

B Physics Observables for New Physics Discoveries

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Contents of this talk

- What has already been done
 - the physics case in the CDR
 - progress since April 2007
- What remains to be done
 - improvements to the physics case
 - refined SuperB sensitivity estimates
 - answers to IRC questions
- Goals for this meeting
 - ... and beyond

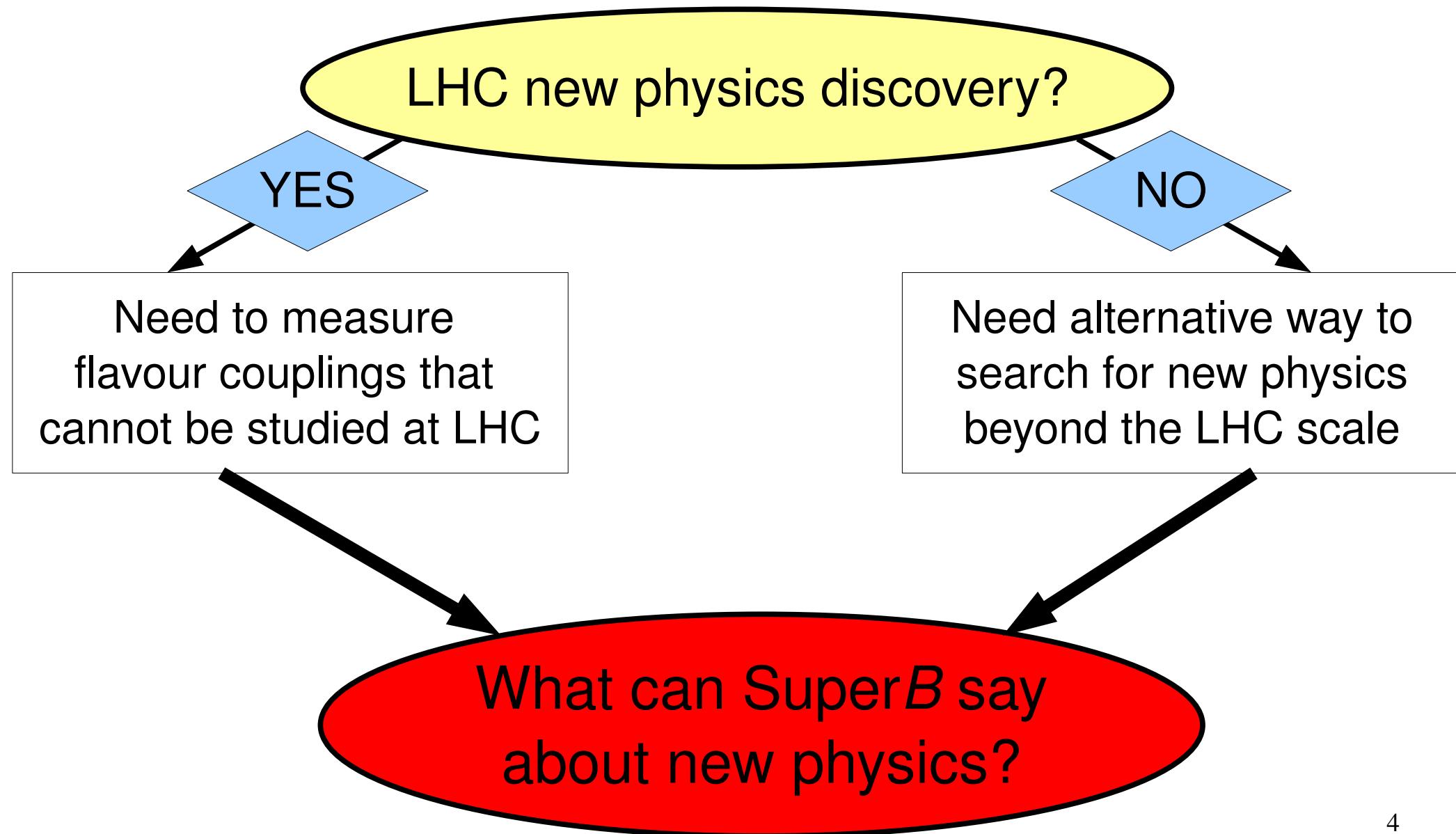


Need more and better
publicity material!

