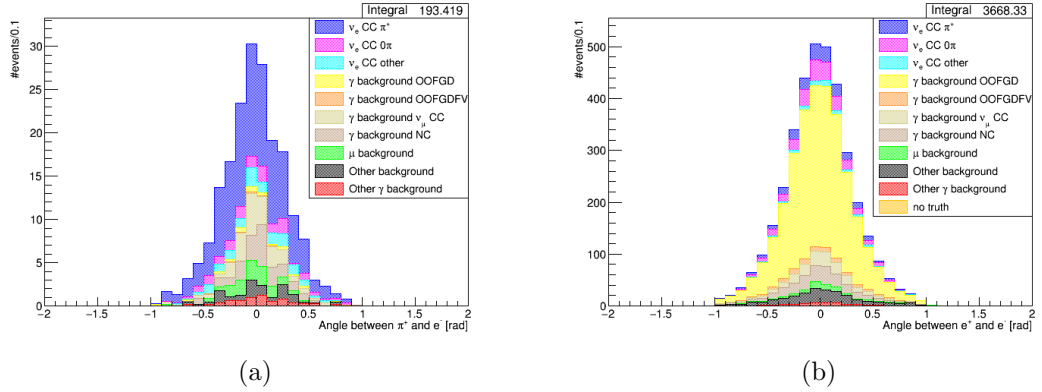


Figure A.2: The angle between the two selected tracks for (a) e^- and π^+ in the ν_e CC π^+ selection sample, and (b) the e^+e^- pair in the vertexing systematic sample.

Sample	PoT	NSpills	Nbunches	NECal	ECal/bunch (%)
Run 2 - Water Out					
Data	3.59337e+19	423187	3385496	562888	16.6265
MC	1.6794e+21	1.97781e+07	1.58225e+08	1.91209e+07	12.0846
Sand	7.05023e+20	8.30297e+06	6.64237e+07	2.81367e+06	4.23594
Run 2 - Water In					
Data	4.33934e+19	598617	4788936	689265	14.3929
MC	1.20375e+21	1.66058e+07	1.32847e+08	1.38547e+07	10.4291
Sand	7.05023e+20	9.72588e+06	7.7807e+07	2.81367e+06	3.61622
Run 3b - Water Out					
Data	2.17273e+19	260193	2081544	315907	15.1766
MC	3.07766e+21	3.68563e+07	2.9485e+08	3.46192e+07	11.7413
Sand	7.05023e+20	8.44295e+06	6.75436e+07	2.81367e+06	4.16571
Run 3c - Water Out					
Data	1.36447e+20	1480300	11842400	2.14796e+06	18.1378
MC	3.07766e+21	3.33893e+07	2.67114e+08	3.46192e+07	12.9605
Sand	7.05023e+20	7.64874e+06	6.119e+07	2.81367e+06	4.59826
Run 4 - Water Out					
Data	1.78319e+20	1529336	12234688	2.74962e+06	22.4739
MC	3.41282e+21	2.92697e+07	2.34157e+08	3.83982e+07	16.3985
Sand	7.05023e+20	6.04656e+06	4.83725e+07	2.81367e+06	5.81668
Run 4 - Water In					
Data	1.64228e+20	1600804	12806432	2.58112e+06	20.1549
MC	3.61215e+21	3.52091e+07	2.81673e+08	4.10491e+07	14.5733
Sand	7.05023e+20	6.87216e+06	5.49773e+07	2.81367e+06	5.11788
Run 8 - Water Out					
Data	4.15013e+20	1766203	14129624	5.61527e+06	39.7411
MC	3.61002e+21	1.53634e+07	1.22908e+08	3.68835e+07	30.0092
Sand	7.05023e+20	3.00042e+06	2.40034e+07	2.81367e+06	11.722
Run 8 - Water In					
Data	1.58053e+20	778207	6225656	2.18188e+06	35.0465
MC	2.71677e+21	1.33766e+07	1.07013e+08	2.8016e+07	26.18
Sand	7.05023e+20	3.47133e+06	2.77707e+07	2.81367e+06	10.1318

Table A.1: Table showing the numbers used to evaluate the correction and systematic uncertainty for ECal pileup affecting FGD1 target selections.

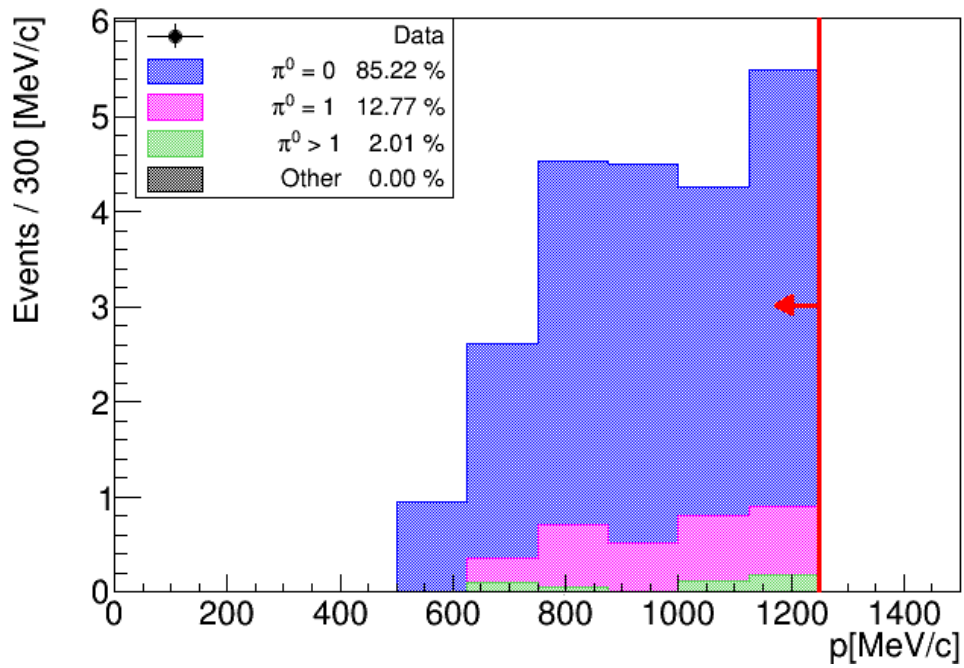


Figure A.3: The number of π^0 particles present in the ν_e CC π^+ signal sample at low momentum regions comparable to Super-K.

¹ **Appendix B**

² **Hyper-K Analysis Appendix**

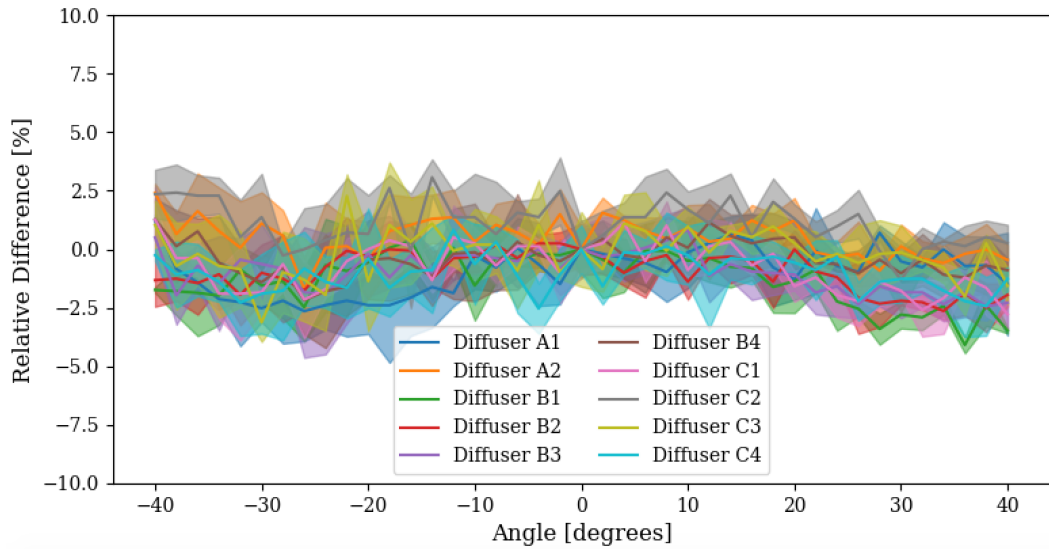


Figure B.1: The relative full width half maximum of the signal pulse, normalised to zero degrees, for PMMA bare diffusers.