Assumptions

- Super*B* will accumulate 75/ab on the Y(4S)
 - beam energies 7 GeV e⁻ on 4 GeV e⁺
 - 5 years operation @ $L_{\text{peak}} \sim 10^{36}/\text{cm}^2/\text{s}$
 - data taking starts ~2014 \Rightarrow 75/ab by ~2020
 - LHC operation will be successful
 - LHCb will accumulate 10/fb before Super*B* starts
 - ATLAS & CMS will have plenty of data
- (no assumption whether or not NP is discovered at LHC)

Two Scenarios



What has already been done?

- Physics case in the CDR builds mainly on extrapolations of existing experiments
 - SuperBaBar:
 - SLAC-R-709 [arXiv:hep-ph/0503261]
 - SuperKEKB:
 - Letter of Intent, KEK Report 04-4 & arXiv:hep-ex/0406071
 - Other extrapolations from (mainly) BaBar results
- Identified need for strong focus on theoretically clean observables

Will be Studied at Super B

Δm_K ϵ_K ϵ'/ϵ_K $B(K_L \rightarrow \pi^0 \nu \bar{\nu})$ $B(K^+ \rightarrow \pi^+ \nu \bar{\nu})$ $B(K^+ \rightarrow l^+ \nu)$

Δm_d

$A_{SL}(B_d)$

$S(B_d \rightarrow J/\psi K_S)$

$S(B_d \rightarrow \phi K_S)$

$\alpha(B \rightarrow \pi \pi, \rho \pi, \rho \rho)$

$\gamma(B \rightarrow D K)$

CKM fits

Δm_s

$A_{SL}(B_s)$

$S(B_s \rightarrow J/\psi \phi)$

$S(B_s \rightarrow \phi \phi)$

$B(b \rightarrow s \gamma)$

$A_{CP}(b \rightarrow s \gamma)$

$S(B^0 \rightarrow K_S \pi^0 \gamma)$

$S(B_s \rightarrow \phi \gamma)$

$B(b \rightarrow d \gamma)$

$A_{CP}(b \rightarrow d \gamma)$

$A_{CP}(b \rightarrow (d+s)\gamma)$

$S(B^0 \rightarrow \rho^0 \gamma)$

$B(b \rightarrow s l^+ l^-)$

$B(b \rightarrow d l^+ l^-)$

$A_{FB}(b \rightarrow s l^+ l^-)$

$B(b \rightarrow s \nu \bar{\nu})$

$B(B_s \rightarrow l^+ l^-)$

$B(B_d \rightarrow l^+ l^-)$

$B(B^+ \rightarrow l^+ \nu)$

$B(\mu \rightarrow e \gamma)$

$B(\mu \rightarrow e^+ e^- e^+)$

$(g-2)_\mu$ μ EDM

$B(\tau \rightarrow \mu \gamma)$

$B(\tau \rightarrow e \gamma)$

$B(\tau^+ \rightarrow l^+ l^- l^+)$

τ CPV

τ EDM

$B(D_{(s)}^+ \rightarrow l^+ \nu)$

x_D y_D

$charm$ CPV

SuperB Physics Strength

- very many observables sensitive to new physics
- maximize sensitivity by combining information
- correlations between results distinguish models

Super Flavour Factory
“treasure chest”
of new physics observables



SuperB is a **general purpose detector** for flavour physics

(beware tendency to focus on “golden channels”)