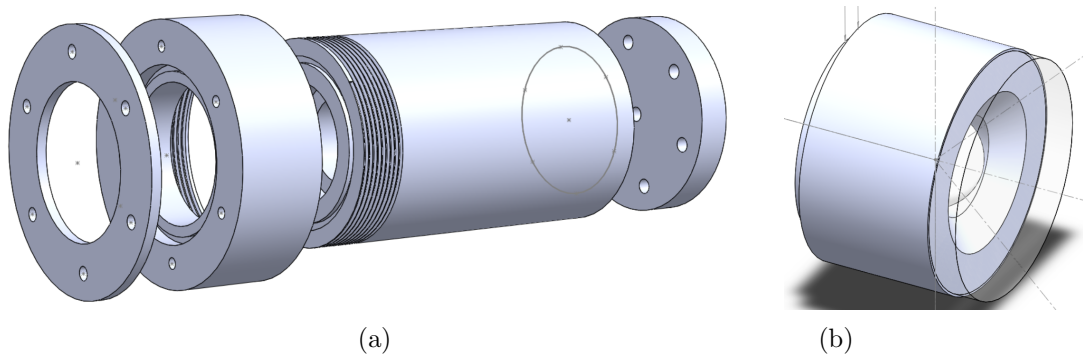


Figure B.2: The intermediate conceptual enclosure designs between V1 and V3. (a) V2 consisted the long main body that was prominent in V1 in combination with the threaded screw design seen in V3. (b) V2a was a singular enclosure design smaller than previous, with a torch-like design at the front. Neither V2 or V2a made it to production.