INCREASING THEORETICAL UNCERTAINTY

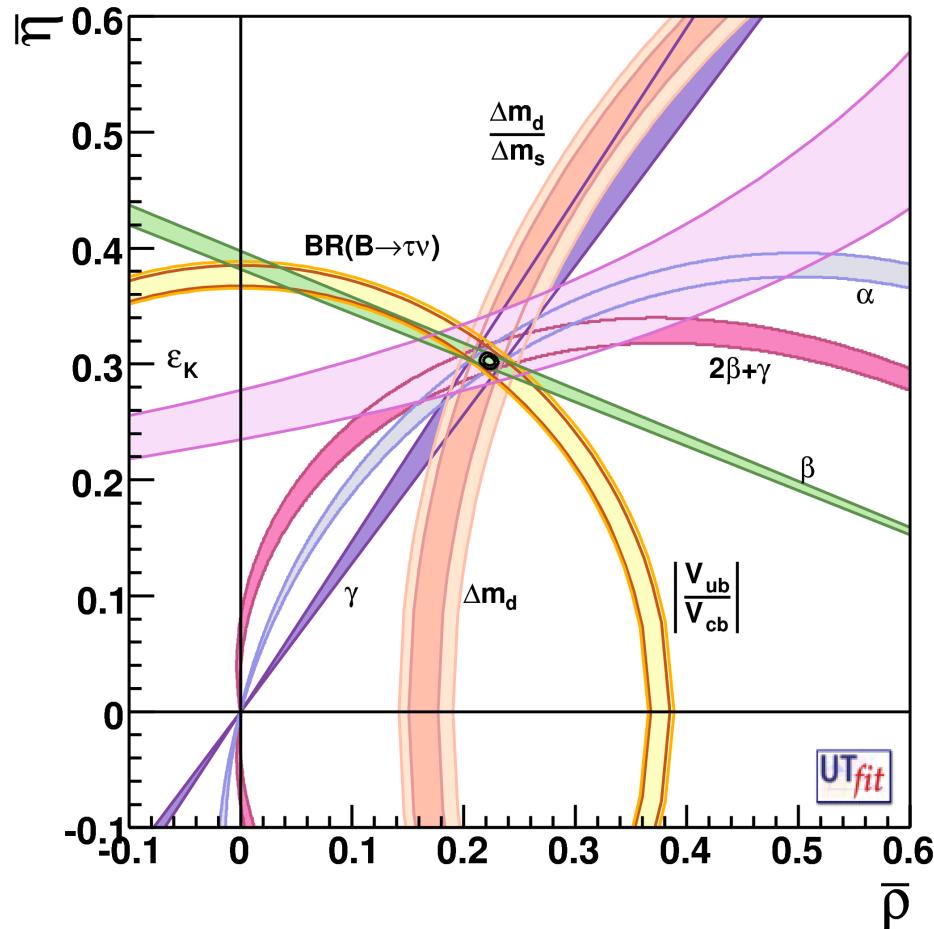
no theory improvements needed	$\beta(J/\psi K)$, $\gamma(DK)$, $\alpha(\pi\pi)^*$, lepton FV and UV, $S(p^0\gamma)$ CPV in $B \rightarrow X\gamma$, D and τ decays zero of FB asymmetry $B \rightarrow X_s l^+l^-$	NP insensitive or null tests of the SM or SM already known with the required accuracy
improved lattice QCD	meson mixing , $B \rightarrow D^{(*)} l\nu$, $B \rightarrow \pi(\rho) l\nu$, $B \rightarrow K^*\gamma$, $B \rightarrow \rho\gamma$, $B \rightarrow l\nu$, $B_s \rightarrow \mu\mu$	target error: ~1-2% Feasible (see below)
improved OPE+HQE	$B \rightarrow X_{u,c} l\nu$, $B \rightarrow X\gamma$	target error: ~1-2% Possibly feasible with SuperB data getting rid of the shape function. Detailed studies required
improved QCDF or SCET or flavour symmetries	S 's from TD A_ϕ in $b \rightarrow s$ transitions	target error: ~2-3% large and hard to improve uncertainties on small corrections. In addition, FS+data can bound the theoretical error