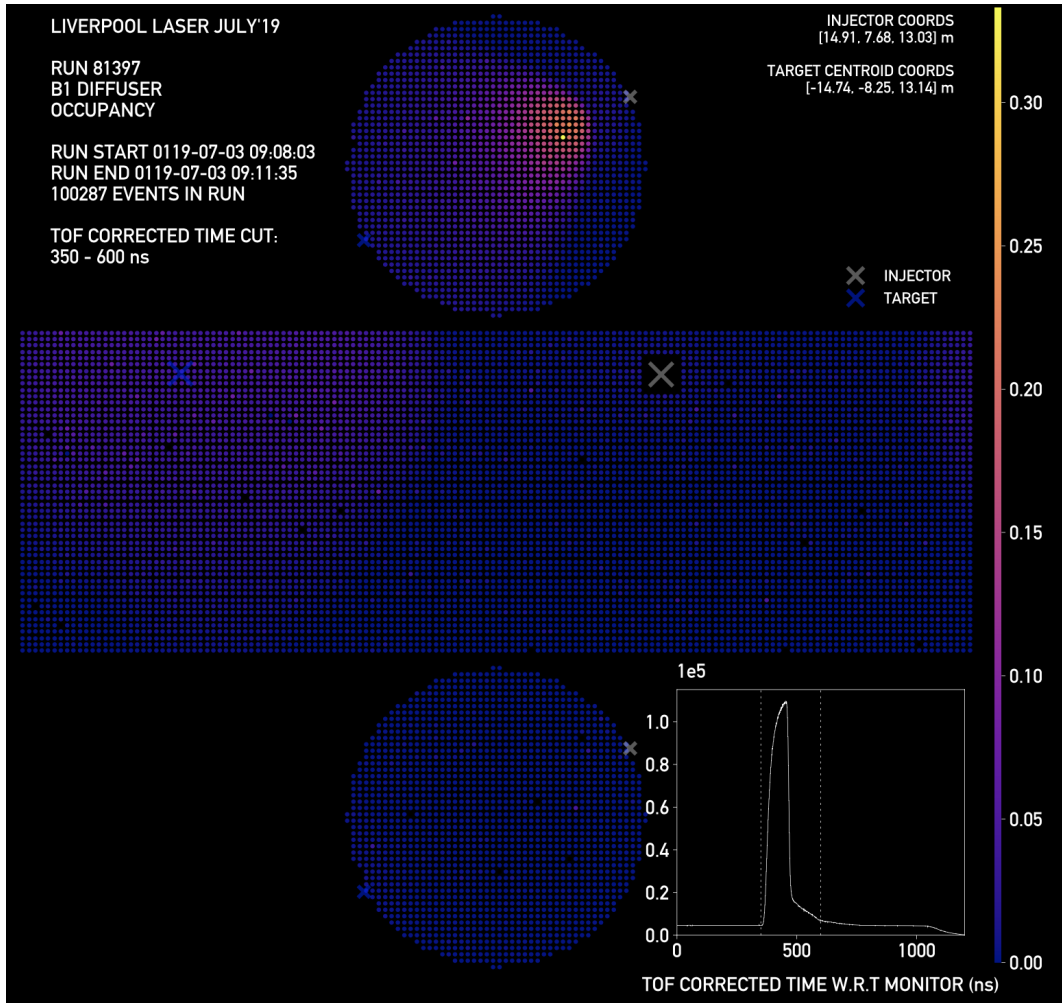


Figure B.3: An example PMT hit occupancy event display for diffuser installed at the B1 injection point over the full detector. The time of flight corrected hits as a function of time is shown on bottom right. Plot modified from [157].