Some theoretically clean channels

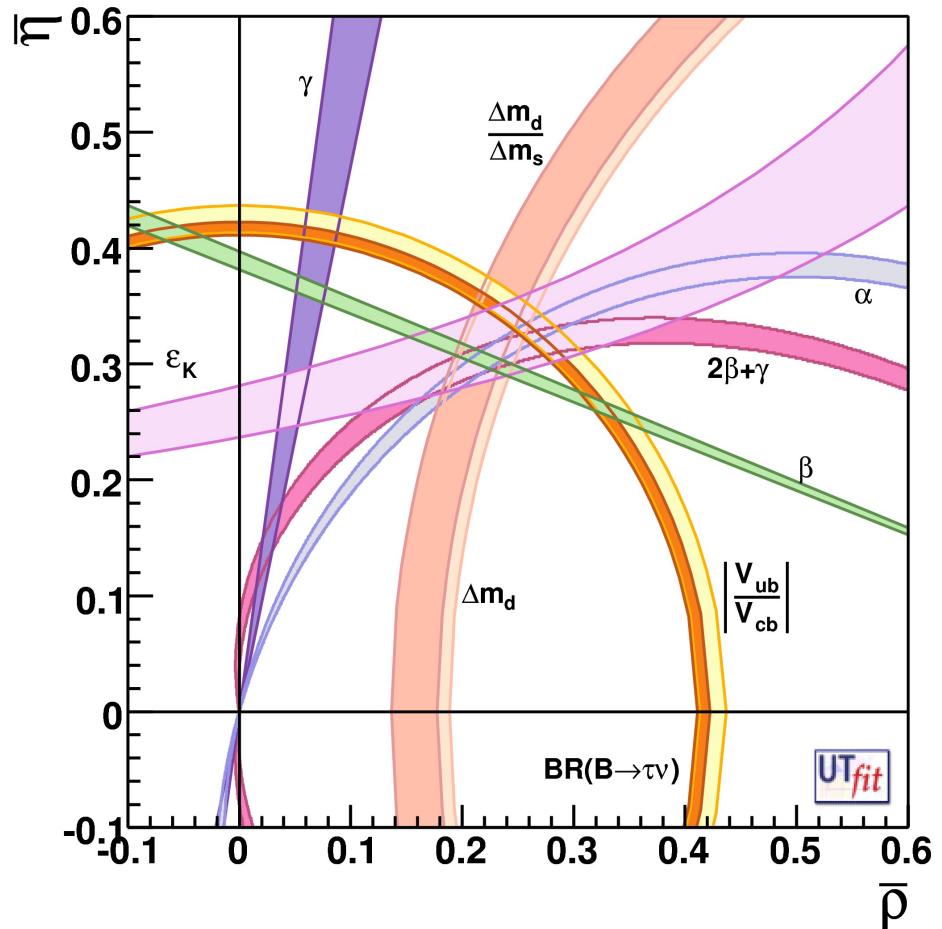
- Unitarity Triangle angles (α, β, γ)
 - Not discussed in this meeting
 - Work required:
 - How to reach 1° precision on α (theory + experiment)
 - Limiting systematics on β
- Unitarity Triangle sides (V_{ub}, V_{cb})
 - Talks of Viaud, Gambino
 - Work required
 - How to reach 1-2% precision on V_{ub}, V_{cb}

SuperB UT fit scenarios

“the nightmare”



“the dream”



- Possible NP discovery from precise CKM metrology
- Precise knowledge of SM parameters essential in any scenario

More theoretically clean channels

- $B \rightarrow X_s \gamma$
 - Talks of Walsh, Hurth
- $B \rightarrow X_s l^+ l^-$
 - Talks of Renga, Hurth
- $B \rightarrow X_s \nu \bar{\nu}$
 - Talk of Renga

Inclusive channels:

- need detailed study of sensitivity for realistic SuperB accelerator & detector
- effect of hermeticity on recoil analyses
- Important to work with tools group

Some exclusive channels

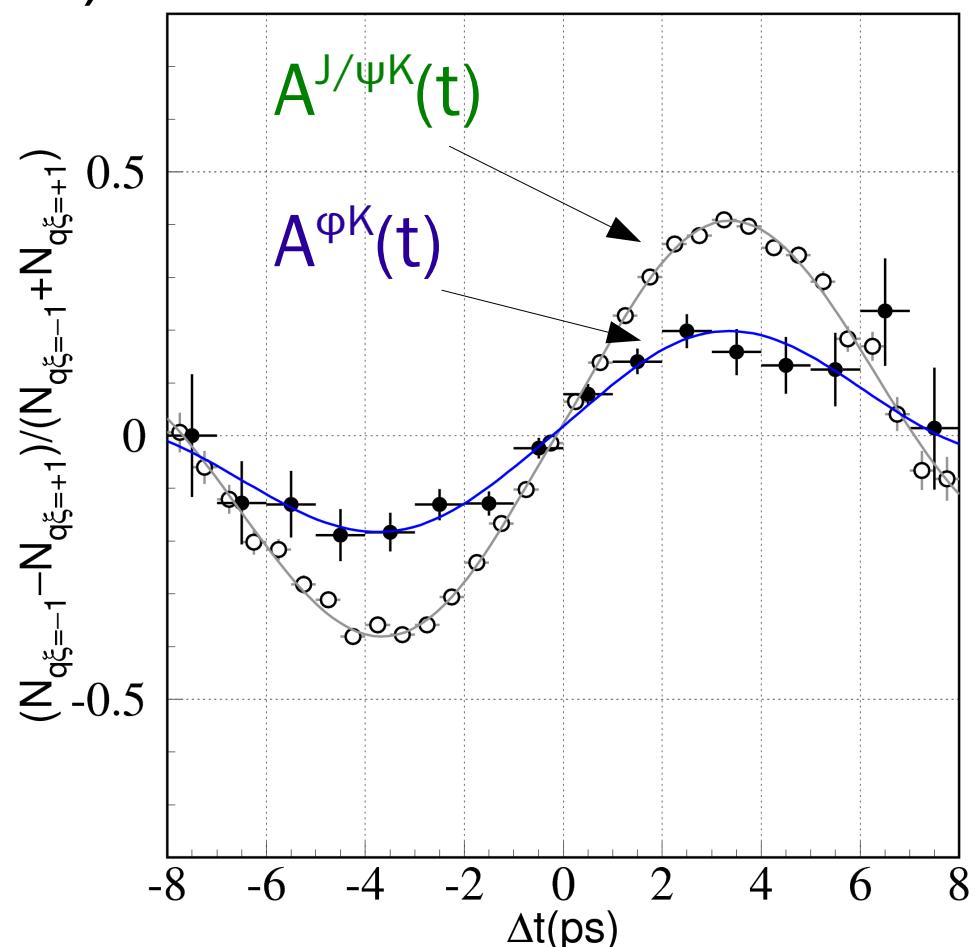
- $B \rightarrow l\nu(\gamma)$
 - Talks of Kou, Bevan/Cavoto
- $B \rightarrow K^{(*)}\tau^+\tau^-$
 - Talk of Robertson

Some modes not discussed in CDR

- Assess interest and feasibility of measurements at Super B

Some topics not covered

- Make Super B version of this plot (from SuperKEKB physics case)
- Focus more on *really* theoretically clean channels
- However, this is a good example of the kind of discovery plot that we need to make



Some topics not covered

- Various exclusive channels, eg. $K^{(*)} l^+ l^-$
- This specific example is sensitive to new physics **BUT**
 - Not as clean as the inclusive channel
 - Will be well measured by LHCb
- There are other similar examples

Progress since the CDR

- Important new experimental results
 - A couple of examples:
 - $B \rightarrow X_d \gamma$ (BaBar arXiv:0708.1652)
 - $B \rightarrow D^{(*)} \tau \nu$ (Belle arXiv:0706.4429 & BaBar arXiv:0709.1698)
 - These have impact on Super*B* sensitivity estimates
- Final “Flavour in the LHC Era” workshop
 - Commencement of follow-up workshops (“Interplay of Collider and Flavour Physics”)
 - A lot of ongoing work in this important area – highly relevant to Super*B* physics case