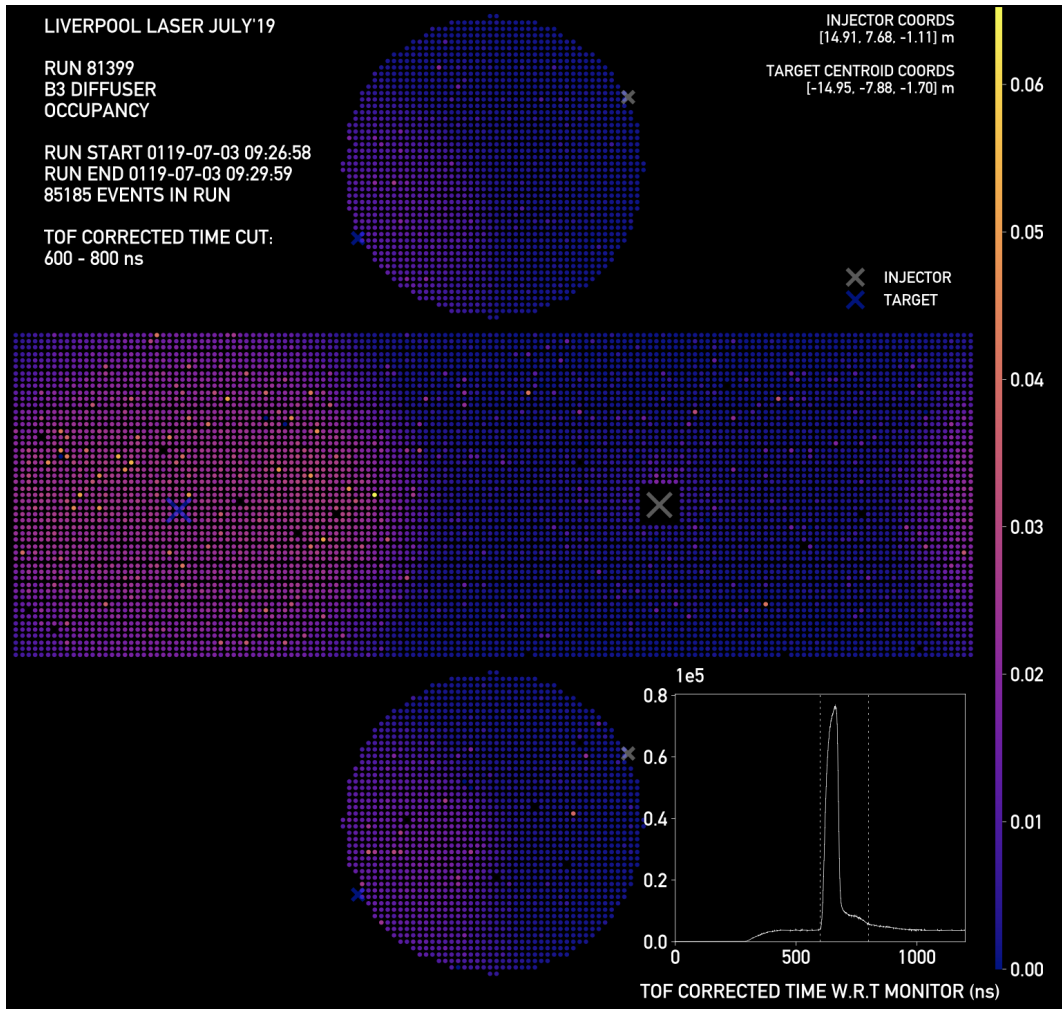


Figure B.4: An example PMT hit occupancy event display for diffuser installed at the B3 injection point over the full detector. The time of flight corrected hits as a function of time is shown on bottom right. Plot modified from [157].

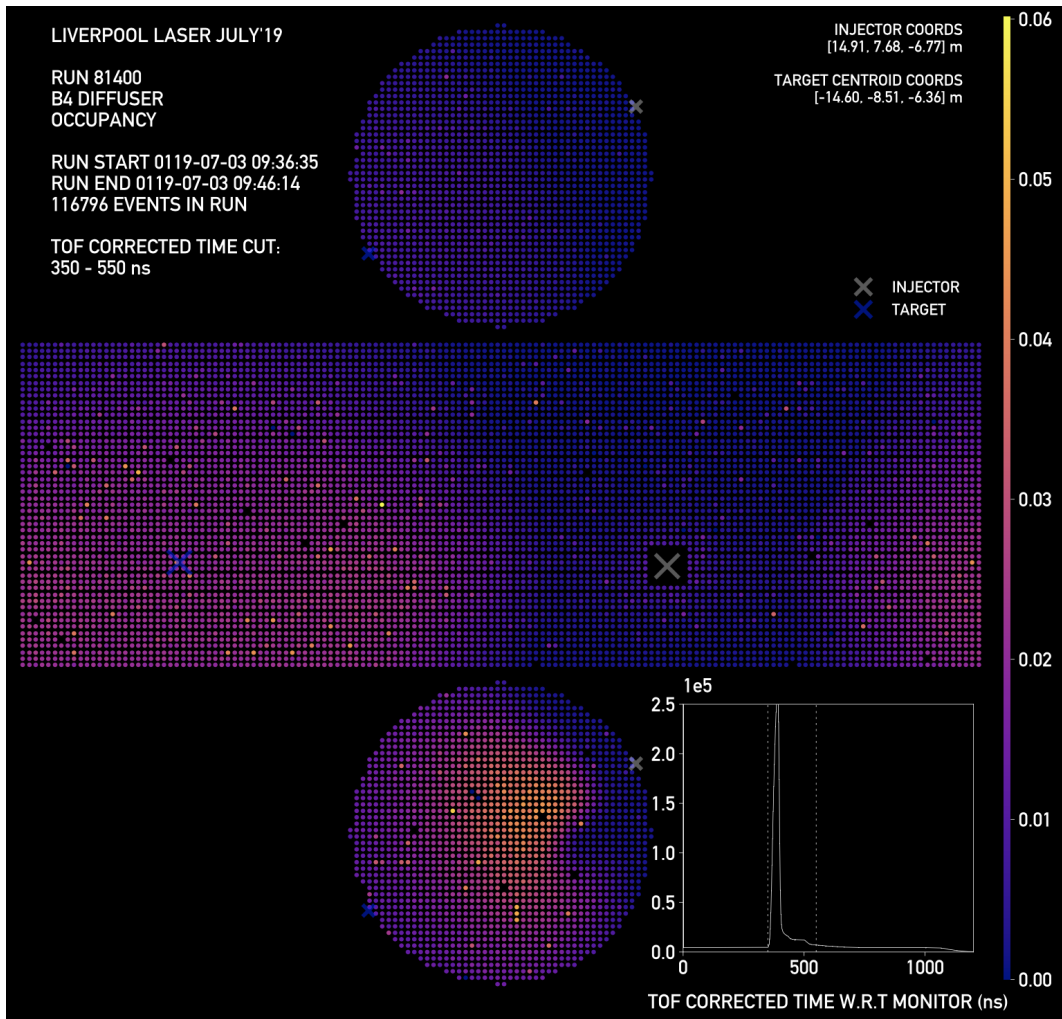


Figure B.5: An example PMT hit occupancy event display for diffuser installed at the B4 injection point over the full detector. The time of flight corrected hits as a function of time is shown on bottom right. Plot modified from [157].