Interplay

Some (not all!) of the papers on flavour-LHC interplay in the last ~2 years

- E. Accomando *et al.*, CERN-2006-009 [arXiv:hep-ph/0608079].
B. C. Allanach and C. G. Lester, Phys. Rev. D **73** (2006) 015013 [arXiv:hep-ph/0507283].
B. C. Allanach, Phys. Lett. B **635** (2006) 123 [arXiv:hep-ph/0601089].
R. R. de Austri, R. Trotta and L. Roszkowski, JHEP **0605** (2006) 002 [arXiv:hep-ph/0602028].
M. S. Carena, A. Menon, R. Noriega-Papaqui, A. Szynkman and C. E. M. Wagner, Phys. Rev. D **74** (2006) 015009 [arXiv:hep-ph/0603106].
J. R. Ellis, S. Heinemeyer, K. A. Olive and G. Weiglein, arXiv:hep-ph/0604180.
G. Isidori and P. Paradisi, Phys. Lett. B **639** (2006) 499 [arXiv:hep-ph/0605012].
E. Lunghi, W. Porod and O. Vives, Phys. Rev. D **74** (2006) 075003 [arXiv:hep-ph/0605177].
B. C. Allanach, C. G. Lester and A. M. Weber, JHEP **0612** (2006) 065 [arXiv:hep-ph/0609295].
M. K. Parida *et al.*, Pramana **67** (2006) 849.
G. Isidori, F. Mescia, P. Paradisi and D. Temes, Phys. Rev. D **75** (2007) 115019 [arXiv:hep-ph/0703035].
M. S. Carena, A. Menon and C. E. M. Wagner, Phys. Rev. D **76** (2007) 035004 [arXiv:0704.1143 [hep-ph]].
P. J. Fox, Z. Ligeti, M. Papucci, G. Perez and M. D. Schwartz, arXiv:0704.1482 [hep-ph].
L. Roszkowski, R. Ruiz de Austri and R. Trotta, JHEP **0707** (2007) 075 [arXiv:0705.2012 [hep-ph]].
J. R. Ellis, S. Heinemeyer, K. A. Olive, A. M. Weber and G. Weiglein, arXiv:0706.0652 [hep-ph].
J. R. Ellis, S. Heinemeyer, K. A. Olive and G. Weiglein, arXiv:0706.0977 [hep-ph].
Y. Grossman, Y. Nir, J. Thaler, T. Volansky and J. Zupan, arXiv:0706.1845 [hep-ph].
O. Buchmueller *et al.*, arXiv:0707.3447 [hep-ph].
S. Dittmaier, G. Hiller, T. Plehn and M. Spannowsky, arXiv:0708.0940 [hep-ph].
J. Ellis, T. Hahn, S. Heinemeyer, K. A. Olive and G. Weiglein, arXiv:0709.0098 [hep-ph].
-

Talks in this meeting on interplay

- Talks in this meeting
 - Tools HEP/Flavour interplay (Ronga)
 - SUSY breaking scenario (Shindou)
 - MFV + Snowmass point (Ciuchini + Silvestrini)
- Slight emphasis on theoretical side at present, but experimentalists should pay close attention (and get involved)

Snowmass points

- From hep-ph/0202233
 - “The question of which parameter choices are useful as benchmark scenarios depends on the purpose of the actual investigation.”
 - “It should be obvious that it is not possible to define a single set of benchmark scenarios that will serve all purposes. The usefulness of a particular scenario will always depend on which sector of the theory and which physics issue is investigated.”
- We should not be afraid to define our own benchmark parameter points if necessary

Digression: Target Luminosity

- $L_{\text{peak}} \sim 10^{36}/\text{cm}^2/\text{s} \Leftrightarrow 75/\text{ab}$ on the Y(4S)
- Most of the theoretically clean observables will still be experimentally limited
- Do we need to aim higher?
- Contrary question: would 10/ ab be enough?