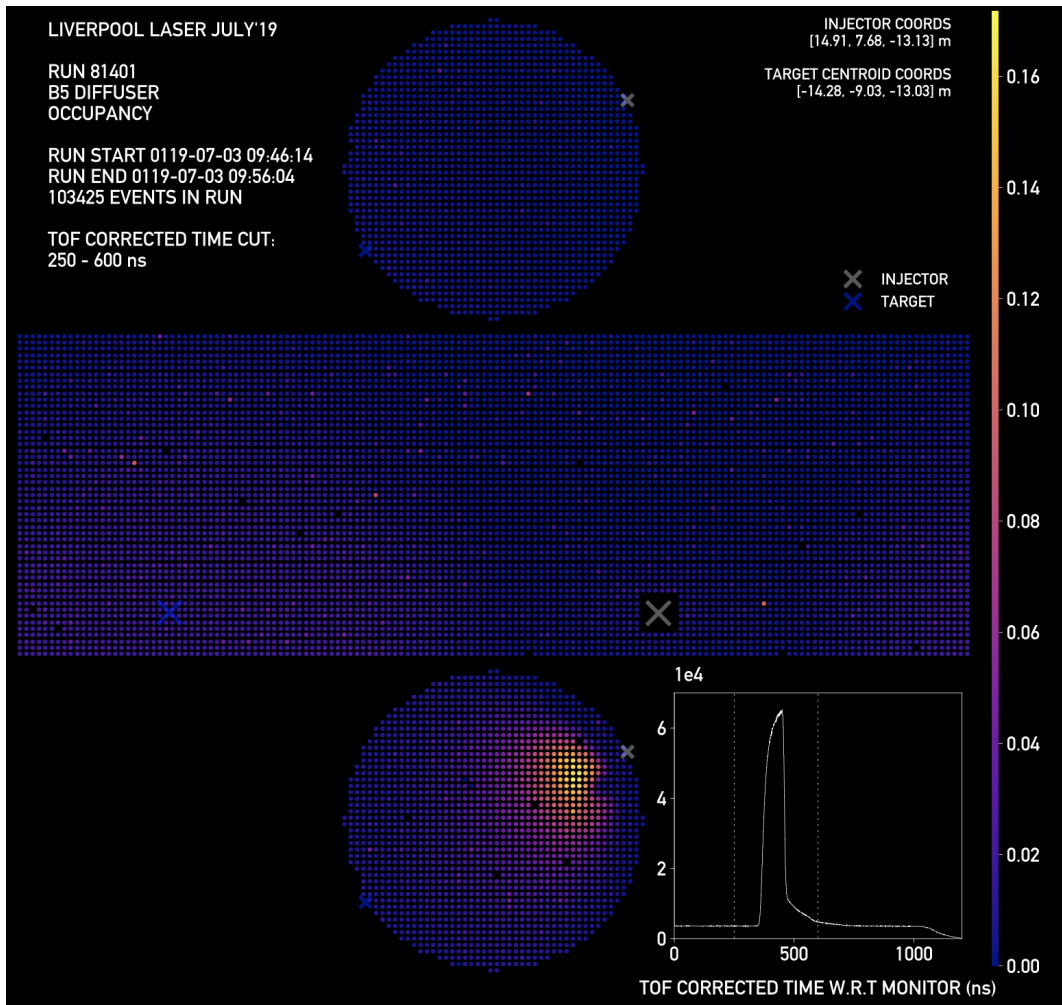


Figure B.6: An example PMT hit occupancy event display for diffuser installed at the B5 injection point over the full detector. The time of flight corrected hits as a function of time is shown on bottom right. Plot modified from [157].