SuperB physics in tables

Observable	B factories (2 ab^{-1})	SuperB (75 ab^{-1})
$\sin(2\beta) (J/\psi K^0)$	0.018	0.005 (\dagger)
$\cos(2\beta) (J/\psi K^{*0})$	0.30	0.05
$\sin(2\beta) (D h^0)$	0.10	0.02
$\cos(2\beta) (D h^0)$	0.20	0.04
$S(J/\psi \pi^0)$	0.10	0.02
$S(D^+ D^-)$	0.20	0.03
$S(\phi K^0)$	0.13	0.02 (\star)
$S(\eta' K^0)$	0.05	0.01 (\star)
$S(K_S^0 K_S^0 K_S^0)$	0.15	0.02 (\star)
$S(K_S^0 \pi^0)$	0.15	0.02 (\star)
$S(\omega K_S^0)$	0.17	0.03 (\star)
$S(f_0 K_S^0)$	0.12	0.02 (\star)
$\gamma (B \rightarrow DK, D \rightarrow CP \text{ eigenstates})$	$\sim 15^\circ$	2.5°
$\gamma (B \rightarrow DK, D \rightarrow \text{suppressed states})$	$\sim 12^\circ$	2.0°
$\gamma (B \rightarrow DK, D \rightarrow \text{multibody states})$	$\sim 9^\circ$	1.5°
$\gamma (B \rightarrow DK, \text{combined})$	$\sim 6^\circ$	$1-2^\circ$
$\alpha (B \rightarrow \pi\pi)$	$\sim 16^\circ$	3°
$\alpha (B \rightarrow \rho\rho)$	$\sim 7^\circ$	$1-2^\circ$ (\star)
$\alpha (B \rightarrow \rho\pi)$	$\sim 12^\circ$	2°
$\alpha (\text{combined})$	$\sim 6^\circ$	$1-2^\circ$ (\star)
$2\beta + \gamma (D^{(*)\pm} \pi^\mp, D^\pm K_S^0 \pi^\mp)$	20°	5°
$ V_{cb} (\text{exclusive})$	4% (\star)	1.0% (\star)
$ V_{cb} (\text{inclusive})$	1% (\star)	0.5% (\star)
$ V_{ub} (\text{exclusive})$	8% (\star)	3.0% (\star)
$ V_{ub} (\text{inclusive})$	8% (\star)	2.0% (\star)
$BR(B \rightarrow \tau\nu)$	20%	4% (\dagger)
$BR(B \rightarrow \mu\nu)$	visible	5%
$BR(B \rightarrow D\tau\nu)$	10%	2%
$BR(B \rightarrow \rho\gamma)$	15%	3% (\dagger)
$BR(B \rightarrow \omega\gamma)$	30%	5%
$A_{CP}(B \rightarrow K^*\gamma)$	0.007 (\dagger)	0.004 (\dagger \star)
$A_{CP}(B \rightarrow \rho\gamma)$	~ 0.20	0.05
$A_{CP}(b \rightarrow s\gamma)$	0.012 (\dagger)	0.004 (\dagger)
$A_{CP}(b \rightarrow (s+d)\gamma)$	0.03	0.006 (\dagger)
$S(K_S^0 \pi^0 \gamma)$	0.15	0.02 (\star)
$S(\rho^0 \gamma)$	possible	0.10
$A_{CP}(B \rightarrow K^*\ell\ell)$	7%	1%
$A^{FB}(B \rightarrow K^*\ell\ell)s_0$	25%	9%
$A^{FB}(B \rightarrow X_s \ell\ell)s_0$	35%	5%
$BR(B \rightarrow K\nu\bar{\nu})$	visible	20%
$BR(B \rightarrow \pi\nu\bar{\nu})$	-	possible

Mode	Observable	B Factories (2 ab^{-1})	SuperB (75 ab^{-1})
$D^0 \rightarrow K^+ K^-$	y_{CP}	$2-3 \times 10^{-3}$	5×10^{-4}
$D^0 \rightarrow K^+ \pi^-$	y'_D	$2-3 \times 10^{-3}$	7×10^{-4}
	$x_D'^2$	$1-2 \times 10^{-4}$	3×10^{-5}
$D^0 \rightarrow K_S^0 \pi^+ \pi^-$	y_D	$2-3 \times 10^{-3}$	5×10^{-4}
	x_D	$2-3 \times 10^{-3}$	5×10^{-4}
Average	y_D	$1-2 \times 10^{-3}$	3×10^{-4}
	x_D	$2-3 \times 10^{-3}$	5×10^{-4}

5-10x
improvement

Process	Sensitivity
$\mathcal{B}(\tau \rightarrow \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow eee)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow \mu\eta)$	4×10^{-10}
$\mathcal{B}(\tau \rightarrow e\eta)$	6×10^{-10}
$\mathcal{B}(\tau \rightarrow \ell K_S^0)$	2×10^{-10}

+ τ FC physics (CPV, ...)

Super Flavour Factory

a "treasure chest"
of new
physics-
sensitive
observables

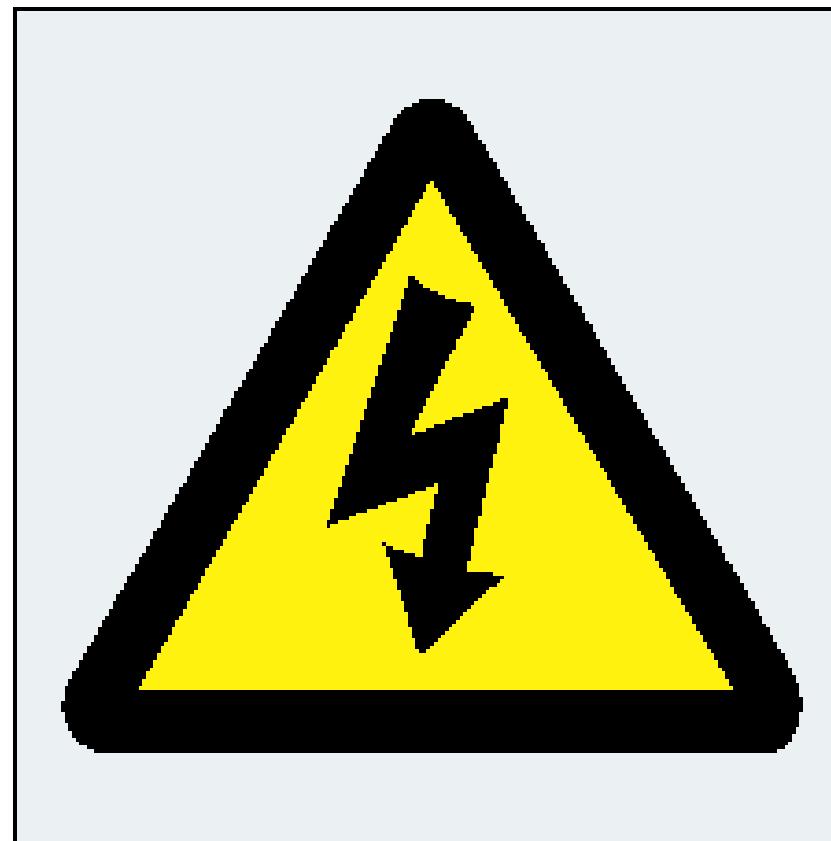


Observable	Error with 1 ab^{-1}
$\Delta\Gamma$	0.16 ps^{-1}
Γ	0.07 ps^{-1}
β_s from angular analysis	20°
A_{SL}^*	0.006
A_{CH}	0.004
$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)$	-
$ V_{cb}/V_{ts} $	0.08
$\mathcal{B}(B_s \rightarrow \gamma\gamma)$	38%
β_s from $J/\psi\phi$	10°

Goals for this meeting (and beyond)

- answers to IRC questions
- produce more and better publicity material
- new ideas always welcome
- refined SuperB sensitivity estimates
 - need to go significantly beyond what is in the CDR
 - realistic machine and detector simulation
 - work closely with tools group

Back Up



SuperB and LHCb (upgrade)

- SuperB will start running >2015
- Results from LHCb will be in
 - Dramatically improve knowledge in B_s sector
 - Much better determination of γ
 - Precise studies of some exclusive modes
 - eg. $B_s \rightarrow \mu\mu$, $B \rightarrow K(*)\mu\mu$
- These improvements on the current knowledge are assumed in SuperB physics case
 - Strong focus on theoretically clean processes