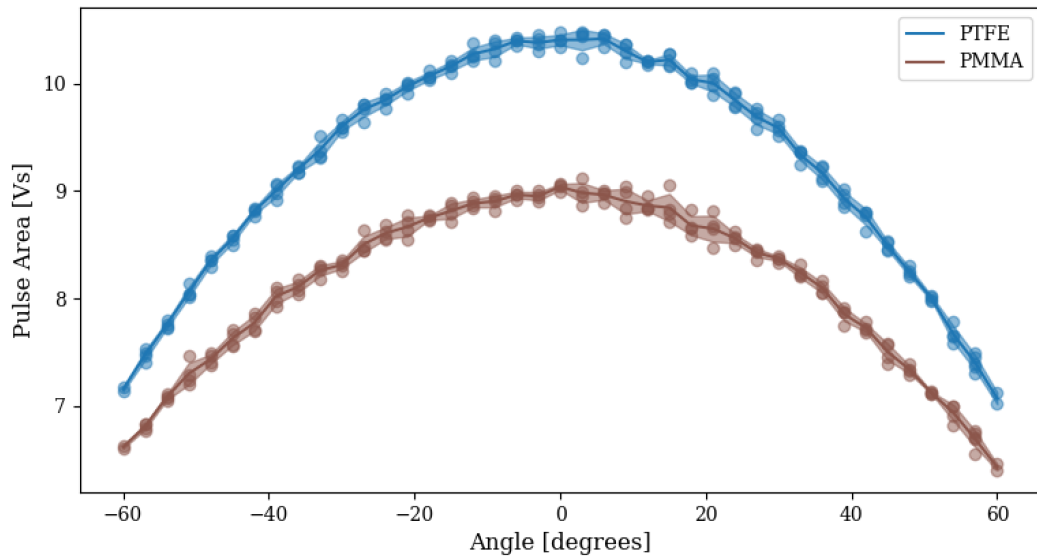


Figure B.7: A comparison of the pulse area as a function of angle for bare PMMA and PTFE diffusers.

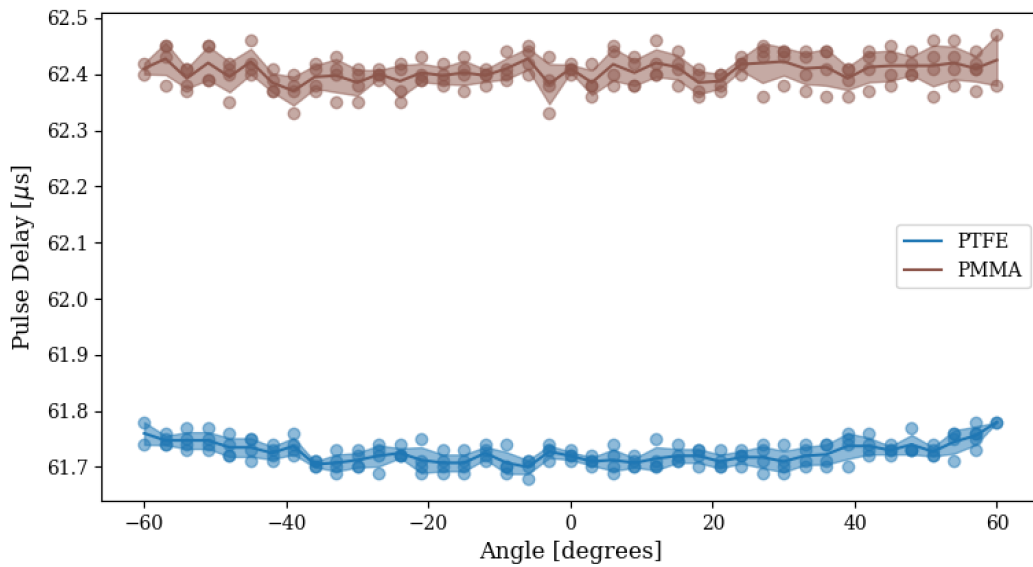


Figure B.8: A comparison of the pulse delay as a function of angle for bare PMMA and PTFE diffusers.

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